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Synthesis and properties of magnetic-optical core-shell nanoparticles

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Due to their high integrity, facile surface chemistry, excellent stability, and dual properties from the core and shell materials, magnetic-plasmonic core-shell nanoparticles are of great interest across a number of science, engineering and biomedical disciplines. They are promising for applications in a broad range of areas including catalysis, energy conversion, biological separation, medical imaging, disease detection and treatment. The technological applications have driven the need for high quality nanoparticles with well controlled magnetic and optical properties. Tremendous progress has been made during past few decades in synthesizing and characterizing magnetic-plasmonic core-shell nanoparticles, mainly iron oxide-gold core-shell nanoparticles. This review introduces various approaches for the synthesis of spherical and anisotropic magnetic-plasmonic core-shell nanoparticles focusing on iron oxide-gold core-shell nanoparticles. Growth mechanisms are discussed to provide understanding of the key factors controlling shape-controlled synthesis. Magnetic and optical properties are summarized from both computational and experimental studies.

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1. Introduction

Nanoscale materials are a topic of considerable interest across a number of science, engineering and biomedical disciplines. The basic rationale is that nanoscale materials, typically 1–100 nm, exhibit exceptional structural and functional properties that are not available in bulk materials or discrete molecules. Two major classes of functional nanoplatforms have been extensively studied and widely used in a variety of fields: plasmonic nanoparticles (NPs) and magnetic NPs (MNPs).

Plasmonic NPs are typically composed of noble metals, generally gold (Au) and silver (Ag). They exhibit unique localized surface plasmon resonance (LSPR), the collective oscillation of the conduction electrons of the NPs in resonance with the electric field of the incident light.^{1–3} This LSPR leads to strongly enhanced radiative (*e.g.* absorption and scattering) and non-radiative (*e.g.* photothermal and energy transfer) properties.^{4,5} Compared to Ag NPs, Au NPs are more stable under ambient conditions. Additionally, the LSPR of Au NPs can be tuned from the visible to near infrared region (NIR) by adjusting the particle's size, shape and structure.^{5–9} These intriguing optical properties have made Au NPs highly favorable for sensing,

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Dr Elise Chaffin received her Ph.D. degree in chemistry from the University of Memphis in 2016. She is currently a post-doctoral fellow at the University of Memphis. Under the mentorship of Dr Wang and Huang, her research has been focused on modelling the optical properties of noble metal nanoparticles using discrete dipole approximation and Mie theory in collaboration with the synthesis and characterization by the Huang group.



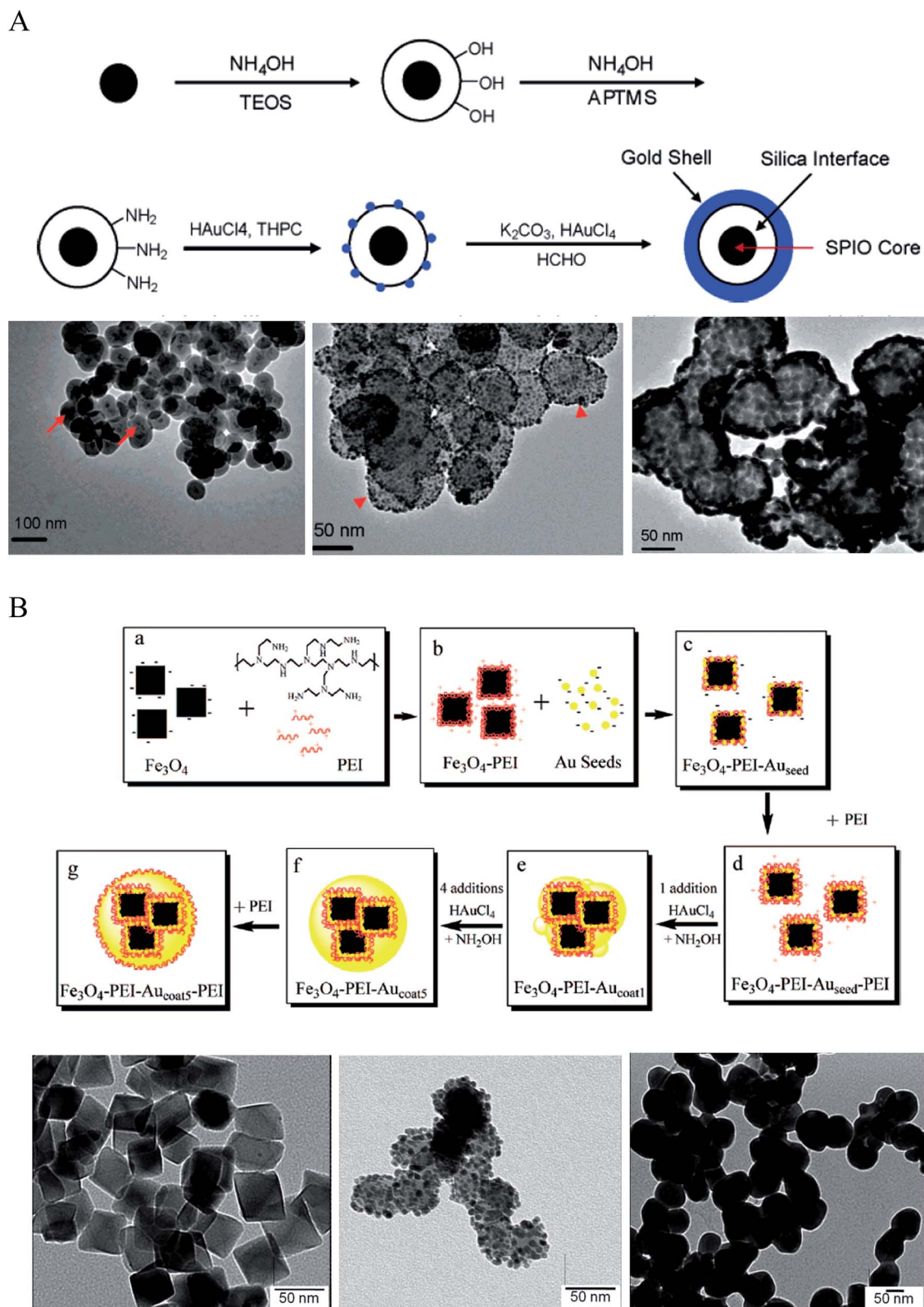


Fig. 2 Synthesis of IO–Au core–shell NPs via Au-seeded growth methods. (A) APTMS and (B) PEI are used as the anchor agent for the adsorption of Au seeds on IO NPs. (A) is reprinted with permission from ref. 23. Copyright (2007) American Chemical Society. (B) is reprinted with permission from ref. 71. Copyright (2009) American Chemical Society.



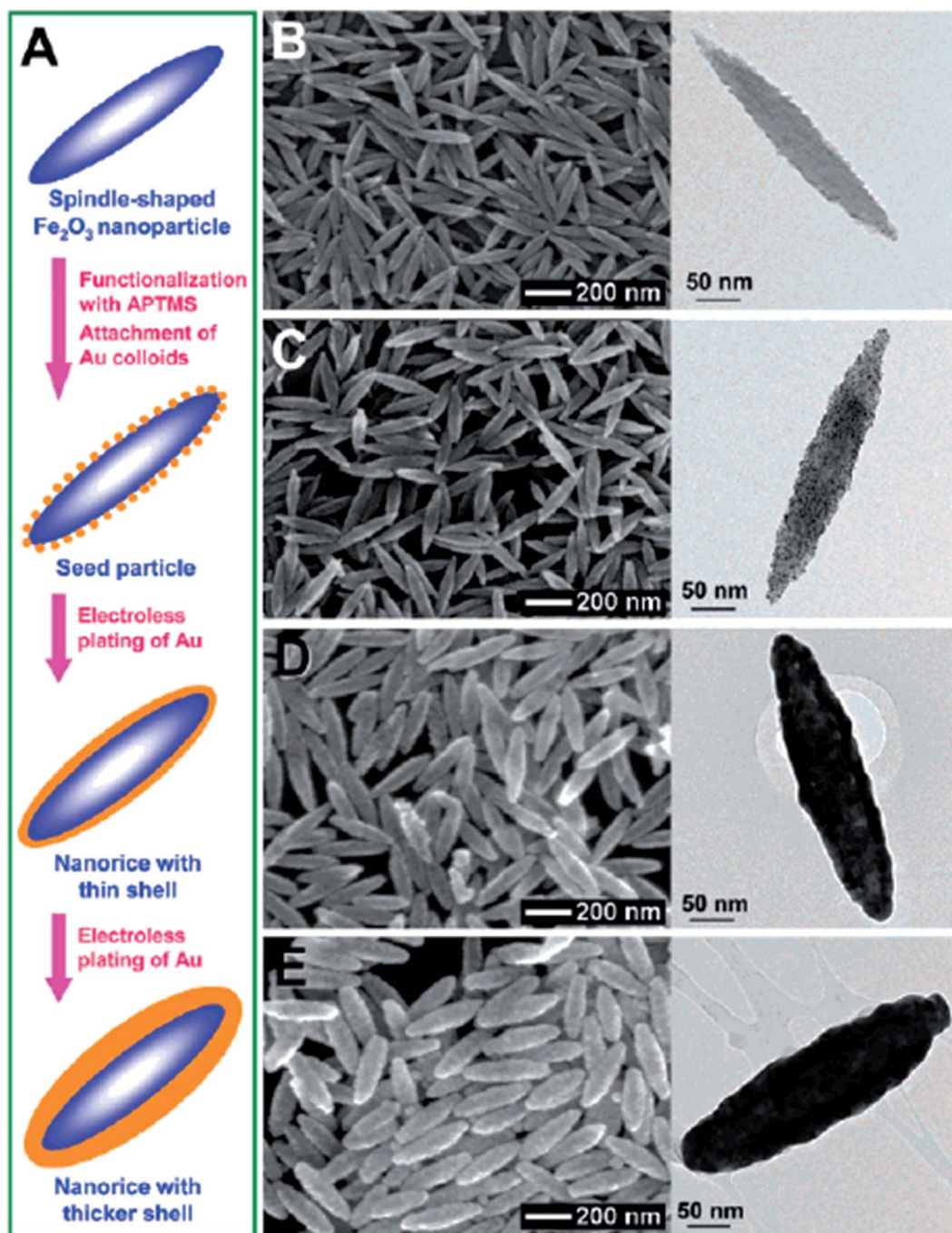


Fig. 3 Synthesis of IO–Au core–shell nanorices *via* an Au-seeded growth method. (A) Schematic of the preparation of IO–Au core–shell nanorices. (B–E) SEM (left) and TEM (right) images of IO–Au core–shell nanorices at different preparation steps. (B) The IO nanorices. (C) Au-seeded IO nanorices. (D) IO–Au core–shell nanorices with thin shells (~ 13 nm). (E) IO–Au core–shell nanorices with thick shells (~ 28 nm). Reprinted with permission from ref. 80. Copyright (2006) American Chemical Society.

a surface-catalysed reducing agent, which prevents formation of solid Au NPs. Thus, anisotropic IO–Au NSTs were obtained. This is the first report of the synthesis of monodisperse IO–Au core–shell NSTs.

We recently developed a facile method for the synthesis of uniform IO–Au core–shell NSTs (Fig. 4B)⁷² with all steps performed in aqueous solutions. In our studies, octahedral Fe_3O_4 NPs capped with PEI (edge length ~ 35 nm) were synthesized by

precipitating ferrous sulfate in a base solution followed by oxidation with KNO_3 at 90°C in the presence of branched PEI (MW ~ 25 000). Citrate-capped Au NPs (<10 nm) were attached to the IO NPs by electrostatic adsorption and the Au-seeded Fe_3O_4 NPs were further stabilized by PEI. The growth of Au shells was initiated by injecting the Au-seeded IO NPs into a growth solution containing HAuCl_2 , CTAB, AgNO_3 and AA. By adjusting the amount of Au-seeded Fe_3O_4 NPs, IO–Au core–shell



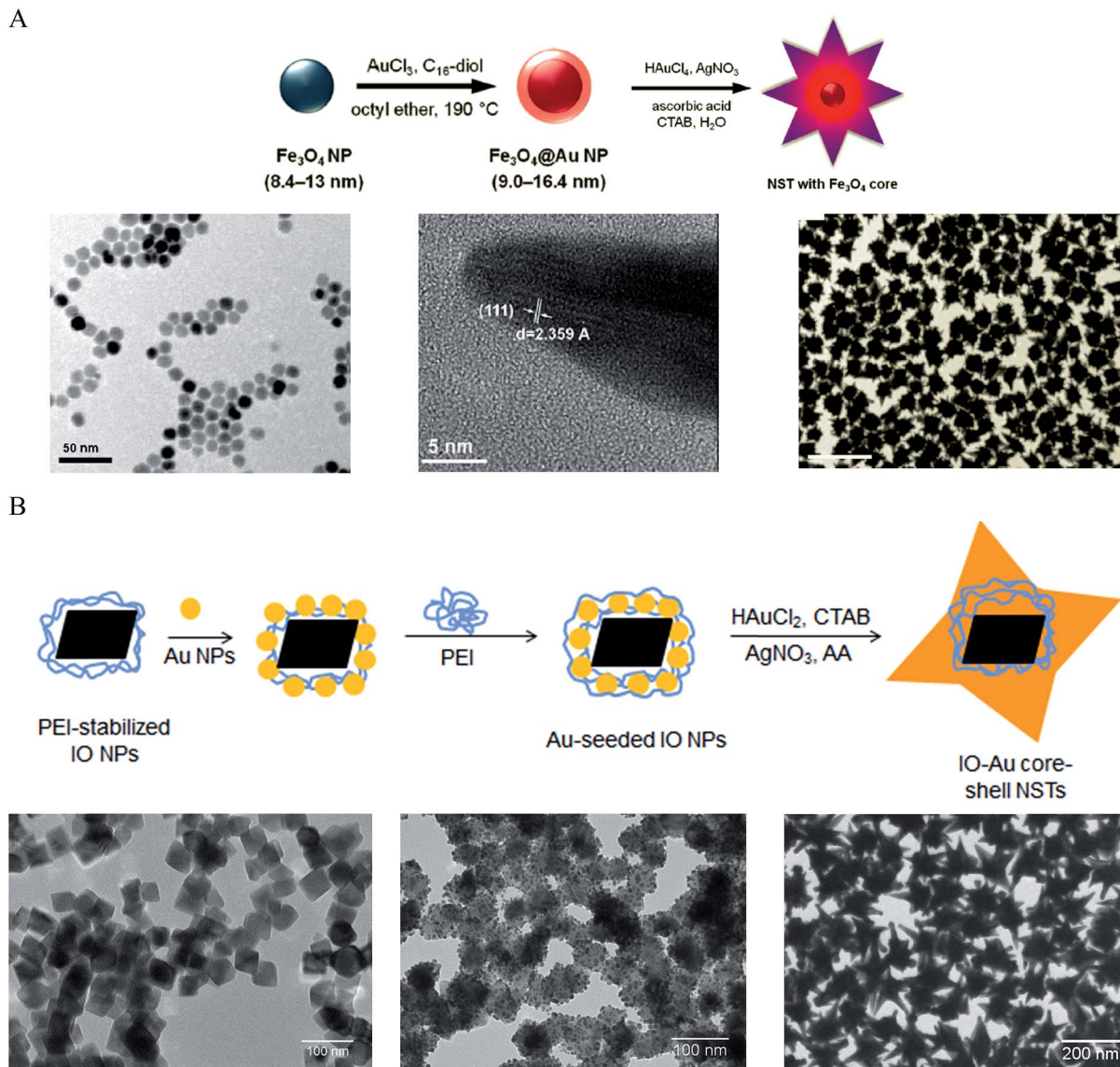


Fig. 4 Synthesis of IO–Au core–shell NSTs. (A) Synthesis of IO–Au core–shell NSTs from small IO–Au core–shell nanospheres. (Top) Schematic of the preparation procedure. (Bottom) TEM image of IO NPs (left), HRTEM image of a tip (middle), and TEM image of IO–Au NSTs. Reprinted with permission from ref. 82. Copyright (2010) American Chemical Society. (B) Synthesis of IO–Au core–shell NSTs by the Au-seeded growth method. (Top) Schematic of the preparation procedure. (Bottom) TEM images of IO NPs (left), Au-seeded IO NPs (middle) and IO–Au NSTs. Reprinted with permission from ref. 72. Copyright (2016) American Chemical Society.

NSTs with base sizes from 70 to 150 nm were obtained. In addition to forming NSTs, this method can be easily adapted to make nanospheres and nanopopcorns by changing the chemical composition of the growth solution. When AgNO_3 was removed from the growth solution, IO–Au core–shell nanospheres were formed. When the concentration of AA was tripled, nanopopcorns were obtained. The sizes of the nanospheres and nanopopcorns can also be tuned without affecting particle quality by simply varying the amount of Au-seeded IO NPs in the growth solution. A major advantage of this method is the capability to tune both the size and the shape of the core–shell particles. Later, Shi and co-workers reported another approach to make IO–Au NSTs using Ag-seeded IO NPs.²⁵ In their study, small Ag NPs were synthesized for the seeds *via* the

reduction of AgNO_3 by NaBH_4 in a partially thiolated PEI aqueous solution. The PEI-stabilized Ag NPs were added into a mixture of FeCl_2 and ammonium hydroxide and autoclaved in a sealed pressure vessel at 134 °C for 3 h to obtain Ag-seeded Fe_3O_4 NPs. This hydrothermal process led to the homogenous coating of Au NPs on the Fe_3O_4 NPs. Addition of the Ag-seeded Fe_3O_4 NPs into a growth solution containing HAuCl_4 , CTAB, AgNO_3 and AA resulted in the formation of Fe_3O_4 –Au core–shell NSTs within 1 h.

Using small Ag NPs (2–5 nm) as the seeds and Fe_3O_4 NPs coated with poly (maleic anhydride-*alt*-1-octadecene) as the core, we have produced IO–Au core–shell NPs in oval and pin shapes.^{37,84} The polymer interacts with OA on IO NPs with the hydrophobic alkyl chain while exposing the hydrophilic



carboxylate ions to anchor positively charged metal ions. Ag seeds were used as the nucleation sites because the Ag-seeded IO NPs can be formed without aggregation by reduction of the purified $\text{Ag}(\text{NH}_3)_2^+$ adsorbed IO NPs. In this way, the formation of un-adsorbed Ag NPs that would lead to solid Au NPs is avoided. Ag has a nearly identical crystal structure to Au, with almost 100% lattice matching. Galvanic replacement between Au ions and Ag seeds during subsequent growth process is prevented by externally introducing a stronger agent, such as ascorbic acid (AA), that is stronger than Ag ($E_{\text{dehydroAA/AA}}^{\circ} = 0.06 \text{ V}$ versus $E_{\text{Ag}^+/\text{Ag}}^{\circ} = 0.80 \text{ V}$). Thus, Au can be uniformly deposited onto the Ag nucleation sites to form Au shell.

2.3 Growth mechanism

A common model to understand the growth mechanism of anisotropic solid Au NPs is that the Au seeds are faceted NPs. The formation of various shapes is the outcome of the interplay between the facet binding tendency of the stabilizing agents and the growth kinetics.⁸⁵ As for the core-shell NPs, Halas and

co-workers have studied the growth mechanism of spherical Au nanoshell on silica core by monitoring the progression of Au nanoshell formation with TEM imaging.⁶⁶ Their studies showed that at the early stages the adsorbed Au seeds gradually grew with time on the silica core, then coalesced on the core surface until finally forming a continuous metallic shell. In the studies of IO–Au NSTs prepared from spherical IO–Au NPs, Wei and co-workers used high resolution TEM (HRTEM) and found multiple twinning defects and overall growth along the (111) directions. As the mismatching of crystal lattices of Au and IO are within 3%, it is believed that the formation of the shell is due to favorable epitaxial growth.

We recently investigated the growth mechanisms for shape-controlled synthesis by studying the growth processes of IO–Au nanospheres, popcorn and stars using both experimental and computational methods (Fig. 5).⁷² By following the growth process with TEM, it was found that the IO–Au core-shell nanospheres grew at a faster rate than the popcorns and stars. For spheres, the Au seeds coalesced within 5 min and formed



Fig. 5 Computational studies on growth mechanisms of IO–Au core-shell nanospheres, nanopopcorns, and nanostars. (A–C) TEM images of IO–Au core-shell nanospheres, nanopopcorns, and nanostars. (D–F) The adsorption of Ag on the Au (100) surface (D), Au (110) surface (E), and Au (111) surface (F). (G) Schematic of the proposed geometry of the tip of the nanostar. (H) HRTEM image of a tip of the nanostar. The inset shows an enlarged view of the lattice, where the position of each atom is visible. (I) The positions of atoms according to the proposed model in (G), which is viewed from the [1–10] direction. (J–L) The growth mechanism of Au nanospheres (J), nanopopcorn (K), and nanostars (L). Reprinted with permission from ref. 72. Copyright (2016) American Chemical Society.



continuous Au shells within 20 min. For the popcorns and stars, no visible changes on the Au seeds were observed after 5 min. The Au seeds coalesced at 20 min, with continuous Au shell formation occurring after 30 min. Differences between popcorns and stars were found at 30 min, with the popcorns exhibiting short and wide protrusions and the stars thin spikes.

The growth of the Au-seeded IO NPs into anisotropic was due to the effects by AgNO_3 . AgNO_3 has been widely used to assist the preparation of Au NRs and other anisotropic metal NPs.^{85–87} It has also been used to form spiky Au nanoshells from an Ag-seeded polymer template.^{88,89} A previous model to explain the role of Ag^+ is silver underpotential deposition, the reduction of Ag^+ to Ag^0 on a metal substrate with a surface potential less than the standard reduction potential.⁹⁰ Based on this model, Ag^+ is reduced and deposited onto the surfaces of Au seeds in the presence of a reducing agent. This may explain the slower growth rate for Au NPCs and NSTs compared with the nanospheres. Using the first principles density functional theory method, we have determined that the adsorption energies (the energy gain from adsorption) of Ag on (110), (100), and (111) are 3.0, 2.7, and 2.2 eV per atom, respectively. Thus, Ag deposition is correspondingly faster on the (110) facet, followed by the (100) and (111) surfaces. With sufficient Ag^+ in the solution, Ag covers the (110) and (100) surfaces of Au seeds, blocking the adsorption of Au on these faces. This blocking effect leaves only the (111) surface for Au growth, which leads to the anisotropic growth of the Au.

The branched growth of the IO–Au core–shell NPs into popcorns or stars is a result of kinetic control on the reductive deposition of Au by adjusting the concentration of AA. When the reduction of Au^+ to Au^0 is fast (higher AA concentration), Au atoms are deposited onto the entire (111) facet, leading to the growth of the Au island along the (111) direction. This mechanism is confirmed by the HRTEM image of the tip of the popcorn, which shows that the tip of popcorn grows along the {111} direction. When the reduction of Au^+ to Au^0 is slow, Au atoms are mainly deposited at the ridge formed by two {111} planes. This is because the Au atoms at the ridge are more under coordinated than that at the facet center, which makes new Au atoms bind the adsorbed Au atoms more strongly. Also, the electrical field at the ridge is stronger than that at the center of the facet, making the reduction of Au^+ faster at the ridge. This cohesive growth of Au atoms from two planes leads to the formation of a twin boundary along the (110) direction. This model was supported by the HRTEM image of the star tip, which shows the tip is made of two of {111} planes growing towards the (110) direction.

3. Properties

3.1 Magnetic properties of IO–Au core–shell NPs

Due to the unpaired electrons in the 3d shell of Fe^{2+} and Fe^{3+} , nanocrystals formed from Fe^{2+} and Fe^{3+} can be in ferromagnetic, ferrimagnetic and antiferromagnetic states. In ferromagnetic materials, the magnetic moments of two sublattices align parallel to each other even without an external field. In ferrimagnetic materials, the magnetic moments align

antiparallel, but do not cancel each other out. This is different from the antiferromagnetic materials where the magnetic moments of two sublattices are equal and align antiparallel. Thus, there is no net magnetic moment in zero magnetic field for antiferromagnetic materials. The magnetic moments lose ordering beyond a specific temperature called the Curie temperature T_C for ferromagnets and ferrimagnets and the Neel temperature T_N for antiferromagnets. Magnetite Fe_3O_4 is a well-known ferrimagnetic material with a T_C of 858 K.⁹¹ Maghemite $\gamma\text{-Fe}_2\text{O}_3$ is ferrimagnetic at room temperature. They are unstable at high temperatures and thus T_C is hard to determine. Both Fe_3O_4 and $\gamma\text{-Fe}_2\text{O}_3$ NPs are superparamagnetic at room temperature when their sizes are sufficiently small (less than 30 nm).^{20,92–94} For the superparamagnetic NPs, the magnetization randomly flips directions due to thermal fluctuations and shows zero net magnetization at zero magnetic field. The particles behave like paramagnetic material under an external field but their magnetic susceptibility is much larger than that of paramagnets.

The magnetic properties of IO NPs depend on the surface. When the size of the NPs decreases, the magnetic properties decrease due to increased surface effects. Passivation with organic and inorganic layer helps decrease the surface effect, but it could also adversely affect the surface magnetic moment of IO NPs. The surface magnetic moments of IO NPs can be disordered *via* the interaction with Au electrons.⁹⁵ The IO–Au core–shell NPs could have larger surface effect than IO NPs by structural distortions that cause spin canting.^{63,96} This leads to decreased magnetic properties of the IO–Au core–shell NPs compared to the IO NPs. Magnetic properties are commonly measured by a superconducting quantum interference device by scaling to the total mass of the materials. Au has 4 times higher density than IO. However, Au is diamagnetic in bulk states. It is ferromagnetic when the size very small (4–5 nm).⁹⁷ Therefore, IO–Au core–shell NPs generally give significantly lower saturation magnetization (M_s) than IO NPs due to the mass contribution from the diamagnetic Au shell. For example, coating the tetracubic IO NPs (~60 nm) with a 10.9 nm Au shell decreased the M_s from 20 to 0.6 emu g^{-1} at both 100 K and 250 K (Fig. 6A and B).⁶⁹ Coating 258 nm Fe_3O_4 NPs with a 7.5 nm Au shell decreased the M_s from 75.6 to 61.0 emu g^{-1} .⁷⁰

The IO–Au core–shell NPs typically exhibit the same magnetic behaviour as the IO core. Increasing Au shell would lead to a decrease in the magnetic properties of the core–shell NPs due to the mass contribution of the diamagnetic Au. However, in our recent studies, we found that increasing the Au shell thickness led to an increase of the magnetic properties of IO–Au core–shell nanopopcorns at 10 K (Fig. 6C).⁷² The particles are capped with CTAB that binds to Au *via* Br^- and is counterbalanced by CTA^+ in a bilayer structure.^{98–100} Thus, charge transfer between Br^- and Au may occur, which would enhance the electron mobility and, thus, the surface induced magnetism of Au. Similarly, studies by Crespo *et al.* showed that the charge transfer between a thiol ligand and an Au surface was a major reason for the observed ferromagnetic properties of thiolated Au NPs.⁹⁷ It has also been recognized that the magnetic material could spin polarize the Au conduction electrons at the IO–Au





Fig. 6 Magnetic properties of IO–Au core–shell NPs. (A and B) Magnetization as a function of applied field at 100 K and 250 K for uncoated and Au coated tetragonal IO NPs. Reprinted with permission from ref. 69. Copyright (2009) American Chemical Society. (C) Magnetization as a function of applied field at 10 K for IO–Au core–shell nanopopcorns of three different sizes. Reprinted with permission from ref. 72. Copyright (2016) American Chemical Society. (D) Magnetization as a function of applied field at 10 K and 300 K for bare IO NPs (sample A), IO–Au core–shell NPs with low Au (sample B) and high Au (sample C). Reprinted with permission from ref. 46. Copyright (2011) American Chemical Society.

interface and thus result in magnetization of the surrounding nonmagnetic Au.¹⁰¹ In addition, a large orbital magnetic moment could be induced at the Fe₃O₄–Au interface. This interface effect leads to the magnetization enhancement of Fe₃O₄ NPs by a factor of six (Fig. 6D).⁴⁶

Research has shown that there is no significant difference in the coercivity between coated and uncoated IO NPs.⁶⁹ However, enhanced coercivity has been reported by Pal *et al.* on small Fe₃O₄ NPs (6 nm) with ultrathin shell (1 nm thickness).⁶³ The Au-coated Fe₃O₄ NPs gave a coercivity H_c of 200 Oe while the uncoated Fe₃O₄ NPs had H_c of 160 Oe at 5 K. The coercivity enhancement is possibly due to the role of spin disorder at the Fe₃O₄–Au interface and weak exchange coupling between surface and core spins.

Separation of IO–Au core–shell NPs from suspension usually takes much longer than uncoated IO NPs because of the massive amount of Au added to the IO NPs by the Au shell. This can be theoretically understood by calculating the particle

terminal velocity under external magnetic field. The particle terminal velocity V is equal to $V = \frac{F_m}{6\pi\eta R}$ where F_m is the magnetic force exerted onto the particle by the external magnetic field, η is the viscosity of the solution, and R is the hydrodynamic radius of the particle. The F_m of the IO–Au core–shell particle is the same as the IO particle if Au coating does not affect the magnetic property of the IO core. Thus, the particle terminal velocity V is inversely proportional to the hydrodynamic radius R of the particle. The particle also takes time to reach the terminal velocity. The time, called relaxation time τ is equal to $\tau = \frac{m}{6\pi\eta R}$ where m is the mass the particle. This means that τ is linearly proportional to the mass of the particle and is inversely proportional to the hydrodynamic radius of the particle. Taking into account of both V and τ from the above two equations, the time taken for a particle to be separated by an external magnetic field is therefore proportional to the mass of



the NP. The mass of IO and IO–Au core–shell NP can be calculated based on their sizes measured by TEM and the densities of IO (5.24 g cm^{-3}) and Au (19.32 g cm^{-3}). If an IO NP is 20 nm in diameter and the Au shell is 10 nm, the mass of IO–Au core–shell NPs is 27 times larger than the IO NP. This means that the separation time for a core–shell particle is roughly 27 times longer than that of the IO NP traveling the same distance. Thus, if an IO NP takes 5 min to be separated from the solution, the separation of an IO–Au core–shell NP requires more than 2 h due to the mass added by the diamagnetic Au shell.

3.2 Optical properties of IO–Au core–shell NPs

The optical properties of core–shell plasmonic NPs have been a topic of great interest in materials science for many years. Their properties are dependent on the dielectric functions of the core and shell materials, the core size and shape, as well as the shell thickness and geometry. An intensely studied core–shell NP is the Au nanoshell with a silica core pioneered by Halas and co-workers.¹⁰² The Si–Au core–shell NPs exhibit strong NIR properties, with increasing thickness of the Au shell leading to an LSPR blue shift.⁶⁶ They proposed a hybridization model to understand the plasmons of concentric nanoshells, where the hybridization of the plasmons of inner and outer nanoshells determines the LSPR of the Si–Au core–shell NPs.¹⁰³

Using Mie theory, we recently calculated the optical properties of spherical Fe_3O_4 –Au core–shell NPs, and compared them with Si–Au, Co–Au, hollow Au, and solid Au NPs (Fig. 7A).¹⁰⁴

Compared with solid Au NPs of the same size, the LSPR peak of IO–Au core–shell NPs is further red shifted than those of silica and hollow NPs. The LSPR peak intensity is comparable to that of solid Au NPs, but much lower than those of the hollow Au and Si–Au core–shell NPs. In contrast, the Co–Au core–shell NPs have very weak plasmon peaks, although cobalt NPs have stronger magnetic properties than magnetite NPs. These optical differences can be traced to the differences in the dielectric properties of the core material, which is related to the refractive index of the material. IO NPs are nontransparent and have complex refractive indices. Silica, on the other hand, is a non-absorbing material and has only a real refractive index. While the LSPR peak wavelength is determined by real component of the refractive index of the material, the peak intensity is determined by the imaginary component of the refractive index. Increasing the real refractive index leads to a red shift of the LSPR peak whereas increasing the imaginary refractive index leads to a reduction of the peak intensity (Fig. 7B). Similar to Si–Au core–shell NPs, the LSPR wavelength blue shifts with the increase of Au shell thickness when the total diameter of the NPs is fixed (Fig. 7C and D). It follows the same universal scaling as the Si–Au core–shell NPs that has been reported by Jain and El-Sayed.¹⁰⁵ The shift in LSPR wavelength is determined by the ratio of the thickness of the shell to the radius of the core (t/R). Compared to the Si–Au system, the Fe_3O_4 –Au NPs show a larger decay constant in the plasmon shift *versus* the ratio of shell thickness. This is consistent with LSPR peaks that are more red-

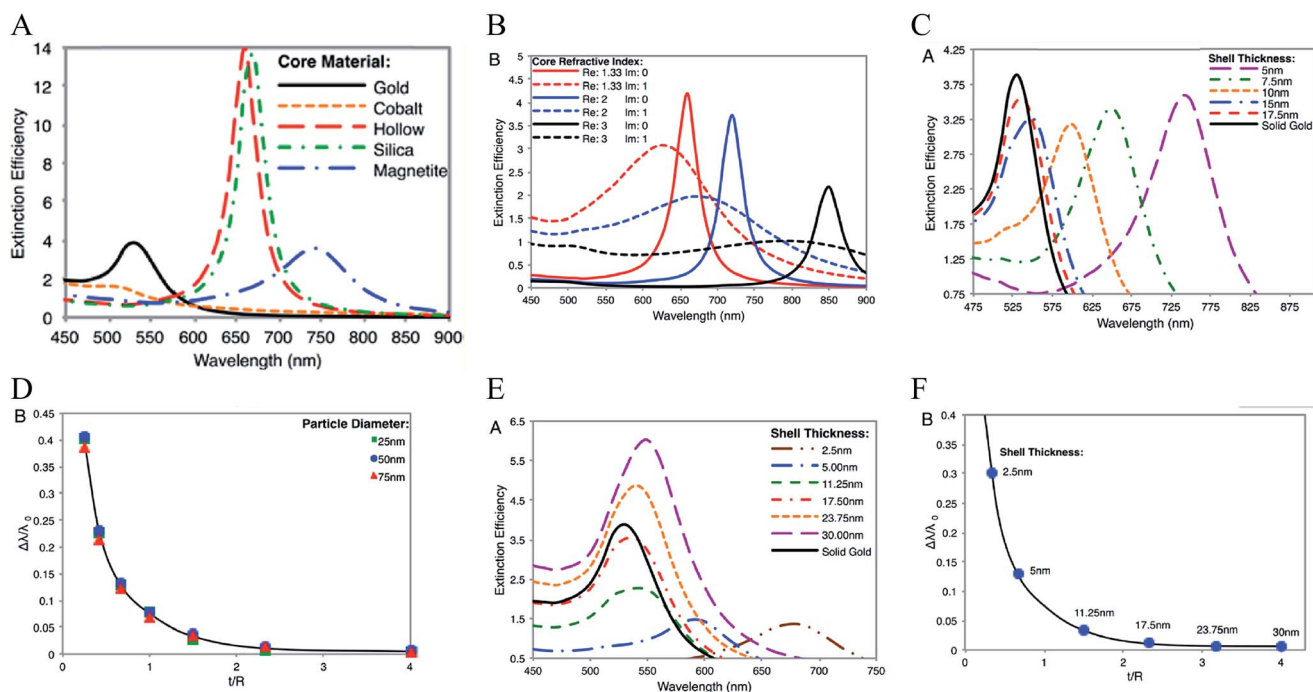


Fig. 7 Calculated optical properties of Fe_3O_4 –Au core–shell NPs. (A) Extinction spectrum of Fe_3O_4 –Au core–shell NPs in comparison with other core–shell NPs and solid Au NPs with the same diameter $D_{\text{total}} = 50 \text{ nm}$ and a shell thickness of 5 nm. (B) Extinction spectra of core–shell NPs with varied core refractive indices. (C) Extinction spectra of Fe_3O_4 –Au core–shell NPs with different shell thickness with fixed total particle size. $D_{\text{total}} = 50 \text{ nm}$. (D) Fractional shifts ($\Delta\lambda/\lambda_0$) of the LSPR peak maximums of the IO–Au NPs from (C). (E) Extinction spectra of Fe_3O_4 –Au core–shell NPs with different shell thickness with fixed IO core size. $D_{\text{core}} = 15 \text{ nm}$. (F) Fractional shifts ($\Delta\lambda/\lambda_0$) of the LSPR peak maximums of the IO–Au NPs from (E). Reprinted with permission from ref. 104. Copyright (2014) American Chemical Society.



shifted than those of Si–Au NPs having identical shell thicknesses and core radii. In typical IO–Au NP synthesis, the IO-core diameter remains constant; as the Au shell thickness increases, the total particle size increases as well. In this case, the LSPR peak is seen to blue shift first with increasing the thickness of the Au shell (Fig. 7E and F). However, as the thickness of the Au shell continues to grow, a red shift of the LSPR peak occurs. In this case, the peak shift seems do not adhere to the universal scaling equation. In fact, the universal scaling still applies. When the Au NPs with the same diameter of each core–shell NP are used as the references, the data follows the universal scaling nicely.

Theoretically, a spherical IO–Au core–shell NP with a 20 nm core and a 15 nm shell has a LSPR peak at 540 nm.¹⁰⁴ When the core is increased to 35 nm and Au shell thickness is decreased to 7.5 nm, the LSPR is red shifted to 650 nm. Most existing IO NPs are 30 nm and less. Thus, they have LSPR in the UV-Vis region after Au coating. To make particles with LSPR over 800 nm, the core needs to be larger than 100 nm assuming shell thickness t is equal to or larger than 5 nm.¹⁰⁵ NIR-absorbing IO–Au NPs less

than 50 nm have been reported by Gao and co-workers (Fig. 8A).⁴⁴ These particles have a core and shell separated by a few nanometer polymer gap. Although the core is only 25 nm in diameter, the particles absorb over 700 nm when the Au shell is 2–3 nm. When the shell is decreased to 1–2 nm, the particles show a broad absorption band with the plasmon resonance approximately 850 nm. This is consistent with our computational studies that demonstrated a polymer gap induce plasmon red shift.

A common way to tune the LSPR of plasmonic NPs is to change the shape. For example, when an Au NP is changed from sphere to rod, the LSPR is split into two peaks.¹⁴ One is around 520 nm (transverse band) resulting from the electron oscillations along the short axis of the rod. The other one (longitudinal band) is at a longer wavelength with much stronger intensity due to the electron oscillations along the long axis. The LSPR of this longitudinal band is very sensitive the particle's aspect ratio (length/width). Increasing the aspect ratio causes large shift of the LSPR wavelength into NIR region. IO–Au core–shell nano-rods have not been experimentally produced. Theoretically,

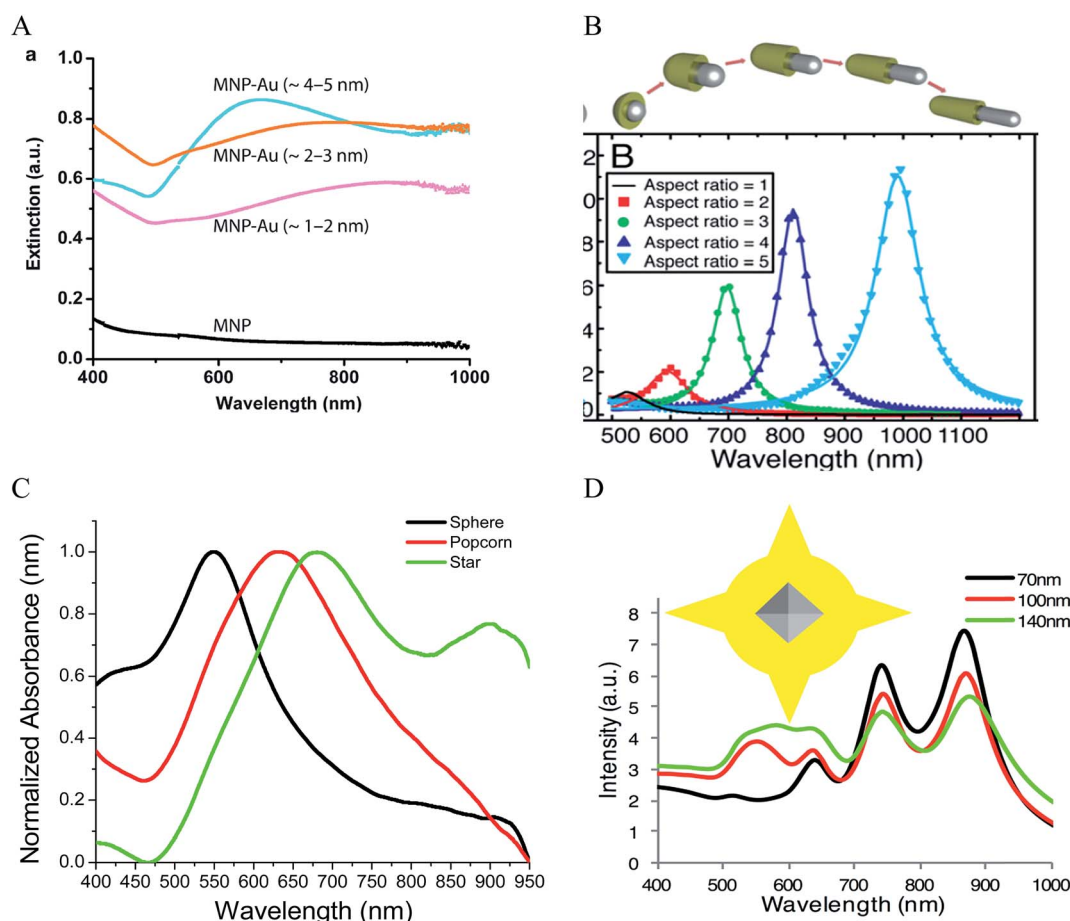


Fig. 8 NIR absorbing IO–Au core–shell NPs. (A) Absorption spectra of Fe_3O_4 –Au core–shell NPs with different shell thickness. $D_{\text{core}} = 25$ nm. Reprinted with permission from Macmillan Publishers Ltd: [Nature Communications] (ref. 44), copyright (2010). (B) Calculated extinction spectra of IO–Au core–shell NRs with different aspect ratios. Reprinted with permission from ref. 106. Copyright (2011) Elsevier. (C) Absorption spectra of Fe_3O_4 –Au core–shell NPs with different shapes. The Fe_3O_4 core is octahedral with an edge length of 35 nm. Reprinted with permission from ref. 72. Copyright (2016) American Chemical Society. (D) Calculated extinction spectra of IO–Au core–shell NSTs with different base size. The Fe_3O_4 core is octahedral with an edge length of 35 nm. Reprinted with permission from ref. 72. Copyright (2016) American Chemical Society.



γ -Fe ₂ O ₃	Maghemite
HAuCl ₄	Chloroauric acid
HRTEM	High resolution transmission electron microscopy
IO	Iron oxide
IO–Au	Iron oxide–gold
LSPR	Localized surface plasmon resonance
MNPs	Magnetic nanoparticles
MRI	Magnetic resonance imaging
M _s	Saturation magnetization
NaBH ₄	Sodium borohydride
NPs	Nanoparticles
NH ₂ OH	Hydroxylamine
NIR	Near infrared region
NR	Nanorod
NST	Nanostar
OA	Oleic acid
OAM	Oleylamine
PAH	Poly(allylamine hydrochloride)
PEI	Polyethyleneimine
PVP	Poly(vinylpyrrolidone)
Si	Silica
TEM	Transmission electron microscopy
THPC	Tetrakis(hydroxymethyl) phosphonium chloride

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