



Cite this: *Phys. Chem. Chem. Phys.*,  
2025, 27, 13861

Received 5th May 2025,  
Accepted 10th June 2025

DOI: 10.1039/d5cp01698c

rsc.li/pccp

# Treatment of spin–orbit coupling with internally contracted multireference coupled cluster theory†

Andreas Köhn<sup>✉</sup> and Julia Netz<sup>✉</sup>

We present the formalism for computing state–interaction matrix elements within the multistate extension of internally contracted multireference coupled-cluster theory. In particular, we focus on the determination of spin–orbit coupling matrix elements. A pilot implementation is presented and tested for the zero-field splitting of atomic  $^2\text{P}$  and  $^3\text{P}$  terms as well as for molecular  $^2\Pi$  terms. We also investigate the impact of the non-Hermiticity of the underlying coupled-cluster theory on spin–orbit couplings between non-equivalent states. For this, we consider the coupling between the  $^3\Sigma^-$  ground state and the  $^1\Pi$  excited state of NH for several interatomic distances. Only a very small asymmetry is found for this case, but it is also seen that the present theory has to be carefully revisited with respect to the treatment of the response of the reference coefficients in order to avoid artefacts.

## 1 Introduction

Spin–orbit interactions<sup>1</sup> are at the heart of many important phenomena, such as intersystem crossing,<sup>1,2</sup> being important for the photophysics of organic dyes, or zero-field splitting of the ground state of transition metal or rare earth ions, giving rise to their unique magnetic properties<sup>3,4</sup> and making them interesting candidates for molecular magnetism, spintronics and quantum technology.<sup>5</sup> The computational modelling of spin–orbit interactions was pursued early on in quantum chemistry, starting with perturbation approaches within configuration interaction (CI) theory.<sup>6–9</sup> There also exist implementations that include the spin–orbit effects directly into the self-consistent solutions<sup>10,11</sup> and full relativistic treatments based on the Dirac equation.<sup>12</sup>

For light and medium-sized nuclei, spin–orbit coupling matrix elements do not exceed  $2000\text{ cm}^{-1}$  (0.25 eV) and the perturbative approach is usually sufficient. While the size of spin–orbit coupling effects are qualitatively well recovered by small CI expansions, a more quantitative treatment calls for the inclusion of dynamic correlation effects. This can of course be accomplished by larger multireference CI expansions,<sup>7,8</sup> or by (multiconfiguration) perturbation theory.<sup>13–17</sup> For large molecules, density functional theory is an alternative option.<sup>18,19</sup>

In terms of accurate wavefunction theory, coupled-cluster theory is the most desirable target due to its high accuracy and size-consistency.<sup>20</sup> However, appropriate extensions of the theory have to be developed in order to make it applicable to near-

degenerate states,<sup>21,22</sup> for which spin–orbit coupling is often most important. Pioneering work has been carried out at the level of equation-of-motion ionisation potential coupled-cluster (EOMIP-CC) theory by Klein and Gauss.<sup>23</sup> This approach was later generalised to further EOM approaches (like electron-attachment or spin-flip) by Epifanovsky *et al.*,<sup>24</sup> which significantly widens the scope of spin-cases that can be treated. Spin–orbit effects for multireference coupled-cluster theory have for the first time been explored by of Mück and Gauss<sup>25</sup> who focused on Mukherjee's state-specific multireference coupled-cluster (Mk-MRCC) theory.<sup>26,27</sup>

In the present work, we seek to integrate spin–orbit effects into internally contracted multireference coupled-cluster (icMRCC) theory.<sup>28–35</sup> While some works exist that combine icMRCC energies and spin–orbit interaction matrix elements from small CI expansions to arrive at spin–orbit coupled states of small molecules,<sup>36,37</sup> the aim of the present work is to also include dynamic correlation effects on the matrix elements. To this end, we exploit that the solutions of the icMRCC equations, despite obtained in a state-specific fashion, can also be viewed as the solutions of a multistate framework<sup>33,35</sup> such that they can be directly used as zeroth-order function for the evaluation of spin–orbit coupling matrix elements.

The present work is organised as follows: in Sections 2.1–2.3, we review the icMRCC theory for state specific energies and first-order properties, as well as the multistate extension of the theory. In Section 2.4 we introduce the new equations for treating state interactions within the multistate framework and its application to spin–orbit coupling (Section 2.5). In Section 3.1, we shortly describe the implementation and show thereafter the first exploratory results obtained by the new approach. These consist of the zero-field splittings of atoms, Section 3.2, and molecules with degenerate ground state,

Institute for Theoretical Chemistry, University of Stuttgart, Pfaffenwaldring 55, DE-70569 Stuttgart, Germany. E-mail: koehn@theochem.uni-stuttgart.de

† Electronic supplementary information (ESI) available. See DOI: <https://doi.org/10.1039/d5cp01698c>



Section 3.3. In addition we will investigate the impact of the non-Hermiticity of the coupled-cluster approach, Section 3.4.

## 2 Theory

### 2.1 Internally contracted multireference coupled-cluster theory

The internally contracted multireference coupled-cluster (icMRCC) wavefunction<sup>30,31</sup> is defined as

$$|\Psi\rangle = e^{\hat{T}}|\psi_0\rangle = e^{\hat{T}}\sum_{\mu}|\Phi_{\mu}\rangle c_{\mu} \quad (1)$$

where  $|\psi_0\rangle$  is the multiconfigurational reference wave function. We assume that it is a complete active space (CAS) expansion in terms of Slater determinants  $|\Phi_{\mu}\rangle$  and corresponding coefficients  $c_{\mu}$ . The index  $\mu$  is assumed to run over the entire CAS manifold.

Dynamical electron correlation is introduced by the cluster operator

$$\hat{T} = \sum_{\rho} t_{\rho} \hat{\tau}_{\rho} \quad (2)$$

consisting of excitation operators  $\hat{\tau}_{\rho}$  and corresponding cluster amplitudes  $t_{\rho}$ . The index  $\rho$  runs over all included excitations, *e.g.* one- and two-body excitations in case of the icMRCC singles and doubles (icMRCCSD) approximation. For the present work, we keep the notation rather abstract, for a more detailed definition of the excitation manifold see for example ref. 34. For the purposes of the present work it suffices to say that the excitation manifold includes excitations that promote electrons from doubly occupied orbitals into partially occupied or unoccupied orbitals in either the active space or the secondary space (virtual orbitals), but they also promote electrons within the active space or from the active space into the secondary space. Purely active excitations that only reshuffle electrons inside the active space are omitted, instead the expansion coefficients  $c_{\mu}$  will be reoptimised. As the excitation spaces overlap, the excitation operators do not commute in general, at variance to the single-reference case. The specific choice of the excitation manifold requires special considerations in order to avoid linear dependencies and to achieve size-consistency, as discussed previously.<sup>31,34,38</sup>

By inserting the icMRCC wavefunction, eqn (1), into the electronic Schrödinger equation with the (non-relativistic) clamped-nuclei electronic Hamiltonian  $\hat{H}$ , the equations to determine the cluster amplitudes  $t_{\rho}$  and coefficients  $c_{\mu}$  can be derived. We use the linked form of the projection approach, where we first premultiply both sides of the Schrödinger equation by  $e^{-\hat{T}}$  before projecting to excited configurations. We obtain the amplitude equations

$$0 = \langle \psi_0 | \hat{\tau}_{\rho}^{\dagger} e^{-\hat{T}} \hat{H} e^{\hat{T}} | \psi_0 \rangle \quad (3)$$

by projecting to the non-redundant internally contracted manifold  $\{\langle \psi_0 | \hat{\tau}_{\rho}^{\dagger} \rangle\}$ .

Projection onto the reference space determinants leads to an eigenvalue problem

$$\sum_{\nu} \langle \Phi_{\mu} | e^{-\hat{T}} \hat{H} e^{\hat{T}} | \Phi_{\nu} \rangle c_{\nu} = E c_{\mu} \quad (4)$$

of the effective Hamiltonian

$$H_{\mu\nu}^{\text{eff}} = \langle \Phi_{\mu} | e^{-\hat{T}} \hat{H} e^{\hat{T}} | \Phi_{\nu} \rangle. \quad (5)$$

The equations, eqn (3) and (4) are solved simultaneously, the lowest eigenvalue of the effective Hamiltonian gives the icMRCC energy.

### 2.2 Expectation values

In analogy to single-reference coupled-cluster theory, molecular properties can be computed within the icMRCC framework by introducing a stationary energy functional.<sup>31,39</sup> For a more compact notation, we define the similarity transformed Hamiltonian

$$\tilde{H} = e^{-\hat{T}} \hat{H} e^{\hat{T}} \quad (6)$$

and write the desired energy functional as:

$$\mathcal{L} = \langle \tilde{\psi}_0 | \tilde{H} | \psi_0 \rangle + \langle \psi_0 | \hat{A} \tilde{H} | \psi_0 \rangle - E(\langle \tilde{\psi}_0 | \psi_0 \rangle - 1). \quad (7)$$

This expression contains additional coefficients, which are contained in the left-hand reference function

$$\langle \tilde{\psi}_0 | = \sum_{\mu} \tilde{c}_{\mu} \langle \Phi_{\mu} | \quad (8)$$

and the  $\hat{A}$  operator

$$\hat{A} = \sum_{\rho} \lambda_{\rho} (\hat{\tau}_{\rho}^{\dagger})^{\dagger}. \quad (9)$$

The energy  $E$  appears here as Lagrange multiplier for the normalization condition  $\langle \tilde{\psi}_0 | \psi_0 \rangle = 1$ .

Taking the derivatives with respect to  $\lambda_{\rho}$  and  $\tilde{c}_{\mu}$  recovers, by construction, the amplitude and coefficient equations, eqn (3) and (4). Requiring stationarity with respect to the cluster amplitudes and the reference coefficients leads to two additional equations. The condition  $\partial_{c_{\mu}} \mathcal{L} = 0$  leads to<sup>39</sup>

$$\langle \tilde{\psi}_0 | (\tilde{H} - E) | \Phi_{\mu} \rangle + \langle \psi_0 | \hat{A} \tilde{H} | \Phi_{\mu} \rangle + \langle \Phi_{\mu} | \hat{A} \tilde{H} | \psi_0 \rangle = 0, \quad (10)$$

and the condition  $\partial_{t_{\rho}} \mathcal{L} = 0$  gives

$$\langle \tilde{\psi}_0 | (\partial_{t_{\rho}} \tilde{H}) | \psi_0 \rangle + \langle \psi_0 | \hat{A} (\partial_{t_{\rho}} \tilde{H}) | \psi_0 \rangle = 0. \quad (11)$$

The last term on the left-hand side of eqn (10) results from the derivative of the internally contracted projection manifold. As this term vanishes in the limit of a complete cluster operator, it has been argued to neglect this term (neglect of internal projection response, NIPR).<sup>39</sup> Eqn (10) and (11) can be summarized as a homogenous linear set of equations

$$\begin{pmatrix} (\tilde{c}_{\mu}) & (\lambda_{\rho}) \end{pmatrix} \begin{pmatrix} (H_{\mu\nu}^{\text{eff}} - E \delta_{\mu\nu}) & (\langle \Phi_{\mu} | (\partial_{t_{\sigma}} \tilde{H}) | \psi_0 \rangle) \\ (\langle \psi_0 | \hat{\tau}_{\rho}^{\dagger} \tilde{H} | \Phi_{\nu} \rangle) & (\langle \psi_0 | \hat{\tau}_{\rho}^{\dagger} (\partial_{t_{\sigma}} \tilde{H}) | \psi_0 \rangle) \end{pmatrix} = 0 \quad (12)$$



Phys. Chem. Chem. Phys., 2025, 27, 13861–13869 | 13863

employ the Lagrange formalism of coupled-cluster response theory and set up the following functional:

$$L_{ab}(\varepsilon) = \langle \bar{\psi}_a | \bar{H}^b(\varepsilon) | \psi_b \rangle + \langle \psi_b | \hat{\Lambda}^{ab}(\varepsilon) \bar{H}^b(\varepsilon) | \psi_b \rangle - \langle \psi_b | \hat{\Lambda}^{ab}(\varepsilon) e^{-\hat{T}^b(\varepsilon)} e^{\hat{T}^a(\varepsilon)} | \psi_a \rangle \langle \bar{\psi}_a | \bar{H}^b(\varepsilon) | \psi_b \rangle. \quad (30)$$

It contains the objective quantity as the first term, the second and third term involve the operator

$$\hat{\Lambda}^{ab} = \sum_{\rho} \lambda_{\rho}^{ab} \hat{\tau}_{\rho}^{\prime b \dagger} \quad (31)$$

which includes the Lagrange multipliers for the constraint that the multistate equations, including the coupling between states *a* and *b*, must be fulfilled for any value of  $\varepsilon$ :

$$\langle \psi_b | \hat{\tau}_{\rho}^{\prime b \dagger} \bar{H}^b(\varepsilon) | \psi_b \rangle - \langle \psi_b | \hat{\tau}_{\rho}^{\prime b \dagger} e^{-\hat{T}^b(\varepsilon)} e^{\hat{T}^a(\varepsilon)} | \psi_a \rangle H_{ab}^{\text{eff}}(\varepsilon) = 0. \quad (32)$$

This functional is required to be stationary with respect to variations of the amplitudes, and for vanishing perturbations we get

$$\begin{aligned} \partial_{\rho} L_{ab}|_{\varepsilon=0} &= \langle \bar{\psi}_a | (\partial_{\rho} \bar{H}^b) | \psi_b \rangle + \langle \psi_b | \hat{\Lambda}^{ab} (\partial_{\rho} \bar{H}^b) | \psi_b \rangle \\ &- \langle \psi_b | \hat{\Lambda}^{ab} (\partial_{\rho} e^{-\hat{T}^b}) e^{\hat{T}^a} | \psi_a \rangle \langle \bar{\psi}_a | \bar{H}^b | \psi_b \rangle \\ &- \langle \psi_b | \hat{\Lambda}^{ab} e^{-\hat{T}^b} e^{\hat{T}^a} | \psi_a \rangle \langle \bar{\psi}_a | (\partial_{\rho} \bar{H}^b) | \psi_b \rangle = 0. \end{aligned} \quad (33)$$

The second term on the right-hand side vanishes due to eqn (24). This leaves us with the following linear set of equations, from which  $\hat{\Lambda}^{ab}$  can be determined:

$$\begin{aligned} &\langle \psi_b | \hat{\Lambda}^{ab} (\partial_{\rho} \bar{H}^b) | \psi_b \rangle - \langle \psi_b | \hat{\Lambda}^{ab} e^{-\hat{T}^b} e^{\hat{T}^a} | \psi_a \rangle \langle \bar{\psi}_a | (\partial_{\rho} \bar{H}^b) | \psi_b \rangle \\ &= -\langle \bar{\psi}_a | (\partial_{\rho} \bar{H}^b) | \psi_b \rangle. \end{aligned} \quad (34)$$

Note that these equations are independent of the specific perturbing operator. The desired coupling matrix elements for any perturbing operator  $\hat{V}$  are given by

$$\begin{aligned} \partial_{\varepsilon} L_{ab}|_{\varepsilon=0} &= \langle \bar{\psi}_a | \bar{V}^b | \psi_b \rangle + \langle \psi_b | \hat{\Lambda}^{ab} \bar{V}^b | \psi_b \rangle \\ &- \langle \psi_b | \hat{\Lambda}^{ab} e^{-\hat{T}^b} e^{\hat{T}^a} | \psi_a \rangle \langle \bar{\psi}_a | \bar{V}^b | \psi_b \rangle. \end{aligned} \quad (35)$$

## 2.5 Spin-orbit coupling matrix elements

As specific perturbation, we consider here the spin-orbit operator. A simple, yet accurate approximation is the mean-field spin-orbit operator<sup>8,41,42</sup>

$$\hat{H}_{\text{SOMF}} = \sum_{\xi} \sum_{pq} \left( f_{\text{SO}}^{(\xi)} \right)_p^q \hat{D}_q^{\rho, \xi} \quad (36)$$

where the summation is restricted to spatial orbitals and we are assuming restricted spin orbitals with equal spatial orbitals for either spin case. This formulation reduces the spin-orbit interaction to a

one-particle operator, while the two-particle contributions are treated as a screening interaction by the mean field of the remaining electrons. Eqn (36) contains the spin excitation operators<sup>8</sup>

$$\hat{D}_q^{\rho, x} = \frac{1}{2} (\hat{a}_q^{\rho} + \hat{a}_q^{\bar{\rho}}), \quad (37)$$

$$\hat{D}_q^{\rho, y} = \frac{1}{2i} (\hat{a}_q^{\rho} - \hat{a}_q^{\bar{\rho}}), \quad (38)$$

$$\hat{D}_q^{\rho, z} = \frac{1}{2} (\hat{a}_q^{\rho} - \hat{a}_q^{\bar{\rho}}). \quad (39)$$

The operator has triplet spin symmetry, which implies that the required multipliers  $\hat{\Lambda}^{ab}$  are also triplet operators.

The  $\langle \bar{\psi}_a |$  states in eqn (35) are formally the left-hand solutions of the effective Hamiltonian, which is a bit inconvenient in general as the left-hand side eigenvalue problem involves similarity transformed matrix elements  $e^{-\hat{T}^i} \hat{H} e^{\hat{T}^i}$  of all model space states. Within the present work, we have chosen the *ad hoc* provision to use the coefficients  $\tilde{c}_{\mu}$  from the state-specific *A* equations, eqn (12). The expression for the spin-orbit interaction matrix elements reads therefore

$$H_{ab}^{\text{SO}} = \langle \bar{\psi}_a | \bar{H}_{\text{SOMF}}^b | \psi_b \rangle + \langle \psi_b | \hat{\Lambda}^{ab} \bar{H}_{\text{SOMF}}^b | \psi_b \rangle \quad (40)$$

$$\begin{aligned} &- \langle \psi_b | \hat{\Lambda}^{ab} e^{-\hat{T}^b} e^{\hat{T}^a} | \psi_a \rangle \langle \bar{\psi}_a | \bar{H}_{\text{SOMF}}^b | \psi_b \rangle \\ &= H_{ab}^{\text{SO, eff}} + H_{ab}^{\text{SO, A}} - S_{ab}^A H_{ab}^{\text{SO, eff}} \end{aligned} \quad (41)$$

The matrix elements can be partitioned into contributions that are very similar to those found by Mück and Gauss for the Mk-MRCC theory.<sup>25</sup>

## 3 Results and discussion

### 3.1 Implementation and computational setup

Integrals and CASSCF orbitals were imported from the Molpro program package.<sup>43–47</sup> CASSCF optimizations always included state-averaging over all relevant orbitals, in particular avoiding symmetry breaking in atomic P or molecular  $\Pi$  states. The spin-orbit mean-field integrals were based on the Breit–Pauli operator,<sup>8</sup> where the two-electron contributions were contracted with the state-averaged CASSCF density.

Based on these integrals, the GeCCo code<sup>48</sup> was employed to carry out state-specific icMRCCSD computations for the relevant states, including reference relaxation. The resulting coefficients  $c_{\mu}$  and amplitudes  $t_{\rho}$  were stored for the subsequent computations. For each state, also the state-specific *A* equations were solved<sup>39</sup> and the  $\tilde{c}_{\mu}$  coefficients were stored, too. All coefficients and amplitudes are spin-adapted, but are stored in their explicit spin-orbital representation for one specific  $M_S$  state. Usually the smallest  $M_S$  (0 or 1/2) is used. For the computation of the coupling between triplet states the  $M_S = 1$  components had to be used, as the SO coupling matrix element between the  $M_S = 0$  components is vanishing.

The spin-orbit computations were implemented by a plug-in code for GeCCo, which sets up the relevant equations in



symbolic form. Details of this are given in the ESI† which also contains sample inputs and program outputs that illustrate the procedure.

The code in particular autogenerates all the required equations. It starts by generating the contributions to the Lagrange expression, eqn (30). In accordance with the general truncation rules employed for icMRCCSD,<sup>31</sup> the energy terms retain the cluster operator terms up to quartic terms and the amplitude equation terms up to quadratic order. The generated terms are then further processed symbolically: by symbolic derivation, the linear set of equations to determine  $\Lambda_{ab}$  is generated, see eqn (34), and by replacing the (unperturbed) Hamiltonian by the spin-orbit operator, the final expressions for the transition matrix elements are obtained, see eqn (35) and (40), respectively. These equations are then translated into a sequence of binary tensor contractions, which are evaluated numerically.

The computational complexity of icMRCCSD formally scales as that of comparable multireference CI methods, that is,  $O(N^6)$  with system size  $N$  for fixed active space size ( $O(M^4)$  with basis set size  $M$ ), and factorially with the number of active orbitals. In comparison to multireference CI computations, the prefactor for icMRCCSD is large for the present non-optimal implementation.<sup>49</sup> The non-linear amplitude and reference coefficient equations, eqn (3) and (4), can sometimes be hard to converge. For the computation of spin-orbit matrix elements, in addition the  $\Lambda$  equations, eqn (12) for each state and the Lagrange multiplier equations, eqn (34), need to be solved. The computational complexity of these equations is the same as for the amplitude equations, but they are usually easier to converge due to their linear nature.

The test computations employed the cc-pCVXZ ( $X = D, T, Q$ ) basis sets.<sup>50,51</sup> Both calculations correlating only valence electrons and calculations including the next-lower electronic shell were run. For comparison, additional computations were carried out at the CASCI and MRCI level, using the Molpro code.<sup>8,43,52</sup>

### 3.2 Atomic zero-field splitting

As a first set of test cases, we consider the spin-orbit coupling in the lowest term of the atoms Al, Si, S, and Cl. The results are summarized in Table 1. As reference, the table reports the effective spin-orbit coupling constant computed from the experimental fine structure splittings given in ref. 53. In Table 1, we particularly list the different contributions that add up to the final MS-icMRCCSD coupling matrix element, see also eqn (41). These results show that the main contribution comes from the effective Hamiltonian term but that the effects of the  $\Lambda$  and overlap terms are significant and improve the result towards the experimental reference value. The improvement of correlated calculations like icMRCCSD and MRCI with respect to CASCI is particularly obvious for the core-correlation effects, which naturally require an extended correlation space. For the icMRCCSD computations, core-correlation effects particularly enter through the effective Hamiltonian term and the  $\Lambda$  term, whereas the overlap term does not change much. Within the very similar expression derived for the Mk-MRCCSD theory, the overlap term

**Table 1** Computed atomic spin-orbit coupling constants ( $\text{cm}^{-1}$ ) of the lowest term of third-row atoms in comparison to experimental references. Negative numbers indicate that the smallest  $J$  quantum number is the lowest state

| System/<br>settings <sup>b</sup> | MS-icMRCC                       |                                     |                                                         | Total | CASCI | MRCI  | Exp. <sup>a</sup> |
|----------------------------------|---------------------------------|-------------------------------------|---------------------------------------------------------|-------|-------|-------|-------------------|
|                                  | $H_{\text{ab}}^{\text{SO,eff}}$ | $H_{\text{ab}}^{\text{SO},\Lambda}$ | $-S_{\text{ab}}^{\Lambda}H_{\text{ab}}^{\text{SO,eff}}$ |       |       |       |                   |
| Al <sup>2</sup> P                |                                 |                                     |                                                         |       |       |       |                   |
| fc cc-pCVDZ                      | −31.8                           | 1.2                                 | 1.4                                                     | −29.1 | −34.0 | −29.7 |                   |
| fc cc-pCVTZ                      | −32.3                           | 1.1                                 | 1.6                                                     | −29.6 | −34.5 | −30.2 |                   |
| fc cc-pCVQZ                      | −32.5                           | 1.1                                 | 1.7                                                     | −29.7 | −34.7 | −30.4 |                   |
| cc cc-pCVDZ                      | −34.5                           | −1.4                                | 1.5                                                     | −34.4 | −34.0 | −37.9 |                   |
| cc cc-pCVTZ                      | −35.8                           | −2.1                                | 1.7                                                     | −36.2 | −34.5 | −39.8 |                   |
| cc cc-pCVQZ                      | −35.9                           | −2.3                                | 1.7                                                     | −36.5 | −34.7 | −40.4 | −37.35            |
| Si <sup>3</sup> P                |                                 |                                     |                                                         |       |       |       |                   |
| fc cc-pCVDZ                      | −62.5                           | 0.6                                 | 1.4                                                     | −60.5 | −63.7 | −61.4 |                   |
| fc cc-pCVTZ                      | −63.7                           | 0.5                                 | 1.6                                                     | −61.5 | −64.6 | −62.5 |                   |
| fc cc-pCVQZ                      | −63.8                           | 0.5                                 | 1.6                                                     | −61.7 | −64.7 | −62.6 |                   |
| cc cc-pCVDZ                      | −67.1                           | −4.0                                | 1.4                                                     | −69.7 | −63.7 | −73.6 |                   |
| cc cc-pCVTZ                      | −69.0                           | −5.1                                | 1.6                                                     | −72.5 | −64.6 | −76.2 |                   |
| cc cc-pCVQZ                      | −69.4                           | −5.3                                | 1.6                                                     | −73.2 | −64.7 | −76.8 | −73.7             |
| S <sup>3</sup> P                 |                                 |                                     |                                                         |       |       |       |                   |
| fc cc-pCVDZ                      | 178.3                           | −3.6                                | −4.1                                                    | 170.7 | 181.5 | 173.0 |                   |
| fc cc-pCVTZ                      | 180.4                           | −3.6                                | −5.8                                                    | 170.9 | 182.9 | 173.9 |                   |
| fc cc-pCVQZ                      | 180.7                           | −3.7                                | −6.1                                                    | 171.0 | 183.3 | 174.2 |                   |
| cc cc-pCVDZ                      | 186.8                           | 5.0                                 | −4.2                                                    | 187.7 | 181.5 | 198.2 |                   |
| cc cc-pCVTZ                      | 189.9                           | 6.9                                 | −6.0                                                    | 190.8 | 182.9 | 202.7 |                   |
| cc cc-pCVQZ                      | 190.8                           | 7.4                                 | −6.2                                                    | 192.0 | 183.3 | 204.3 | 194.61            |
| Cl <sup>2</sup> P                |                                 |                                     |                                                         |       |       |       |                   |
| fc cc-pCVDZ                      | 272.3                           | −2.6                                | −5.3                                                    | 264.3 | 272.4 | 265.4 |                   |
| fc cc-pCVTZ                      | 274.5                           | −2.9                                | −7.7                                                    | 263.8 | 273.7 | 265.9 |                   |
| fc cc-pCVQZ                      | 275.2                           | −2.9                                | −8.1                                                    | 264.3 | 274.4 | 266.6 |                   |
| cc cc-pCVDZ                      | 283.5                           | 9.0                                 | −5.5                                                    | 287.1 | 272.4 | 288.6 |                   |
| cc cc-pCVTZ                      | 287.0                           | 11.0                                | −7.9                                                    | 290.1 | 273.7 | 292.7 |                   |
| cc cc-pCVQZ                      | 288.3                           | 11.8                                | −8.2                                                    | 291.8 | 274.4 | 294.5 | 294.12            |

<sup>a</sup> Computed from energy splittings given in ref. 53. <sup>b</sup> System: atom and term as indicated; settings: 'fc' = frozen core, 'cc' = core-correlation, and basis sets used.

was also interpreted as a dimensionless screening factor for the effective Hamiltonian matrix element:<sup>25</sup>

$$H_{ab}^{\text{SO,eff}} - S_{ab}^{\Lambda}H_{ab}^{\text{SO,eff}} = \xi H_{ab}^{\text{SO,eff}} \quad (42)$$

The values for  $\xi = (1 - S_{ab}^{\Lambda})$  are all in the range of 0.95–0.98 for the atomic examples.

Basis set effects are not strongly pronounced for the considered cases, the differences between the cc-pCVDZ and cc-pCVQZ results are at most 5 per cent (for silicon). Concerning a comparison of the MRCI and the MS-icMRCCSD results, the two methods appear to be en par for the considered cases. There is a slight tendency for MRCI to overshoot the experimental references for Al, Si, and S by 3 to 10  $\text{cm}^{-1}$  but this is partially an effect of using the mean-field approach for the spin-orbit operator. In fact, using the full two-electron operator improves the MRCI results<sup>8</sup> (see ESI† for detailed results). Effects of similar size have previously been observed for spin-orbit coupling matrix elements in transition metal compounds.<sup>54</sup> Our present MS-icMRCCSD implementation, however, is restricted to effective one-electron operators, therefore we could not investigate this effect further for the coupled-cluster case.



### 3.3 Spin-orbit splitting in molecular $^2\Pi$ states

The spin-orbit splittings of  $^2\Pi$  radicals have been investigated in a number of previous works<sup>8,23,25</sup> as test cases with well-established reference values. In Table 2 we report results using large cc-pCVQZ basis sets and correlated core electrons. We consider both runs with minimal CAS and full-valence CAS. More detailed results on the different contributions to the MS-icMRCCSD values can be found in the ESI.† The MS-icMRCCSD results are compared to values obtained from CASCI and MRCI calculations, and literature results for Mk-MRCCSD<sup>25</sup> (minimal CAS) and EOMIP-CCSD.<sup>24</sup>

For the homologous series OH, SH, and SeH, the spin-orbit splitting increases strongly, as expected, where SeH is certainly a borderline case for the application of the Breit–Pauli operator. For the lightest case, OH, all methods and choices of active space agree within 3 cm<sup>−1</sup> and are also close to the experimental reference. For SH, core-correlation effects become more important and CASCI clearly underestimates the coupling. As for the case of OH, the spin-orbit couplings of SH computed by MS-icMRCCSD and Mk-MRCCSD are very similar, and rather close to the experimental reference result (deviation of approximately 2 cm<sup>−1</sup>). Larger deviations from experiment (up to 90 cm<sup>−1</sup>) are found for SeH. Part of the deviations may come from the use of the Breit–Pauli operator, EOMIP-CCSD computations reported by Cheng *et al.*<sup>58</sup> also report a value of −1673 cm<sup>−1</sup> for the Breit–Pauli operator, while employing the spin-orbit operator from exact two-component theory (X2C) gives −1717 cm<sup>−1</sup>. In addition, there is also an impact of the correlation of additional inner core shells.<sup>58</sup> In the present work, we only considered the correlation of the 3s3p3d semi-core of Se. For instance, the Mk-MRCCSD computations of ref. 25 correlate all electrons, leading to the slightly larger value of −1708 cm<sup>−1</sup> quoted in Table 2.

**Table 2** Computed<sup>a</sup> spin-orbit splittings (cm<sup>−1</sup>) of the  $^2\Pi$  ground states of selected molecules. A negative value indicates that the  $\Omega = 1/2$  component is the lowest state

| Method                  | CAS <sup>b</sup>       | OH     | SH     | SeH     | ClO                 | SN                 |
|-------------------------|------------------------|--------|--------|---------|---------------------|--------------------|
| MS-icMRCCSD             | Min. <sup>b</sup>      | −135.8 | −375.7 | −1672.3 | −308.1              | 230.1              |
|                         | Full val. <sup>c</sup> | −135.5 | −374.9 | n/a     | n/a                 | 221.5              |
| CASCI                   | Min. <sup>b</sup>      | −138.8 | −351.6 | −1617.7 | −202.2              | 195.0              |
|                         | Full val. <sup>c</sup> | −138.8 | −350.5 | −1612.0 | −252.0 <sup>d</sup> | 213.1              |
| MRCI                    | Min. <sup>b</sup>      | −136.8 | −380.9 | −1709.0 | −225.8              | 218.0              |
|                         | Full val. <sup>c</sup> | −136.7 | −380.3 | −1704.3 | −276.2 <sup>d</sup> | 226.2              |
| Mk-MRCCSD <sup>e</sup>  | Min. <sup>b</sup>      | −135.1 | −375.2 | −1707.9 | −312.3              | n/a                |
| EOMIP-CCSD <sup>f</sup> |                        | −136.7 | −371.6 | −1680.4 | −308.3              | 242.9 <sup>g</sup> |
| Exp. <sup>h</sup>       |                        | −139.2 | −377.0 | −1764.4 | −320.3              | 222.9              |

<sup>a</sup> All computations employ the cc-pCVQZ basis set, core electrons were correlated (except CASCI), 1s cores of S and Cl and 1s2s2p core of Se was kept frozen. Used equilibrium distances – OH: 0.9697 Å, SH: 1.3409 Å, SeH: 1.5811 Å, ClO: 1.5696 Å, SN: 1.4940 Å; the same structures were used in ref. 8 and 25. <sup>b</sup> Minimal active space: (3,2) for OH, SH, SeH, and ClO; (1,2) for SN. <sup>c</sup> Full valence active space: (7,5) for OH, SH, SeH; (13,8) for ClO; (11,8) for SN. <sup>d</sup> Extending the active space by a 4π orbital improves CASCI to −301.7 cm<sup>−1</sup> and MRCI to −317.4 cm<sup>−1</sup>, see ref. 8. <sup>e</sup> Values taken from ref. 25. <sup>f</sup> Values taken from ref. 24. <sup>g</sup> A better value is obtained by EOMEA-CCSD: 226.8 cm<sup>−1</sup>. <sup>h</sup> Values taken from ref. 55 (OH, SH, SN), ref. 56 (SeH), and ref. 57 (ClO).

For the three hydrides, so far no significant differences between employing a minimal CAS and a full valence CAS was found. This is clearly different for the case of ClO. Here, the CASCI and MRCI results are clearly different for the two choices of active space, and yet both are not sufficient for a quantitative description. The MRCI value for the full-valence CAS is −276 cm<sup>−1</sup>, which is still more than 40 cm<sup>−1</sup> off the experimental result of −320 cm<sup>−1</sup>. As shown by Berning *et al.*,<sup>8</sup> the MRCI description can be improved by adding a set of 4π orbitals to the active space, resulting in nearly quantitative agreement with experiment.

In coupled-cluster theory, already the minimal active space gives a comparatively good result. The values for MS-icMRCCSD, Mk-MRCCSD and EOMIP-CCSD range between −308 and −312 cm<sup>−1</sup>. Unfortunately, the MS-icMRCCSD computation with a large active space failed due to memory problems and we cannot report a value here.

The results for the last case, SN, however, appear promising for MS-icMRCCSD theory. Here, we also find significant effects of going beyond the minimal active space and the MS-icMRCCSD result is very close to the experimental reference.

### 3.4 Non-Hermiticity of coupled-cluster theory

All the previous examples have in common that they concern spin-orbit coupling between equivalent, degenerate states. Therefore, the pair of adjoint matrix elements fulfils the required symmetry condition  $H_{ab}^{SO} = H_{ba}^{SO*}$ . However, the non-Hermiticity of coupled-cluster theory does not guarantee this in general.

Here, we investigate the spin-orbit coupling matrix elements between two non-equivalent states of the NH radical, its  $X^3\Sigma^-$  ground state and the  $c^1\Pi$  excited state. The molecule has been recently studied in detail due to its relevance for astrochemistry, particularly its photodissociation cross sections in the UV range were investigated, with special focus on the  $c^1\Pi$  state.<sup>59</sup>

With this example, we also want to demonstrate the performance of MS-icMRCCSD in the bond-breaking regime, where multireference effects become strong. We therefore have covered distances between 0.5 Å (repulsive tail) and 3.2 Å (dissociative tail). Of course, the full treatment of the system requires much more states to be investigated (ref. 59 considers 12 low-lying electronic states), which is beyond our present scope; we will focus on the coupling of this single pair of states.

Table 3 summarizes the results, further details are given in the ESI.† The two states couple rather strongly at short distances and also at their equilibrium distances at approximately 1.0 to 1.1 Å. The excited  $^1\Pi$  state has a broad avoided crossing at around 1.8 to 2.0 Å (see ESI† for the energy curves), and for longer distances, the spin-orbit coupling drops rapidly towards zero.

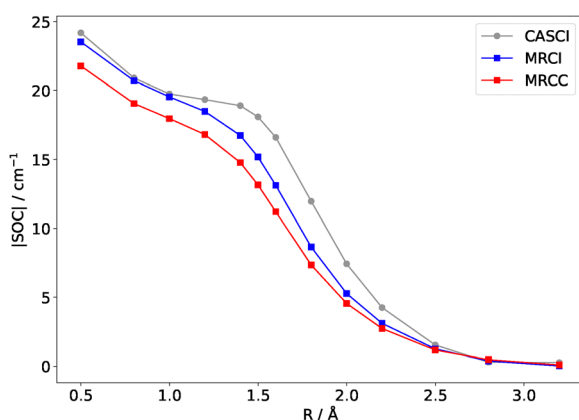
The run of the coupling is also shown in Fig. 1. MRCI and icMRCCSD are in good agreement, the CASCI curves behaves slightly differently at intermediate distances, most likely due to a different position of the avoided crossing in the absence of dynamical correlation. This again demonstrates the utility of spin-orbit coupling matrix elements from correlated computations that include state interactions (which is ensured for icMRCCSD with reference relaxation<sup>33</sup>).



**Table 3** Computed<sup>a</sup> spin–orbit coupling matrix elements ( $\text{cm}^{-1}$ ) between the states  $X^3\Sigma^-$  (state 'a') and the  $c^1\Pi$  (state 'b'). We show the matrix element of the  $M_s = 1$  component of  $^3\Sigma^-$  with the x component of the  $^1\Pi$  state which couples via the y component of the spin–orbit operator and gives real values. We fixed the global phase factor to positive values for the matrix elements

| <i>R</i> | CASCI                                     | MRCI                                      | MS-icMRCC using $\langle\psi $ |                      | MS-icMRCC using $\langle\psi $ |                      |                                                |
|----------|-------------------------------------------|-------------------------------------------|--------------------------------|----------------------|--------------------------------|----------------------|------------------------------------------------|
|          | $H_{ab}^{\text{SO}} = H_{ba}^{\text{SO}}$ | $H_{ab}^{\text{SO}} = H_{ba}^{\text{SO}}$ | $H_{ab}^{\text{SO}}$           | $H_{ba}^{\text{SO}}$ | $H_{ab}^{\text{SO}}$           | $H_{ba}^{\text{SO}}$ | $\sqrt{H_{ab}^{\text{SO}} H_{ba}^{\text{SO}}}$ |
| 0.5      | 24.18                                     | 23.52                                     | 21.72                          | 21.82                | 21.71                          | 21.84                | 21.78                                          |
| 0.8      | 20.91                                     | 20.70                                     | 19.05                          | 19.14                | 19.02                          | 19.10                | 19.06                                          |
| 1.0      | 19.75                                     | 19.53                                     | 18.00                          | 18.03                | 17.92                          | 18.01                | 17.96                                          |
| 1.2      | 19.34                                     | 18.48                                     | 16.81                          | 16.68                | 16.72                          | 16.90                | 16.81                                          |
| 1.4      | 18.90                                     | 16.74                                     | 14.69                          | 14.63                | 14.65                          | 14.89                | 14.77                                          |
| 1.5      | 18.09                                     | 15.17                                     | 13.06                          | 12.98                | 13.03                          | 13.25                | 13.14                                          |
| 1.6      | 16.60                                     | 13.12                                     | 11.12                          | 11.07                | 11.11                          | 11.29                | 11.20                                          |
| 1.8      | 11.96                                     | 8.65                                      | 7.30                           | 7.37                 | 7.29                           | 7.40                 | 7.35                                           |
| 2.0      | 7.43                                      | 5.27                                      | 4.51                           | 4.75                 | 4.51                           | 4.59                 | 4.55                                           |
| 2.2      | 4.25                                      | 3.10                                      | 2.71                           | 3.17                 | 2.70                           | 2.76                 | 2.73                                           |
| 2.5      | 1.53                                      | 1.25                                      | 1.14                           | 2.36                 | 1.14                           | 1.19                 | 1.17                                           |
| 2.8      | 0.30                                      | 0.38                                      | 0.45                           | 2.66                 | 0.45                           | 0.50                 | 0.47                                           |
| 3.2      | 0.26                                      | 0.02                                      | 0.09                           | 0.01                 | 0.08                           | 0.11                 | 0.09                                           |

<sup>a</sup> All calculations use the cc-pCVDZ basis set, the 1s core of nitrogen was kept frozen.



**Fig. 1** Run of the computed spin–orbit coupling matrix elements between the  $X^3\Sigma^-$  and the  $c^1\Pi$  state of NH. The computations use the cc-pCVDZ basis set, all electrons were correlated. For MRCC, the averaged value is plotted, see text.

With respect to the matrix elements from MS-icMRCCSD we make one observation that is documented in Table 3 and which calls for further development of the theory concerning the proper Lagrange multipliers for the reference coefficients. In eqn (40) we have proposed to use the  $\tilde{c}_\mu$  coefficient from the  $\Lambda$  equations from the individual states. Using this definition, we get the matrix elements listed in Table 3 under 'MS-icMRCC using  $\langle\psi|$ '. For  $R < 1.8$  Å this leads to very similar matrix elements  $H_{ab}^{\text{SO}}$  and  $H_{ba}^{\text{SO}}$  with deviations of less than  $0.3 \text{ cm}^{-1}$ . For larger distances, however, a problem appears; while the  $H_{ab}^{\text{SO}}$  matrix elements (which use the cluster amplitudes from the  $^1\Pi$  state and the  $\tilde{c}_\mu$  coefficients from the  $^3\Sigma^-$  state) still follow the trend of the MRCI computations, the  $H_{ba}^{\text{SO}}$  matrix elements do not. The latter are computed from the cluster amplitudes of the

$^3\Sigma^-$  state and the  $\tilde{c}_\mu$  coefficients of the  $^1\Pi$  state. A close analysis of the contributions (see also ESI†) shows that the latter are the main cause of the deviations. Possibly, this effect is connected to the avoided state crossing, which also leads to slow convergence of the icMRCCSD amplitude and coefficient equations. The convergence of the (linear)  $\Lambda$  equations is less problematic and inspection of the amplitudes does not indicate a major numerical problem. Nevertheless, the  $\langle\psi_b|\hat{H}_{\text{SOMF}}|\psi_a\rangle$  CASCI-like contribution is clearly the main term that is responsible for the non-physical result.

If we replace the  $\tilde{c}_\mu$  amplitudes by the icMRCC reference state amplitudes  $c_\mu$ , we get the results listed in Table 3 under 'MS-icMRCC using  $\langle\psi|$ '. In fact, the alternative  $H_{ab}^{\text{SO}}$  matrix elements deviate only marginally from the initial formulation and so do the  $H_{ba}^{\text{SO}}$  for short distances. In the critical region beyond  $1.8$  Å, however, the alternative formulation avoids the artefacts described above, and the  $H_{ab}^{\text{SO}}$  and  $H_{ba}^{\text{SO}}$  matrix elements agree within  $0.3 \text{ cm}^{-1}$  for all tested bond distances. In case of a pure two-state interaction, the resulting energy splitting is proportional to  $\sqrt{H_{ab}^{\text{SO}} H_{ba}^{\text{SO}}}$  and we use this procedure to compute an 'averaged' coupling matrix element. These are also the values used for Fig. 1.

## Conclusions

We have presented the theory of interstate coupling matrix elements within a multistate framework for internally contracted multireference coupled-cluster (MS-icMRCC) theory. The equations were implemented in a pilot code based on our specialised symbolic algebra program GeCCo and tested for the computation of spin–orbit coupling matrix elements, using a mean-field Breit–Pauli operator. Apart from computing the state-specific icMRCC wavefunctions for the individual states, the formalism requires solving an additional linear system of equations for each coupled pair. The theory is very similar to that presented by Mück and Gauss<sup>25</sup> for state-specific Mk-MRCCSD theory, but can be more easily applied to larger active spaces.

Demonstrative results were obtained for the zero-field splitting of the lowest terms ( $^2P$  or  $^3P$ ) of the atoms Al, Si, S, and Cl, as well as for the  $^2\Pi$  terms of OH, SH, SeH, ClO, and SN. Active spaces of different sizes (minimal and full valence) were used, and the impact of core-correlation contributions was investigated; these are particularly significant for the heavier atoms. The MS-icMRCCSD results are very accurate and match those of multireference configuration interaction (MRCI). The advantage of the MS-icMRCCSD approach with respect to MRCI is the extensivity of the theory, which holds the promise to keep the same accuracy also for large systems. The currently investigated systems, however, are clearly too small to demonstrate this.

We have also investigated the impact of the non-Hermiticity of coupled-cluster theory on the spin–orbit coupling matrix elements between non-equivalent states. To this end, we investigated the coupling between the  $^3\Sigma^-$  ground state and the  $^1\Pi$  excited state of NH for several interatomic distances. The



impact of non-Hermiticity was not very large for this case, the differences between the adjoint matrix elements are less than  $0.3\text{ cm}^{-1}$  for matrix elements of an average size of  $10\text{ cm}^{-1}$ . It was found, however, that the treatment of the response of the reference coefficient has to be carefully reviewed in order to avoid artefacts.

## Author contributions

The research reported in this work was conceived and carried out by A. K., based on an initial investigation carried out by J. N. under the supervision of A. K.

## Conflicts of interest

There are no conflicts to declare.

## Data availability

The data supporting this article have been included as part of the ESI.†

## Acknowledgements

The authors dedicate this work to Professor Christel Marian on the happy occasion of her 70th birthday and in recognition of her pioneering contributions to the quantum chemical treatment of spin-orbit interactions. A. K. also acknowledges the support by the Stuttgart Center for Simulation Science (SimTech).

## Notes and references

- 1 C. M. Marian, *Wiley Interdiscip. Rev.:Comput. Mol. Sci.*, 2011, **2**, 187–203.
- 2 T. J. Penfold, E. Gindensperger, C. Daniel and C. M. Marian, *Chem. Rev.*, 2018, **118**, 6975–7025.
- 3 M. Feng and M.-L. Tong, *Chem. – Eur. J.*, 2018, **24**, 7574–7594.
- 4 S. T. Liddle and J. V. Slagere, *Chem. Soc. Rev.*, 2015, **44**, 6655–6669.
- 5 E. Coronado, *Nat. Rev. Mater.*, 2020, **5**, 87–104.
- 6 B. A. Hess, R. J. Buenker, C. M. Marian and S. D. Peyerimhoff, *Chem. Phys.*, 1982, **71**, 79–85.
- 7 D. R. Yarkony, *Int. Rev. Phys. Chem.*, 1992, **11**, 195–242.
- 8 A. Berning, M. Schweizer, H.-J. Werner, P. J. Knowles and P. Palmieri, *Mol. Phys.*, 2000, **98**, 1823–1833.
- 9 P. k Malmqvist, B. O. Roos and B. Schimmelpfennig, *Chem. Phys. Lett.*, 2002, **357**, 230–240.
- 10 D. Ganyushin and F. Neese, *J. Chem. Phys.*, 2013, **138**, 104113.
- 11 T. Bodenstein, A. Heimermann, K. Fink and C. Wüllen, *ChemPhysChem*, 2022, **23**, e202100648.
- 12 A. Almoukhalalati, S. Knecht, H. J. A. Jensen, K. G. Dyall and T. Saue, *J. Chem. Phys.*, 2016, **145**, 074104.
- 13 V. Jovanović, I. Lyskov, M. Kleinschmidt and C. M. Marian, *Mol. Phys.*, 2017, **115**, 109–137.
- 14 L. Lang and F. Neese, *J. Chem. Phys.*, 2019, **150**, 104104.
- 15 R. Majumder and A. Y. Sokolov, *J. Phys. Chem. A*, 2023, **127**, 546–559.
- 16 R. Majumder and A. Y. Sokolov, *J. Chem. Theory Comput.*, 2024, **20**, 4676–4688.
- 17 Z. Zhao and F. A. Evangelista, *J. Phys. Chem. Lett.*, 2024, **15**, 7103–7110.
- 18 S. Schmitt, P. Jost and C. V. Wüllen, *J. Chem. Phys.*, 2011, **134**, 194113.
- 19 F. Dinkelbach, M. Kleinschmidt and C. M. Marian, *J. Chem. Theory Comput.*, 2017, **13**, 749–766.
- 20 R. J. Bartlett and M. Musiał, *Rev. Mod. Phys.*, 2007, **79**, 291–352.
- 21 D. I. Lyakh, M. Musiał, V. F. Lotrich and R. J. Bartlett, *Chem. Rev.*, 2012, **112**, 182–243.
- 22 A. Köhn, M. Hanauer, L. A. Mück, T.-C. Jagau and J. Gauss, *Wiley Interdiscip. Rev.:Comput. Mol. Sci.*, 2013, **3**, 176–197.
- 23 K. Klein and J. Gauss, *J. Chem. Phys.*, 2008, **129**, 194106.
- 24 E. Epifanovsky, K. Klein, S. Stopkiewicz, J. Gauss and A. I. Krylov, *J. Chem. Phys.*, 2015, **143**, 064102.
- 25 L. A. Mück and J. Gauss, *J. Chem. Phys.*, 2012, **136**, 111103.
- 26 U. S. Mahapatra, B. Datta and D. Mukherjee, *J. Chem. Phys.*, 1999, **110**, 6171–6188.
- 27 F. A. Evangelista, W. D. Allen and H. F. Schaefer, *J. Chem. Phys.*, 2006, **125**, 154113.
- 28 A. Banerjee and J. Simons, *Int. J. Quantum Chem.*, 1981, **19**, 207–216.
- 29 A. Banerjee and J. Simons, *J. Chem. Phys.*, 1982, **76**, 4548–4559.
- 30 F. A. Evangelista and J. Gauss, *J. Chem. Phys.*, 2011, **134**, 114102.
- 31 M. Hanauer and A. Köhn, *J. Chem. Phys.*, 2011, **134**, 204111.
- 32 M. Hanauer and A. Köhn, *J. Chem. Phys.*, 2012, **136**, 204107.
- 33 Y. A. Aoto and A. Köhn, *J. Chem. Phys.*, 2016, **144**, 074103.
- 34 A. Köhn, J. A. Black, Y. A. Aoto and M. Hanauer, *Mol. Phys.*, 2020, **118**, e1743889.
- 35 R. G. Adam, A. Waigum and A. Köhn, *Wiley Interdiscip. Rev.:Comput. Mol. Sci.*, 2025, **15**, e70023.
- 36 M. M. F. de Moraes and Y. A. Aoto, *Theor. Chem. Acc.*, 2020, **139**, 71.
- 37 M. M. F. de Moraes and Y. A. Aoto, *J. Mol. Spectrosc.*, 2022, **385**, 111611.
- 38 M. Hanauer and A. Köhn, *J. Chem. Phys.*, 2012, **137**, 131103.
- 39 P. K. Samanta and A. Köhn, *J. Chem. Phys.*, 2018, **149**, 064101.
- 40 B. Jeziorski and H. J. Monkhorst, *Phys. Rev. A:At., Mol., Opt. Phys.*, 1981, **24**, 1668–1681.
- 41 B. Hess, C. M. Marian, U. Wahlgren and O. Gropen, *Chem. Phys. Lett.*, 1996, **251**, 365–371.
- 42 F. Neese, *J. Chem. Phys.*, 2005, **122**, 034107.
- 43 H.-J. Werner, P. J. Knowles, P. Celani, W. Györfy, A. Hesselmann, D. Kats, G. Knizia, A. Köhn, T. Korona, D. Kreplin, R. Lindh, Q. Ma, F. R. Manby, A. Mitrushenkov, G. Rauhut, M. Schütz, K. R. Shamasundar, T. B. Adler, R. D. Amos, S. J. Bennie, A. Bernhardsson, A. Berning, J. A. Black, P. J. Bygrave, R. Cimiraglia, D. L. Cooper, D. Coughtrie, M. J. O. Deegan, A. J. Dobbyn, K. Doll, M. Dornbach, F. Eckert, S. Erfort, E. Goll, C. Hampel, G. Hetzer, J. G. Hill, M. Hodges, T. Hrenar, G. Jansen, C. Köppl, C. Kollmar, S. J. R. Lee, Y. Liu, A. W. Lloyd, R. A. Mata, A. J. May, B. Mussard, S. J. McNicholas, W. Meyer,



- T. F. Miller III, M. E. Mura, A. Nicklass, D. P. O'Neill, P. Palmieri, D. Peng, K. A. Peterson, K. Pflüger, R. Pitzer, I. Polyak, M. Reiher, J. O. Richardson, J. B. Robinson, B. Schröder, M. Schwilk, T. Shiozaki, M. Sibaev, H. Stoll, A. J. Stone, R. Tarroni, T. Thorsteinsson, J. Toulouse, M. Wang, M. Welborn and B. Ziegler, *MOLPRO, version 2024.1, a package of ab initio programs*, see <https://www.molpro.net>.
- 44 H.-J. Werner, P. J. Knowles, F. R. Manby, J. A. Black, K. Doll, A. Hesselmann, D. Kats, A. Köhn, T. Korona, D. A. Kreplin, Q. Ma, T. F. Miller, A. Mitrushchenkov, K. A. Peterson, I. Polyak, G. Rauhut and M. Sibaev, *J. Chem. Phys.*, 2020, **152**, 144107.
  - 45 P. J. Knowles and H.-J. Werner, *Chem. Phys. Lett.*, 1985, **115**, 259–267.
  - 46 H.-J. Werner and P. J. Knowles, *J. Chem. Phys.*, 1985, **82**, 5053–5063.
  - 47 D. A. Kreplin, P. J. Knowles and H.-J. Werner, *J. Chem. Phys.*, 2019, **150**, 194106.
  - 48 A. Köhn, R. G. Adam, Y. A. Aoto, A. Bargholz, J. Black, M. Hanauer, P. Samanta *et al.* GitHub, see <https://github.com/ak-ustutt/GeCCo-public>.
  - 49 J. A. Black, A. Waigum, R. G. Adam, K. R. Shamasundar and A. Köhn, *J. Chem. Phys.*, 2023, **158**, 134801.
  - 50 D. E. Woon and T. H. Dunning Jr., *J. Chem. Phys.*, 1995, **103**, 4572–4585.
  - 51 K. A. Peterson and T. H. Dunning Jr., *J. Chem. Phys.*, 2002, **117**, 10548–10560.
  - 52 H.-J. Werner and P. J. Knowles, *J. Chem. Phys.*, 1988, **89**, 5803–5814.
  - 53 A. Kramida, Yu. Ralchenko, J. Reader and NIST ASD Team, NIST Atomic Spectra Database (ver. 5.12), [Online]. Available: <https://physics.nist.gov/asd> [2025, April 29]. National Institute of Standards and Technology, Gaithersburg, MD., 2024.
  - 54 J. Netz, A. O. Mitrushchenkov and A. Köhn, *J. Chem. Theory Comput.*, 2021, **17**, 5530–5537.
  - 55 K. P. Huber and G. Herzberg, *Molecular Spectra and Molecular Structure. IV. Constants of diatomic molecules*, Van Nostrand Reinhold Company, 1979.
  - 56 P. Bollmark, B. Lindgren, B. Rydh and U. Sassenberg, *Phys. Scr.*, 1978, **17**, 561–564.
  - 57 J. A. Coxon, *Can. J. Phys.*, 1979, **57**, 1538–1552.
  - 58 L. Cheng, F. Wang, J. F. Stanton and J. Gauss, *J. Chem. Phys.*, 2018, **148**, 044108.
  - 59 T. Uhlíková, S. N. Yurchenko, A. N. Perri, J. Tennyson and G.-S. Kim, *J. Chem. Phys.*, 2025, **162**, 144108.

