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## Peroxy radical kinetics and new particle formation

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Chamber experiments showing “pure biogenic nucleation” have shown an important role for covalently bound organic association products (“dimers”). These form from peroxy-radical ( $RO_2$ ) cross reactions. Chamber experiments at low- $NO_x$  conditions often have quite high hydrocarbon reactant concentrations and relatively low concentrations of oxygenated volatile organic compounds (OVOCs). This can skew the radical chemistry in chambers relative to the real atmosphere, favoring  $RO_2$  and disfavoring  $HO_2$  radicals.  $RO_2$  cross reaction kinetics are in turn highly uncertain. Here we explore the implications of the  $RO_2$  to  $HO_2$  ratio in chamber experiments as well as the implications of uncertain  $RO_2$  cross reaction kinetics and the potential for added CO to mimic more atmospheric radical conditions. We treat a plausible range of  $RO_2$  rate coefficients under both typical chamber conditions and atmospheric conditions to see how dimerization is affected by high concentrations of OVOCs, and thus lower  $RO_2 : HO_2$  relative to smog chamber experiments. We find that if  $RO_2$  reactions are fast, relatively high yields of low volatility dimers can participate in new particle formation. The results are highly sensitive to both the (uncertain)  $RO_2$  kinetics as well as  $RO_2 : HO_2$ , suggesting both that low- $NO_x$  chamber results should be extrapolated to the atmosphere with caution but also that the atmosphere itself may be highly sensitive to the specific (and rich) mixture of organic compounds and thus peroxy radicals.

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### Environmental significance

New particle formation is a major contributor to atmospheric aerosol number and is enhanced by low volatility organics. Understanding the conditions under which these organics are formed is crucial to our understanding of new particle formation. These processes are often studied in laboratory smog chamber experiments, which simulate atmospheric conditions without the complication of the many species present in the actual atmosphere. In addition, parameterizations derived from those experiments may have limited applicability if they lack molecular detail; understanding this chemistry at the molecular level allows for a broader extension of experimental results. In this work we look at how species present in abundance in the atmosphere, but not in smog chambers, affect the chemistry that leads to these low volatility products. This can cause a high bias in new-particle formation rates observed in smog chamber experiments. Thus we stress the importance of knowing the exact conditions of smog chamber experiments and building molecular level understanding before extending results to atmospheric conditions.

## 1 Introduction

Our understanding of organic oxidation in the atmosphere has been transformed by recent smog-chamber<sup>1,2</sup> and flow-reactor experiments,<sup>3–5</sup> enabled by extraordinary advances in mass spectrometry<sup>6,7</sup> and particle detection.<sup>8</sup> Especially interesting is a process known as “autoxidation”, the self oxidation of organic compounds in the presence of only molecular oxygen.<sup>9–11</sup> Within seconds after an oxidant (OH,  $O_3$ , Cl,  $NO_3$ , etc.) reacts with a stable volatile organic compound (VOC), autoxidation can generate highly oxygenated organic molecules (HOMs) with multiple oxygen containing functional groups. The HOMs in turn can have very low vapor pressures and thus condense to existing particles as secondary organic aerosol, even driving

growth of sub-10 nm particles.<sup>12,13</sup> In some cases the HOMs can even nucleate new particles, with<sup>14</sup> or without<sup>15</sup> sulfuric acid.

New-particle formation in the atmosphere contributes up to half of the number concentration of global cloud condensation nuclei and there are huge uncertainties associated with the mechanisms of nucleation and growth and their sensitivities to different conditions.<sup>16–18</sup> Because of its contribution to cloud condensation nuclei, accurately representing new-particle formation in models is critical for understanding aerosol cloud interactions and thus aerosol climate forcing.<sup>16,19–21</sup> One of the largest sources of uncertainty in anthropogenic aerosol forcing is the pre-industrial baseline;<sup>22</sup> this is because the relationship between new-particle formation and the concentration of cloud condensation nuclei is highly non-linear and so the forcing – the difference between the present day and the pre-industrial – depends strongly on the pre-industrial initial condition.

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Our understanding of pre-industrial aerosol in turn depends on our understanding of the contribution of biogenic vapors to aerosols. Biogenic volatile organic compounds (VOCs) dominate global VOC concentrations, especially in clean environments.<sup>23,24</sup> Monoterpenes especially have been widely studied because of their role in secondary aerosol formation, both through growth of primary aerosol and their contribution to new particle formation. Monoterpenes, including  $\alpha$ -pinene, are a large source of atmospheric precursor VOCs; they can be oxidized to form low-volatility organics and have been shown to contribute to growth of new particles.<sup>12,15</sup> Organic peroxy radicals ( $\text{RO}_2$ ) formed after an initial step of monoterpene oxidation can autoxidize rapidly and potentially multiple times to form HOMs.

However, it appears that HOMs are always relatively minor products of VOC oxidation, with total molar yields  $\lesssim 0.1$  in most cases.<sup>25,26</sup> Furthermore, HOMs do not necessarily have “low volatility” in a sense that is meaningful in the atmosphere. We define volatility classes based on saturation mass concentrations ( $C^*$ , in  $\mu\text{g m}^{-3}$ )<sup>27,28</sup> according to the behavior of vapors in those classes, and there are a succession of low volatility classes with  $C^* \leq 0.1 \mu\text{g m}^{-3}$  ( $p^* \sim 10^{-6}$  Pa). Ultra Low Volatility Organic Compounds (ULVOC) can nucleate – they will often stick to anything, including other ULVOCs. Extremely Low Volatility Organic Compounds (ELVOC) will stick to any existing particle, no matter how small, but are not efficient nucleators. Low Volatility Organic Compounds (LVOC) stick to most particles but are sensitive to curvature and the Kelvin effect, so may not stick to particles smaller than 5–10 nm diameter. Semi-Volatile Organic Compounds (SVOC) exist in significant quantities in both the gas phase and in particles at equilibrium under typical atmospheric conditions. Even SVOC have very low vapor pressures no greater than  $10^{-3}$  Pa.

While many HOMs have saturation concentrations in the LVOC, ELVOC, and ULVOC ranges due to their high oxygen content, some HOMs are in the SVOC range. This brings in an important distinction between HOMs and the volatility classes. HOMs have at least 6 oxygen atoms and are formed in the gas phase *via* autoxidation, but the definition places no restrictions on volatility.<sup>11</sup> The volatility classes describe saturation concentration ranges, but do not restrict the chemistry that formed the species.

Further, substantial evidence suggests that “pure biogenic” nucleation is driven by covalently bound “dimers” (ROOR) formed when two autoxidized peroxy radicals react with each other.<sup>28–32</sup> These dimers are formed *via* reaction of two peroxy radical species and result in large molecules that may be ULVOCs and thus contribute to nucleation. However, these dimers do not necessarily have to be HOMs, as two peroxy radicals that have not undergone any autoxidation may form a dimer, but most dimers are either ULVOCs or LVOCs and thus contribute to nucleation and growth of new particles. This raises a concern that applying smog-chamber or flow-reactor data to the atmosphere may not be straightforward. These are complex reaction systems, and radical termination (especially dimer formation) is a non-linear (even explicitly quadratic) process.

One specific concern is that  $\text{RO}_2$  to  $\text{HO}_2$  ratios ( $\text{RO}_2 : \text{HO}_2$ ) in chambers may be artificially high due to low concentrations of oxygenated carbon-containing molecules like  $\text{CO}$ ,  $\text{CH}_2\text{O}$ , *etc.* (OVOCs for shorthand). Small OVOCs promote the direct conversion of  $\text{OH}$  to  $\text{HO}_2$ . In the atmosphere, where these OVOCs are typically abundant,  $\text{RO}_2 : \text{HO}_2 \lesssim 1$ ,<sup>33,34</sup> but when no OVOCs are present (which is typical of many laboratory experiments),  $\text{RO}_2 : \text{HO}_2$  can be very large. This means that chambers with relatively low OVOC levels likely will not have atmospherically relevant  $\text{RO}_2 : \text{HO}_2$ . If this alters the radical termination and thus stable product formation (including HOMs and especially dimers), direct projections to atmospheric conditions will not be straightforward. Furthermore, the absolute radical concentrations (and thus overall termination rates) may also compete with autoxidation, which may again affect projections to atmospheric conditions.

We show a simplified chemical topology of this system in Fig. 1. This emphasizes the  $\text{HO}_x\text{--RO}_x$  radical cycle, progressing from the reaction of a VOC with an oxidant ( $\text{OH}$  here) to  $\text{RO}_2$ , then an alkoxy radical ( $\text{RO}$ ), then  $\text{HO}_2$ , and finally back to  $\text{OH}$ . We omit many important termination steps as well as the specific conversion reactions to focus on the key dependencies. The central issue is  $\text{RO}_2$  termination, which depends on  $\text{RO}_2^- : \text{HO}_2$  as well as the associated radical–radical kinetics. At sufficiently high  $\text{RO}_2 : \text{HO}_2$ ,  $\text{RO}_2$  cross reactions will dominate termination, either *via* dimer formation as shown or *via* the so-called “molecular” pathway.<sup>35</sup> At sufficiently low  $\text{RO}_2 : \text{HO}_2$ ,  $\text{RO}_2^- + \text{HO}_2$  may dominate termination *via* hydroperoxide (ROOH) formation. VOC oxidation will immediately yield  $\text{RO}_2$ , as shown, but OVOC oxidation (using  $\text{CO}$  as a surrogate in the figure as well as the rest of this study) will often bypass  $\text{RO}_2$  formation and lead straight to  $\text{HO}_2$ .<sup>36</sup> Because atmospheric VOC reaction mechanisms inevitably involve multiple oxidation steps and multiple intermediate OVOC products, the atmosphere is relatively rich in OVOCs (with fast OVOC reaction rates) compared to VOCs (and VOC reaction rates) and so tends to have  $\text{RO}_2^- : \text{HO}_2 \lesssim 1$ .

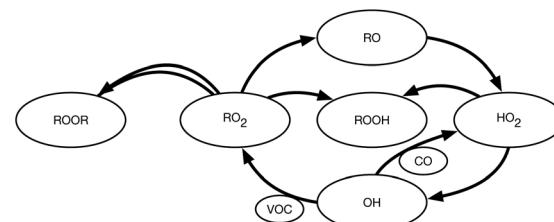


Fig. 1 Phenomenology for  $\text{HO}_x\text{--RO}_x$  radical cycling chemistry and volatile organic hydrocarbon (VOC) oxidation (without  $\text{NO}_x$ ). VOC oxidation leads to organo-peroxy radicals ( $\text{RO}_2$ ). If the  $\text{RO}_2$  do not terminate, they produce RO radicals, some of which rapidly convert to hydroperoxy radicals ( $\text{HO}_2$ ).  $\text{HO}_2$  can be converted back to OH, completing the cycle. For hydrocarbons,  $\text{RO}_2$  is almost inevitably produced before  $\text{HO}_2$  and termination, *etc.*, can make  $\text{RO}_2 : \text{HO}_2 \gg 1$ . However, oxygenated volatile organic carbon (OVOCs), or most simply  $\text{CO}$ , will convert OH to  $\text{HO}_2$  directly. Because the atmosphere is rich in  $\text{CO}$  and OVOCs,  $\text{RO}_2 : \text{HO}_2 \lesssim 1$ ; however, smog chambers often have very low OVOCs and so can have  $\text{RO}_2 : \text{HO}_2 \gg 1$ . This can bias “chamber” chemistry versus chemistry typical of the atmosphere.



In broad terms,  $\text{HO}_2$  and  $\text{RO}_2$  react with each other roughly 1 in every 10–30 collisions ( $k \approx 10^{-11} \text{ cm}^3 \text{ mol}^{-1} \text{ s}^{-1}$ ),<sup>37</sup> depending only weakly on the identity of the R group. A further simplification is that the  $\text{HO}_2 + \text{NO}$  reaction and all  $\text{RO}_2 + \text{NO}$  reactions (not shown) also react with about this frequency. Thus “high  $\text{NO}_x$ ” (really high NO) conditions occur when  $\text{NO} : \text{HO}_2 \gg 1$  and “low- $\text{NO}_x$ ” conditions when  $\text{NO} : \text{HO}_2 \ll 1$ .<sup>38</sup> Here we focus on low-NO conditions. A huge complication is that  $\text{RO}_2 + \text{RO}_2$  cross reaction kinetics are rich and variable, ranging from extremely slow to extremely fast.<sup>35,39,40</sup>

Many  $\text{RO}_2$  species are formed even during oxidation of a single VOC, and so there are many squared cross reactions; the real atmosphere is quadratically richer still. The rich and variable cross-reaction kinetics and relative paucity of kinetics data require parameterization of most  $\text{RO}_2$  cross reactions. Evidence suggests that electron-donating groups near the peroxy moiety (*i.e.* multiple  $\text{CH}_3$  groups such as *t*-butyl peroxy radicals) tend to raise the cross reaction barrier and reduce the rate coefficient precipitously, while electron-withdrawing groups (*i.e.* containing oxygen, such as peroxy acyl) tend to lower the barrier and raise the rate coefficient. Autoxidation involves successive internal H-atom transfers and oxygen addition in  $\text{RO}_2$ , and so the autoxidized  $\text{RO}_2$  (we shall refer to these as “ $\text{Ox}_n\text{RO}_2$ ” with  $n$  corresponding to the number of autoxidation steps) have multiple oxygenated functional groups and thus may have high cross reaction rate coefficients. There is still significant uncertainty as to just how fast the cross reactions get for relevant  $\text{Ox}_n\text{RO}_2$ . The peroxy acyl self reaction is fast, occurring in 1 out of 10 collisions. Some direct measurements and model simulations indicate that some  $\text{Ox}_n\text{RO}_2$  can react with each other on every collision,<sup>4,30</sup> while other studies suggest reaction probabilities closer to 1 in  $100^3$ . This may be modeling or measurement error, or it may simply reflect the enormous diversity of  $\text{RO}_2$  species and their cross reactions. In light of this, we follow Schervish and Donahue<sup>28</sup> using the general scheme first proposed by Madronich and Calvert.<sup>39</sup> We define “anchor” self-reaction rate coefficients for classes starting with tertiary- $\text{RO}_2$  (*t*-butyl  $\text{RO}_2$ ), extending through primary  $\text{RO}_2$  and then progressively more oxidized  $\text{RO}_2$ . The rate coefficients increase dramatically through this range. The cross-reaction rate coefficients are then given by the geometric mean, including a symmetry factor ( $k_{n,m} = 2\sqrt{k_{n,n}k_{m,m}}$ ).<sup>28,39</sup>

Schervish and Donahue<sup>28</sup> showed that the  $\text{RO}_2 + \text{HO}_2$  reaction was not competitive under any of the conditions of temperature or  $\text{NO}_x$  when simulating a chamber experiment without added OVOCs (and thus very high  $\text{RO}_2 : \text{HO}_2$ , as we will show). While under low  $\text{NO}_x$  conditions the  $\text{RO}_2 + \text{HO}_2$  reaction does occur, under the conditions studied in that work, it never built up appreciable yields. However, oxidation of CO to  $\text{CO}_2$  (as a surrogate for all OVOCs) converts OH directly to  $\text{HO}_2$ ; this reaction can become increasingly important, potentially dominating the yields. Therefore just as  $\text{NO}_x$  has been shown to suppress dimer yields, increasing  $\text{HO}_2$  is likely to have a similar effect.

Prior studies have shown that dimerization can be important under chamber conditions, but as the  $\text{RO}_2$  rate coefficients (and dimerization branching ratios) are a huge source of uncertainty

in this chemistry, it is difficult to say with certainty that dimerization is important in the real atmosphere.<sup>28,30</sup> Zhao *et al.*<sup>3</sup> suggest a minimum rate coefficient of  $10^{-12} \text{ cm}^3 \text{ mol}^{-1} \text{ s}^{-1}$  for relevant  $\text{RO}_2$  cross reactions while Berndt *et al.*<sup>4</sup> report some rate coefficients close to the kinetic limit of  $10^{-10} \text{ cm}^3 \text{ mol}^{-1} \text{ s}^{-1}$  but also find rate coefficients that span orders of magnitude for different dimers. Thus likely some dimers form very quickly while others form much more slowly. Those that form quickly may form under any relevant  $\text{RO}_2 : \text{HO}_2$  conditions in  $\alpha$ -pinene ozonolysis experiments (both in the chamber and in the atmosphere), while those that form slowly may appear when  $\text{RO}_2 : \text{HO}_2$  is high (in the chamber, especially in dark ozonolysis experiments), but not when  $\text{RO}_2 : \text{HO}_2$  is low (in the atmosphere).

Relatively low  $\text{RO}_2 : \text{HO}_2$  values in the atmosphere are due to high concentrations of OVOCs that promote direct conversion of OH to  $\text{HO}_2$ . In general,  $\text{RO}_2 : \text{HO}_2 \lesssim 1$  in the atmosphere, with higher values observed in low- $\text{NO}_x$  conditions where radical cycling and radical chain lengths tend to be lower. For example, Tan *et al.* reported  $\text{RO}_2 : \text{HO}_2 = 0.66$  in Beijing.<sup>33</sup> Stevens *et al.* report  $\text{RO}_2 : \text{HO}_2 > 1$  higher (around 15 for pristine conditions and 4 for slightly polluted conditions) at a remote site in Colorado, but still much lower than what we see in the model when no OVOCs are present.<sup>41</sup> Modeling done under low  $\text{NO}_x$  conditions for the marine atmosphere<sup>40</sup> and in rural Alabama<sup>42</sup> has also reported  $1 < \text{RO}_2 : \text{HO}_2 < 5$ , also much lower than our modeled  $\text{RO}_2 : \text{HO}_2$  without OVOCs.

Under chamber conditions with no or very low OVOC concentrations, OH reactivity is low because its main reaction is with the hydrocarbon of interest (for example  $\alpha$ -pinene). Since the chamber experiments we will consider here are designed to mimic the remote atmosphere, we are especially interested in atmospheric OH reactivity measurements under those conditions. Ferracci *et al.*<sup>43</sup> report measured OH reactivity from a suite of field campaigns across the world including many remote areas that span a range from  $4 \text{ s}^{-1}$  to almost  $50 \text{ s}^{-1}$ , the average being around  $15 \text{ s}^{-1}$ . Measurements of OH reactivity from the boreal forest site in Hyttiälä, Finland are also consistent with these values.<sup>44,45</sup> In a remote forest in the Great Lakes Region in northern Michigan, di Carlo *et al.*<sup>46</sup> estimate that only about 2% of a total OH reactivity of approximately  $13 \text{ s}^{-1}$  is attributable to monoterpenes.

A further complication is that in ozonolysis chamber experiments, the alkene + ozone reaction governs the overall radical production. This is rarely if ever true during daytime in the real atmosphere, where photolytic sources of  $\text{HO}_x$  are substantial.<sup>47</sup> Most alkenes also react with OH on almost every collision, whereas the background atmosphere has a richer distribution of OH reactivity. For all of these reasons, adding an OVOC surrogate such as CO to a chamber experiment is a potentially viable method to mimic the influence of OVOCs on the radical balance, but measures such as OH reactivity (and branching) as well as the actual  $\text{RO}_2 : \text{HO}_2$  and not simply reproducing ambient CO may better meet the experimental design objective.

This has significance when considering nucleation. Current knowledge suggests that ULVOC dimers are the species responsible for pure biogenic nucleation and lower



atmospheric  $\text{RO}_2$  to  $\text{HO}_2$  ratios may help explain why pure biogenic nucleation occurs in chamber experiments, but not always under seemingly similar atmospheric conditions.

Additionally, because the autoxidation pathway (and thus HOM production) is competitive with bimolecular termination reactions, it is strongly dependent on the conditions of the experiment or the atmosphere, including both the absolute radical concentrations as well as the radical ratios. Here we focus on the termination pathway of peroxy radicals with  $\text{HO}_2$  in order to show conditions under which this pathway becomes critical that are common in the atmosphere, but may not be represented in chamber experiments. However, the actual atmosphere rarely fits into what we here consider our more atmospherically relevant regime. This is because most often all of the potential peroxy-radical reactant pathways are available and competitive in the actual atmosphere (including reaction of acyl peroxy radicals with  $\text{NO}_2$  to form peroxy acyl nitrates<sup>48</sup> – which we omit for simplicity here). This speaks to how important it is for mechanistic representations of the atmosphere to be built based on experiments under which conditions are well-known so their extrapolation to more atmospheric conditions can be evaluated.

The major issue is this. If  $\text{RO}_2$  cross reactions are relatively slow (slower than  $\text{RO}_2 + \text{HO}_2$ ), then chambers with unrealistically high  $\text{RO}_2 : \text{HO}_2$  will misrepresent the atmosphere and possibly overproduce dimers and other ELVOC products. However, if  $\text{RO}_2$  cross reactions are fast (potentially 30 times faster than  $\text{RO}_2 + \text{HO}_2$ ), then they can still dominate termination for  $\text{RO}_2 : \text{HO}_2 \approx 1$ . Under those conditions,  $\text{RO}_2 : \text{HO}_2$  is less of a concern and chamber ULVOC yields, as well as nucleation and growth rates, are more likely to be directly relevant to the atmosphere. As it happens, a diversity of cross reactions may span these conditions, and reality likely lies somewhere in the middle. If this is so, it is especially important to gain a molecular level understanding to accurately predict atmospheric behavior.

## 2 The model

Our objective is to model the effect of OVOCs on  $\text{RO}_2 : \text{HO}_2$  and the associated oxidation products of a VOC such as  $\alpha$ -pinene. We are specifically interested in whether typical chamber conditions result in unrealistic net yields of the lowest volatility products such as the ROOR dimers (some of which are “Ultra Low Volatility Organic Compounds”, ULVOCs, in the volatility basis set classification<sup>28</sup>). High  $\text{RO}_2 : \text{HO}_2$  might even promote an unrealistic extent of autoxidation due to insufficient  $\text{HO}_2$  termination and lead to elevated yields of “Extremely Low Volatility Organic Compounds”, ELVOCs (which also include many less oxidized dimers). This could cause chamber experiments to show either higher nucleation rates (driven by ULVOCs) or higher sub 3 nm growth rates (driven by ELVOCs)<sup>13</sup> than would occur in the atmosphere, leading to an overall high bias in new-particle formation rates.

Any model sensitivity to chamber conditions (specifically  $\text{RO}_2 : \text{HO}_2$  and  $\text{RO}_2$  kinetics) would also suggest a corresponding sensitivity in the real atmosphere. It is possible that the real

world (especially under low- $\text{NO}_x$  conditions) is very sensitive to these kinetics, making this an important topic for future research.

To simulate OVOCs in the atmosphere we will simulate adding CO (also a potentially effective experimental technique) and explore the sensitivity of  $\text{RO}_2 : \text{HO}_2$  and thus the ULVOC and ELVOC yields to CO levels. Because both the autoxidation and  $\text{RO}_2$  (cross) termination kinetics are uncertain, we shall explore a phase space spanning potential rate coefficients.

### 2.1 $\text{RO}_2$ cross reactions and dimerization

We represent variability in  $\text{RO}_2$  cross reaction kinetics following the scheme first suggested by Madronich and Calvert.<sup>39</sup> The rate coefficient for any  $\text{RO}_2$  cross reaction is parameterized as the geometric mean of the self-reaction rate coefficients (multiplied by 2 because of the difference in symmetry) defined for a sequence of  $\text{RO}_2$  ranging from slow tertiary- $\text{RO}_2$  to fast hydroxy- $\text{RO}_2$ . We augment this by defining self-reaction rate coefficients for  $\text{Ox}_n\text{RO}_2$  representing 1, 2, and 3 steps of autoxidation, assuming that this will generally introduce electron withdrawing functional groups near the peroxy radical moiety. Here we investigate different scenarios of rate coefficients by changing the rate coefficient associated with  $\text{Ox}_1\text{RO}_2$ ,  $\text{Ox}_2\text{RO}_2$  and  $\text{Ox}_3\text{RO}_2$ , but leaving the  $\text{RO}_2$  and  $\text{Ox}_0\text{RO}_2$  the same in every scenario. The  $\text{RO}_2$  and the  $\text{Ox}_1\text{RO}_2$  rate coefficients are slow so that in every scenario, enough autoxidation occurs to allow the opportunity for dimers from the  $\text{Ox}_2\text{RO}_2$  and the  $\text{Ox}_3\text{RO}_2$  to form (consistent with observations).

We simulate 3 cases for the  $\text{Ox}_{1,2,3}\text{RO}_2$  cross reactions: a “fast” case, “middle” case, and a “slow” case that are described in Table 1. The rate coefficients range from  $10^{-13} \text{ cm}^3 \text{ mol}^{-1} \text{ s}^{-1}$  to  $10^{-10} \text{ cm}^3 \text{ mol}^{-1} \text{ s}^{-1}$ , which reflects the range of measured rate coefficients as well as the rate coefficients used in the other modeling studies.<sup>3,4,30</sup> The fast case here is the same case used in Schervish and Donahue.<sup>28</sup> Even in the fast case, the rate coefficients associated with  $\text{RO}_2$ ,  $\text{Ox}_0\text{RO}_2$ , and  $\text{Ox}_1\text{RO}_2$  are kept slow so that they don’t dominate over autoxidation preventing any  $\text{Ox}_2\text{RO}_2$  and  $\text{Ox}_3\text{RO}_2$  from forming.

### 2.2 Differences from previous implementation

Our model is described in detail in Schervish and Donahue,<sup>28</sup> however, we provide a brief description here and discuss some changes in the following sections. The model simulates the gas-phase chemistry, culminating in formation (and loss) of  $\text{RO}_2$  species, with  $\text{RO}_2$  termination mapping into the 2D-VBS *via*

**Table 1** Self reaction rate coefficients ( $\text{cm}^3 \text{ mol}^{-1} \text{ s}^{-1}$ ) for the three  $\text{RO}_2$  reactivity cases investigated here

Reaction	Fast $k$	Middle $k$	Slow $k$
$k(\text{RO}_2)$	$10^{-13}$	$10^{-13}$	$10^{-13}$
$k(\text{Ox}_0\text{RO}_2)$	$10^{-13}$	$10^{-13}$	$10^{-13}$
$k(\text{Ox}_1\text{RO}_2)$	$10^{-12}$	$10^{-13}$	$10^{-13}$
$k(\text{Ox}_2\text{RO}_2)$	$10^{-11}$	$10^{-12}$	$10^{-13}$
$k(\text{Ox}_3\text{RO}_2)$	$10^{-10}$	$10^{-12}$	$10^{-13}$



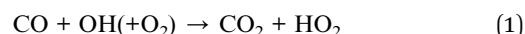
a succession of kernels specific to each termination step. The gas-phase chemical mechanism includes small-molecule  $\text{HO}_x$  and  $\alpha$ -pinene chemistry. The reactions begin with oxidation of a precursor, in this case  $\alpha$ -pinene *via* dark ozonolysis, producing peroxy radicals which then undergo two types of reactions: autoxidation or radical termination. The rate-limiting step in autoxidation is an internal H-atom transfer, so it is represented as a unimolecular process with a relatively large activation energy and thus a strong temperature dependence, immediately followed by  $\text{O}_2$  addition. We allow peroxy radicals ( $\text{Ox}_n\text{RO}_2$ ) to undergo up to three stages of autoxidation with  $n$  being the number of times the peroxy radical has autoxidized from 0 to 3. Each of these peroxy radicals is represented as a point in the 2D-VBS with more autoxidized peroxy radicals being progressively lower in volatility and higher in O : C. Here we will look at only cases with no  $\text{NO}_x$  present so the radical termination options available to each peroxy radical are unimolecular termination, reaction with  $\text{HO}_2$  to form a hydroperoxide, and dimerization.

Once a simulation has reached steady state with a stable production distribution, these products are mapped into the 2D-VBS using kernels anchored to the peroxy radical surrogate ( $\text{C}^*$  and O : C) that produced them. These kernels allow us to represent a wide variety of stabilization pathways producing a wide variety of different products through one surrogate species; the variety of species that the surrogate represents are instead mapped to a distribution of products within the 2D-VBS defined by a transformation relative to the surrogate  $\text{RO}_2$  volatility and O : C. Overall, the chemical mechanism will tell us how much, for example,  $\text{Ox}_2\text{ROOH}$  (termination with  $\text{HO}_2$  of a peroxy radical that has gone through two steps of autoxidation) is produced in the simulation and the kernels tell us where in the 2D-VBS to distribute all of that product.

**2.2.1  $\text{RO}_2$  kinetics.** As in Schervish and Donahue,<sup>28</sup> we represent the  $\text{RO}_2$  cross reactions as a branching between dimer formation and monomer formation following an  $\text{RO}_2$  association reaction. In that work we modeled the branching ratio based on the  $\text{RO}_2$  volatility. Here we assume any reaction between two autoxidized peroxy radicals ( $\text{Ox}_n\text{RO}_2$ ) will form a dimer 100% of the time. We also assume a dimer branching ratio for reactions between peroxy radicals that have not autoxidized ( $\text{RO}_2$ ) to be 0.05. For cross reactions between  $\text{Ox}_n\text{RO}_2$  and  $\text{RO}_2$ , we assume a branching ratio of 0.5. There is some evidence that the dimerization formation rate (the association rate coefficient times the branching ratio toward dimers) is representing average behavior that is slower than what we see in chamber experiments. This can be seen through measured rate coefficients for dimerization being around  $10^{-13} \text{ cm}^3 \text{ mol}^{-1} \text{ s}^{-1}$  at the slowest,<sup>49</sup> while we use that as the association rate coefficient, the actually dimerization rate coefficient is slowed down quite a bit for the least autoxidized peroxy radicals by the volatility dependence of the branching ratio. There is also evidence for higher branching toward dimers based on higher dimer yields in experiments under high  $\text{NO}_x$ <sup>50</sup> and low temperature<sup>31</sup> conditions than we see in the version of the model used in Schervish and Donahue.<sup>28</sup>

**2.2.2 Inclusion of CO.** When none of the OVOCs that promote conversion of OH to  $\text{HO}_2$  are present, the main

mechanism for  $\text{HO}_x$  cycling is through the reactions of  $\text{HO}_x$  with ozone. Under these conditions the conversion of OH to  $\text{HO}_2$  is slow leading to  $\text{HO}_2 : \text{OH} \approx 2$  and an  $\text{RO}_2 : \text{HO}_2 \approx 1000$ . When OVOCs are present we expect to see higher  $\text{HO}_2$  concentrations and a lower  $\text{RO}_2 : \text{HO}_2$ . We add CO as a proxy for all of these OVOCs and add the following reaction allowing another pathway for conversion of OH to  $\text{HO}_2$ .



**2.2.3  $\alpha$ -Pinene oxidation by OH.** Reflective of more recent work on the products of OH oxidation of  $\alpha$ -pinene, we have increased the fraction of peroxy radicals formed from OH reactions that are able to autoxidize to 0.2 from our previous value of 0.1.<sup>51,52</sup>

### 2.3 CLOUD CSTR simulation

In our previous study we simulated a chamber as a batch reactor and explored the time evolution of radicals and reaction products for different temperatures and  $\text{NO}_x$  concentrations. Here we simulate a Continuously Stirred Tank Reactor (CSTR, or steady-state chamber) with conditions that closely mimic those of the CERN CLOUD chamber. Specifically, we simulate a CSTR with a flushing timescale of 3 h and a steady inflow of VOC ( $\alpha$ -pinene), and ozone. All simulations are at 298 K. The CO concentration is fixed at concentrations between 0 and 10 000 ppb. For all of the results discussed below we allow the system to run to a steady state ( $\approx 5$  h after starting reagent inflow) and discuss only those steady-state values. The model parameters are in Table 2.

In chambers with fairly low aerosol loading (such as the CLOUD chamber), the major loss of condensable vapors is (presumably irreversible) deposition to the chamber walls. Under typical fan conditions in CLOUD, the  $\text{H}_2\text{SO}_4$  deposition timescale is roughly 4 minutes.<sup>53</sup> The organic vapors are heavier, with correspondingly lower diffusion constants, and so have deposition timescales of order 10 min.<sup>54,55</sup> This correspondence of vapor deposition timescales between CLOUD and remote ambient conditions is a design feature, though caution is required because turbulent deposition to the chamber walls scales as  $\sqrt{D}$  whereas laminar condensation to particles scales as  $D$  (or molecular speed in the kinetic regime).<sup>56,57</sup> In order to compare modeled ULVOC collision frequencies and modeled size-dependent growth rates with chamber data, we therefore add a wall-loss term to our simulation of the CSTR. Here that is  $0.0017 \text{ s}^{-1}$  (a 10 min lifetime).

Table 2 Model parameters to simulate typical  $\alpha$ -pinene + ozone CLOUD experiment

Parameter	Value
$\phi_{\alpha\text{-pinene}}$	$0.0018 \text{ ppb s}^{-1}$
$\phi_{\text{O}_3}$	$0.072 \text{ ppb s}^{-1}$
$k_{\text{flush}}$	$9.26 \times 10^{-5} \text{ s}^{-1} (\tau = 3 \text{ h})$
$k_{\text{wall}}$	$1.70 \times 10^{-3} \text{ s}^{-1} (\tau = 10 \text{ min})$



Here we simulate a dark ozonolysis experiment driven by the  $\alpha$ -pinene + ozone reaction. Unlike the atmosphere, where photochemical  $\text{HO}_x$  production is a major radical source during the day, in chamber experiments such as this the “heartbeat” of the chemistry – the intrinsic rate – is controlled by the ozonolysis rate. Alkene ozonolysis produces  $\text{OH}$ , and here we assume an  $\text{OH}$  yield of 0.8.<sup>58,59</sup> The critical  $\text{OH}$  branching shown in Fig. 1 reduces to a competition between  $\alpha$ -pinene ( $k_{\text{OH}} \simeq 5.4 \times 10^{-11} \text{ cm}^3 \text{ mol}^{-1} \text{ s}^{-1}$ ) and  $\text{CO}$  ( $k_{\text{OH}} \simeq 2.4 \times 10^{-13} \text{ cm}^3 \text{ mol}^{-1} \text{ s}^{-1}$ ). Thus we expect that  $\text{CO}$  will need to exceed  $\alpha$ -pinene by a factor of approximately 500 for  $\text{RO}_2 : \text{HO}_2$  to approach 1.

### 3 Results

#### 3.1 CO concentration and OH reactivity

Fig. 2 shows radical levels and  $\text{RO}_2 : \text{HO}_2$  vs.  $\text{CO}$  concentration, which are essentially identical for the three different  $\text{RO}_2$  reactivity scenarios (the figure shows the middle case).  $\text{RO}_2 : \text{HO}_2$  is inversely proportional to  $\text{CO}$  for much of the range (log–log slope = −1) and ultimately approaches a value near 2. For low

$\text{CO}$  it exceeds 100; these very large values are partly a consequence of the simplified  $\alpha$ -pinene mechanism that resolves  $\text{RO}_2$  products into the 2D-VBS instead of treating subsequent  $\text{RO}$  chemistry explicitly. The ratio is slightly lower when  $\text{RO}_2$  reactions are faster because the sink for  $\text{RO}_2$  via the cross reactions increases, leading to much lower steady-state  $\text{RO}_2$  concentrations, but only slightly lower  $\text{HO}_2$  concentrations. The model shows that very high  $\text{CO}$  concentrations are required to reach atmospherically relevant  $\text{RO}_2 : \text{HO}_2$ , but as  $\text{CO}$  is acting as a proxy for any compound whose oxidation would promote the conversion of  $\text{OH}$  to  $\text{HO}_2$ , these high concentrations merely represent what is necessary to reach that conversion rate. Also, at these high  $\text{CO}$  levels, the absolute  $\text{RO}_2$  and  $\text{HO}_2$  concentrations of  $2\text{--}5 \times 10^8 \text{ mole cm}^{-3}$  are reasonable for low-NO atmospheric conditions, suggesting that the termination chemistry will be appropriate. However, the  $\text{OH}$  has been almost completely scavenged by  $\text{CO}$ , and so multi-generation oxidation (aging) by  $\text{OH}$  will be negligible. This suggests that experiments conducted under these (high  $\text{CO}$ ) conditions would likely provide a good representation of first-generation atmospheric chemistry but miss any multi-generation chemistry important to the atmosphere. This is both a feature and a limitation.

Fig. 3 shows both the total  $\text{OH}$  reactivity and the fraction of  $\text{OH}$  reacting with the parent hydrocarbon ( $\alpha$ -pinene) vs. added  $\text{CO}$  in this simulated chamber experiment. Looking at  $\text{OH}$  reactivity we can see that when  $\text{CO}$  is absent, the  $\text{OH}$  reactivity is very low (around  $0.5 \text{ s}^{-1}$ ), much lower than reported  $\text{OH}$  reactivity in the remote atmosphere. About 1000–3000 ppb of  $\text{CO}$  is required to reproduce an  $\text{OH}$  reactivity of  $10\text{--}15 \text{ s}^{-1}$  typical of measurements in the remote atmosphere. In this range, the fraction of this reactivity that is due to  $\alpha$ -pinene, the only

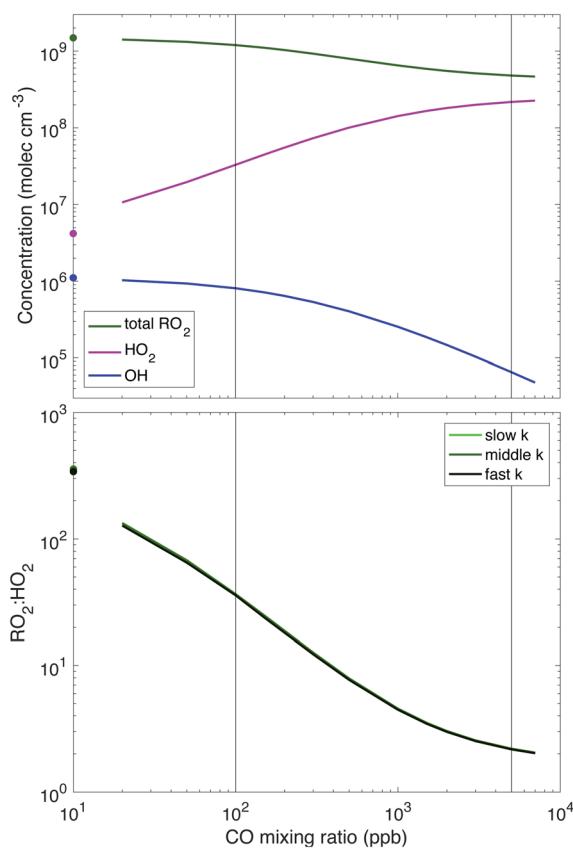


Fig. 2 Key radical levels vs.  $\text{CO}$  concentration in dark  $\alpha$ -pinene ozonolysis. Top: concentrations of total  $\text{RO}_2$ ,  $\text{HO}_2$ , and  $\text{OH}$ . Bottom:  $\text{RO}_2 : \text{HO}_2$ . All  $\text{RO}_2$  cross reaction scenarios give nearly identical results. Symbols on the  $y$ -axis indicate the values without  $\text{CO}$ . Added  $\text{CO}$  suppresses  $\text{OH}$  and enhances  $\text{HO}_2$  while slightly reducing  $\text{RO}_2$  (because  $\text{HO}_2$  decreases the  $\text{RO}_2$  lifetime). The vertical lines are at  $\text{CO}$  concentrations of 100 ppbv and 5000 ppbv. Values of  $\text{RO}_2 : \text{HO}_2$  below 10 (roughly the largest  $\text{RO}_2 : \text{HO}_2$  observed in the atmosphere) require  $\text{CO} \geq 1000 \text{ ppb}$ .

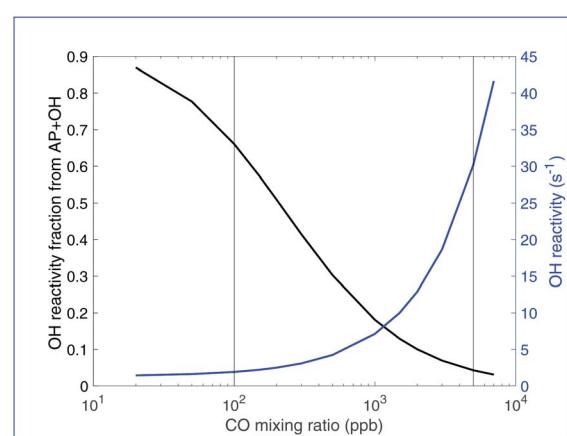


Fig. 3 OH reactivity vs.  $\text{CO}$ . The left  $y$ -axis and black curve show the fraction due to the  $\alpha$ -pinene +  $\text{OH}$  reaction and the right  $y$ -axis and blue curve show the absolute reactivity. At low  $\text{CO}$ ,  $\text{OH}$  reactivity is low and over 80% is attributable to the  $\alpha$ -pinene +  $\text{OH}$  reaction. As  $\text{CO}$  increases, the  $\text{OH}$  reactivity increases and the fraction due to  $\alpha$ -pinene decreases.  $\text{OH}$  reactivity in the remote atmosphere can range from  $4\text{--}50 \text{ s}^{-1}$ , but  $10\text{--}15 \text{ s}^{-1}$  is typical, with about 2% of the reactivity attributable to  $\alpha$ -pinene. In this plot that occurs in the range of 1000–3000 ppb of  $\text{CO}$ . The vertical lines at 100 ppb and 5000 ppb are the conditions of the steady-state VBS distributions in Fig. 4 and 5.

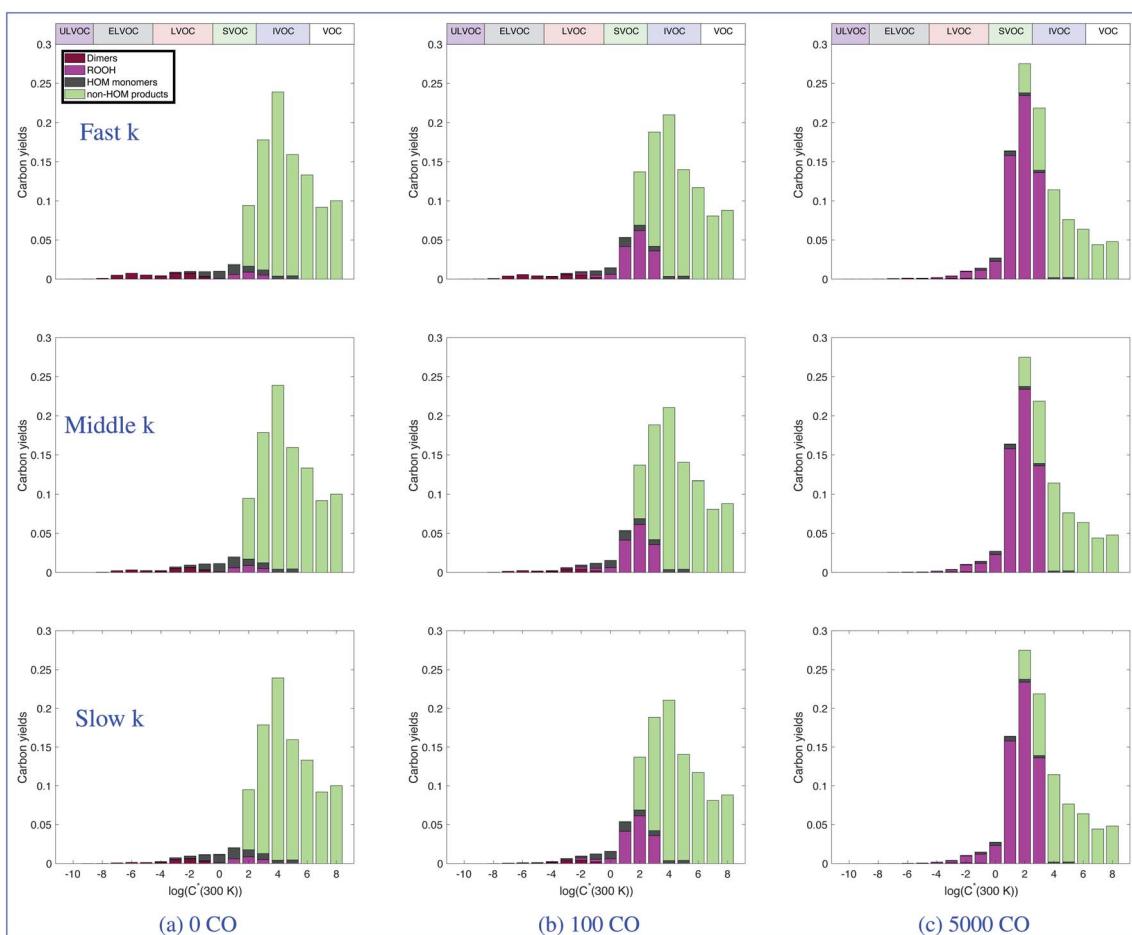


primary hydrocarbon in the model, is less than 10%. While primary hydrocarbons can account for up to 50% of the total OH reactivity in the atmosphere, according to di Carlo *et al.*<sup>46</sup> much of this is due to more abundant hydrocarbons like isoprene and a much lower fraction when only  $\alpha$ -pinene is present at a concentration of approximately 800 ppt, a lower percentage is completely reasonable. In this simulated experiment, especially because the governing  $\alpha$ -pinene + ozone reaction automatically produces RO<sub>2</sub> with a stoichiometry of 1, only a low fraction of OH reactivity can go to the  $\alpha$ -pinene to mimic the reaction topology shown in Fig. 1. This again indicates added CO of several ppm.

The vertical lines in Fig. 2 and 3 for CO = 100 and 5000 ppbv represent two cases we explore in greater depth below, along with the 0 CO case indicated with symbols on the y-axes. Relatively low CO between 0 and 100 ppbv is likely to represent typical conditions in chamber experiments without intentionally added CO (but instead just what is present in the balance air). The higher CO case, as shown, leads to OH reactivity and RO<sub>2</sub> : HO<sub>2</sub> that is more consistent with typical atmospheric values.

### 3.2 Dimers and ULVOCs

**3.2.1 Volatility basis set distributions.** Fig. 4 shows distributions of products for the three RO<sub>2</sub> rate coefficient cases at three different CO mixing ratios as a 1-dimensional volatility basis set. The various volatility classes described above are shown in the color-coded horizontal bar across the top of the figure. The vertical bars are color coded by the RO<sub>2</sub> termination pathway, as indicated in the legend. Each row shows a different RO<sub>2</sub> reactivity case – fast on top, middle in the middle, and slow on the bottom – and each column is a different steady-state CO concentration. The two non-zero CO concentrations are shown as vertical dashed lines in Fig. 2; as such the two left-hand columns represent a potential range for “unmodified” chambers with no-to-low CO in the carrier gas, while the right-hand column represents conditions with a more atmospheric RO<sub>2</sub> : HO<sub>2</sub>. Looking across any row of shows the change in product distributions with increasing CO, once again as a proxy for any OVOC that will promote direct conversion of OH to HO<sub>2</sub>. For zero CO (and thus low HO<sub>2</sub>), most of the RO<sub>2</sub> that does not undergo autoxidation terminates to molecules *via* a simplified



**Fig. 4** Volatility basis set distributions of the products from  $\alpha$ -pinene ozonolysis (including autoxidation and dimerization) at different CO mixing ratios for three different peroxy-radical reactivity scenarios. Each row shows a different reactivity case: the top row is the fast case, the middle row is the middle case, and the bottom row is the slow case. Each column shows a different CO mixing ratio in ppb, as indicated. Bar colors show the RO<sub>2</sub> termination pathway, while different volatility classes are indicated *via* the horizontal bars along the top. CO enhances HO<sub>2</sub> and tends to replace dimers (crimson) with hydroperoxides (magenta), reducing the yields of the lowest volatility products.



pathway shown in light green labeled “non-HOM products”. In every case, the product distributions are different at different CO concentrations. Most notably, as CO increases for every case there is a marked increase in the products from the  $\text{RO}_2 + \text{HO}_2$  reaction, shown in pink, and a decrease in the dimers, shown in crimson. When we look down a column, we see the effect of slowing down the  $\text{RO}_2$  kinetics has on the product distributions. Unsurprisingly we see a decrease in dimers as those reactions are slowed.

Notably, the “traditional”  $\alpha$ -pinene oxidation products (light green) predominate in all cases. Modifying  $\text{RO}_2 : \text{HO}_2$  changes the distribution across the board, but the yield changes in the SVOC region and above are modest, with some redistribution of products toward hydroperoxides (dark magenta) overall. Consequently, chamber experiments focused on mass yields in the  $1\text{--}1000 \mu\text{g m}^{-3}$  would see modest changes in overall mass yields with increased hydroperoxide content as  $\text{RO}_2 : \text{HO}_2$  decreases. This is consistent with observations.<sup>60,61</sup> Further, the very low yields of the entire “xLVOC” ensemble is notable and the variability with  $\text{RO}_2 : \text{HO}_2$  serves to emphasize that this tail

is highly sensitive to conditions. This serves as a cautionary tale on the one hand but also suggests that corresponding yields in the real atmosphere are likely also quite sensitive to conditions, and that this variability might well be important to atmospheric behavior.

Fig. 5 focuses on the lower volatility products. Here we can more clearly see the suppression of dimers both by increasing CO as well as by decreasing the  $\text{RO}_2$  rate coefficient. We can also very clearly see the effect on the ULVOCs. The only products that fall in the ULVOC range are the dimers formed from the self reaction of two  $\text{Ox}_3\text{RO}_2$  radicals and a small fraction of the cross reaction between an  $\text{Ox}_3\text{RO}_2$  and an  $\text{Ox}_2\text{RO}_2$ . Therefore it makes sense that the ULVOCs follow the same trend as the dimers in general. Looking across the top row (the fast case), we see low, but visible yields of ULVOCs with very low CO, but at the highest CO mixing ratio shown (representing more ambient  $\text{RO}_2 : \text{HO}_2$ ), the yield is approximately halved. In the middle case, the  $\text{RO}_2$  kinetics are still fast enough to sustain visible ULVOC yields at low CO, but the influence of CO is much more pronounced at 100 ppb of CO and there are almost no ULVOCs present at

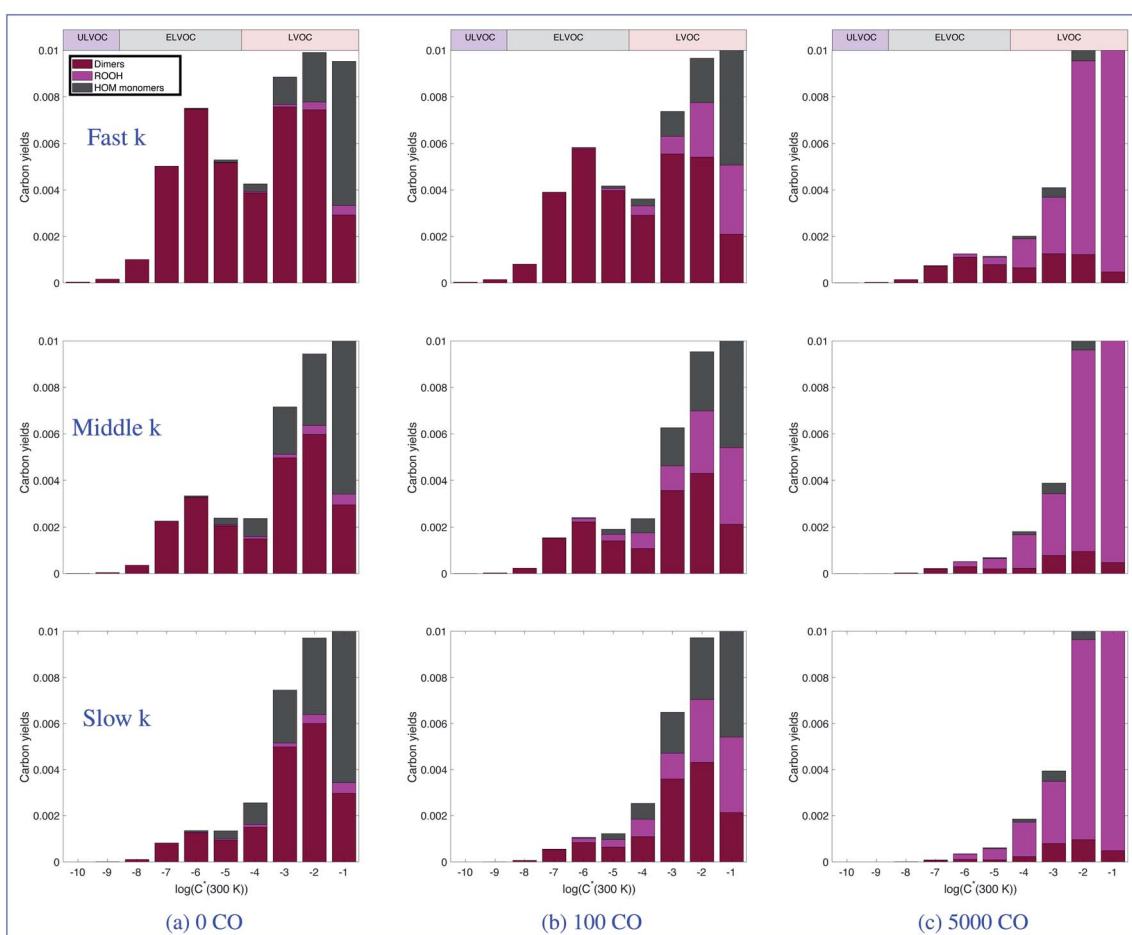
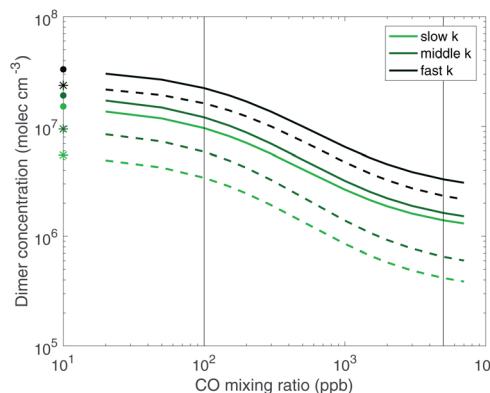


Fig. 5 Volatility basis set distributions of the products from  $\alpha$ -pinene ozonolysis (including autoxidation and dimerization) at different CO mixing ratios for three different peroxy-radical reactivity scenarios, focusing on the lowest volatility products. Each row shows a different reactivity case: the top row is the fast case, the middle row is the middle case, and the bottom row is the slow case. Each column shows a different CO mixing ratio in ppb, as indicated. Bar colors show the  $\text{RO}_2$  termination pathway, while different volatility classes are indicated via the horizontal bars along the top. The suppression of dimer formation by  $\text{HO}_2$  (promoted by CO) is clearly evident.



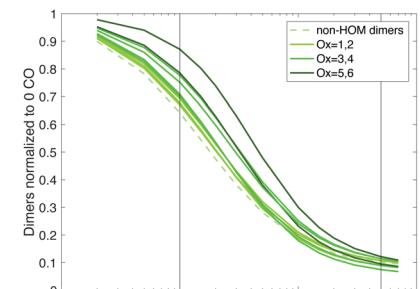


**Fig. 6** Total steady-state dimer concentrations vs. CO from  $\alpha$ -pinene ozonolysis simulation. The dashed lines show dimers formed from  $\text{RO}_2$  that have undergone at least one step of autoxidation. Adding CO to match atmospheric  $\text{RO}_2 : \text{HO}_2$  (the second vertical line) sharply reduces dimer concentrations in all cases. This shows that dimer yields are likely to be highly sensitive to  $\text{RO}_2 : \text{HO}_2$  over the range of values found in chambers and the real atmosphere, and that chamber dimer yields should be applied to atmospheric simulations with caution.

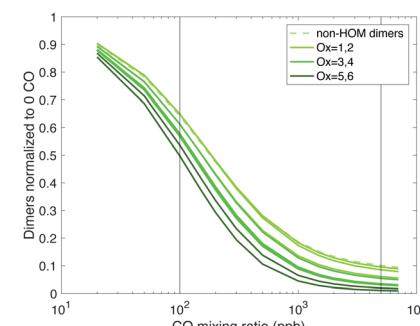
5000 ppb of CO. In the slow case, while the dimer trend is still noticeable, the ULVOC yields are notably lower even at low CO, so there is little visible influence of CO on the ULVOC yields. As nucleation is dependent on the amount of ULVOCs, we expect to see different nucleation rates at each of the different CO mixing ratios and for each of the  $\text{RO}_2$  reactivity scenarios.

**3.2.2 Overall yield sensitivity.** Fig. 6 shows the total dimers for each of the  $\text{RO}_2$  reactivity cases as well as the total oxidized dimers (total dimers minus the non-HOM dimers, formed from the reaction of peroxy radicals that do not undergo autoxidation) over a wide range of CO mixing ratios. The total dimer formation for the fast case is a factor of 7–8 lower at very high CO compared to low CO; most of the dimers in this case are HOM dimers. In the other cases, the dimers are reduced by about an order of magnitude with added CO. Broadly, adding CO and thus decreasing  $\text{RO}_2 : \text{HO}_2$  from an unrealistic to a more atmospherically relevant value can decrease the total dimer concentrations (yields) by almost an order of magnitude, regardless of the  $\text{RO}_2$  reactivity case considered. The total dimer concentrations are however also sensitive to the  $\text{RO}_2$  reactivity as well.

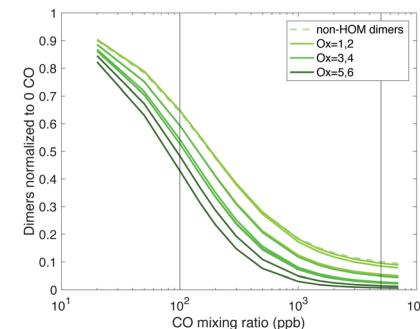
Total dimers are not a sufficiently precise measure of the actual situation; Fig. 5 shows that the dimers are distributed over almost 10 orders of magnitude in volatility. Thus, in Fig. 7 we show the relative change of each dimer class normalized by the zero CO case, where the dimer class is the total autoxidation number (*i.e.*  $\text{Ox} = 6$  is a dimer produced from  $\text{Ox}_3\text{RO}_2 + \text{Ox}_3\text{RO}_2$ ). We use a linear y-axis scale to emphasize the strength of these effects. Looking at the suppression of dimers due to CO (actually  $\text{HO}_2$ ), we see in the fast case, the oxidized dimers are suppressed far less at low CO mixing ratios when their cross reactions are fast. The dimers in the middle and slow cases are also consistently suppressed by CO, with the most oxidized dimers being suppressed slightly more aggressively by CO when their rate coefficient is slower. The concentration of the most



(a) Fast k



(b) Middle k



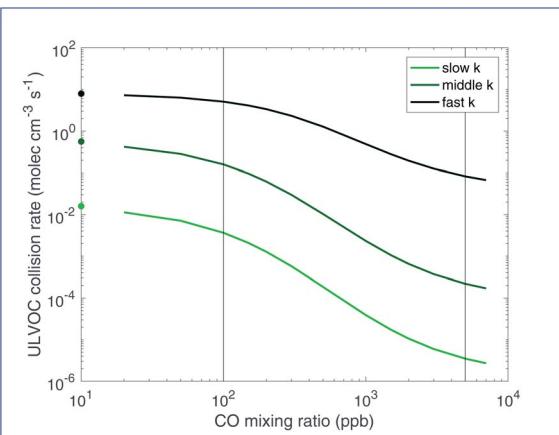
(c) Slow k

**Fig. 7** Dimer levels vs. CO (normalized to zero CO) from  $\alpha$ -pinene ozonolysis for different  $\text{RO}_2$  reactivity cases. CO suppresses dimer formation by enhancing  $\text{HO}_2$  concentrations and thus lowering  $\text{RO}_2 : \text{HO}_2$ . The effect of increasing CO is greatest on the slowest forming dimers and is very different (and less extreme) in the fast  $\text{RO}_2$  reactivity case, when the peroxy radical reactions with each other are faster than their reactions with  $\text{HO}_2$ .

oxidized dimers in the slow and middle cases is comparable to the fast case when no CO is present, but they are suppressed more aggressively by  $\text{HO}_2$  in the slow and middle cases, creating situations where we might measure appreciable dimers under low CO chamber conditions, but not under atmospheric conditions with lower  $\text{RO}_2 : \text{HO}_2$ .

**3.2.3 Nucleation.** Dimers are critical to nucleation, but not all dimers can contribute to nucleation as even some HOM dimers only reach the LVOC or ELVOC classes. Schervish and Donahue<sup>28</sup> proposed the ULVOC volatility class ( $\log C^* < -8.5$ ) based on evidence that compounds in this range are highly associated with observed “pure organic” nucleation. Cluster formation must be related to the collision rate of these ULVOC





**Fig. 8** ULVOC collision rates vs. CO for different peroxy radical reactivity cases following  $\alpha$ -pinene ozonolysis in a steady-state chamber such as the CLOUD chamber. This is the maximum likely nucleation rate. Symbols on the y-axis show the ULVOC collision rate at zero CO. For rapid  $\text{RO}_2$  reactions the ULVOC collision rate (and thus nucleation) is only weakly sensitive to CO (and thus  $\text{RO}_2 : \text{HO}_2$ ). However, for the slower  $\text{RO}_2$  reactivity cases the nucleation rate is highly sensitive to CO, suggesting that control of  $\text{RO}_2$  branching is important in chamber experiments.

species, and so we can compare modeled ULVOC collision rates with measured nucleation rates. To calculate the ULVOC collision rate we assume collision rate coefficient of  $k_{\text{coll}} = 3 \times 10^{-10} \text{ cm}^3 \text{ mol}^{-1} \text{ s}^{-1}$ , which we then multiply by the ULVOC concentration squared. In Fig. 8 we show the collision frequency of ULVOCs for the representative  $\text{RO}_2$  kinetic cases as a function of added CO (thus  $\text{RO}_2 : \text{HO}_2$ ).

## 4 Discussion

It is not news that low- $\text{NO}_x$  chamber experiments can deviate from atmospheric conditions by favoring  $\text{RO}_2$  cross reactions over  $\text{HO}_2$  reactions. However, the recent emergence of autoxidation and the potential for very rapid  $\text{RO}_2$  cross reactions involving highly oxidized  $\text{RO}_2$  radicals substantially complicates this issue. Further, new-particle formation appears to be especially sensitive to the covalently bound dimers formed from those cross reactions, because the high molecular weight combined with significant polar functionality can drive the products into the Ultra Low Volatility Organic Carbon (ULVOC) class. Our simulations show that yields of these dimers are sensitive to  $\text{RO}_2 : \text{HO}_2$  over the range of conditions likely to be found in chambers and in the atmosphere, and also that they are sensitive to a plausible range of  $\text{RO}_2$  cross reaction rate coefficients, which are currently not well constrained. The simulations also suggest that manipulating (and measuring)  $\text{RO}_2 : \text{HO}_2$  is chamber experiments is important. However, these dimers are also clearly visible in ambient measurements,<sup>62</sup> and so complete elimination of dimer formation under ambient conditions is unlikely.

Our results are generally consistent with the results from McFiggans *et al.*<sup>63</sup> McFiggans *et al.*<sup>63</sup> show that both monomer and dimers are drastically reduced in the presence of small

molecules that compete for  $\text{RO}_2$ , including due to the increase in  $\text{HO}_2$  caused by increasing CO. However, our simulation and the experiments presented in McFiggans *et al.*<sup>63</sup> differ significantly, making a direct comparison difficult. Specifically, we simulate dark ozonolysis with some OH chemistry at low CO, whereas OH is the primary oxidant in McFiggans *et al.*<sup>63</sup>; this different chemistry could cause the products to be different enough to lead to different volatility distributions when compared with our dark ozonolysis simulation. Notably, McFiggans *et al.*<sup>63</sup> see a decrease in SOA yields (with organic aerosol concentrations of  $1\text{--}10 \mu\text{g m}^{-3}$ ) in the  $\alpha$ -pinene + CO experiments while our simulation shows an increase in the ROOH yields in the SVOC range.

We have maintained the same reaction kernels from the original work in Schervish and Donahue;<sup>28</sup> we are focused on the dimer and ULVOC products. The major take-home message of our work is that we must consider the radical termination reactions carefully before jumping to extend results to the atmosphere, and so we elect not to “tune” the termination kernels to observed yields here. As all of the chemistry is resolved before the kernels are implemented to distribute the products in the VBS, the  $\text{RO}_2 + \text{HO}_2$  kernel does not affect any other products. One attempt to “tune” radical-VBS dimer yields to match both chamber and flow-tube dimer yields has been applied to a chemical transport model simulation of nucleation over the Amazon; there Zhao *et al.*<sup>64</sup> required a mixture of “high” and “low” autoxidation barrier  $\text{OxRO}_2$  in order to match experimental dimer yields. McFiggans *et al.*<sup>63</sup> also provide gas-phase mass spectrometry data that show large decreases in the  $\alpha$ -pinene produced monomers and dimers. Our dimer results are consistent with this, but we see little to no change in the LVOOC region where HOM monomers dominate. This discrepancy may be due to the limitations (specificity) of the nitrate chemical ionization source, which clusters less efficiently with less oxygenated and thus higher volatility products.<sup>12,13</sup> However, our focus here is on nucleation and especially dimer formation. In this our model is consistent with the results of McFiggans *et al.*<sup>63</sup> where in a mixture including  $\alpha$ -pinene and CO, dimers were reduced by a factor of 2 compared to a pure  $\alpha$ -pinene case.

We are especially interested in how these parameters affect nucleation, and how this might vary between chambers and the real atmosphere. Fig. 8 shows our modeled ULVOC collision rates are indeed highly sensitive to both  $\text{RO}_2 : \text{HO}_2$  and  $\text{RO}_2$  reactivity. The conditions simulated here are comparable to conditions in CLOUD where nucleation rates are observed between  $0.1$  and  $10 \text{ s}^{-1}$ .<sup>15</sup> Our modeled ULVOC collision frequencies at low CO are consistent with ULVOC nucleation efficiency for all  $\text{RO}_2$  reactivity cases considered, indicating that any of these association rate scenarios would be consistent with experimental findings at low CO mixing ratios. This is sensible as  $\text{RO}_2 : \text{HO}_2$  is high and  $\text{HO}_2$  termination is minor. However, extending to more atmospheric conditions, the system is very sensitive to added CO. When association is slowed down for the less oxidized peroxy radicals, high concentrations of  $\text{Ox}_2\text{RO}_2$  and  $\text{Ox}_3\text{RO}_2$  can be achieved and produce ULVOC dimers. With about 7000 ppb of CO present, the ULVOC collision rate is about 3 orders of magnitude lower than for zero CO in the middle



case. The slow case shows similar sensitivity. Overall, Fig. 8 shows that nucleation rates measured in chambers may extend to the real atmosphere, but that this is highly sensitive to the (uncertain)  $\text{RO}_2$  cross reaction kinetics.

The actual nucleation (stable cluster formation) rate is likely to depend on volatility within the ULVOC class (and also temperature, as the class itself moves with temperature-dependent volatility<sup>28</sup>). Observations also show that pure biogenic nucleation is highly sensitive to charge, with ions increasing observed new-particle formation rates by up to two orders of magnitude.<sup>15</sup> Just as lowered temperature stabilizes clusters, ions can as well; consequently it is plausible that the compounds at the lower end of the ELVOC class in Fig. 5 are more effective in nucleation within charged clusters. Broadly, we expect the nucleation efficiency to be a function of the xLVOC volatility distribution (predominantly ULVOC and ELVOC), with efficiency near 1 for collisions of the lowest volatility species and falling off toward higher volatility, modulated by effects such as charge. Implementation of schemes based on this phenomenology suggest that biogenic nucleation (especially ion induced) may be very important in the present-day tropical atmosphere over the Amazon.<sup>64</sup>

## 5 Conclusion

Chamber experiments are often run under conditions leading to very low  $\text{HO}_2$  concentrations and thus high  $\text{RO}_2$  to  $\text{HO}_2$  ratios ( $\text{RO}_2 : \text{HO}_2$ ) due to the low concentrations of oxidized volatile organic compounds that can convert OH directly to  $\text{HO}_2$ . High  $\text{RO}_2 : \text{HO}_2$  in turn can lead to high concentrations of dimers due to the lack of competition for  $\text{RO}_2$  from  $\text{HO}_2$ . Depending on the association rate coefficient for forming these dimers, they may or may not form when  $\text{HO}_2$  concentrations are higher in the real atmosphere. When  $\text{RO}_2 + \text{RO}_2$  rate coefficients are fast, dimer concentrations, and the resulting ULVOC dimer yields, stay high at all the CO concentrations (used as a proxy for all oxidized volatile organic compounds), but when  $\text{RO}_2 + \text{RO}_2$  rate coefficients are slow, dimers are formed in high concentrations when CO concentrations are low, but fall off quickly as CO increases. Therefore some dimers may be forming under high  $\text{RO}_2 : \text{HO}_2$  in chamber experiments, contributing to nucleation and/or growth that will exceed values in the real atmosphere. However, we also conclude that the system is likely sensitive to these same elements ( $\text{RO}_2 : \text{HO}_2$  as well as the  $\text{RO}_2$  cross reaction kinetics) in the real atmosphere, thus real-world new-particle formation is likely to be sensitive to these same factors. Thus it is crucial to constrain  $\text{RO}_2 : \text{HO}_2$  and  $\text{RO}_2$  branching in order to accurately extrapolate their results to the atmosphere. At a minimum it is highly advisable to manipulate  $\text{RO}_2 : \text{HO}_2$  by adding a reagent such as CO (potentially to quite high levels) in order to experimentally explore the chamber sensitivity to these conditions.

## Author contributions

MS and NMD conceived the project and designed the overall scope. MS implemented and executed the radical VBS model. MS and NMD wrote the manuscript.

## Conflicts of interest

The authors declare no conflicts.

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## Notes and references

- 1 M. Shrivastava, C. D. Cappa, J. Fan, A. H. Goldstein, A. B. Guenther, J. L. Jimenez, C. Kuang, A. Laskin, S. T. Martin, N. L. Ng, T. Petaja, J. R. Pierce, P. J. Rasch, P. Roldin, J. H. Seinfeld, J. Shilling, J. N. Smith, J. A. Thornton, R. Volkamer, J. Wang, D. R. Worsnop, R. A. Zaveri, A. Zelenyuk and Q. Zhang, *Rev. Geophys.*, 2017, **55**, 509–559.
- 2 D. Pagonis and P. J. Ziemann, *Aerosol Sci. Technol.*, 2018, **52**, 1178–1193.
- 3 Y. Zhao, J. A. Thornton and H. O. T. Pye, *Proc. Natl. Acad. Sci. U. S. A.*, 2018, **115**, 12142–12147.
- 4 T. Berndt, W. Scholz, B. Mentler, L. Fischer, H. Herrmann, M. Kulmala and A. Hansel, *Angew. Chem., Int. Ed.*, 2018, **57**, 3820–3824.
- 5 Z. Peng and J. L. Jimenez, *Chem. Soc. Rev.*, 2020, **49**, 2570–2616.
- 6 F. D. Lopez-Hilfiker, C. Mohr, M. Ehn, F. Rubach, E. Kleist, J. Wildt, T. F. Mentel, A. Lutz, M. Hallquist, D. Worsnop and J. A. Thornton, *Atmos. Meas. Tech.*, 2014, **7**, 983–1001.
- 7 T. Jokinen, M. Sipilä, H. Junninen, M. Ehn, G. Lönn, J. Hakala, T. Petäjä, R. L. Mauldin III, M. Kulmala and D. R. Worsnop, *Atmos. Chem. Phys.*, 2012, **12**, 4117–4125.
- 8 J. Kontkanen, K. Lehtipalo, L. Ahonen, J. Kangasluoma, H. E. Manninen, J. Hakala, C. Rose, K. Sellegrí, S. Xiao, L. Wang, X. Qi, W. Nie, A. Ding, H. Yu, S. Lee, V.-M. Kerminen, T. Petäjä and M. Kulmala, *Atmos. Chem. Phys.*, 2017, **17**, 2163–2187.
- 9 J. D. Crounse, L. B. Nielsen, S. Jørgensen, H. G. Kjaergaard and P. O. Wennberg, *J. Phys. Chem. Lett.*, 2013, **4**, 3513–3520.
- 10 M. Ehn, J. A. Thornton, E. Kleist, M. Sipila, H. Junninen, I. Pullinen, M. Springer, F. Rubach, R. Tillmann, B. Lee, F. Lopez-Hilfiker, S. Andres, I.-H. Acir, M. Rissanen, T. Jokinen, S. Schobesberger, J. Kangasluoma, J. Kontkanen, T. Nieminen, T. Kurten, L. B. Nielsen, S. Jørgensen, H. G. Kjaergaard, M. Canagaratna, M. D. Maso, T. Berndt, T. Petaja, A. Wahner, V.-M. Kerminen, M. Kulmala, D. R. Worsnop, J. Wildt and T. F. Mentel, *Nature*, 2014, **506**, 476–479.
- 11 F. Bianchi, T. Kurtén, M. Riva, C. Mohr, M. Rissanen, R. Pontus, T. Berndt, J. Crounse, P. Wennberg, T. F. Mentel, J. Wildt, H. Junninen, T. Jokinen, M. Kulmala, D. Worsnop, J. Thornton, N. M. Donahue, H. G. Kjaergaard and M. Ehn, *Chem. Rev.*, 2019, **119**, 3472–3509.
- 12 J. Tröstl, W. K. Chuang, M. Heinritzi, C. Yan, U. Molteni, L. Ahlm, C. Frege, F. Bianchi, R. Wagner, M. Simon,



- K. Lehtipalo, C. Williamson, J. S. Craven, J. Duplissy, A. Adamov, J. Almeida, A.-K. Bernhammar, M. Breitenlechner, S. Brilke, A. Dias, S. Ehrhart, R. C. Flagan, A. Franchin, C. Fuchs, H. Gordon, R. Guida, M. Gysel, A. Hansel, C. R. Hoyle, T. Jokinen, H. Junninen, J. Kangasluoma, H. Keskinen, J. Kim, M. Krapf, A. Kürten, A. Laaksonen, M. Lawler, M. Leiminger, S. Mathot, O. Möhler, T. Nieminen, A. Onnela, T. Petäjä, F. M. Piel, P. Miettinen, M. P. Rissanen, L. Rondo, N. Sarnela, S. Schobesberger, K. Sengupta, M. Sipilä, J. N. Smith, G. Steiner, A. Tomé, A. Virtanen, A. C. Wagner, E. Weingartner, D. Wimmer, P. M. Winkler, P. Ye, K. S. Carslaw, J. Curtius, J. Dommen, J. Kirkby, M. Kulmala, I. Riipinen, D. R. Worsnop, N. M. Donahue and U. Baltensperger, *Nature*, 2016, **530**, 527–531.
- 13 D. Stolzenburg, L. Fischer, A. L. Vogel, M. Heinritzi, M. Schervish, M. Simon, A. C. Wagner, L. Dada, L. R. Ahonen, A. Amorim, A. Baccarini, P. S. Bauer, B. Baumgartner, A. Bergen, F. Bianchi, M. Breitenlechner, S. Brilke, S. Buenrostro Mazon, D. Chen, A. Dias, D. C. Draper, J. Duplissy, I. El Haddad, H. Finkenzeller, C. Frege, C. Fuchs, O. Garmash, H. Gordon, X. He, J. Helm, V. Hofbauer, C. R. Hoyle, C. Kim, J. Kirkby, J. Kontkanen, A. Kürten, J. Lampilahti, M. Lawler, K. Lehtipalo, M. Leiminger, H. Mai, S. Mathot, B. Mentler, U. Molteni, W. Nie, T. Nieminen, J. B. Nowak, A. Ojdanic, A. Onnela, M. Passananti, T. Petäjä, L. L. J. Quéléver, M. P. Rissanen, N. Sarnela, S. Schallhart, C. Tauber, A. Tomé, R. Wagner, M. Wang, L. Weitz, D. Wimmer, M. Xiao, C. Yan, P. Ye, Q. Zha, U. Baltensperger, J. Curtius, J. Dommen, R. C. Flagan, M. Kulmala, J. N. Smith, D. R. Worsnop, A. Hansel, N. M. Donahue and P. M. Winkler, *Proc. Natl. Acad. Sci. U. S. A.*, 2018, **115**, 9122–9127.
- 14 F. Riccobono, S. Schobesberger, C. E. Scott, J. Dommen, I. K. Ortega, L. Rondo, J. Almeida, A. Amorim, F. Bianchi, M. Breitenlechner, A. David, A. Downard, E. Dunne, J. Duplissy, S. Ehrhart, R. C. Flagan, A. Franchin, A. Hansel, H. Junninen, M. Kajos, H. Keskinen, A. Kupc, O. Kupiainen, A. Kürten, T. Kurtén, A. N. Kvashin, A. Laaksonen, K. Lehtipalo, V. Makhmutov, S. Mathot, T. Nieminen, T. Olenius, A. Onnela, T. Petäjä, A. P. Praplan, F. D. Santos, S. Schallhart, J. H. Seinfeld, M. Sipilä, D. V. Spracklen, Y. Stozhkov, F. Stratmann, A. Tomé, G. Tsagkogeorgas, P. Vaattovaara, H. Vehkamäki, Y. Viisanen, A. Vrtala, P. E. Wagner, E. Weingartner, H. Wex, D. Wimmer, K. S. Carslaw, J. Curtius, N. M. Donahue, J. Kirkby, M. Kulmala, D. R. Worsnop and U. Baltensperger, *Science*, 2014, **344**, 717–721.
- 15 J. Kirkby, J. Duplissy, K. Sengupta, C. Frege, H. Gordon, C. Williamson, M. Heinritzi, M. Simon, C. Yan, J. Almeida, J. Tröstl, T. Nieminen, I. K. Ortega, R. Wagner, A. Adamov, A. Amorim, A.-K. Bernhammar, F. Bianchi, M. Breitenlechner, S. Brilke, X. Chen, J. Craven, A. Dias, S. Ehrhart, R. C. Flagan, A. Franchin, C. Fuchs, R. Guida, J. Hakala, C. R. Hoyle, T. Jokinen, H. Junninen, J. Kangasluoma, J. Kim, M. Krapf, A. Kürten, A. Laaksonen, K. Lehtipalo, V. Makhmutov, S. Mathot, U. Molteni, S. A. Monks, A. Onnela, O. Peräkylä, F. Piel, T. Petäjä, A. P. Praplan, K. J. Pringle, N. A. D. Richards, M. P. Rissanen, L. Rondo, N. Sarnela, S. Schobesberger, C. E. Scott, J. H. Seinfeld, S. Sharma, M. Sipilä, G. Steiner, Y. Stozhkov, F. Stratmann, A. Tomé, A. Virtanen, A. L. Vogel, A. C. Wagner, P. E. Wagner, E. Weingartner, D. Wimmer, P. M. Winkler, P. Ye, X. Zhang, A. Hansel, J. Dommen, N. M. Donahue, D. R. Worsnop, U. Baltensperger, M. Kulmala, J. Curtius and K. S. Carslaw, *Proc. Natl. Acad. Sci. U. S. A.*, 2016, **113**, 12053–12058.
- 22 K. S. Carslaw, H. Gordon, D. S. Hamilton, J. S. Johnson, L. A. Regayre, M. Yoshioka and K. J. Pringle, *Curr. Clim. Change Rep.*, 2017, **3**, 1–15.
- 23 A. Guenther, T. Karl, P. Harley, C. Wiedinmyer, P. I. Palmer and C. Geron, *Atmos. Chem. Phys.*, 2006, **6**, 3181–3210.
- 24 S. H. Chung and J. H. Seinfeld, *J. Geophys. Res.: Atmos.*, 2002, **107**, 4407.
- 25 T. Jokinen, M. Sipilä, S. Richters, V.-M. Kerminen, P. Paasonen, F. Stratmann, D. Worsnop, M. Kulmala, M. Ehn, H. Herrmann and T. Berndt, *Angew. Chem., Int. Ed.*, 2014, **53**, 14596–14600.
- 26 M. P. Rissanen, T. Kurtén, M. Sipilä, J. A. Thornton, J. Kangasluoma, N. Sarnela, H. Junninen, S. Jørgensen, K. Lehtipalo, V. Makhmutov, S. Mathot, U. Molteni, A. Onnela, O. Peräkylä, F. Piel, T. Petäjä, A. P. Praplan, K. Pringle, A. Rap, N. A. Richards, I. Riipinen, M. P. Rissanen, L. Rondo, N. Sarnela, S. Schobesberger, C. E. Scott, J. H. Seinfeld, M. Sipilä, G. Steiner, Y. Stozhkov, F. Stratmann, A. Tomé, A. Virtanen, A. L. Vogel, A. Wagner, P. E. Wagner, E. Weingartner, D. Wimmer, P. M. Winkler, P. Ye, X. Zhang, A. Hansel, J. Dommen, N. M. Donahue, D. R. Worsnop, U. Baltensperger, M. Kulmala, K. S. Carslaw and J. Curtius, *Nature*, 2016, **530**, 521–526.
- 16 J. R. Pierce and P. J. Adams, *Atmos. Chem. Phys.*, 2009, **9**, 1339–1356.
- 17 J. Merikanto, D. V. Spracklen, G. W. Mann, S. J. Pickering and K. S. Carslaw, *Atmos. Chem. Phys.*, 2009, **9**, 8601–8616.
- 18 D. V. Spracklen, K. S. Carslaw, M. Kulmala, V.-M. Kerminen, S.-L. Sihto, I. Riipinen, J. Merikanto, G. W. Mann, M. P. Chipperfield, A. Wiedensohler, W. Birmili and H. Lihavainen, *Geophys. Res. Lett.*, 2008, **35**, L06808.
- 19 M. Wang and J. E. Penner, *Atmos. Chem. Phys.*, 2009, **9**, 239–260.
- 20 J. Kazil, P. Stier, K. Zhang, J. Quaas, S. Kinne, D. O'Donnell, S. Rast, M. Esch, S. Ferrachat, U. Lohmann and J. Feichter, *Atmos. Chem. Phys.*, 2010, **10**, 10733–10752.
- 21 H. Gordon, K. Sengupta, A. Rap, J. Duplissy, C. Frege, C. Williamson, M. Heinritzi, M. Simon, C. Yan, J. Almeida, J. Tröstl, T. Nieminen, I. K. Ortega, R. Wagner, E. M. Dunne, A. Adamov, A. Amorim, A.-K. Bernhammar, F. Bianchi, M. Breitenlechner, S. Brilke, X. Chen, J. S. Craven, A. Dias, S. Ehrhart, L. Fischer, R. C. Flagan, A. Franchin, C. Fuchs, R. Guida, J. Hakala, C. R. Hoyle, T. Jokinen, H. Junninen, J. Kangasluoma, J. Kim, J. Kirkby, M. Krapf, A. Kürten, A. Laaksonen, K. Lehtipalo, V. Makhmutov, S. Mathot, U. Molteni, S. A. Monks, A. Onnela, O. Peräkylä, F. Piel, T. Petäjä, A. P. Praplan, K. J. Pringle, N. A. D. Richards, M. P. Rissanen, L. Rondo, N. Sarnela, S. Schobesberger, C. E. Scott, J. H. Seinfeld, S. Sharma, M. Sipilä, G. Steiner, Y. Stozhkov, F. Stratmann, A. Tomé, A. Virtanen, A. L. Vogel, A. C. Wagner, P. E. Wagner, E. Weingartner, D. Wimmer, P. M. Winkler, P. Ye, X. Zhang, A. Hansel, J. Dommen, N. M. Donahue, D. R. Worsnop, U. Baltensperger, M. Kulmala, J. Curtius and K. S. Carslaw, *Proc. Natl. Acad. Sci. U. S. A.*, 2016, **113**, 12053–12058.

- S. Schallhart, M. K. Kajos, R. Taipale, M. Springer, T. F. Mentel, T. Ruuskanen, T. Petäjä, D. R. Worsnop, H. G. Kjaergaard and M. Ehn, *J. Am. Chem. Soc.*, 2014, **136**, 15596–15606.
- 27 N. M. Donahue, S. A. Epstein, S. N. Pandis and A. L. Robinson, *Atmos. Chem. Phys.*, 2011, **11**, 3303–3318.
- 28 M. Schervish and N. M. Donahue, *Atmos. Chem. Phys.*, 2020, **20**, 1183–1199.
- 29 C. Frege, I. K. Ortega, M. P. Rissanen, A. P. Praplan, G. Steiner, M. Heinritzi, L. Ahonen, A. Amorim, A.-K. Bernhammer, F. Bianchi, S. Brilke, M. Breitenlechner, L. Dada, A. Dias, J. Duplissy, S. Ehrhart, I. El-Haddad, L. Fischer, C. Fuchs, O. Garmash, M. Gonin, A. Hansel, C. R. Hoyle, T. Jokinen, H. Junninen, J. Kirkby, A. Kürten, K. Lehtipalo, M. Leiminger, R. L. Mauldin, U. Molteni, L. Nichman, T. Petäjä, N. Sarnela, S. Schobesberger, M. Simon, M. Sipilä, D. Stolzenburg, A. Tomé, A. L. Vogel, A. C. Wagner, R. Wagner, M. Xiao, C. Yan, P. Ye, J. Curtius, N. M. Donahue, R. C. Flagan, M. Kulmala, D. R. Worsnop, P. M. Winkler, J. Dommen and U. Baltensperger, *Atmos. Chem. Phys.*, 2018, **18**, 65–79.
- 30 U. Molteni, M. Simon, M. Heinritzi, C. R. Hoyle, A.-K. Bernhammer, F. Bianchi, M. Breitenlechner, S. Brilke, A. Dias, J. Duplissy, C. Frege, H. Gordon, C. Heyn, T. Jokinen, A. Kürten, K. Lehtipalo, V. Makhmutov, T. Petäjä, S. M. Pieber, A. P. Praplan, S. Schobesberger, G. Steiner, Y. Stozhkov, A. Tomé, J. Tröstl, A. C. Wagner, R. Wagner, C. Williamson, C. Yan, U. Baltensperger, J. Curtius, N. M. Donahue, A. Hansel, J. Kirkby, M. Kulmala, D. R. Worsnop and J. Dommen, *Earth Space Chem.*, 2019, **3**, 873–883.
- 31 Q. Ye, M. Wang, V. Hofbauer, D. Stolzenburg, D. Chen, M. Schervish, A. Vogel, R. L. Mauldin, R. Baalbaki, S. Brilke, L. Dada, A. Dias, J. Duplissy, I. El Haddad, H. Finkenzeller, L. Fischer, X. He, C. Kim, A. Kürten, H. Lamkaddam, C. P. Lee, K. Lehtipalo, M. Leiminger, H. E. Manninen, R. Marten, B. Mentler, E. Partoll, T. Petäjä, M. Rissanen, S. Schobesberger, S. Schuchmann, M. Simon, Y. J. Tham, M. Vazquez-Pufleau, A. C. Wagner, Y. Wang, Y. Wu, M. Xiao, U. Baltensperger, J. Curtius, R. Flagan, J. Kirkby, M. Kulmala, R. Volkamer, P. M. Winkler, D. Worsnop and N. M. Donahue, *Environ. Sci. Technol.*, 2019, **53**, 12357–12365.
- 32 M. Simon, L. Dada, M. Heinritzi, W. Scholz, D. Stolzenburg, L. Fischer, A. C. Wagner, A. Kürten, B. Rörup, X.-C. He, J. Almeida, R. Baalbaki, A. Baccarini, P. S. Bauer, L. Beck, A. Bergen, F. Bianchi, S. Bräkling, S. Brilke, L. Caudillo, D. Chen, B. Chu, A. Dias, D. C. Draper, J. Duplissy, I. El Haddad, H. Finkenzeller, C. Frege, L. Gonzalez-Carracedo, H. Gordon, M. Granzin, J. Hakala, V. Hofbauer, C. R. Hoyle, C. Kim, W. Kong, H. Lamkaddam, C. P. Lee, K. Lehtipalo, M. Leiminger, H. Mai, H. E. Manninen, G. Marie, R. Marten, B. Mentler, U. Molteni, L. Nichman, W. Nie, A. Ojdanic, A. Onnela, E. Partoll, T. Petäjä, J. Pfeifer, M. Philippov, L. L. J. Quéléver, A. Ranjithkumar, M. Rissanen, S. Schallhart, S. Schobesberger, S. Schuchmann, J. Shen, M. Sipilä, G. Steiner, Y. Stozhkov, C. Tauber, Y. J. Tham, A. R. Tomé, M. Vazquez-Pufleau, A. Vogel, R. Wagner, M. Wang, D. S. Wang, Y. Wang, S. K. Weber, Y. Wu, M. Xiao, C. Yan, P. Ye, Q. Ye, M. Zauner-Wieczorek, X. Zhou, U. Baltensperger, J. Dommen, R. C. Flagan, A. Hansel, M. Kulmala, R. Volkamer, P. M. Winkler, D. R. Worsnop, N. M. Donahue, J. Kirkby and J. Curtius, *Atmos. Chem. Phys.*, 2020, **20**, 9183–9207.
- 33 Z. Tan, K. Lu, A. Hofzumahaus, H. Fuchs, B. Bohn, F. Holland, Y. Liu, F. Rohrer, M. Shao, K. Sun, Y. Wu, L. Zeng, Y. Zhang, Q. Zou, A. Kiendler-Scharr, A. Wahner and Y. Zhang, *Atmos. Chem. Phys.*, 2019, **19**, 7129–7150.
- 34 Z. Tan, F. Rohrer, K. Lu, X. Ma, B. Bohn, S. Broch, H. Dong, H. Fuchs, G. I. Gkatzelis, A. Hofzumahaus, F. Holland, X. Li, Y. Liu, Y. Liu, A. Novelli, M. Shao, H. Wang, Y. Wu, L. Zeng, M. Hu, A. Kiendler-Scharr, A. Wahner and Y. Zhang, *Atmos. Chem. Phys.*, 2018, **18**, 12391–12411.
- 35 G. S. Tyndall, R. A. Cox, C. Granier, R. Lesclaux, G. K. Moortgat, M. J. Pilling, A. R. Ravishankara and T. J. Wallington, *J. Geophys. Res.*, 2001, **106**, 12157.
- 36 R. Atkinson, D. L. Baulch, R. A. Cox, J. N. Crowley, R. F. Hampson, R. G. Hynes, M. E. Jenkin, M. J. Rossi, J. Troe and T. J. Wallington, *Atmos. Chem. Phys.*, 2008, **8**, 4141–4496.
- 37 R. Atkinson, D. L. Baulch, R. A. Cox, J. N. Crowley, R. F. Hampson, R. G. Hynes, M. E. Jenkin, M. J. Rossi, J. Troe and I. Subcommittee, *Atmos. Chem. Phys.*, 2006, **6**, 3625–4055.
- 38 P. O. Wennberg, *IGAC Newsletter No. 50 - July 2013*, 2013, <https://issuu.com/igacnews/docs/igac/3>.
- 39 S. Madronich and J. G. Calvert, *J. Geophys. Res.: Atmos.*, 1990, **95**, 5697.
- 40 N. M. Donahue and R. G. Prinn, *J. Geophys. Res.: Atmos.*, 1990, **95**, 18387–18411.
- 41 P. S. Stevens, J. H. Mather, W. H. Brune, F. Eisele, D. Tanner, A. Jefferson, C. Cantrell, R. Shetter, S. Sewall, A. Fried, B. Henry, E. Williams, K. Baumann, P. Goldan and W. Kuster, *J. Geophys. Res.: Atmos.*, 1997, **102**, 6379–6391.
- 42 C. A. Cantrell, J. A. Lind, R. E. Shetter, J. G. Calvert, P. D. Goldan, W. Kuster, F. C. Fehsenfeld, S. A. Montzka, D. D. Parrish, E. J. Williams, M. P. Buhr, H. H. Westberg, G. Allwine and R. Martin, *J. Geophys. Res.: Atmos.*, 1992, **97**, 20671–20686.
- 43 V. Ferracci, I. Heimann, N. L. Abraham, J. A. Pyle and A. T. Archibald, *Atmos. Chem. Phys.*, 2018, **18**, 7109–7129.
- 44 V. Sinha, J. Williams, J. Lelieveld, T. Ruuskanen, M. Kajos, J. Patokoski, H. Hellen, H. Hakola, D. Mogensen, M. Boy, J. Rinne and M. Kulmala, *Environ. Sci. Technol.*, 2010, **44**, 6614–6620.
- 45 A. C. Nölscher, J. Williams, V. Sinha, T. Custer, W. Song, A. M. Johnson, R. Axinte, H. Bozem, H. Fischer, N. Pouvesle, G. Phillips, J. N. Crowley, P. Rantala, J. Rinne, M. Kulmala, D. Gonzales, J. Valverde-Canossa, A. Vogel, T. Hoffmann, H. G. Ouwersloot, J. Vilà-Guerau de Arellano and J. Lelieveld, *Atmos. Chem. Phys.*, 2012, **12**, 8257–8270.
- 46 P. di Carlo, W. H. Brune, M. Martinez, H. Harder, R. Lesher, X. Ren, T. Thornberry, M. A. Carroll, V. Young, P. B. Shepson,



- D. Riemer, E. Apel and C. Campbell, *Science*, 2004, **304**, 722–725.
- 47 J. A. Logan, M. J. Prather, S. C. Wofsy and M. B. McElroy, *J. Geophys. Res.: Atmos.*, 1981, **86**, 7210–7254.
- 48 H. L. Singh, L. J. Salas, B. A. Ridley, J. D. Shetter, N. M. Donahue, F. C. Fehsenfeld, D. W. Fahey, D. D. Parish, E. J. Williams, S. C. Liu, G. Hübner and P. C. Murphy, *Nature*, 1985, **318**, 347–349.
- 49 H. O. T. Pye, E. L. D'Ambro, B. H. Lee, S. Schobesberger, M. Takeuchi, Y. Zhao, F. Lopez-Hilfiker, J. Liu, J. E. Shilling, J. Xing, R. Mathur, A. M. Middlebrook, J. Liao, A. Welti, M. Graus, C. Warneke, J. A. de Gouw, J. S. Holloway, T. B. Ryerson, I. B. Pollack and J. A. Thornton, *Proc. Natl. Acad. Sci. U. S. A.*, 2019, **116**, 6641–6646.
- 50 C. Yan, W. Nie, A. L. Vogel, L. Dada, K. Lehtipalo, D. Stolzenburg, R. Wagner, M. P. Rissanen, M. Xiao, L. Ahonen, L. Fischer, C. Rose, F. Bianchi, H. Gordon, M. Simon, M. Heinritzi, O. Garmash, P. Roldin, A. Dias, P. Ye, V. Hofbauer, A. Amorim, P. S. Bauer, A. Bergen, A.-K. Bernhammer, M. Breitenlechner, S. Brilke, A. Buchholz, S. B. Mazon, M. R. Canagaratna, X. Chen, A. Ding, J. Dommen, D. C. Draper, J. Duplissy, C. Frege, C. Heyn, R. Guida, J. Hakala, L. Heikkilä, C. R. Hoyle, T. Jokinen, J. Kangasluoma, J. Kirkby, J. Kontkanen, A. Kürten, M. J. Lawler, H. Mai, S. Mathot, R. L. Mauldin, U. Molteni, L. Nichman, T. Nieminen, J. Nowak, A. Ojdanic, A. Onnela, A. Pajunoja, T. Petäjä, F. Piel, L. L. J. Quéléver, N. Sarnela, S. Schallhart, K. Sengupta, M. Sipilä, A. Tomé, J. Tröstl, O. Väistö, A. C. Wagner, A. Ylirsirniö, Q. Zha, U. Baltensperger, K. S. Carslaw, J. Curtius, R. C. Flagan, A. Hansel, I. Riipinen, J. N. Smith, A. Virtanen, P. M. Winkler, N. M. Donahue, V.-M. Kerminen, M. Kulmala, M. Ehn and D. R. Worsnop, *Sci. Adv.*, 2020, **6**, eaay4945.
- 51 T. Berndt, S. Richters, T. Jokinen, N. Hyttinen, T. Kurtén, R. V. Otkjær, H. G. Kjaergaard, F. Stratmann, H. Herrmann, M. Sipilä, M. Kulmala and M. Ehn, *Nat. Commun.*, 2016, **7**, 13677.
- 52 T. Berndt, S. Richters, R. Kaethner, J. Voigtländer, F. Stratmann, M. Sipilä, M. Kulmala and H. Herrmann, *J. Phys. Chem. A*, 2015, **119**, 10336–10348.
- 53 J. Duplissy, J. Merikanto, A. Franchin, G. Tsagkogeorgas, J. Kangasluoma, D. Wimmer, H. Vuollekoski, S. Schobesberger, K. Lehtipalo, R. C. Flagan, D. Brus, N. M. Donahue, H. Vehkämäki, J. Almeida, A. Amorim, P. Barmet, F. Bianchi, M. Breitenlechner, E. M. Dunne, R. Guida, H. Henschel, H. Junninen, J. Kirkby, A. Kürten, A. Kupc, A. Määttänen, V. Makhmutov, S. Mathot, T. Nieminen, A. Onnela, A. P. Praplan, F. Riccobono, L. Rondo, G. Steiner, A. Tome, H. Walther, U. Baltensperger, K. S. Carslaw, J. Dommen, A. Hansel, T. Petäjä, M. Sipilä, F. Stratmann, A. Vrtala, P. E. Wagner, D. R. Worsnop, J. Curtius and M. Kulmala, *J. Geophys. Res.: Atmos.*, 2016, **212**, 1752–1775.
- 54 J. E. Krechmer, D. Pagonis, P. J. Ziemann and J. L. Jimenez, *Environ. Sci. Technol.*, 2016, **50**, 5757–5765.
- 55 P. Ye, X. Ding, J. Hakala, V. Hofbauer, E. S. Robinson and N. M. Donahue, *Aerosol Sci. Technol.*, 2016, **50**, 822–834.
- 56 A. Kürten, C. Williamson, J. Almeida, J. Kirkby and J. Curtius, *Atmos. Chem. Phys.*, 2015, **15**, 4063–4075.
- 57 N. Donahue, W. Chuang and M. Schervish, *Advances in Atmospheric Chemistry: Organic Oxidation and Multiphase Chemistry*, World Scientific, 2019, ch. 4, vol. 2, pp. 199–317.
- 58 A. A. Presto and N. M. Donahue, *J. Phys. Chem. A*, 2004, **108**, 9096–9104.
- 59 S. E. Paulson, M. Chung, A. D. Sen and G. Orzechowska, *J. Geophys. Res.: Atmos.*, 1998, **103**, 25533–25539.
- 60 K. M. Henry and N. M. Donahue, *Aerosol Sci. Technol.*, 2011, **45**, 686–690.
- 61 M. Krapf, I. E. Haddad, E. A. Bruns, U. Molteni, K. R. Daellenbach, A. S. Prévôt, U. Baltensperger and J. Dommen, *Chem.*, 2016, **1**, 603–616.
- 62 T. Jokinen, J. Kontkanen, K. Lehtipalo, H. E. Manninen, J. Aalto, A. Porcar-Castell, O. Garmash, T. Nieminen, M. Ehn, J. Kangasluoma, H. Junninen, J. Levula, J. Duplissy, L. R. Ahonen, P. Rantala, L. Heikkilä, C. Yan, M. Sipilä, D. R. Worsnop, J. Bäck, T. Petäjä, V.-M. Kerminen and M. Kulmala, *Sci. Rep.*, 2017, **7**, 45707.
- 63 G. McFiggans, T. F. Mentel, J. Wildt, I. Pullinen, S. Kang, E. Kleist, S. Schmitt, M. Springer, R. Tillmann, C. Wu, D. Zhao, M. Hallquist, C. Faxon, M. Le Breton, Å. M. Hallquist, D. Simpson, R. Bergström, M. E. Jenkins, M. Ehn, J. A. Thornton, M. R. Alfarra, T. J. Bannan, C. J. Percival, M. Priestley, D. Topping and A. Kiendler-Scharr, *Nature*, 2019, **565**, 587–593.
- 64 B. Zhao, M. Shrivastava, N. M. Donahue, H. Gordon, M. Schervish, J. E. Shilling, R. A. Zaveri, J. Wang, M. O. Andreae, C. Zhao, B. Gaudet, Y. Liu, J. Fan and J. D. Fast, *Proc. Natl. Acad. Sci. U. S. A.*, 2020, **117**, 25344–25351.

