



Cite this: *Environ. Sci.: Atmos.*, 2025, 5, 848

## Formation of the aminoperoxy radical in the atmospheric oxidation of ammonia†

Vili-Taneli Salo, <sup>ab</sup> Jing Chen <sup>b</sup> and Henrik G. Kjaergaard <sup>\*b</sup>

Atmospheric oxidation of ammonia is initiated by its reaction with the hydroxyl radical, producing the aminyl radical ( $\text{NH}_2$ ). Thus far, it has been believed that the subsequent fate of  $\text{NH}_2$  is to react bimolecularly with other atmospheric trace gases like  $\text{NO}$ ,  $\text{NO}_2$ , or  $\text{O}_3$ . Its reaction with  $\text{O}_2$  has been considered insignificant under atmospheric conditions. However, this is based on a rate coefficient that is orders of magnitude smaller than those known for analogous reactions of  $\text{O}_2$  with carbon-, sulfur-, and other nitrogen-centered radicals. We demonstrate by multireference calculations and kinetic modelling that the reaction of  $\text{NH}_2$  and  $\text{O}_2$  leading to the formation of the aminoperoxy radical ( $\text{NH}_2\text{O}_2$ ) occurs with a rate coefficient similar to those of the aforementioned analogous radicals. We show that the previously estimated small rate coefficient is due to an unimolecular rate limiting step in the formation of measured products rather than the initial  $\text{NH}_2 + \text{O}_2$  reaction. The lack of experimental detection of  $\text{NH}_2\text{O}_2$  in the existing literature is likely due to the experiments being conducted at either high temperature or low pressure. We show that the atmospheric presence of  $\text{NH}_2\text{O}_2$  depends greatly on atmospheric conditions. Its formation is an important, yet previously overlooked pathway in atmospheric ammonia oxidation, especially at low temperatures.

Received 7th April 2025  
Accepted 6th June 2025

DOI: 10.1039/d5ea00042d

[rsc.li/esatmospheres](https://rsc.li/esatmospheres)

## 1 Introduction

Ammonia ( $\text{NH}_3$ ) is one of the most abundant nitrogen-containing compounds in the atmosphere. It is emitted into the atmosphere from agricultural activities, wild fires, and in polar regions also from penguin and seabird colonies, with a current estimated annual global emission of 58 Tg  $\text{N}$ .<sup>1–5</sup> Ammonia has an atmospheric lifetime from hours to a few days.<sup>6–8</sup> Its most significant loss mechanisms are acid–base reactions, wet deposition, and gas-phase oxidation reactions.<sup>7,9</sup> The branching of these routes is uncertain, but the lifetime of  $\text{NH}_3$  with respect to oxidation is estimated to be ten times longer than its other loss mechanisms.<sup>10</sup> This suggests that only a minor fraction of ammonia is oxidized in the gas phase. However, at high altitudes, above the boundary layer, this oxidation is likely more important. Moreover, because the transient concentration of  $\text{NH}_3$  is high,  $\text{NH}_3$  oxidation may lead

to significant production of  $\text{NO}_x$  and other nitrogen-containing species in the atmosphere.

The atmospheric oxidation reactions of ammonia are known to be initiated *via* hydrogen abstraction by the OH radical<sup>11,12</sup> or by halogen radicals,<sup>13</sup> yielding the aminyl radical ( $\text{NH}_2$ ). The  $\text{NH}_2$  radical is known to react bimolecularly with  $\text{NO}_2$ ,  $\text{NO}$ ,  $\text{O}_3$ , and  $\text{O}_2$ .<sup>14</sup> The reactions with  $\text{NO}_2$  and  $\text{NO}$  are believed to be dominant since the IUPAC recommended rate coefficient for its reaction with  $\text{O}_2$  is smaller than  $6 \times 10^{-21} \text{ cm}^3$  per molecule per s.<sup>14</sup> It has been argued in earlier literature that if the rate coefficient of the  $\text{NH}_2 + \text{O}_2$  reaction is larger than  $3 \times 10^{-18} \text{ cm}^3$  per molecule per s, then it would be the dominant bimolecular reaction of  $\text{NH}_2$  under atmospheric conditions.<sup>15</sup>

A wide range of rate coefficients for the  $\text{NH}_2 + \text{O}_2$  reaction ( $10^{-15}$ – $10^{-21} \text{ cm}^3$  per molecule per s) have been determined using different experimental setups.<sup>15–21</sup> The largest rate coefficient ( $10^{-15} \text{ cm}^3$  per molecule per s) was determined in 1972 by measuring the  $\text{NH}_2$  decay, which was revised in 1979 to  $10^{-17} \text{ cm}^3$  per molecule per s by the same authors.<sup>19</sup> The basis of the current IUPAC recommendation is a more recent experimental study from 1991,<sup>15</sup> where the rate coefficient was inferred indirectly by measuring the  $\text{NO}_x$  and  $\text{N}_2\text{O}$  formation in experiments of  $\text{NH}_2$  in excess molecular oxygen with and without added  $\text{NO}$ . The product distributions were explained by the known kinetics of other competing reactions in their reaction system; therefore, the rate coefficient was given as an upper limit.

<sup>a</sup>Department of Chemistry, University of Helsinki, P.O. Box 55, Helsinki 00014, Finland

<sup>b</sup>Department of Chemistry, University of Copenhagen, Copenhagen 2100, Denmark. E-mail: [hgk@chem.ku.dk](mailto:hgk@chem.ku.dk)

† Electronic supplementary information (ESI) available: Further details and discussion of the  $\text{CH}_3 + \text{O}_2 \rightarrow \text{CH}_3\text{O}_2$  calculations, CVT calculations, Gibbs energy surface calculations, master equation simulations and sensitivity tests,  $\text{NH}_2 + \text{O}_2 \rightarrow \text{NH}_2\text{O}_2$  reaction potentials calculated with CASPT2 and CASPT2-IPEA methods, structural parameters and visualizations of CAS orbitals of relevant structures obtained at the NEVPT2 level of theory, and xyz-coordinates of the studied molecules. See DOI: <https://doi.org/10.1039/d5ea00042d>



For methane,  $\text{CH}_4$ , the reaction with OH and  $\text{O}_2$  leads to the formation of  $\text{CH}_3\text{O}_2$ , which is the most abundant alkylperoxy radical in the atmosphere.<sup>22</sup> This peroxy radical has been detected in laboratory experiments by *e.g.* cavity ring down spectroscopy.<sup>23</sup> As far as we know, clear evidence of the formation of the corresponding peroxy radical,  $\text{NH}_2\text{O}_2$ , has not been reported in any of the experimental studies of  $\text{NH}_3$  oxidation.<sup>15–21</sup>

The thermostability of the aminoperoxy radical,  $\text{NH}_2\text{O}_2$ , has been investigated theoretically in multiple studies.<sup>24–28</sup> These previous studies give a somewhat conflicting picture of the reaction enthalpy for the  $\text{NH}_2\text{O}_2$  radical formation. According to earlier studies, the formation of  $\text{NH}_2\text{O}_2$  is endothermic by 10 kcal mol<sup>−1</sup>, while later studies found it to be exothermic by 3–6 kcal mol<sup>−1</sup>.<sup>24–28</sup> The latter assessments are likely more reasonable for the radical addition reaction. These types of reactions are usually exoergic processes, because the reaction only involves the formation of a bond and does not require breaking of any covalent bonds. Despite the discrepancies between the available values, the  $\text{NH}_2 + \text{O}_2 \rightarrow \text{NH}_2\text{O}_2$  reaction appears to be close to thermoneutral. This suggests that unless the reaction is prevented by an insurmountable barrier, the reaction is reversible; thus, the branching between  $\text{NH}_2 + \text{O}_2$  and  $\text{NH}_2\text{O}_2$  is sensitive to the accuracy of the value of the reaction enthalpy and also likely to the specific reaction conditions.

In this work, we study the  $\text{NH}_2 + \text{O}_2 \rightleftharpoons \text{NH}_2\text{O}_2$  reaction using multireference electronic structure methods. The interacting  $\text{NH}_2 + \text{O}_2$  system consists of three unpaired electrons, coupled to an overall doublet state, so multireference methods are necessary to obtain reasonable predictions of the shape of the reaction potential. We assess the reversibility of this reaction and estimate the branching ratios between  $\text{NH}_2 + \text{O}_2$  and  $\text{NH}_2\text{O}_2$  under a broad set of relevant atmospheric conditions. We estimate the high pressure limit reaction rate coefficient using canonical variational theory (CVT)<sup>29</sup> and model the temperature- and pressure-dependent rate coefficients with Rice–Ramsberger–Kassel–Marcus (RRKM/ME)<sup>30</sup> and inverse Laplace transform (ILT/ME) master equation models.<sup>31</sup> We also apply the multireference methodologies to the analogous  $\text{CH}_3 + \text{O}_2 \rightarrow \text{CH}_3\text{O}_2$  reaction for comparison.

## 2 Methods

We obtained the starting geometries for various multireference calculations by carrying out geometry optimizations at the  $\omega\text{B97X-D3/aug-cc-pVTZ}$  level.<sup>32–34</sup> In some instances, we also used B3LYP(D3BJ),<sup>35–37</sup> M06-2X,<sup>38</sup> and CAM-B3LYP<sup>39</sup> functionals for comparison purposes. All DFT calculations were done using ORCA version 5.0.3,<sup>40,41</sup> in which the used DFT functionals were implemented with analytical energy gradients for geometry optimizations, while the second derivatives of the potentials (Hessian) were calculated numerically.

In all multireference calculations, we used the complete active space self-consistent field (CASSCF) method for constructing the zeroth-order wave functions for the subsequent multireference calculations.<sup>42</sup> The used active spaces are

denoted using the general  $n$ -electrons in  $m$ -orbitals, CAS( $ne,mo$ ) notation, where the configuration state function space is constructed from the Full-CI expansion within the CAS subspace. We used the perturbation-based Super-CI SCF optimizer [SuperCI(PT)] in the CASSCF calculations.<sup>43</sup> The natural orbitals corresponding to the optimized CAS orbitals for each relevant obtained stationary structure are visualized in the ESI (Section S8).<sup>†</sup> All CASSCF and subsequent multireference (MR) calculations were done using the aug-cc-pVTZ basis set.

In the calculations with the  $\text{NH}_2\text{O}_2$  molecule and related bimolecular reactants, we used two active spaces. The first, which was used for all the geometry optimizations and frequency calculations, is a CAS(13e,11o) space, which includes all valence orbitals, except the 2s orbitals of the oxygen atoms and the lone pair of the nitrogen atom. The second active space CAS(19e,14o) corresponds to the full-valence CAS of  $\text{NH}_2\text{O}_2$  and was used in single-point energy correction calculations. Calculations concerning the  $\text{NH}_2$  and  $\text{O}_2$  separately were done using full-valence CAS, which are CAS(7e,6o) and CAS(12e,8o), respectively.

In the calculations of further unimolecular reactions of  $\text{NH}_2\text{O}_2$ , the geometries and frequencies of all intermediate species were obtained using the CAS(13e,11o) active space, but the electronic energies were corrected with the full-valence CAS(19e,14o) active space. The HNO, OH, NO, and  $\text{H}_2\text{O}$  calculations were done with their respective full-valence active spaces.

For geometry optimizations and frequency calculations concerning the  $\text{CH}_3\text{O}_2$  molecule and related bimolecular reactants, we used a CAS(13e,9o) active space, which consists of all oxygen valence orbitals, as well as the C–O bonding and antibonding orbitals, and the C–H bonding and antibonding orbitals are not included in the CAS. The geometry and frequencies of the isolated  $\text{CH}_3$  molecule were calculated using the full-valence CAS(7e,7o) active space.

We used  $N$ -electron valence state second-order perturbation theory (NEVPT2)<sup>44,45</sup> to treat the dynamical electron correlation in the studied systems. All NEVPT2 calculations in this work were done using the fully internally contracted (FIC) variants of theory (in earlier literature, they were referred to as partial contraction, PC). Strict pre-screening criteria were used for higher order reduced density matrices, 3-RDM and 4-RDM ( $1 \times 10^{-16}$  cutoff for configuration weights), to prevent false intruder states arising from approximated RDM.<sup>46,47</sup> NEVPT2 geometry optimizations and frequency calculations were done with ORCA-5.0.3,<sup>41</sup> using numerical gradients and numerical Hessians. We studied the  $\text{NH}_2 + \text{O}_2 \rightarrow \text{NH}_2\text{O}_2$  reaction also with CASPT2 and CASPT2-IPEA methods.<sup>48,49</sup> The results obtained with these methods are discussed in the ESI Section S6.<sup>†</sup>

We also carried out benchmark calculations with the complete basis set extrapolated CCSD(T),<sup>50–53</sup> using two-point extrapolation with cc-pVTZ and cc-pVQZ basis sets,<sup>54</sup> denoted as CCSD(T)/CBS. Additionally, we conducted W2X and W3X-L composite method calculations,<sup>55</sup> using Molpro and MRCC programs.<sup>56,57</sup> The W2X extends the CCSD(T)/CBS method by also considering core–valence correlation and scalar-relativistic



effects, upon which the W3X-L method further adds post-CCSD(T) contributions up to CCSDT(Q).

## 2.1 Thermodynamics and kinetics

In the thermochemical analyses, standard approaches were used for obtaining rigid-rotor, translation, and electronic contributions to their respective thermodynamic quantities. By default, all vibrational analyses were based on the harmonic approximation, but at non-stationary points of potential energy surfaces the vibrations were obtained orthogonal to the gradient of the energy (see ESI Section S2.1.1† for details). The entropy contributions from low-frequency modes were scaled with the qRRHO method,<sup>58</sup> using 100 cm<sup>-1</sup> reference value for the weighting function between vibrational and rotational entropies. Furthermore, the vibration corresponding to the internal rotation around the N–O bond in NH<sub>2</sub>⋯O<sub>2</sub> was treated with a one-dimensional rigid hindered rotor model (details are provided in the ESI Section S2.1.2†).

We used three methodologies to assess the kinetics of the studied reaction systems: canonical variational theory for the high-pressure limit thermal rate coefficients,  $k_{\infty}^{\text{CVT}}(T)$  (CVT details are provided in ESI Section S2†), and Rice–Ramsberger–Kassel–Marcus (RRKM/ME) and inverse Laplace transform (ILT/ME) master equation models to calculate temperature- and pressure-dependent rate coefficients  $k(p, T)$ .<sup>30,31</sup> The CVT rate coefficient equation is given by

$$k_{\infty}^{\text{CVT}}(T, s) = \kappa \frac{k_{\text{B}}T}{h} \left( \frac{p^{\ominus}}{k_{\text{B}}T} \right)^{1-M} \exp\left(-\Delta G(s)^{\ddagger, \text{CVT}}/k_{\text{B}}T\right) \quad (1)$$

where  $\kappa$  is the tunneling coefficient,  $k_{\text{B}}$  is the Boltzmann constant,  $T$  is the absolute temperature,  $h$  is the Planck constant,  $p^{\ominus}$  is the reference pressure of 1 bar,  $M$  is the reaction molecularity,  $\Delta G(s)^{\ddagger, \text{CVT}}$  is the quasi-thermodynamic Gibbs energy of activation, and  $s$  is the reaction coordinate. The tunneling coefficient  $\kappa$  was assumed as one, except for the NH<sub>2</sub>O<sub>2</sub> → HNOOH isomerization reaction, where the coefficient was calculated using the Eckart potential approximation.<sup>59</sup> For the most part, the CVT rate coefficient equation is identical to the conventional transition state theory rate coefficient equation. The only difference is the  $\Delta G(s)^{\ddagger, \text{CVT}}$  term, which is defined as the energy difference of the maximum of the Gibbs energy along the minimum energy path of reaction coordinate  $\left(\max_s G(s)^{\ddagger}\right)$  and the Gibbs energy of the reactants ( $G_{\text{reac}}$ ):

$$\Delta G(s)^{\ddagger, \text{CVT}} = \max_s G(s)^{\ddagger} - G_{\text{reac}} \quad (2)$$

The Gibbs energy of activation for the NH<sub>2</sub> + O<sub>2</sub>  $\xrightleftharpoons[k_r]{k_f}$  NH<sub>2</sub>O<sub>2</sub> association reaction ( $\Delta G_{\text{f}}^{\ddagger, \text{CVT}}$ ) was calculated indirectly from the reaction Gibbs energy change ( $\Delta G$ ) and the reverse dissociation Gibbs energy of activation ( $\Delta G_{\text{r}}^{\ddagger, \text{CVT}}$ ):

$$\Delta G_{\text{f}}^{\ddagger, \text{CVT}} = \Delta G + \Delta G_{\text{r}}^{\ddagger, \text{CVT}} \quad (3)$$

where

$$\Delta G = G(\text{NH}_2\text{O}_2) - G(\text{NH}_2 + \text{O}_2) \quad (4)$$

and

$$\Delta G_{\text{r}}^{\ddagger, \text{CVT}} = G_{\text{TS}}^{\ddagger, \text{CVT}} - G(\text{NH}_2\text{O}_2) \quad (5)$$

We believe that W3X-L is the most accurate method in the present study for estimating the reaction energy [ $E(\text{NH}_2) + E(\text{O}_2) - E(\text{NH}_2\text{O}_2)$ ]. Thus, we calculated the  $\Delta G$  from the W3X-L electronic energies of NH<sub>2</sub>, O<sub>2</sub>, and NH<sub>2</sub>O<sub>2</sub> and combined with the thermodynamic corrections obtained at full-valence NEVPT2 for NH<sub>2</sub> and O<sub>2</sub> and at NEVPT2(13e,11o) for NH<sub>2</sub>O<sub>2</sub>. We were not able to calculate the W3X-L energy in the transition state geometry, so the  $\Delta G_{\text{r}}^{\ddagger, \text{CVT}}$  was calculated using NEVPT2(19e,14o) electronic energies with thermodynamic corrections obtained at the NEVPT2(13e,11o) level.

In both the RRKM/ME and ILT/ME calculations, we used a temperature independent collisional energy transfer model, where the average energy lost due to collisions with the N<sub>2</sub> bath was set to  $\langle \Delta E \rangle_{\text{d}} = 100 \text{ cm}^{-1}$ . The high-pressure limit Arrhenius parameters,  $E_{\text{a}}$  and  $A$ , which are required for the ILT/ME simulations, were derived from Arrhenius plots of the  $k_{\infty}^{\text{CVT}}(T)$  values at different temperatures (see ESI Section S3†). We used the standard Arrhenius equation with a temperature independent pre-exponential factor in the fitting of the  $k_{\infty}^{\text{CVT}}(T)$  against the temperature:

$$k_{\infty}^{\text{CVT}}(T) = A \times e^{-E_{\text{a}}/RT} \quad (6)$$

In the RRKM/ME calculations, the NH<sub>2</sub> + O<sub>2</sub> reaction is assumed to form a pre-reactive complex NH<sub>2</sub>⋯O<sub>2</sub> with a collision limited rate coefficient (10<sup>-10</sup> cm<sup>3</sup> per molecule per s in our calculations), whereafter the isomerization of the complex through a transition state to the NH<sub>2</sub>O<sub>2</sub> minimum was calculated with RRKM. In RRKM/ME and ILT/ME simulations, the rate coefficients of the reverse dissociation ( $k_{\text{r}}$ ) were obtained from the equilibrium constant ( $K_{\text{eq}}$ ) and the forward reaction pseudo-first order rate coefficient ( $k'_{\text{f}} = k_{\text{f}}[\text{O}_2]$ ) with the following detailed balance condition:

$$K_{\text{eq}} = \frac{k'_{\text{f}}}{k_{\text{r}}} = \frac{k'_{\text{f}}}{K_{\text{eq}}} \quad (7)$$

All ILT/ME and RRKM/ME simulations were done using the MESMER program version 7.0.<sup>60</sup> The MESMER simulations are discussed in more detail in the ESI Section S3.†

## 3 Results and discussion

The formation of peroxy radicals from alkyl radicals reacting with O<sub>2</sub> is well known.<sup>61,62</sup> Small peroxy radicals like CH<sub>3</sub>O<sub>2</sub> have been observed with cavity ring-down experiments.<sup>23,63</sup> In contrast, the observation of the analogous NH<sub>2</sub>O<sub>2</sub> radical has been elusive. In light of this, we have investigated the NH<sub>2</sub> + O<sub>2</sub> → NH<sub>2</sub>O<sub>2</sub> reaction. First, the potential energy surface (PES) is explored, second the kinetics of the reaction, third the



formation of  $\text{NH}_2\text{O}_2$  under different atmospheric conditions, and finally we discuss the lack of detection of  $\text{NH}_2\text{O}_2$  in previous experiments.

### 3.1 Potential energy surface of $\text{NH}_2 + \text{O}_2$

In Fig. 1, we show that formation of the peroxy radical is exoergic for both the  $\text{CH}_3$  and  $\text{NH}_2$  radicals reacting with  $\text{O}_2$ . Both the  $\text{CH}_3\text{O}_2$  and  $\text{NH}_2\text{O}_2$  formation reactions proceed *via* a barrierless potential, with the alkyl peroxy radical bound much stronger. Thus, it appears that  $\text{NH}_2\text{O}_2$  formation is not hindered or prevented by unfavorable reaction energetics. A similar potential has also been found for the  $\text{CH}_3\text{S} + \text{O}_2 \rightarrow \text{CH}_3\text{SO}_2$  reaction.<sup>64</sup>

In Fig. 2, we show the reaction potential of the  $\text{NH}_2 + \text{O}_2 \rightarrow \text{NH}_2\text{O}_2$  reaction with different methods. The potential energy surface of the  $\text{NH}_2 + \text{O}_2 \rightarrow \text{NH}_2\text{O}_2$  reaction was obtained in reverse *via* dissociation of  $\text{NH}_2\text{O}_2$  to  $\text{NH}_2$  and  $\text{O}_2$  by using the N–O distance as the reaction coordinate and carrying out a relaxed scan with respect to this coordinate from 1.45 Å to 3.00 Å distance with a step size of 0.05 Å (Fig. 2a). The geometries along the reaction potential were optimized at the multi-reference NEVPT2(13e,11o) level (black line in Fig. 2a), which were further corrected with full-valence NEVPT2(19e,14o) single-point energy calculations (red line in Fig. 2a). The relative energy difference between  $\text{NH}_2 + \text{O}_2$  and  $\text{NH}_2\text{O}_2$  was obtained by increasing the distance between  $\text{NH}_2$  and  $\text{O}_2$  to 30 Å and optimizing the geometry of the resulting structure with respect to this distance constraint and then comparing the energy with that of the minimum. In addition, the relative energy of the reactants with respect to  $\text{NH}_2\text{O}_2$  was calculated with a range of other methods. These results are shown on the left-side of Fig. 2a by the scattered symbols at the  $\infty$  mark on the horizontal axis.

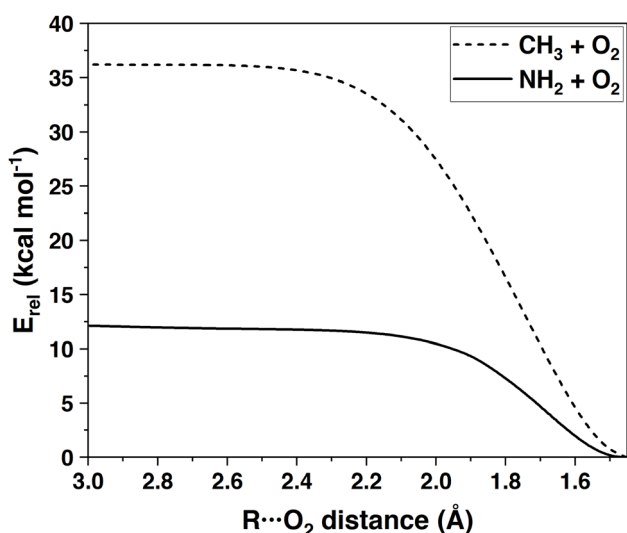


Fig. 1 Electronic potential energy curves of the  $\text{R} + \text{O}_2 \rightarrow \text{RO}_2$  reactions as a function of the  $\text{R}-\text{O}_2$  distance for  $\text{R}=\text{CH}_3$  (dashed) and  $\text{R}=\text{NH}_2$  (solid), obtained with NEVPT2(13e,9o)/aug-cc-pVTZ and NEVPT2(19e,14o)/aug-cc-pVTZ, respectively.

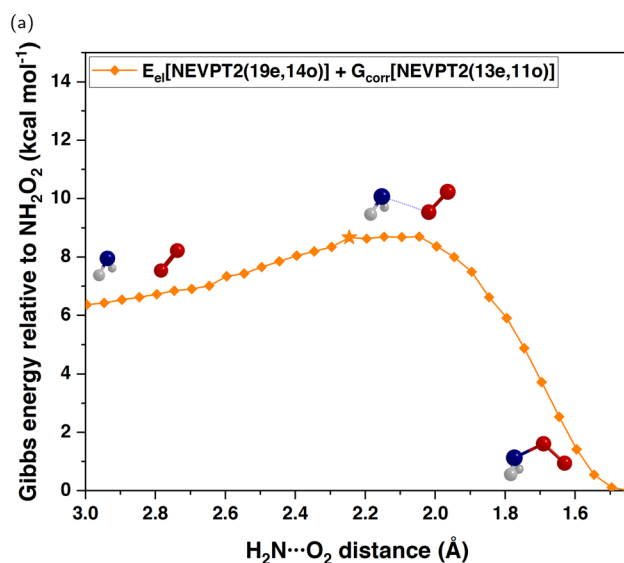
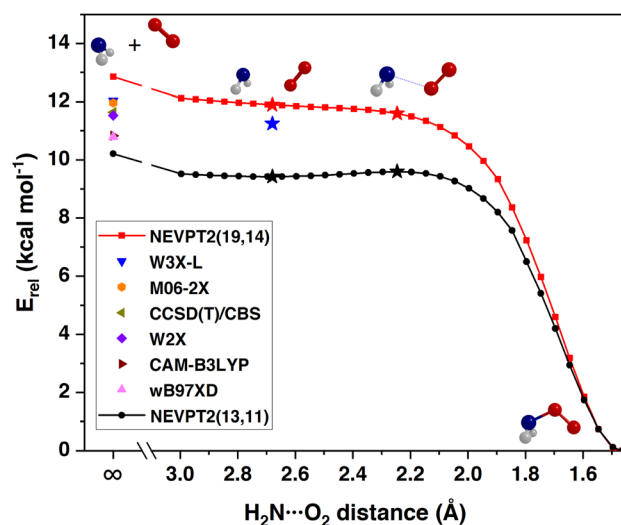


Fig. 2 Minimum energy path of the  $\text{NH}_2 + \text{O}_2 \rightarrow \text{NH}_2\text{O}_2$  reaction as a function of the  $\text{H}_2\text{N}-\text{O}_2$  distance. (a) Electronic potential energy surface with NEVPT2(13e,11o) geometry optimizations (black) and NEVPT2(19e,14o) single-point energy corrections (red). The black  $\star$  symbols correspond to stationary points obtained at the NEVPT2(13e,11o) level, and the blue  $\star$  is calculated with W3X-L. (b) The Gibbs energy surface of the reaction around the Gibbs energy saddle point, at 298 K and 1 atm, with electronic energies calculated at the NEVPT2(19e,14o) level and thermodynamic corrections at the NEVPT2(13e,11o) level. The orange  $\star$  highlights the position of the NEVPT2(13e,11o) saddle point.

The relative energy at the NEVPT2(19e,14o) level was obtained by optimizing the  $\text{NH}_2$  and  $\text{O}_2$  geometries separately at their respective full-valence active space NEVPT2 levels, summing up their energies, and then subtracting the NEVPT2(19e,14o)/NEVPT2(13e,11o) energy of  $\text{NH}_2\text{O}_2$ . The NEVPT2(13e,11o) and NEVPT2(19e,14o) reaction potentials show *ca.* 3  $\text{kcal mol}^{-1}$  difference in the estimated association energies (Fig. 2a). The relative energies obtained with other methods are scattered between the NEVPT2(13e,11o) and

NEVPT2(19e,14o) relative energies. W3X-L may be regarded as the most accurate method used in the present work, and therefore the NEVPT2(19e,14o) relative energies and PES appear more reasonable than the NEVPT2(13e,11o) energies.

The NEVPT2(13e,11o) reaction curve (black line in Fig. 2a) shows a pre-reactive van der Waals complex and a shallow submerged barrier at around 2.65 Å and 2.2 Å N–O distances, respectively (★ symbols in Fig. 2a). Despite the presence of the saddle point in the NEVPT2(13e,11o) PES, the height of the barrier is very low and is not present in the NEVPT2(19e,14o) corrected surface (red line in Fig. 2a). This means that using this point as a transition state may be inaccurate. The region around this point was inspected more closely in the corresponding Gibbs energy surface (Fig. 2b) to better define the location of the transition state of the reaction. Details of how the Gibbs energy surface was obtained are discussed in the ESI Section S2.1.† There are three distinct features in the Gibbs energy surface compared to the PES: the saddle point is more pronounced, it is shifted to a shorter N–O distance, and the pre-reactive complex is absent from the Gibbs energy surface.

Overall, the  $\text{NH}_2 + \text{O}_2 \rightarrow \text{NH}_2\text{O}_2$  reaction does not appear to be prevented by potential energy barriers; in contrast, the reaction PES shows either a negligible barrier or no barrier. The reaction is exoergic but considerably less than the analogous  $\text{CH}_3 + \text{O}_2 \rightarrow \text{CH}_3\text{O}_2$  reaction (Fig. 1), so a smaller fraction of the  $\text{NH}_2\text{O}_2$  product is expected to be present under ambient conditions compared to  $\text{CH}_3\text{O}_2$ , due to a faster back reaction. At finite temperature, the entropy penalty associated with two molecules reacting to form one molecule introduces a clear saddle point in the Gibbs energy surface of the  $\text{NH}_2 + \text{O}_2 \rightarrow \text{NH}_2\text{O}_2$  reaction, which controls the rate of the reaction.

### 3.2 Kinetics of $\text{NH}_2 + \text{O}_2$

We have estimated the kinetics of the  $\text{NH}_2 + \text{O}_2 \xrightleftharpoons[k_r]{k_f} \text{NH}_2\text{O}_2$  reaction with CVT, ILT/ME, and RRKM/ME (Table 1). The thermal high-pressure limit (HPL) rate coefficients  $k_\infty(T)$  for the forward association ( $k_f$ ) and reverse dissociation ( $k_r$ ) were calculated with the CVT method, using the maximum of the Gibbs energy surface as the transition state structure (Fig. 2b). The  $k_\infty(298\text{ K})$  value,  $1.1 \times 10^{-13}$  cm<sup>3</sup> per molecule per s, is eight orders of magnitude larger than the current IUPAC recommended value ( $10^{-21}$  cm<sup>3</sup> per molecule per s). However, reactions involving small molecules often do not exhibit HPL kinetics at ambient pressure.<sup>65</sup> Therefore, we studied the pressure and temperature dependence of the reaction system with the ILT and RRKM master equation models. As expected, the

ILT/ME and RRKM/ME rate coefficients at 298 K and 1 atm are smaller than the HPL values obtained with CVT. The pressure dependence of  $k_f$  at varying temperatures, obtained with the ILT/ME method, is illustrated in Fig. S7.† At 298 K and 1 atm, the ILT/ME and RRKM/ME rate coefficients are roughly  $10^{-15}$  cm<sup>3</sup> per molecule per s, suggesting the rapid formation of  $\text{NH}_2\text{O}_2$  under atmospheric conditions.

Based on our ME calculations, we find that the rate coefficient at 1 atm pressure is about two orders of magnitude lower, and at 0.1 atm pressure about 3 orders of magnitude lower, than the HPL value (Table 1) (ESI Section S3).†

Independent of the methodology used, our calculated rate coefficients of  $\text{NH}_2\text{O}_2$  formation are larger than those obtained by experimentally measuring the  $\text{NH}_2$  decay rate<sup>18–21</sup> and several orders of magnitude larger than the currently recommended rate coefficient value for the  $\text{NH}_2 + \text{O}_2$  reaction by IUPAC.<sup>15,66</sup> We hesitate to give absolute recommendations for the rate coefficient because the methods and approximations we have used also have their uncertainties. However, we carried out various sensitivity tests (ESI Section S3†) and found the rate coefficients to vary by less than two orders of magnitude.

In addition, we performed similar multireference electronic structure calculations and ME simulations for the analogous  $\text{CH}_3 + \text{O}_2 \rightarrow \text{CH}_3\text{O}_2$  system (see ESI Sections S1 and S3†), for which a lot of experimental kinetics data are available (see ref. 66 for example). Our  $k(p, T)$  values for that system are in reasonable agreement with the currently recommended values.<sup>66–68</sup> The experimental rate coefficient for the  $\text{CH}_3 + \text{O}_2 \rightarrow \text{CH}_3\text{O}_2$  reaction at 298 K and 1 atm is  $8.1 \times 10^{-13}$  cm<sup>3</sup> per molecule per s,<sup>68</sup> while our calculated values under these conditions are  $2.1 \times 10^{-13}$  and  $1.1 \times 10^{-12}$  cm<sup>3</sup> per molecule per s, with ILT/ME and RRKM/ME, respectively (ESI Table S2†), both of which are in good agreement with the experimental value.

### 3.3 Atmospheric presence of $\text{NH}_2\text{O}_2$

Whether formation of  $\text{NH}_2\text{O}_2$  has any effect on atmospheric processes depends on the stability of  $\text{NH}_2\text{O}_2$  with respect to back-dissociation, *i.e.* its fraction at equilibrium compared with the loss rates of  $\text{NH}_2$  and  $\text{NH}_2\text{O}_2$  with respect to other competing reactions.

The competition of bimolecular reactions of  $\text{NH}_2$  with NO, NO<sub>2</sub>, and O<sub>3</sub> can be estimated by comparing the pseudo-first order rate coefficients of these reactions. An upper limit for these competing bimolecular reactions would be that they occur with a rate coefficient near the collision limit of around  $10^{-10}$  cm<sup>3</sup> per molecule per s and that the combined concentration of these reactants is 40 ppb, which leads to a pseudo-first order

Table 1 Calculated rate coefficients of the  $\text{NH}_2 + \text{O}_2 \rightleftharpoons \text{NH}_2\text{O}_2$  reaction at 298 K

| $\text{NH}_2 + \text{O}_2 \xrightleftharpoons[k_r]{k_f} \text{NH}_2\text{O}_2$ | CVT (HPL)             | ILT/ME (1 atm)        | RRKM/ME (1 atm)       |
|--------------------------------------------------------------------------------|-----------------------|-----------------------|-----------------------|
| $k_f$ (cm <sup>3</sup> per molecule per s)                                     | $1.1 \times 10^{-13}$ | $1.7 \times 10^{-15}$ | $2.5 \times 10^{-15}$ |
| $k_r$ (s <sup>-1</sup> )                                                       | $8.6 \times 10^6$     | $1.3 \times 10^5$     | $1.8 \times 10^5$     |



rate coefficient of  $100 \text{ s}^{-1}$ .<sup>69</sup> The actual pseudo-first order rate coefficient is likely much smaller than this upper limit.

The reactions of  $\text{NH}_2$  with  $\text{NO}_x$  and  $\text{O}_3$  are considerably faster than the reaction with  $\text{O}_2$ , but the concentration of  $\text{O}_2$  is substantially higher than the  $\text{NO}_x$  and  $\text{O}_3$  concentrations. With our calculated rate coefficients in Table 1 ( $2 \times 10^{-15} \text{ cm}^3$  per molecule per s) and the atmospheric  $\text{O}_2$  concentration, the pseudo-first order rate coefficient of  $\text{NH}_2 + \text{O}_2 \rightarrow \text{NH}_2\text{O}_2$  is  $10^4 \text{ s}^{-1}$ , thus the reaction is much faster than other bimolecular reactions. The back reaction  $\text{NH}_2\text{O}_2 \rightarrow \text{NH}_2 + \text{O}_2$  is even faster than the forward reaction (Table 1). This suggests that the  $\text{NH}_2 + \text{O}_2 \rightleftharpoons \text{NH}_2\text{O}_2$  equilibrium is established before the competing bimolecular reactions start consuming either  $\text{NH}_2$  or  $\text{NH}_2\text{O}_2$ . Even if the calculated rate coefficients in Table 1 were 1–2 orders of magnitude too high, this would still hold.

Then, if the  $\text{NH}_2 + \text{O}_2 \rightleftharpoons \text{NH}_2\text{O}_2$  reaction has reached equilibrium and there are no reactions affecting the equilibrium, the fraction of  $\text{NH}_2\text{O}_2$  can be easily calculated. At equilibrium,  $k_f[\text{NH}_2][\text{O}_2] = k_r[\text{NH}_2\text{O}_2]$ , hence the ratio  $[\text{NH}_2\text{O}_2]/[\text{NH}_2] = (k_f/k_r)[\text{O}_2]$ . Our calculations with ILT/ME and RRKM/ME suggest that at 298 K and 1 atm (with 21%  $\text{O}_2$  partial pressure), the fraction of  $\text{NH}_2\text{O}_2$  is 6.4%. We carried out ILT/ME simulations of  $\text{NH}_2 + \text{O}_2 \rightleftharpoons \text{NH}_2\text{O}_2$  with a broad set of conditions to show how the equilibrium fraction of  $\text{NH}_2\text{O}_2$  varies with temperature and pressure. In Fig. 3a, we show the fraction of  $\text{NH}_2\text{O}_2$  as a function of temperature (230–330 K) at four total pressures: 700, 570, 400, and 200 torr, with 21%  $\text{O}_2$  partial pressure. Fig. 3b shows the  $\text{NH}_2\text{O}_2$  equilibrium fraction as a function of altitude (0–10 km), for global (288 K, solid line)<sup>70</sup> and polar (263 K, dashed) average surface temperatures. The temperatures and pressures at various altitudes were calculated with the barometric formula, using the temperature lapse rate of  $-6.5 \text{ K km}^{-1}$  (ESI Section S4†).

The formation of  $\text{NH}_2\text{O}_2$  is more pronounced at lower temperatures and higher pressures, which is explained by the entropy effect (Fig. 3a). The average temperature of Earth's atmosphere at the ground level is 288 K, where 10% of formed  $\text{NH}_2$  will add  $\text{O}_2$  forming  $\text{NH}_2\text{O}_2$  (Fig. 3b). In polar regions, where the mean surface temperature is 263 K, the  $\text{NH}_2\text{O}_2$  fraction is much larger at the ground level (32%); therefore, it is likely an important pathway in  $\text{NH}_3$  oxidation. At high altitudes (Fig. 3b), where both the temperature and pressure are low, the fraction of  $\text{NH}_2\text{O}_2$  becomes substantial.

### 3.4 Unimolecular reactions of $\text{NH}_2\text{O}_2$

It is critical to assess whether there are any unimolecular reactions of  $\text{NH}_2\text{O}_2$  that would affect the  $\text{NH}_2 + \text{O}_2 \rightleftharpoons \text{NH}_2\text{O}_2$  equilibrium. In previous literature, it has been demonstrated that the most exothermic products of the  $\text{NH}_2 + \text{O}_2$  reaction would be  $\text{NO} + \text{H}_2\text{O}$  and  $\text{HNO} + \text{OH}$ , the prior being the thermodynamically preferred product channel.<sup>15</sup> We calculated the unimolecular reaction pathways that connect the  $\text{NH}_2\text{O}_2$  to these product channels. The Gibbs energy profile of these reactions at 298 K and 1 atm is shown in Fig. 4. We also validated some of the calculated energy asymptotes in Fig. 4 against reaction enthalpies in the Active Thermochemical Tables database<sup>71</sup> (see ESI Section S7† for details).

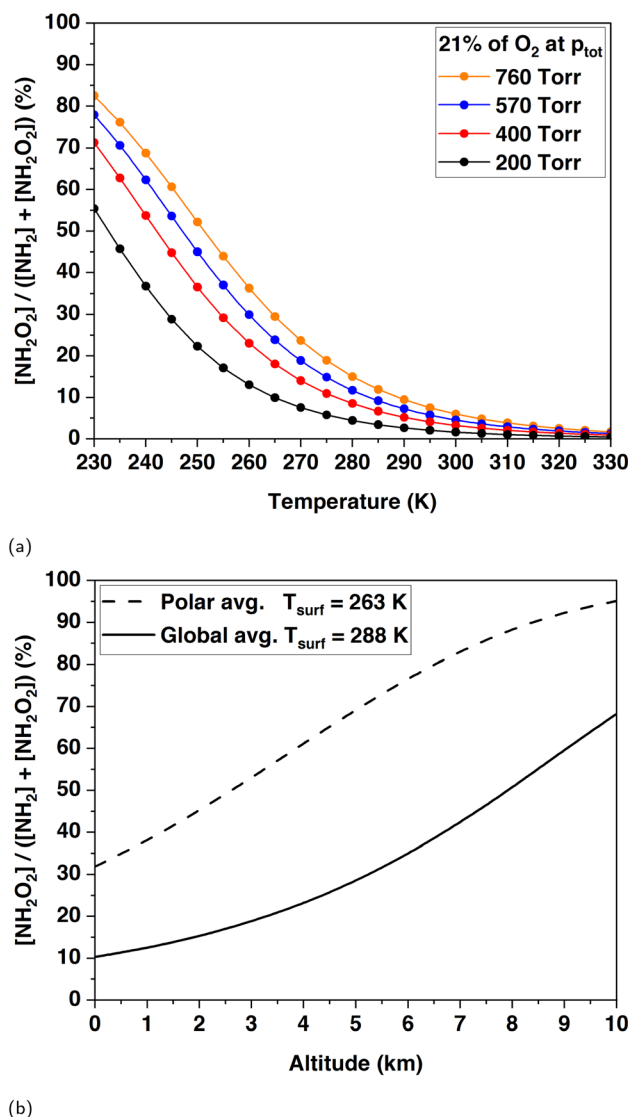


Fig. 3 Fraction of  $\text{NH}_2\text{O}_2$  under typical tropospheric conditions: (a) at total pressures of 200 torr (black dotted line), 400 torr (blue dotted line), 570 torr (red dotted line), and 760 torr (orange dotted line), in the temperature range 230–330 K; (b) as a function of altitude of 0–10 km, with the global average surface temperature 288 K (solid line) and polar average surface temperature 263 K (dashed line). All calculations were carried out with 21%  $\text{O}_2$  concentration, using the ILT/ME method.

After the initial association of  $\text{NH}_2$  and  $\text{O}_2$ , the formed  $\text{NH}_2\text{O}_2$  has three accessible unimolecular reactions: the reverse dissociation to  $\text{NH}_2 + \text{O}_2$ , isomerization to the  $\text{HNOOH}$  radical *via* intramolecular hydrogen atom transfer, and dissociation to the aminoxyl radical and atomic oxygen ( $\text{NH}_2\text{O} + \text{O}$ ). The last reaction has been deemed unlikely to be competitive under ambient conditions,<sup>28</sup> which is also supported by our findings: the reaction is endergonic by  $31.4 \text{ kcal mol}^{-1}$  at 298 K and 1 atm (Fig. 4). The isomerization of  $\text{NH}_2\text{O}_2$  to  $\text{HNOOH}$ , which ultimately leads to the formation of  $\text{NO}$  and  $\text{HNO}$ , occurs *via* a tight transition state, with a Gibbs energy barrier height of  $28.5 \text{ kcal mol}^{-1}$  (Fig. 4).



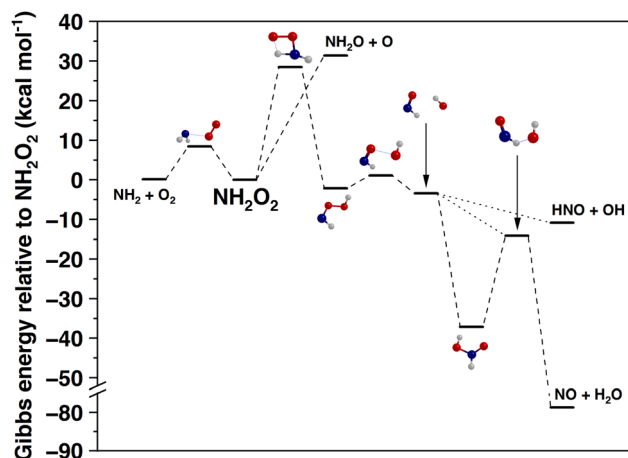


Fig. 4 Gibbs energy diagram of the unimolecular reactions of  $\text{NH}_2\text{O}_2$ , including those leading to  $\text{HNO}$  and  $\text{NO}$  formation, calculated at the NEVPT2 level (see details in the Methods section), at 298 K and 1 atm. Red = oxygen, blue = nitrogen, and gray = hydrogen. Zero energy corresponds to the Gibbs energy of  $\text{NH}_2\text{O}_2$ .

The hydrogen atom transfer reaction (H-shift) is accelerated by quantum mechanical tunneling, which we estimated using the Eckart potential approximation.<sup>59</sup> We simulated the hydrogen atom transfer reaction at 298 K and 1 atm pressure using RRKM/ME and determined a rate coefficient of  $7.6 \times 10^{-6} \text{ s}^{-1}$  including tunneling. The effective rate coefficient for the overall reaction from  $\text{NH}_2 + \text{O}_2$  to  $\text{HNOOH}$ , which we use as a proxy for  $\text{NO}$  production, is calculated by combining ILT/ME for the  $\text{NH}_2 + \text{O}_2 \rightarrow \text{NH}_2\text{O}_2$  reaction with a RRKM/ME for the  $\text{NH}_2\text{O}_2 \rightarrow \text{HNOOH}$  reaction. This gives  $5.8 \times 10^{-25} \text{ cm}^3 \text{ per molecule per s}$  at 298 K and 1 atm. Thus, this pathway is insignificant under ambient conditions. This small rate coefficient is in qualitative agreement with the IUPAC recommendation.<sup>15,66</sup>

It is evident that under atmospheric conditions, the only plausible reactions of  $\text{NH}_2\text{O}_2$  are the reverse dissociation back to  $\text{NH}_2 + \text{O}_2$  or further bimolecular reactions. In this work, we did not explicitly model any of these further bimolecular reactions. Because  $\text{NH}_2\text{O}_2$  is a peroxy radical, its most likely further bimolecular reactions are with  $\text{NO}$ ,  $\text{NO}_2$ ,  $\text{HO}_2$ , and other  $\text{RO}_2$ . If we estimate that the bimolecular rate coefficients of  $\text{NH}_2\text{O}_2$  are comparable to those of alkyl  $\text{RO}_2$ , which are known to be in the range of  $10^{-10}$ – $10^{-13} \text{ cm}^3 \text{ per molecule per s}$ ,<sup>72</sup> the lifetimes of  $\text{NH}_2$  (excluding the reaction with  $\text{O}_2$ ) and  $\text{NH}_2\text{O}_2$  are similar.

It is known that peroxy radicals ( $\text{RO}_2$ ) can react bimolecularly with  $\text{NO}_2$ ,  $\text{NO}$ ,  $\text{HO}_2$  and other  $\text{RO}_2$  to form *e.g.* peroxy nitrates ( $\text{ROONO}_2$ ), organic nitrates ( $\text{RONO}_2$ ), hydroperoxides ( $\text{ROOH}$ ), alcohols ( $\text{ROH}$ ), carbonyl compounds ( $\text{R-H=O}$ ), alkoxy radicals ( $\text{RO}$ ), and organic peroxides ( $\text{ROOR}$ ).<sup>72</sup> For the aminoperoxy radical  $\text{NH}_2\text{O}_2$ , the corresponding products would be  $\text{NH}_2\text{OONO}_2$ ,  $\text{NH}_2\text{ONO}_2$ ,  $\text{NH}_2\text{OOH}$ ,  $\text{NH}_2\text{OH}$ , and  $\text{NH}_2\text{OOR}$ , which are different compound classes to those currently expected as products of  $\text{NH}_3$  oxidation ( $\text{NO}$ ,  $\text{N}_2\text{O}$ , and  $\text{N}_2$ ).<sup>9</sup> The bimolecular reactions of  $\text{NH}_2\text{O}_2$  will likely also produce the  $\text{NH}_2\text{O}$  radical, which is also known to form in the

bimolecular reactions of  $\text{NH}_2$  with  $\text{O}_3$ ,  $\text{NO}_2$ , and  $\text{HO}_2$ .<sup>14,21,73</sup> Therefore, the substantial branching ratio of  $\text{NH}_2\text{O}_2$  may lead to previously undiscovered nitrogen-containing compounds in the atmosphere and influence the current atmospheric modeling.

### 3.5 Previous experiments

There have been several experimental investigations of the reaction between  $\text{NH}_2$  and  $\text{O}_2$  by measuring the  $\text{NH}_2$  decay rate in the presence of  $\text{O}_2$ , and the results suggest that the reaction between  $\text{NH}_2$  and  $\text{O}_2$  is negligible. However, these experiments have been conducted largely under either low-pressure or high-temperature conditions, where the reaction between  $\text{NH}_2$  and  $\text{O}_2$  forming  $\text{NH}_2\text{O}_2$  is not expected to play a substantial role based on our calculations. The calculated fraction of  $\text{NH}_2\text{O}_2$  under different experimental conditions is given in the ESI Section S5.†

The pulse radiolysis experiments of  $\text{NH}_3$  by Pagsberg *et al.*<sup>19</sup> showed that, at 350 K, the  $\text{NH}_2$  decay rate is independent of  $\text{O}_2$  concentration, even at their highest  $\text{O}_2$  concentration experiment (42%  $\text{O}_2$  at 705 torr). In agreement, our results show that the fraction of  $\text{NH}_2\text{O}_2$  is 1.5% under these conditions (ESI Table S4†), and thus the decay of  $\text{NH}_2$  would be largely independent of  $\text{O}_2$ .

Cheskis and Sarkisov<sup>18</sup> carried out room temperature flash photolysis experiments of  $\text{NH}_3$  at 100 and 570 torr total pressures, with varying  $\text{O}_2$  percentage. They observed significant enhancement in the  $\text{NH}_2$  decay at 1 torr  $\text{O}_2$  partial pressure and 570 torr total pressure ( $\approx 0.2\%$   $\text{O}_2$ ) compared to a similar experiment without oxygen. The authors explained that the increased decay rate was unlikely to be due to the reaction between  $\text{NH}_2$  and  $\text{O}_2$  but instead was due to their experimental setup. In the experiment, the  $\text{NH}_2$  radicals were generated *via*  $\text{NH}_3$  photolysis, yielding H radicals that can react with  $\text{O}_2$  to form  $\text{HO}_2$ , which reacted irreversibly with  $\text{NH}_2$ , resulting in the decay of  $\text{NH}_2$ . This is in agreement with our results, which show that at such low  $\text{O}_2$  percentage, the formation of  $\text{NH}_2\text{O}_2$  was insignificant (0.05%, ESI Table S4†). However, no increase in the  $\text{NH}_2$  decay rate was observed when further increasing the  $\text{O}_2$  pressure to 100 and 570 torr in the experiments (18% and 100%  $\text{O}_2$  at 570 torr total pressure). These observations are in contrast to our calculations, which suggest that the fraction of  $\text{NH}_2\text{O}_2$  was 4% and 20% under these conditions, respectively (ESI Table S4†), which would affect the  $\text{NH}_2$  decay rate. The  $\text{NH}_2$  decay was saturated because significant  $\text{NH}_2$  decay rate was also observed upon increasing the  $\text{N}_2$  pressure from 100 torr to 570 torr without any  $\text{O}_2$  in the experiments.

Patrick and Golden<sup>21</sup> conducted experiments in the temperature range 272–348 K where  $\text{NH}_2$  was generated *via*  $\text{O}_3$  photolysis forming atomic O, which reacted with  $\text{NH}_3$  to form  $\text{NH}_2$ . The experiments with  $\text{O}_2$  were carried out at low total pressures of 230–240 torr and  $\text{O}_2$  partial pressures of only 0–14 torr (max. 6%  $\text{O}_2$ ). Our calculations show that the  $\text{NH}_2\text{O}_2$  fraction is less than 3% at such low  $\text{O}_2$  pressures (ESI Table S4†). No significant reactivity between  $\text{NH}_2$  and  $\text{O}_2$  was observed in the experiments, in agreement with our results. Experiments with higher  $\text{O}_2$  partial pressures cannot be conducted in this setup because the reaction between  $\text{O}_2$  and atomic O would suppress the formation of  $\text{NH}_2$ .



## 4 Conclusions

We have carried out high-level multireference calculations and kinetic modeling for studying the formation of the  $\text{NH}_2\text{O}_2$  radical *via* the association reaction between  $\text{NH}_2$  and  $\text{O}_2$ , which is a crucial step in the atmospheric oxidation of  $\text{NH}_3$ . We find that the association reaction is much faster than the current IUPAC estimate and that  $\text{NH}_2 + \text{O}_2$  is the dominant bimolecular reaction of  $\text{NH}_2$  in the atmosphere. Our calculations encompass a broad range of atmospherically relevant conditions, and we show that  $\text{NH}_2\text{O}_2$  formation can play an important role in  $\text{NH}_3$  oxidation, especially under low-temperature and high-pressure conditions, where a substantial equilibrium fraction of  $\text{NH}_2\text{O}_2$  is expected to be present. We encourage laboratory experiments to directly detect the  $\text{NH}_2\text{O}_2$  radical and determine the rate of the  $\text{NH}_2 + \text{O}_2$  reaction under relevant conditions.

## Data availability

Calculation output files related to the study are available in the ESI† and at <http://doi.org/10.5281/zenodo.12657070>.

## Conflicts of interest

There are no conflicts to declare.

## Acknowledgements

We thank the Research Council of Finland Centre of Excellence: Virtual Laboratory for Molecular-Level Atmospheric Transformations (grant 346369), the Jane and Aatos Erkko (JAES) Foundation, and the Villum Fonden (VIL50443) for funding. We also thank the CSC – IT Center of Science (Finland) for providing the supercomputing platforms and computational resources.

## References

- L. Liu, W. Xu, X. Lu, B. Zhong, Y. Guo, X. Lu, Y. Zhao, W. He, S. Wang, X. Zhang, *et al.*, *Proc. Natl. Acad. Sci. U. S. A.*, 2022, **119**, e2121998119.
- R. Zhu, J. Sun, Y. Liu, Z. Gong and L. Sun, *Antarct. Sci.*, 2011, **23**, 78–92.
- J. Lindaas, I. B. Pollack, J. J. Calahorrano, K. O'Dell, L. A. Garofalo, M. A. Pothier, D. K. Farmer, S. M. Kreidenweis, T. Campos, F. Flocke, *et al.*, *J. Geophys. Res.: Atmos.*, 2021, **126**, e2020JD033730.
- S. Riddick, T. Blackall, U. Dragosits, F. Daunt, M. Newell, C. Braban, Y. Tang, J. Schmale, P. Hill, S. Wanless, *et al.*, *Atmos. Environ.*, 2016, **134**, 40–50.
- J. Plautz, *Science*, 2018, **361**, 1060–1063.
- R. Pinder, A. Gilliland and R. Dennis, *Geophys. Res. Lett.*, 2008, **35**, L12808.
- S. N. Behera, M. Sharma, V. P. Aneja and R. Balasubramanian, *Environ. Sci. Pollut. Res.*, 2013, **20**, 8092–8131.
- N. Evangeliou, Y. Balkanski, S. Eckhardt, A. Cozic, M. Van Damme, P.-F. Coheur, L. Clarisse, M. W. Shephard, K. E. Cady-Pereira and D. Hauglustaine, *Atmos. Chem. Phys.*, 2020, **2020**, 1–41.
- S. J. Pai, C. L. Heald and J. G. Murphy, *ACS Earth Space Chem.*, 2021, **5**, 1674–1685.
- J. M. Roberts, A. O. Langford, P. D. Goldan and F. C. Fehsenfeld, *J. Atmos. Chem.*, 1988, **7**, 137–152.
- R. D. Stephens, *J. Phys. Chem.*, 1984, **88**, 3308–3313.
- C. P. Ennis, J. R. Lane, H. G. Kjaergaard and A. J. McKinley, *J. Am. Chem. Soc.*, 2009, **131**, 1358–1359.
- M. Monge-Palacios and J. Espinosa-Garcia, *J. Phys. Chem. A*, 2010, **114**, 4418–4426.
- R. Atkinson, D. L. Baulch, R. A. Cox, J. N. Crowley, R. F. Hampson, R. G. Hynes, M. E. Jenkin, M. J. Rossi and J. Troe, *Atmos. Chem. Phys.*, 2004, **4**, 1461–1738.
- G. Tyndall, J. Orlando, K. E. Nickerson, C. Cantrell and J. Calvert, *J. Geophys. Res.: Atmos.*, 1991, **96**, 20761–20768.
- H. Gesser, *J. Am. Chem. Soc.*, 1955, **77**, 2626–2629.
- R. Jayanty, R. Simonaitis and J. Heicklen, *J. Phys. Chem.*, 1976, **80**, 433–437.
- S. Cheskis and O. Sarkisov, *Chem. Phys. Lett.*, 1979, **62**, 72–76.
- P. B. Pagsberg, J. Eriksen and H. Christensen, *J. Phys. Chem.*, 1979, **83**, 582–590.
- W. Hack, O. Horie and H. G. Wagner, *J. Phys. Chem.*, 1982, **86**, 765–771.
- R. Patrick and D. M. Golden, *J. Phys. Chem.*, 1984, **88**, 491–495.
- R. L. Caravan, M. A. H. Khan, J. Zádor, L. Sheps, I. O. Antonov, B. Rotavera, K. Ramasesha, K. Au, M.-W. Chen, D. Rösch, *et al.*, *Nat. Commun.*, 2018, **9**, 4343.
- M. B. Pushkarsky, S. J. Zalyubovsky and T. A. Miller, *J. Chem. Phys.*, 2000, **112**, 10695–10698.
- C. Pouchan and M. Chaillet, *Chem. Phys. Lett.*, 1982, **90**, 310–316.
- C. F. Melius and J. S. Binkley, *ACS Symp. Ser.*, 1984, 103–115.
- J. W. Bozzelli and A. M. Dean, *J. Phys. Chem.*, 1989, **93**, 1058–1065.
- R. Sumathi and S. D. Peyerimhoff, *J. Chem. Phys.*, 1998, **108**, 5510–5521.
- S. J. Klippenstein, L. B. Harding, P. Glarborg and J. A. Miller, *Combust. Flame*, 2011, **158**, 774–789.
- J. L. Bao and D. G. Truhlar, *Chem. Soc. Rev.*, 2017, **46**, 7548–7596.
- R. A. Marcus, *J. Chem. Phys.*, 1952, **20**, 359–364.
- J. W. Davies, N. J. Green and M. J. Pilling, *Chem. Phys. Lett.*, 1986, **126**, 373–379.
- S. Grimme, J. Antony, S. Ehrlich and H. Krieg, *J. Chem. Phys.*, 2010, **132**, 154104.
- J.-D. Chai and M. Head-Gordon, *Phys. Chem. Chem. Phys.*, 2008, **10**, 6615–6620.
- R. A. Kendall, T. H. Dunning Jr and R. J. Harrison, *J. Chem. Phys.*, 1992, **96**, 6796–6806.
- A. D. Becke, *J. Chem. Phys.*, 1992, **96**, 2155–2160.
- C. Lee, W. Yang and R. G. Parr, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 1988, **37**, 785.



- 37 A. D. Becke and E. R. Johnson, *J. Chem. Phys.*, 2005, **123**, 154101.
- 38 Y. Zhao and D. G. Truhlar, *Theor. Chem. Acc.*, 2008, **120**, 215–241.
- 39 T. Yanai, D. P. Tew and N. C. Handy, *Chem. Phys. Lett.*, 2004, **393**, 51–57.
- 40 F. Neese, *Wiley Interdiscip. Rev.: Comput. Mol. Sci.*, 2012, **2**, 73–78.
- 41 F. Neese, *Wiley Interdiscip. Rev.: Comput. Mol. Sci.*, 2022, **12**, e1606.
- 42 B. O. Roos, P. R. Taylor and P. E. Sigbahn, *Chem. Phys.*, 1980, **48**, 157–173.
- 43 C. Kollmar, K. Sivalingam, B. Helmich-Paris, C. Angeli and F. Neese, *J. Comput. Chem.*, 2019, **40**, 1463–1470.
- 44 C. Angeli, R. Cimiraglia, S. Evangelisti, T. Leininger and J.-P. Malrieu, *J. Chem. Phys.*, 2001, **114**, 10252–10264.
- 45 C. Angeli, R. Cimiraglia and J.-P. Malrieu, *J. Chem. Phys.*, 2002, **117**, 9138–9153.
- 46 Y. Guo, K. Sivalingam and F. Neese, *J. Chem. Phys.*, 2021, **154**, 214111.
- 47 C. Kollmar, K. Sivalingam, Y. Guo and F. Neese, *J. Chem. Phys.*, 2021, **155**, 234104.
- 48 K. Andersson, P. A. Malmqvist, B. O. Roos, A. J. Sadlej and K. Wolinski, *J. Phys. Chem.*, 1990, **94**, 5483–5488.
- 49 G. Ghigo, B. O. Roos and P.-Å. Malmqvist, *Chem. Phys. Lett.*, 2004, **396**, 142–149.
- 50 K. Raghavachari, G. W. Trucks, J. A. Pople and M. Head-Gordon, *Chem. Phys. Lett.*, 1989, **157**, 479–483.
- 51 S. Zhong, E. C. Barnes and G. A. Petersson, *J. Chem. Phys.*, 2008, **129**, 184116.
- 52 F. Neese and E. F. Valeev, *J. Chem. Theory Comput.*, 2011, **7**, 33–43.
- 53 T. Helgaker, W. Klopper, H. Koch and J. Noga, *J. Chem. Phys.*, 1997, **106**, 9639–9646.
- 54 T. H. Dunning Jr, *J. Chem. Phys.*, 1989, **90**, 1007–1023.
- 55 B. Chan and L. Radom, *J. Chem. Theory Comput.*, 2015, **11**, 2109–2119.
- 56 H.-J. Werner, P. J. Knowles, F. R. Manby, J. A. Black, K. Doll, A. Heßelmann, D. Kats, A. Köhn, T. Korona, D. A. Kreplin, Q. Ma, I. Miller, F. Thomas, A. Mitrushchenkov, K. A. Peterson, I. Polyak, G. Rauhut and M. Sibaev, *J. Chem. Phys.*, 2020, **152**, 144107.
- 57 M. Kállay, P. R. Nagy, D. Mester, Z. Rolik, G. Samu, J. Csontos, J. Csóka, P. B. Szabó, L. Gyevi-Nagy, B. Hégyel, I. Ladjánszki, L. Szegedy, B. Ladóczki, K. Petrov, M. Farkas, P. D. Mezei and Á. Ganyecz, *J. Chem. Phys.*, 2020, **152**, 074107.
- 58 S. Grimme, *Chem.–Eur. J.*, 2012, **18**, 9955–9964.
- 59 R. L. Brown, *J. Res. Natl. Bur. Stand.*, 1981, **86**, 357–359.
- 60 D. R. Glowacki, C.-H. Liang, C. Morley, M. J. Pilling and S. H. Robertson, *J. Phys. Chem. A*, 2012, **116**, 9545–9560.
- 61 P. D. Lightfoot, R. Cox, J. Crowley, M. Destriau, G. Hayman, M. Jenkin, G. Moortgat and F. Zabel, *Atmos. Environ., Part A*, 1992, **26**, 1805–1961.
- 62 J. J. Orlando and G. S. Tyndall, *Chem. Soc. Rev.*, 2012, **41**, 6294–6317.
- 63 A. Bossolasco, E. P. Faragó, C. Schoemaeker and C. Fittschen, *Chem. Phys. Lett.*, 2014, **593**, 7–13.
- 64 J. Chen, J. R. Lane, K. H. Bates and H. G. Kjaergaard, *Environ. Sci. Technol.*, 2023, **57**, 21168–21177.
- 65 B. M. Wong, D. M. Matheu and W. H. Green, *J. Phys. Chem. A*, 2003, **107**, 6206–6211.
- 66 R. Atkinson, D. L. Baulch, R. A. Cox, J. N. Crowley, R. F. Hampson, R. G. Hynes, M. E. Jenkin, M. J. Rossi, J. Troe and IUPAC Subcommittee, *Atmos. Chem. Phys.*, 2006, **6**, 3625–4055.
- 67 R. X. Fernandes, K. Luther and J. Troe, *J. Phys. Chem. A*, 2006, **110**, 4442–4449.
- 68 J. B. Burkholder, S. P. Sander, J. Abbatt, J. R., C. Cappa, J. D. Crounse, T. S. Dibble, R. E. Huie, C. E. Kolb, M. J. Kurylo, V. L. Orkin, C. J. Percival, D. M. Wilmouth and P. H. Wine, *Chemical Kinetics and Photochemical Data for Use in Atmospheric Studies*, JPL Publication, Jet Propulsion Laboratory, Pasadena, 19, 2019.
- 69 J. H. Seinfeld and S. N. Pandis, *Atmospheric Chemistry and Physics: From Air Pollution to Climate Change*, John Wiley & Sons, 2016.
- 70 Atmospheric Administration and United States Committee on Extension to the Standard Atmosphere, *US Standard Atmosphere*, National Oceanic and Atmospheric Administration, 1976.
- 71 B. Ruscic, R. E. Pinzon, M. L. Morton, G. von Laszewski, S. J. Bittner, S. G. Nijsure, K. A. Amin, M. Minkoff and A. F. Wagner, *J. Chem. Phys. A*, 2004, **108**, 9979–9997.
- 72 F. Bianchi, T. Kurtén, M. Riva, C. Mohr, M. P. Rissanen, P. Roldin, T. Berndt, J. D. Crounse, P. O. Wennberg, T. F. Mentel, J. Wildt, H. Junninen, T. Jokinen, M. Kulmala, D. R. Worsnop, J. A. Thornton, N. Donahue, H. G. Kjaergaard and M. Ehn, *Chem. Rev.*, 2019, **119**, 3472–3509.
- 73 P. Glarborg, H. Hashemi, S. Cheskis and A. W. Jasper, *J. Phys. Chem. A*, 2021, **125**, 1505–1516.

