



Cite this: *Phys. Chem. Chem. Phys.*,
2022, 24, 1367

Auger electron angular distributions following excitation or ionization from the Xe 3d and F 1s levels in xenon difluoride†

Ruaridh Forbes,^a Paul Hockett,^b Ivan Powis,^c John D. Bozek,^d
Stephen T. Pratt^e and David M. P. Holland^{*f}

Linearly polarized synchrotron radiation has been used to record polarization dependent, non-resonant Auger electron spectra of XeF₂, encompassing the bands due to the xenon M₄₅N₁N₄₅, M₄₅N₂₃N₄₅, M₄₅N₄₅N₄₅ and M₄₅N₄₅V and fluorine KVV transitions. Resonantly excited Auger spectra have been measured at photon energies coinciding with the Xe 3d_{5/2} → σ* and the overlapped Xe 3d_{3/2}/F 1s → σ* excitations in XeF₂. The non-resonant and resonantly excited spectra have enabled the Auger electron angular distributions, as characterized by the β_A parameter, to be determined for the M₄₅N₄₅N₄₅ transitions. In the photon energy range over which the Auger electron angular distributions were measured, theoretical results indicate that transitions into the εf continuum channel dominate the Xe 3d photoionization in XeF₂. In this limit, the theoretical value of the atomic alignment parameter (A₂₀) characterizing the core ionized state becomes constant. This theoretical value has been used to obtain the Auger electron intrinsic anisotropy parameters (α₂) from the β_A parameters extracted from our non-resonant Auger spectra. For a particular Auger transition, the electron kinetic energy measured in the resonantly excited spectrum is higher than that in the directly ionized spectrum, due to the screening provided by the electron promoted into the σ* orbital. The interpretation of the F KVV Auger band in XeF₂ has been discussed in relation to previously published one-site populations of the doubly charged ions (XeF₂²⁺). The experimental results show that the ionization energies of the doubly charged states predominantly populated in the decay of a vacancy in the F 1s orbital in XeF₂ tend to be higher than those populated in the decay of a vacancy in the Xe 4d level in XeF₂.

Received 20th October 2021,
Accepted 16th December 2021

DOI: 10.1039/d1cp04797c

rsc.li/pccp

1. Introduction

Recent theoretical studies,^{1,2} employing methods that allow the inclusion of many-body phenomena, have investigated the single and double ionization spectra associated with the valence and Xe 4d orbitals in the xenon fluorides (XeF₂, XeF₄, and XeF₆). The influence of intra- and interatomic relaxation processes on the decay of a vacancy in the Xe 4d level, through

the emission of an Auger electron, was also studied. Subsequent work³ enabled relativistic effects and electron correlation to be considered. Of particular significance to the present work on the Xe 3d and F 1s Auger spectra of XeF₂ are the calculated one-site and two-site populations of each dicationic state in the double ionization spectra.² Such populations have been discussed previously in relation to Auger spectra.^{4,5} Here, we employ the calculated populations for XeF₂ to assess the doubly ionized final states reached in the Auger decay of a vacancy in the F 1s shell and show that they differ from those populated in the Auger decay of a vacancy in the Xe 4d level of XeF₂.⁶

Our studies of the Auger processes initiated by the creation of a vacancy in the Xe 3d level of XeF₂ are timely as they add to the fundamental spectroscopic information required to interpret recent time-resolved X-ray pump/X-ray probe molecular dynamics experiments on this molecule.⁷

Several previous experimental investigations^{6,8–11} have studied the effect of the electronegative fluorine ligands on the central xenon atom in XeF₂ by measuring the ionization energies of the Xe 4d, 4p, 4s and 3d levels. A comparison between these

^a PULSE Institute, SLAC National Accelerator Laboratory, 2575 Sand Hill Road, Menlo Park, CA 94025, USA

^b National Research Council of Canada, 100 Sussex Dr. Ottawa, ON K1A 0R6, Canada

^c School of Chemistry, University of Nottingham, Nottingham NG7 2RD, UK

^d Synchrotron SOLEIL, l'Orme des Merisiers, Saint-Aubin, BP 48 91192 Gif-sur-Yvette, France

^e Chemical Sciences and Engineering Division, Argonne National Laboratory, Lemont, IL 60439, USA

^f Daresbury Laboratory, Daresbury, Warrington, Cheshire WA4 4AD, UK
E-mail: david.holland@stfc.ac.uk

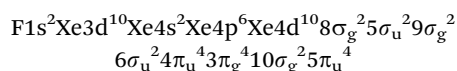
† Electronic supplementary information (ESI) available. See DOI: 10.1039/d1cp04797c

values and those of the corresponding levels in the atomic species shows that withdrawal of valence electron density from the xenon atom leads to an increase in the ionization energies. The Xe 4d and 3d shake-up spectra have also been recorded, and assignments have been proposed for some of the satellite states.^{6,10} Southworth *et al.*¹² have measured the total and partial ion yields of XeF₂ associated with the Xe 3d_{5/2}, Xe 3d_{3/2} and F 1s levels in the photon energy range 660–740 eV. Their results suggest that the molecule with the 3d core hole may dissociate through Coulomb repulsion. The M₄₅N₄₅N₄₅ Auger spectrum of Xe in XeF₂ has been recorded by Aksela *et al.*¹³ using Mg K α (1.254 keV) radiation but the electron angular distributions could not be measured.

In the present work, linearly polarized synchrotron radiation has been employed to record polarization dependent electron spectra encompassing all the Auger decay transitions associated with an initial Xe 3d or F 1s vacancy in XeF₂. Non-resonant (normal) Auger spectra of the M₄₅N₄₅N₄₅ transition have been measured at several photon energies between 687 and 800 eV, thereby allowing the electron angular distributions, as characterized by the β_A parameter, to be determined. In addition, a wide kinetic energy range spectrum, encompassing the structure due to the M₄₅N₁N₄₅, M₄₅N₂₃N₄₅, M₄₅N₄₅N₄₅ and M₄₅N₄₅V transitions, has been measured at 705 eV, using two polarization orientations. Resonantly excited Auger spectra, covering a similar kinetic energy range, have been recorded at 669.9 and 682.8 eV, coinciding with the Xe 3d_{5/2} $\rightarrow$ σ^* and the unresolved Xe 3d_{3/2} $\rightarrow$ σ^* and F 1s $\rightarrow$ σ^* transitions, respectively.¹² These resonant spectra have enabled the shifts in the Auger electron kinetic energy, due to the screening provided by the electron excited into the σ^* orbital, to be studied. Finally, the spectrum associated with the F KVV transition has been recorded at a photon energy of 800 eV. The structure observed in this spectrum is discussed in relation to the one-site and two-site populations of the doubly charged final states.²

Our discussion of the Auger electron angular distributions follows those given recently for the iodine M₄₅N₄₅N₄₅¹⁴ and N₄₅VV¹⁵ transitions in CH₃I. The continuum multiple scattering – X α (CMS-X α) approach¹⁶ has been used to determine the spatial distribution of molecular axes (which depends upon the relative values of the dipole-allowed, symmetry-resolved, photoionization cross sections^{17,18}) following Xe 3d ionization in XeF₂, as well as to calculate the partial *l*-wave contributions to the Xe 3d photoionization cross section. We use these theoretical predictions to help interpret the measured Auger electron angular distributions.

According to our calculations (ESI,† Table S1), the ground state molecular orbital configuration of XeF₂ (*D* _{∞ h symmetry) may be given as}



where the K, L, M₁ and M₂₃ subshells of Xe have not been listed. In this paper we will use atomic or molecular type notation according to the bonding character and localization of the orbital concerned. Photoelectron spectra have established the valence orbital ionization energies.^{9,19–22} The lowest unoccupied molecular orbital (LUMO) is a delocalized antibonding orbital

having σ_u symmetry,^{12,23} and transitions into this orbital have been observed in valence shell absorption spectra.^{24,25}

2. Experimental apparatus and procedure

The Auger electron spectra were recorded with a VG Scienta R4000 hemispherical electron energy analyser mounted on the soft X-ray undulator-based PLÉIADES beamline at the SOLEIL synchrotron radiation facility. Detailed descriptions of the beamline and station instrumentation have been reported previously^{14,26,27} so only a summary is given here.

The undulator allows the electric vector of the linearly polarized synchrotron radiation to lie either parallel or perpendicular to the acceptance axis of the electron analyser. The beamline employs a modified Petersen type monochromator. The spectra were recorded using a grating having 600 lines per mm together with an exit slit width of 100 μ m. This combination results in a theoretical optical resolution of $\sim$ 480 meV at a photon energy of 670 eV and 640 meV at 800 eV.

The electron spectrometer was mounted in a fixed position and photoionization occurred within an in-house designed cell⁶ to minimize the throughput of the sample. The inlet system and the spectrometer were extensively passivated prior to the measurement of spectra to minimize the contribution from atomic xenon. The Auger spectra were recorded with an analyser pass energy of 50 eV and a 0.5 mm curved entrance slit, resulting in a spectrometer resolution of 62.5 meV.

The translational Doppler broadening contributes to the overall line width. For electron kinetic energies of 520 eV, which is appropriate for the Xe M₄₅N₄₅N₄₅ transition in XeF₂, and 640 eV, which is appropriate for the F KVV transition, the corresponding Doppler width contributions are $\sim$ 22 and 24 meV, respectively.

The Auger electron angular distribution, from excitation of randomly oriented molecules or unpolarized atoms using 100% linearly polarized radiation, is given by^{28–30}

$$\frac{dW_{J_i \rightarrow J_f}(\theta)}{d\Omega} = \frac{W_{J_i \rightarrow J_f}^T}{4\pi} [1 + \beta_A P_2(\cos \theta)] \quad (1)$$

where $W_{J_i \rightarrow J_f}^T$ is the total Auger decay rate between initial and final states having total angular momentum J_i and J_f respectively, θ is the angle between the momentum of the ejected Auger electron and the electric vector of the linearly polarized radiation (the alignment axis), β_A is a parameter describing the Auger electron angular distribution for that particular transition, $P_2(\cos \theta)$ is the Legendre polynomial of second degree, and $d\Omega$ is the differential solid angle element in the direction specified by the polar angle θ . Eqn (1) can be rearranged into the more convenient form²⁶

$$\beta_A = \frac{2(I_{\text{par}} - I_{\text{perp}})}{(I_{\text{par}} + 2I_{\text{perp}})} \quad (2)$$

where I_{par} and I_{perp} are the normalized electron intensities for parallel and perpendicular polarization orientations, respectively, relative to the electron detector axis. All the spectra were normalized to the sample pressure, the accumulation time, and the

photon flux prior to processing. The transmission efficiency of the electron analyser was determined as described by Jauhianinen *et al.*,³¹ and all the spectra were corrected for this variation.

The electron kinetic energy scale was calibrated by recording a spectrum encompassing the Xe $M_{45}N_{45}N_{45}$ Auger decay in XeF₂ and using the value of 530.88 eV for the peak due to the unresolved 1D_2 and 1G_4 final states in the $M_4N_{45}N_{45}$ transition.¹³

3. Fitting of the $M_{45}N_{45}N_{45}$ Auger spectrum

The procedure used to fit the peaks due to the xenon $M_{45}N_{45}N_{45}$ Auger decay in XeF₂ was similar to that already described in detail in relation to the iodine $M_{45}N_{45}N_{45}$ transitions in CH₃I.¹⁴ Hence only a brief summary, together with information specific to XeF₂, is given here.

The two-hole final state configuration $4d^{-2}$ gives rise to nine terms¹³ (Table 1), but not all of these transitions were resolved in our spectra. The three pairs of states that were not resolved correspond to (1D_2 , 1G_4), (3P_1 , 3P_0) and (3F_3 , 3F_2). Thus, six peaks (numbered 1–6, Table 1) were fitted to the structure due to the $M_5N_{45}N_{45}$ transition, and another six peaks (numbered 7–12) were fitted to the $M_4N_{45}N_{45}$ transition. The fitting allows the electron kinetic energy, the angular distribution parameter β_A , and the intensity, of each peak to be extracted from the experimental spectra. Aksela *et al.*¹³ recorded spectra, at a photon energy of 1.254 keV, encompassing the $M_{45}N_{45}N_{45}$ Auger decay in atomic Xe and in XeF₂, and found that the width of the

sharpest component, due to the 3F_4 final state in the $M_4N_{45}N_{45}$ transition, was 0.15–0.2 eV broader in XeF₂ than in Xe. Recent measurements of the inherent lifetime widths of the Xe $3d^{-1}$ states in atomic xenon³² have determined widths of ~ 500 meV for the Xe $3d_{5/2}^{-1}$ and $3d_{3/2}^{-1}$ states. We adopt a lifetime width of 750 meV for the Xe $3d$ initial hole state in XeF₂. Furthermore, we set the total instrumental broadening (representing the combined spectrometer broadening and the Doppler contribution) to 75 meV. The photon resolution does not affect the fitting of the non-resonant Auger spectra. In our fitting, the initial core-hole lifetime and the total instrumental broadening were held at fixed values of 750 and 75 meV, respectively.

For the resonantly excited Auger spectra, if the photon bandwidth is narrower than the level width of the neutral excited state (Xe $3d_{5/2}^{-1}\sigma^*$ or Xe $3d_{3/2}^{-1}\sigma^*$ in our case), then the initial state width can be neglected and replaced by a Gaussian representing the photon bandwidth. Our fitting used fixed Gaussian widths of 485 and 500 meV at photon energies of 669.9 and 682.8 eV, respectively, corresponding to the Xe $3d_{5/2}^{-1}\sigma^*$ and Xe $3d_{3/2}^{-1}\sigma^*$ resonances.

4. Calculation of the Xe $3d$ in XeF₂ photoionization partial cross section and the spatial distribution of molecular axes using the CMS- $X\alpha$ approach

The Xe $3d$ partial cross section in XeF₂ photoionization and the spatial distribution of molecular axes in the Xe $3d^{-1}$ ionized state were calculated using methods described previously.^{15,16}

Table 1 Experimental Auger electron kinetic energies and intrinsic anisotropy parameters for the xenon $M_{45}N_{45}N_{45}$ transitions in XeF₂, and calculated intrinsic anisotropy parameters for the $M_{45}N_{45}N_{45}$ transitions in atomic Xe

Transition	Final state	Peak number	Auger electron kinetic energy ^a (eV)	β_A (expt.) ^{bc}	Experimental			Calculated	
					α_2 (expt) ^{cd}	α_2^e	α_2^f	α_2^g	α_2^h
$M_5N_{45}N_{45}$	1S_0	1	512.97 $\pm$ 0.01	0.106 $\pm$ 0.059	−0.505 $\pm$ 0.281	−0.71 $\pm$ 0.15		−1.069	−1.069
	1D_2	2	518.34 $\pm$ 0.01	0.086 $\pm$ 0.025	−0.410 $\pm$ 0.119	−0.7 $\pm$ 0.3	−0.17	−0.131	−0.124
	1G_4							−0.719	−0.710
	3P_1	3	519.21 $\pm$ 0.03	0.129 $\pm$ 0.042	−0.614 $\pm$ 0.200	−0.87		−0.748	−0.749
	3P_0							−1.069	−1.069
	3P_2	4	519.91 $\pm$ 0.02	0.075 $\pm$ 0.017	−0.357 $\pm$ 0.081	−0.2 $\pm$ 0.4		−0.389	−0.371
$M_4N_{45}N_{45}$	3F_3	5	521.44 $\pm$ 0.01	0.014 $\pm$ 0.013	−0.067 $\pm$ 0.062	−0.25 $\pm$ 0.12	−0.1	0.331	0.336
	3F_2							−0.222	−0.226
	3F_4	6	523.06 $\pm$ 0.01	−0.065 $\pm$ 0.011	0.310 $\pm$ 0.052	0.3 $\pm$ 0.2		0.378	0.386
	1S_0	7	525.52 $\pm$ 0.01	0.167 $\pm$ 0.028	−0.835 $\pm$ 0.140	−1.0	−1.0	−1.000	
	1D_2	8	530.88 $\pm$ 0.01	0.052 $\pm$ 0.010	−0.260 $\pm$ 0.050	−0.6 $\pm$ 0.4	−0.33 $\pm$ 0.1	0.050	
	1G_4							−0.653	
$M_4N_{45}N_{45}$	3P_1	9	531.71 $\pm$ 0.02	0.113 $\pm$ 0.023	−0.565 $\pm$ 0.115	0.5 $\pm$ 0.4	−0.43 $\pm$ 0.1	−0.799	
	3P_0							−1.000	
	3P_2	10	532.63 $\pm$ 0.01	−0.032 $\pm$ 0.011	0.160 $\pm$ 0.055	0.39 $\pm$ 0.13	0.24 $\pm$ 0.16	0.378	
	3F_3	11	533.95 $\pm$ 0.01	−0.083 $\pm$ 0.016	0.415 $\pm$ 0.080	0.49 $\pm$ 0.15	0.53 $\pm$ 0.22	0.432	
	3F_2							0.738	
	3F_4	12	535.65 $\pm$ 0.03	0.256 $\pm$ 0.175	−1.280 $\pm$ 0.875	−0.9 $\pm$ 0.5		−0.826	

^a This work. Auger electron kinetic energies obtained by fitting the spectrum recorded at a photon energy of 710 eV. ^b This work. The average β_A parameter for the M_4 ($3d_{3/2}^{-1}$) core hole was obtained by averaging the values obtained at four photon energies (701, 705, 710 and 800 eV). ^c This work. The quoted uncertainty is due only to electron counting statistics and peak fitting, and does not take any systematic errors into account. ^d This work. The experimental α_2 parameters were obtained from the experimental β_A parameters using theoretical values $A_{20}(D_{5/2}) = -0.21$ and $A_{20}(D_{3/2}) = -0.20$. ^e Experimental results on atomic xenon obtained at a photon energy of 697.75 eV.⁴⁰ ^f Experimental results on atomic xenon using proton impact at an energy of 3.6 MeV.³⁹ ^g Calculated (Chen⁵⁶) α_2 parameters for the $M_{45}N_{45}N_{45}$ transitions in atomic xenon. ^h Calculated (Tulkki *et al.*⁵⁷) α_2 parameters for the $M_5N_{45}N_{45}$ transitions in atomic xenon.

Briefly, electric dipole photoionization matrix elements were computed using the static-exchange CMS- $X\alpha$ method.^{16,33,34} A model molecular potential, with the Xe-F separation set to be 1.9754 Å as determined by a MP2/def2-QZVP geometry optimisation, was constructed by placing overlapping spherical regions on the atomic centres, all enclosed within an outer spherical region placed at the centre of mass. Bound state eigenvalues and eigenfunctions within this potential were found, utilising an angular basis of spherical harmonic functions (ranging up to $l_{\max} = 7, 5$ and 3 for respectively the outer-, Xe- and F- regions), and numerical integration to obtain the radial terms. Then, using the $X\alpha$ exchange potential to relate electron density to potential, the trial model potential was iterated to self-consistency. After adapting the asymptotic potential region to have the correct long range Coulombic interaction appropriate for a separating ion-electron system, continuum electron wavefunctions were generated using multiple scattering equations with an expanded angular basis having $l_{\max} = 10, 8, 6$ (outer-, Xe-, F-regions respectively). Hence, utilising orthogonal one-electron initial (bound) orbital and continuum functions from this potential, electric dipole photoionization matrix elements were calculated, permitting the Xe 3d photoionization partial cross section and the spatial distribution of molecular axes, $I(\theta_{\text{ion}})$, in the Xe 3d⁻¹ ionized state to be evaluated.

The alignment of molecular axes can be expressed as a distribution having the form^{15,35}

$$I(\theta_{\text{ion}}) = 1 + \beta_{\text{ion}} P_2(\cos \theta_{\text{ion}}) \quad (3)$$

where β_{ion} is the normalized angular distribution parameter, $P_2(\cos \theta_{\text{ion}})$ is the Legendre polynomial of second order, and θ_{ion} is the angle between the electric vector of the linearly polarized radiation and the molecular axis. The calculation of β_{ion} from the dipole matrix elements was more fully described in connection with our previous investigation of CH₃I.¹⁵

5. Results and discussion

5.1. Overview

An electron spectrum recorded at a photon energy of 705 eV, using perpendicularly polarized radiation, is plotted in Fig. 1. This spectrum encompasses the structure due to the principal Auger transitions ($M_{45}N_1N_{45}$, $M_{45}N_{23}N_{45}$, $M_{45}N_{45}N_{45}$ and $M_{45}N_{45}V$) associated with the decay of the Xe 3d_{5/2}⁻¹ or 3d_{3/2}⁻¹ hole states in XeF₂. Apart from the bands due to the $M_{45}N_{45}V$ transition, these Auger bands strongly resemble the corresponding bands in atomic xenon.³⁶⁻⁴¹ As will be discussed in sections 5.2.4 and 5.2.5, it appears likely that structure arising from the $M_{45}N_{45}V$ transition in XeF₂ occurs in the same electron kinetic energy range as that associated with the fluorine KVV transition in XeF₂. In particular, our experimental results suggest that the peak observed at a kinetic energy of 601.53 eV (Fig. 1) should be attributed to the F KVV decay.

The calculations performed by Partanen *et al.*⁴¹ indicate that, in atomic xenon, satellite Auger transitions originating

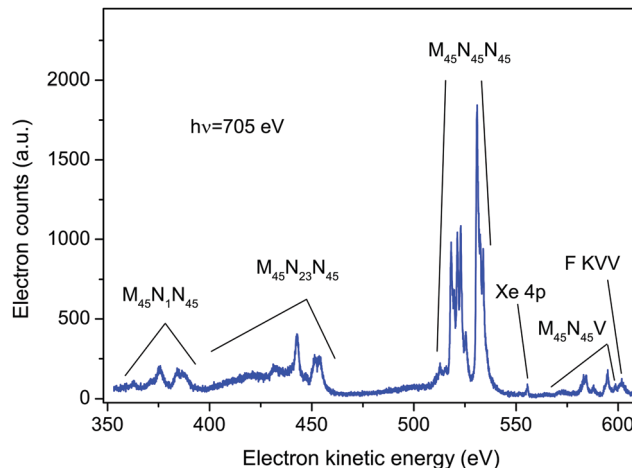


Fig. 1 The non-resonant Auger electron spectrum of XeF₂ recorded at a photon energy of 705 eV using perpendicularly polarized radiation. Most of the structure is associated with the Xe $M_{45}N_1N_{45}$, $M_{45}N_{23}N_{45}$, $M_{45}N_{45}N_{45}$, and $M_{45}N_{45}V$ Auger transitions, although a weak peak arises from direct ionization of the Xe 4p level in XeF₂. In addition, the peak at a kinetic energy of 601.53 eV is attributed to a F KVV transition in XeF₂.

from the 3d⁻¹5p⁻¹ state also contribute in this electron kinetic energy range. Fig. S1 in ESI† shows a spectrum recorded at a photon energy of 705 eV using parallel polarized radiation. In that spectrum an additional peak appears at a binding energy of 216.5 ± 0.2 eV (corresponding to a kinetic energy of 488.5 eV) due to photoionization from the Xe 4s orbital in XeF₂. As the photoelectron angular distribution parameter β_e for the 4s orbital is ~ 2 , a peak due to this orbital is not discernible in the spectrum recorded with perpendicularly polarized radiation. Direct photoionization from the Xe 4p orbital in XeF₂ gives rise to the weak peak occurring at a kinetic energy of ~ 556 eV in Fig. 1. Photoionization from the Xe 4s and Xe 4p orbitals in XeF₂ has been a subject of considerable interest^{9,42} because electron correlation strongly affects the spectrum. However, this topic will not be discussed further in the present work.

Polarization dependent, non-resonant Auger spectra encompassing the $M_{45}N_{45}N_{45}$ transition were measured at photon energies of 687, 689, 701 705, 710 and 800 eV. In addition, polarization dependent resonantly excited Auger spectra, covering a wide kinetic energy range, were measured at photon energies of 669.9 and 682.8 eV. The lower of these energies coincides with the Xe 3d_{5/2} $\rightarrow$ σ^* transition.¹² The higher energy (682.8 eV) corresponds to the excitation energy of a single peak observed in the ion yield.¹² However, that peak was attributed to the unresolved Xe 3d_{3/2} $\rightarrow$ σ^* and F 1s $\rightarrow$ σ^* transitions, with the calculated oscillator strength of the latter transition being ~ 20 times larger than that of the former. Furthermore, the peak is superimposed upon a continuum mainly due to ionization of the Xe 3d_{5/2} level.

5.2. Auger electron band structure

5.2.1. $M_{45}N_{45}N_{45}$ transitions. Polarization dependent electron spectra encompassing the structure due to the $M_{45}N_{45}N_{45}$ and $M_{55}N_{45}N_{45}$ Auger transitions in XeF₂ are plotted

in Fig. 2 (a non-resonant spectrum recorded at a photon energy of 710 eV, and resonantly excited spectra recorded at photon energies of 669.9 and 682.8 eV). Each of the $M_4N_{45}N_{45}$ and $M_5N_{45}N_{45}$ Auger electron bands was fitted with 6 peaks, as shown in Fig. 3 for the non-resonant spectrum.

At these three photon energies, and in the kinetic energy range over which each spectrum is plotted, direct photoionization should not give rise to any structure. However, Partanen *et al.*⁴¹ have calculated that features due to the $3d_{5/2}^{-1}5p^{-1} \rightarrow 4d^{-2}5p^{-1}$ transitions in atomic xenon should occur in the kinetic energy range slightly lower than that encompassing the $M_5N_{45}N_{45}$ Auger bands, and presumably similar transitions arising from the $3d_{3/2}^{-1}5p^{-1}$ initial state would lie at slightly higher kinetic energies. The analogous satellite Auger decays in XeF_2 may be responsible for the background underlying the bands due to the $M_{45}N_{45}N_{45}$ transitions. Our fitted spectrum (Fig. 3) shows that the background contribution is taken into account in a satisfactory manner. The $M_4N_{45}N_{45} : M_5N_{45}N_{45}$ intensity branching ratio, obtained from the fits of the non-resonant Auger spectra, is shown in Fig. S2 (ESI†).

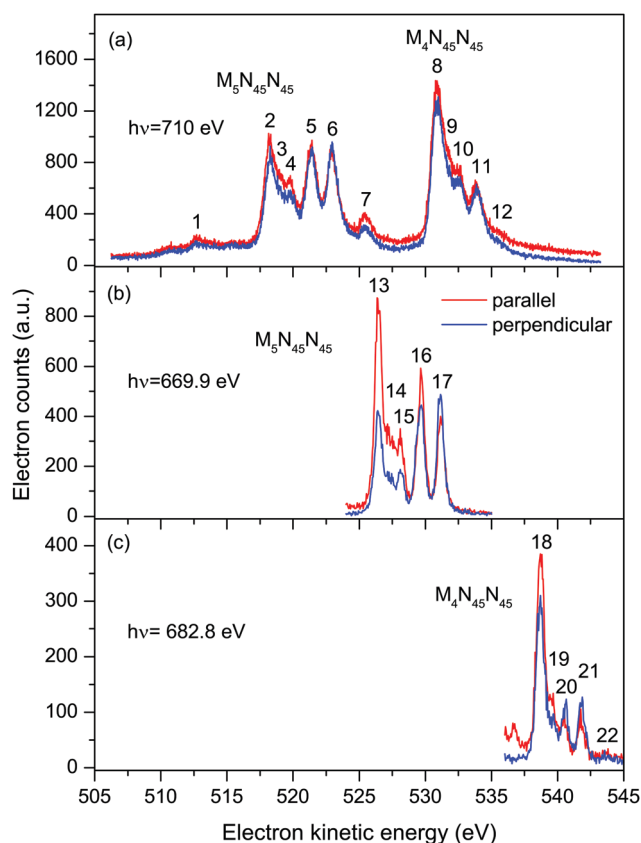


Fig. 2 Auger electron bands due to the $M_{45}N_{45}N_{45}$ transitions in XeF_2 recorded with parallel (red) and perpendicularly (blue) polarized radiation. The non-resonant spectrum (a) was measured at a photon energy of 710 eV. The resonantly excited spectra [(b) and (c)] were measured at photon energies of 669.9 eV ($3d_{5/2} \rightarrow \sigma^*$ transition) and 682.8 eV ($3d_{3/2} \rightarrow \sigma^*$ transition). The kinetic energies of the peaks numbered 1–22 are given in Tables 1 and 2.

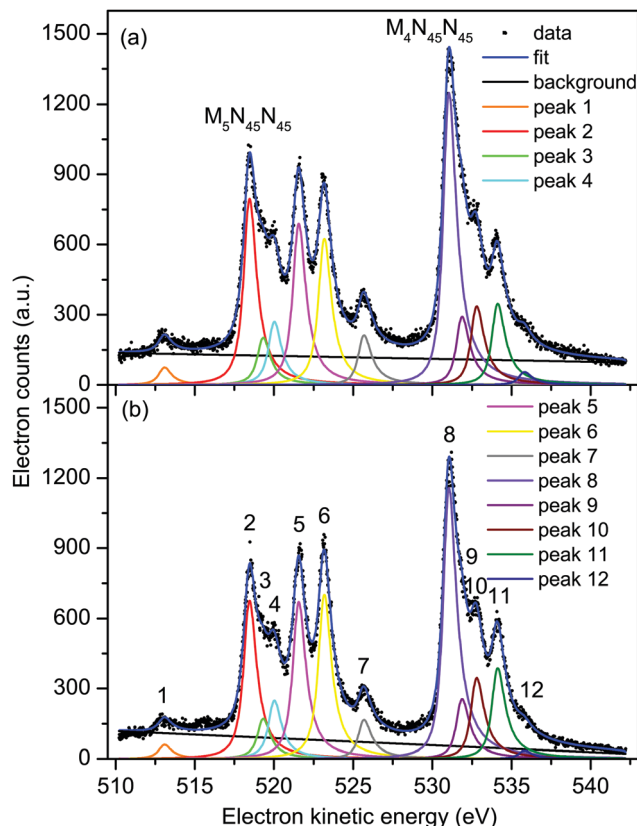


Fig. 3 Non-resonant Auger electron bands due to the $M_{45}N_{45}N_{45}$ transitions in XeF_2 recorded with parallel (a) and perpendicularly (b) polarized radiation at a photon energy of 710 eV. The raw and the fitted data are shown (see text for details). The kinetic energies of the peaks numbered 1–12 are given in Table 1.

Table 1 contains the results extracted from the fitting of the non-resonant $M_{45}N_{45}N_{45}$ spectra. The peaks are assigned by analogy with the corresponding peaks in atomic xenon.³⁷ Our analysis yields an electron kinetic energy of 518.34 eV for peak 2, due to the unresolved 1D_2 and 1G_4 states belonging to the $M_5N_{45}N_{45}$ transition. Using this energy, together with our calibration value of 530.88 eV for the corresponding states in the $M_4N_{45}N_{45}$ transition, results in a M_4 – M_5 spin-orbit splitting of 12.54 eV. Bancroft *et al.*⁹ have reported a spin-orbit splitting of 12.68 eV for the $\text{Xe } 3d_{5/2}^{-1}$ and $3d_{3/2}^{-1}$ hole states in XeF_2 .

The Auger electron kinetic energies for the corresponding states extracted from the fitting of the resonantly excited spectra (Figs. S3 and S4) at 669.9 eV ($3d_{5/2} \rightarrow \sigma^*$) and 682.8 eV ($3d_{3/2} \rightarrow \sigma^*$) are listed in Table 2. In these spectra, the final states are singly ionized with an electron promoted into the σ^* LUMO. Thus, in Table 2, the states are denoted, for example, as $4d^{-2}(^3F_4)\sigma^*$. If the weak peak associated with the 1S_0 state is ignored, the Auger electron kinetic energies, following resonant excitation at 669.9 or 682.8 eV, are shifted, on average, to higher energies by 8.15 and 7.90 eV, respectively, compared to the corresponding values in the non-resonant spectrum (Table 1).

5.2.2. $M_{45}N_1N_{45}$ transitions. Although single configuration calculations have been found to provide a satisfactory description of the $M_{45}N_{45}N_{45}$ Auger transitions in atomic xenon,^{38,41} this is

Table 2 Experimental Auger electron kinetic energies and intrinsic anisotropy parameters for the xenon $M_{45}N_{45}N_{45}$ transitions in XeF_2 following resonant excitation at 669.9 eV ($3d_{5/2} \rightarrow \sigma^*$) or 682.8 eV ($3d_{3/2} \rightarrow \sigma^*$)

Transition and final state	Peak number	Auger electron kinetic energy (eV)	β_A (expt) ^a	α_2 (expt) ^{ab}
$\text{Xe } 3d_{5/2}^{-1}\sigma^* \rightarrow 4d^{-2}\sigma^*$ ^c				
$4d^{-2}(^1D_2 + ^1G_4)\sigma^*$	13	526.45 $\pm$ 0.01	0.498 $\pm$ 0.02	−0.352 $\pm$ 0.01
$4d^{-2}(^3P_{0,1})\sigma^*$	14	527.32 $\pm$ 0.01	0.564 $\pm$ 0.05	−0.399 $\pm$ 0.04
$4d^{-2}(^3P_2)\sigma^*$	15	528.12 $\pm$ 0.01	0.355 $\pm$ 0.04	−0.251 $\pm$ 0.03
$4d^{-2}(^3F_2 + ^3F_3)\sigma^*$	16	529.65 $\pm$ 0.01	0.060 $\pm$ 0.02	−0.042 $\pm$ 0.01
$4d^{-2}(^3F_4)\sigma^*$	17	531.17 $\pm$ 0.01	−0.143 $\pm$ 0.02	0.101 $\pm$ 0.01
$\text{Xe } 3d_{3/2}^{-1}\sigma^* \rightarrow 4d^{-2}\sigma^*$ ^d				
$4d^{-2}(^1D_2 + ^1G_4)\sigma^*$	18	538.70 $\pm$ 0.01	0.132 $\pm$ 0.02	−0.093 $\pm$ 0.01
$4d^{-2}(^3P_{0,1})\sigma^*$	19	539.59 $\pm$ 0.02	0.182 $\pm$ 0.06	−0.129 $\pm$ 0.04
$4d^{-2}(^3P_2)\sigma^*$	20	540.53 $\pm$ 0.01	−0.373 $\pm$ 0.04	0.264 $\pm$ 0.03
$4d^{-2}(^3F_2 + ^3F_3)\sigma^*$	21	541.81 $\pm$ 0.01	−0.406 $\pm$ 0.04	0.287 $\pm$ 0.03
$4d^{-2}(^3F_4)\sigma^*$	22	543.69 $\pm$ 0.07	−0.164 $\pm$ 0.3	0.116 $\pm$ 0.21

^a The quoted uncertainty is due only to electron counting statistics and peak fitting, and does not take any systematic errors into account. ^b The experimental α_2 parameters were obtained from the experimental β_A parameters using $A_{20} = -\sqrt{2}$. ^c Auger electron kinetic energies and β_A parameters obtained by fitting the resonantly excited spectrum recorded at a photon energy of 669.9 eV. ^d Auger electron kinetic energies and β_A parameters obtained by fitting the resonantly excited spectrum recorded at a photon energy of 682.8 eV.

not the case for the $M_{45}N_1N_{45}$ and $M_{45}N_{23}N_{45}$ transitions where electron correlation affects the Auger lines involving the 4s and 4p subshells. Partanen *et al.*⁴¹ have performed configuration interaction calculations, incorporating final ionic state interactions, which take electron correlation into account. For the $M_{45}N_1N_{45}$ Auger decay, the configuration mixing involved the $4s^{-1}4d^{-1}$, $4p^{-1}4d^{-2}4f$ and $4d^{-4}4f^2$ states. The simulated spectra, generated from the calculated kinetic energies and transition probabilities, showed good agreement with the experimental spectrum. As the experimental spectrum for the $M_{45}N_1N_{45}$ transition in atomic xenon is very similar to the corresponding spectrum in XeF_2 , we use the predictions from Partanen *et al.*⁴¹ to discuss the present results.

Six intense Auger transitions, together with some weak correlation satellites at lower kinetic energies, are predicted for the $M_{45}N_1N_{45}$ Auger decay.⁴¹ For the M_5 hole, one of these transitions is associated with peak 23 (Fig. 4(a) and Table 3), and the other two transitions with peak 24. Likewise, for the M_4 hole, peaks 27 and 28 are attributed to one and two transitions, respectively. Table 3 lists the kinetic energies of these peaks in XeF_2 , obtained from the non-resonantly excited spectrum recorded at a photon energy of 705 eV. In the resonantly excited spectrum obtained at a photon energy of 669.9 eV, where only those peaks due to transitions involving an M_5 hole will be observed, the two shifted peaks appear at 379.25 eV (peak 25) and 383.78 eV (peak 26). Similarly, in the resonantly excited spectrum obtained at a photon energy of 682.8 eV, where only those peaks due to transitions involving an M_4 hole will be observed, the two shifted peaks appear at 391.76 eV (peak 29) and 395.11 eV (peak 30). Thus, in the resonantly excited spectra encompassing the $M_{45}N_1N_{45}$ transition, the peaks are shifted by 8.01 eV towards higher kinetic energy.

5.2.3. $M_{45}N_{23}N_{45}$ transitions. Partanen *et al.*⁴¹ showed that mixing amongst the $4p^{-1}4d^{-1}$, $4d^{-3}4f$ and $4s^{-1}4p^{-1}$ configurations was able to produce simulated spectra for the $M_5N_{23}N_{45}$ and $M_4N_{23}N_{45}$ transitions in atomic xenon which were in satisfactory agreement with the experimental results. The calculations

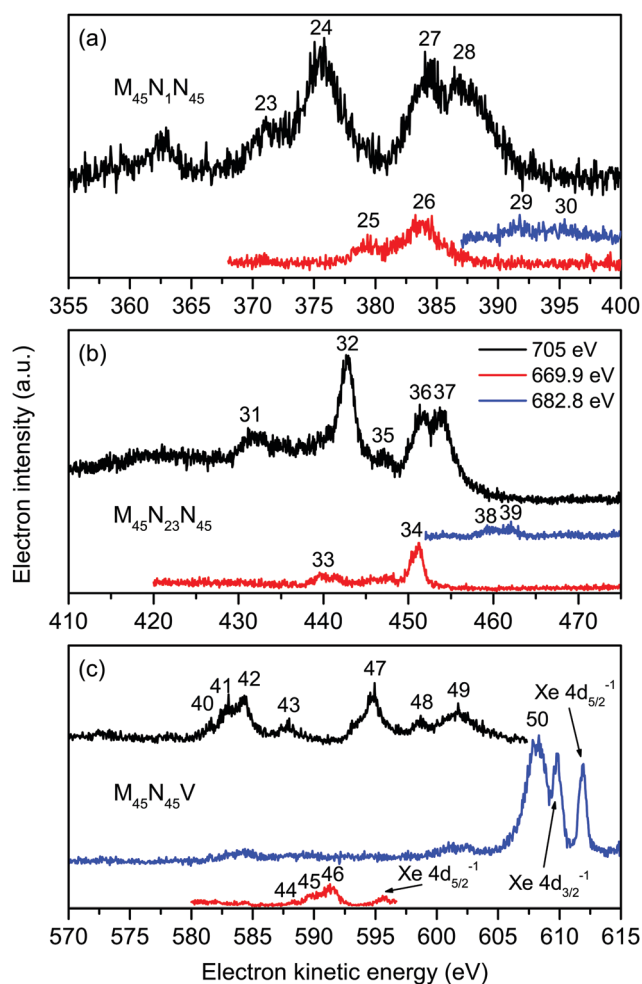


Fig. 4 The non-resonant ($h\nu = 705$ eV), and resonantly excited ($h\nu = 669.9$ and 682.8 eV) Auger electron spectra of XeF_2 showing bands due to the $\text{Xe } M_{45}N_1N_{45}$ (a), $\text{Xe } M_{45}N_{23}N_{45}$ (b) and $\text{Xe } M_{45}N_{45}V$ and F KVV (c) transitions. All of these spectra were measured with perpendicularly polarized radiation. The kinetic energies of the peaks numbered 23–50 are given in Table 3.

Table 3 Auger electron kinetic energies for the xenon $M_{45}N_1N_{45}$, $M_{45}N_{23}N_{45}$ and $M_{45}N_{45}V$, and the fluorine KVV transitions in XeF_2

Transition	Peak number	Auger electron kinetic energy (eV)		
		$h\nu = 705$ eV	$h\nu = 669.9$ eV	$h\nu = 682.8$ eV
$M_5N_1N_{45}$	23	371.03		
	24	375.46		
	25		379.25	
	26		383.78	
$M_4N_1N_{45}$	27	384.23		
	28	387.13		
	29			391.76
	30			395.11
$M_5N_{23}N_{45}$	31	431.28		
	32	442.83		
	33		439.46	
	34		451.10	
$M_4N_{23}N_{45}$	35	446.93		
	36	451.63		
	37	453.73		
	38			459.93
	39			462.03
$M_5N_{45}V$	40	581.53		
	41	582.98		
	42	584.33		
	43	587.93		
	44		588.35	
	45		589.65	
	46		591.30	
$M_4N_{45}V$	47	594.78		
	48	598.68		
F KVV	49	601.53		
	50			608.16

indicated that the overall shape of the $M_{45}N_{23}N_{45}$ Auger band arose from contributions from many individual transitions, and that in most cases several final state configurations contributed to each line.⁴¹

The structure associated with the $M_{45}N_{23}N_{45}$ Auger decay in XeF_2 strongly resembles that observed in the analogous band in atomic xenon although the kinetic energies of the corresponding features are shifted by ~ 2 eV to lower energy. We discuss only five of the numerous peaks associated with the $M_{45}N_{23}N_{45}$ Auger band (Fig. 4(b)). According to Partanen *et al.*,⁴¹ peaks 31 and 32 involve an M_5 hole, whereas peaks 35–37 involve an M_4 hole (Table 3). The kinetic energies of these peaks obtained in the non-resonantly excited spectrum are listed in Table 3, together with the energies of peaks 33 and 34, and the energies of peaks 38 and 39, obtained from the resonantly excited spectra recorded at photon energies of 669.9 and 682.8 eV, respectively. The kinetic energies of the peaks in the resonantly excited spectra are shifted by ~ 8.3 eV to higher energy compared to the corresponding values in the non-resonant spectrum.

5.2.4. $M_{45}N_{45}V$ transitions. The $M_4N_{45}O_1$ and $M_{45}N_{45}O_{23}$ Auger spectra in atomic xenon have been measured and discussed by Kivimäki *et al.*⁴³ The structure observed in those spectra differs from that occurring in the kinetic energy range

570–615 eV in the spectrum of XeF_2 recorded at a photon energy of 705 eV (Fig. 4(c)). The calculations on XeF_2 performed by Buth *et al.*² predict numerous doubly ionized states, having energies between 90 and 100 eV, with 2h configurations in which one hole is formed in the Xe 4d level and another in a valence orbital. It appears likely that many of the Auger bands observed in XeF_2 (Fig. 4(c)) arise from transitions into these final states.

In the absence of detailed theoretical guidance, we assume that peaks 40–43 in the Auger spectrum of XeF_2 arise from an M_5 hole (Fig. 4(c)). The kinetic energies associated with these peaks are given in Table 3. In the resonantly excited spectrum recorded at a photon energy of 669.9 eV, where structure involving an M_5 initial hole is expected to appear, we observe peaks 44–46, which seem to correspond to peaks 40–42, shifted by ~ 6.8 eV towards higher kinetic energy (Table 3). Thus, a shifted $M_5N_{45}V$ Auger band is observed in the resonantly excited spectrum although the shift is smaller than those measured for the $M_{45}N_1N_{45}$, $M_{45}N_{23}N_{45}$ and $M_{45}N_{45}N_{45}$ transitions.

The resonantly excited spectrum recorded at a photon energy of 682.8 eV is dominated by a broad and intense peak occurring at a kinetic energy of 608.16 eV. A very weak feature also appears between ~ 601 –603 eV. The kinetic energy of the intense peak is slightly lower than those due to the Xe $4d_{5/2,3/2}$ main-lines in XeF_2 which are observed at kinetic energies of ~ 612 and ~ 610 eV (Fig. 4(c)).

Fig. 5 shows sections of the resonantly excited, polarization dependent spectra plotted as a function of binding energy. In the spectrum recorded at 669.9 eV, the shake-up structure associated with the Xe 4d main-lines in XeF_2 resembles that measured and discussed previously.^{6,10} The broad peak observed in the spectra recorded at 682.8 eV does not appear to be part of the expected shake-up structure. The intensity of

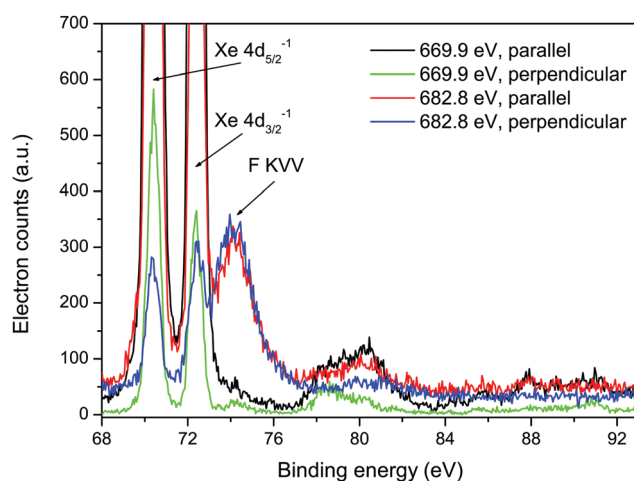


Fig. 5 Resonantly excited electron spectra recorded at photon energies of 669.9 and 682.8 eV, using parallel and perpendicularly polarized radiation, plotted as a function of binding energy. The two intense peaks at binding energies of 70.38 and 72.39 eV arise from the Xe $4d_{5/2}^{-1}$ and $4d_{3/2}^{-1}$ states, and the structure observed between ~ 77 and 83 eV is due to the associated shake-up satellites.⁶

this peak in the spectrum measured with parallel polarized radiation is approximately the same as that in the spectrum measured with perpendicularly polarized radiation. Hence, the electron angular distribution is isotropic. Since the broad peak does not seem to result from a transition involving the Xe atom in XeF_2 , an alternative explanation needs to be sought.

The F KVV Auger spectrum in HF^{44} occurs in the electron kinetic energy range 590–650 eV, and the peak having the lowest energy appears at ~ 595 eV. Auger electron peaks arising from similar transitions and hence possessing similar energies can be anticipated in XeF_2 . Two additional sources of information have been used to guide our interpretation of the broad peak. First, as already mentioned, the calculations performed by Southworth *et al.*¹² showed that the single peak observed in the ion yield at 682.8 eV should be attributed to the overlapped and unresolved $\text{Xe } 3d_{3/2} \rightarrow \sigma^*$ and $\text{F } 1s \rightarrow \sigma^*$ transitions, with the oscillator strength for the latter transition being ~ 20 times larger than that of the former. Second, owing to the shielding provided by the electron promoted into the σ^* orbital, the Auger electron band resulting from resonant excitation ($\text{F } 1s \rightarrow \sigma^*$) will be shifted by a few eV to higher energy compared to the corresponding band due to non-resonant (direct) ionization.

If we assume that peak 49 in the non-resonant ($h\nu = 705$ eV) Auger spectrum of XeF_2 , having an electron kinetic energy of 601.53 eV (Table 3) results from a transition originating from the $(\text{F } 1s)^{-1}$ ionic state, and that the broad peak observed at a kinetic energy of 608.16 eV in the resonantly excited spectrum ($h\nu = 682.8$ eV) corresponds to the same Auger decay process but with the initial neutral excited state being $(\text{F } 1s)^{-1}\sigma^*$, then the shift in kinetic energy is 6.63 eV. Such a shift to higher kinetic energy is consistent with expectations.

Further support for the assignment of peak 49 to a transition involving the F 1s level is provided by the Auger electron angular distribution. For the broad peak at 608.16 eV, the electron angular distribution is isotropic (Fig. 5) as would be expected for excitation from an atomic 1s orbital, since the excited state cannot be aligned. Our Auger electron angular distribution measurements for processes involving the Xe 3d level in XeF_2 show that, although the β_A values are not large, they are nonetheless non-zero. This is consistent with the small, but finite, values of atomic alignment predicted for ionization from the Xe 3d level.³⁰

The absence of any intense peaks in the resonantly excited spectrum recorded at a photon energy of 682.8 eV that could be attributed to peaks 47 and 48 shifted to higher kinetic energies suggests that neither of these peaks is associated with an initial state having a vacancy in the F 1s level. Since the calculated oscillator strength for the $\text{Xe } 3d_{3/2} \rightarrow \sigma^*$ transition is much lower than that for the $\text{F } 1s \rightarrow \sigma^*$ transition,¹² an Auger electron peak associated with the $\text{M}_4\text{N}_{45}\text{V}$ decay, but shifted to higher energy due to the shielding, would be expected to be of low intensity in the resonantly excited spectrum. Based on the limited evidence available, we tentatively attribute peaks 47 and 48 to the $\text{M}_4\text{N}_{45}\text{V}$ Auger transitions. The very weak feature, observed at kinetic energies between 601 and 603 eV in the

resonantly excited spectrum, might correspond to peak 47, shifted to slightly higher kinetic energy.

5.2.5. F KVV transition. Theoretical investigations of the F KLL Auger spectra of BF_3 ⁴ and SiF_4 ⁵ have discussed the assignment of the peaks observed in the experimental spectra in relation to whether the two valence electron vacancies are predicted to be located on the same fluorine atom (so-called one-site population) or are located on different fluorine atoms (two-site population). Auger decay is often considered as an essentially intra-atomic relaxation process based upon the one-centre approximation.⁴⁵ Within this model, the dicationic states associated with the prominent peaks in the Auger spectrum will be those possessing a high one-site population (two-hole density) on the atomic site containing the initial core hole.

Such considerations allowed satisfactory assignments to be proposed for the F KLL spectrum in SiF_4 . The theoretical results showed that three separate groups of dicationic states, having $2p^{-2}$, $2s^{-1}2p^{-1}$, and $2s^{-2}$ character (where s and p refer to the fluorine atomic orbitals), located on one-site (one fluorine atom) dominate the Auger spectrum. The F KVV in XeF_2 Auger spectrum (Fig. 6) can be interpreted in a similar manner. The labelling of the Auger electron peaks in XeF_2 , shown in Fig. 6, follows that predicted for SiF_4 .⁵ A similar interpretation has been given for the structure observed in the Auger electron spectrum of HF .⁴⁴ As would be expected, the Auger electron spectra resulting from a vacancy in the F 1s level of HF , SiF_4 and XeF_2 exhibit similar characteristic peaks.

Based on the theoretical predictions for SiF_4 ,⁵ the intense peak at 647.7 eV in XeF_2 and the weaker peak at 644.9 eV (Fig. 6) are most likely the triplet and singlet states, respectively, associated with the $2p^{-2}$ one-site population. Likewise, the peaks at 629.1 and 620.8 eV may correspond to the theoretically predicted triplet and singlet states, respectively, associated with the $2s^{-1}2p^{-1}$ one-site populations. In SiF_4 , the $2s^{-2}$ dicationic states were associated with structure observed in the kinetic

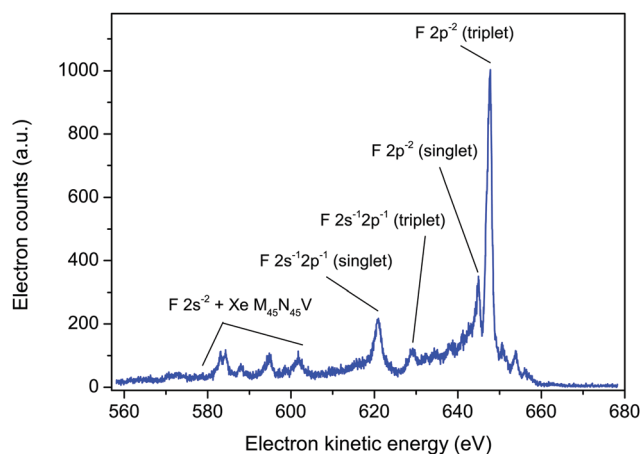


Fig. 6 The fluorine KVV and xenon $\text{M}_4\text{N}_{45}\text{V}$ Auger electron spectrum in XeF_2 , recorded at a photon energy of 800 eV, using perpendicularly polarized radiation. The peaks have been labelled as $\text{F } 2p^{-2}$, $2s^{-1}2p^{-1}$ and $2s^{-2}$ (triplet or singlet) in a manner analogous to those predicted for SiF_4 .⁵

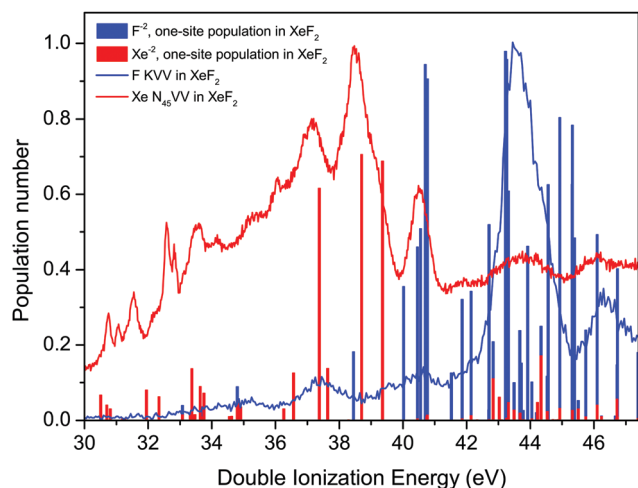


Fig. 7 The experimental Auger electron spectra, plotted as a function of the energy of the final doubly ionized state, due to the decay of a vacancy in the F 1s orbital of XeF₂, and to the decay of a vacancy in the Xe 4d level in XeF₂.⁶ The calculated one-site populations (F²⁻ and Xe²⁻) of the doubly ionized states in XeF₂ have been taken from Buth *et al.*² Most of the sticks showing the populations typically contain contributions from more than one state.

energy range 590–600 eV. However, in XeF₂ the features observed in this energy region appear more complex as a consequence of the additional, overlapping peaks due to the M₄₅N₄₅V transitions.

The F KVV Auger spectrum of XeF₂ has not been studied theoretically but Buth *et al.*² have calculated the one-site and the two-site populations of the double ionization spectra and we make use of these predictions to discuss our experimental results. Fig. 7 shows the F KVV Auger spectrum plotted as a function of the double ionization energy, where we have used the reported value of 691.23 eV for the F 1s in XeF₂ ionization energy.⁸ Also plotted in Fig. 7 is the Xe 4d in XeF₂ Auger spectrum,⁶ where we have used the Xe 4d_{5/2} in XeF₂ ionization energy of 70.38 eV.⁶ It is evident that the final doubly ionized states predominantly populated in the decay of a hole in the Xe 4d shell have energies which are lower than those predominantly populated in the decay of a hole in the F 1s orbital. Such a pattern is consistent with our Mulliken atomic populations (Table S1, ESI†) which predict that the HOMO (5p_u) has a significant Xe 5p lone-pair character.

The calculated one-site populations of the dicationic states of XeF₂ are plotted in Fig. 7. These show that the doubly ionized states at energies below ~40 eV possess a significant Xe²⁻ population, and hence would be expected to play a key role in the decay of a vacancy in the Xe 4d orbital. Note that we adopt the notations, given by Buth *et al.*,² of labelling the one-site states possessing a large dicationic population number on either xenon or a single fluorine as Xe²⁻ or F²⁻, respectively. Doubly ionized states possessing a high F²⁻ one-site population are evident at energies above 40 eV and a dense cluster of such states is predicted between 42 and 46 eV, which coincides with the most prominent peak in the experimental spectrum. The weak structure observed in the F KVV spectrum in the

energy range below ~42 eV may be attributed to states predicted to have a small F²⁻ character.

As discussed previously in our work on the 4d⁻¹ ionization of XeF₂,⁶ the experimental assignments of the XeF₂²⁺ final states with double valence holes in Fig. 7 are not possible at this time. Calculations of the doubly ionized states by Buth *et al.*² show a large number of possible states (18 within 5 eV of threshold), and the experimental double ionization threshold differs from the theoretical value by 2.32 eV. The density of states problem becomes even greater at the higher energies of the F²⁺ states. Finally, although the experimental spectrum appears relatively simple, the broad structure at higher binding energies obscures the fact that multiple states are contributing to each feature.

5.3. Auger electron angular distributions

In discussing the Auger electron angular distributions, we adopt the two-step formulation of the Auger process⁴⁶ in which the initial core excitation or core ionization is assumed to be independent of the subsequent decay. Within this framework, for processes occurring in an atom, the Auger electron angular distribution parameter, β_A , as defined by eqn (1), can be written with β_A factored into two components^{29,30}

$$\beta_A = A_{20}\alpha_2 \quad (4)$$

where A_{20} is the atomic alignment parameter that describes the population of magnetic JM sublevels of the core excited/ionized state and α_2 is the intrinsic anisotropy parameter for the Auger transition. J and M denote the total angular momentum of the intermediate state, and the particular magnetic sublevel, respectively. The degree of alignment can be obtained in terms of the photoexcitation dipole matrix elements, and may vary with the photon energy. The intrinsic anisotropy parameter for a specific Auger transition depends on the total angular momentum of the initial and final states, and on the monopole Coulomb operator matrix elements for the Auger decay, and is independent of the photon energy.^{29,47}

Of particular relevance to the present work are the theoretical studies of photoinduced alignment in core hole states in the rare gases,³⁰ and subsequent work⁴⁸ covering a wider range of elements. These investigations³⁰ show that, in the non-relativistic limit, the alignment $A_{20}(D_{5/2})$ and $A_{20}(D_{3/2})$ resulting from inner shell photoionization of an atomic d orbital, depend upon the ratio

$$R_{3d \rightarrow \epsilon p}^2 / R_{3d \rightarrow \epsilon f}^2 \quad (5)$$

where $R_{3d \rightarrow \epsilon p}$ and $R_{3d \rightarrow \epsilon f}$ are the radial dipole matrix elements for the photoionization of the 3d electron into the p and f continua, respectively.

Auger electron angular distributions have also been studied in several small linear molecules [ref. 49–51, and references therein]. For the molecular case, where the quantization axis is taken as the molecular symmetry axis, the alignment of the state emitting the Auger electron arises from the anisotropic spatial distribution of the molecular ion axes in the lab frame following photoionization, as expressed in eqn (3). The Auger electron angular distribution parameter, β_A , may be factored as^{15,18,52}

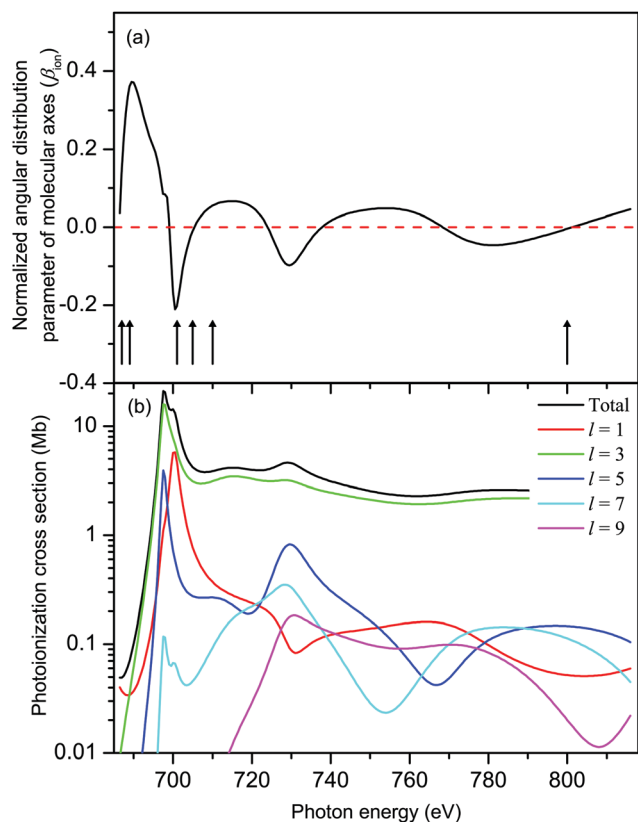


Fig. 8 (a) The calculated, normalized, angular distribution parameter (β_{ion}) characterizing the spatial distribution of molecular axes in XeF_2 in the $\text{Xe } 3\text{d}^{-1}$ ionized state. The vertical bars with arrowheads indicate the photon energies at which measurements of the Auger electron angular distributions associated with the $\text{M}_{45}\text{N}_{45}\text{N}_{45}$ transition were made. The β_{A} parameters derived from these measurements are plotted in Fig. 9. (b) The calculated (CMS- $\text{X}\alpha$) $\text{Xe } 3\text{d}$ photoionization cross section of XeF_2 decomposed into l -wave components.

$$\beta_{\text{A}} = \beta_{\text{ion}} c_{\text{A}} \quad (6)$$

analogous to the atomic case, eqn (4). The intrinsic anisotropy parameter for a particular Auger decay, c_{A} , is independent of the photon energy, and will have a constant value restricted to the range $-\frac{1}{2} \leq c_{\text{A}} \leq 1$.¹⁸

The results from our CMS- $\text{X}\alpha$ calculations (Fig. 8(a)) show that, in the $\text{Xe } 3\text{d}^{-1}$ ionized state in XeF_2 , the molecular axes are not aligned to any significant degree in the photon energy range covered in our measurements of the Auger electron angular distributions. The parameter β_{ion} does, however, display weak oscillations around zero with increasing photon energy. Since c_{A} has a fixed value for a given Auger transition, eqn (6) implies that the predicted oscillations in sign of β_{ion} should be matched by corresponding changes in sign of β_{A} . Experimentally (Fig. 9) there is no evidence of these anticipated variations. Moreover, since the calculated values of β_{ion} , in the photon energy range of interest, are small, the use of eqn (6) to match the experimentally observed β_{A} values requires very large values of c_{A} which lie outside the permitted range. It is concluded that the molecular limit represented by eqn (6) does not provide an adequate description for these XeF_2 Auger transitions.

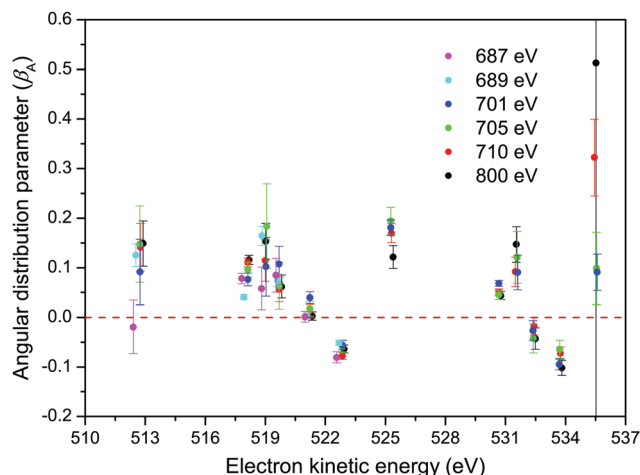


Fig. 9 The angular distribution parameters (β_{A}) for the $\text{M}_{45}\text{N}_{45}\text{N}_{45}$ transitions in XeF_2 derived from the non-resonant Auger spectra recorded with parallel and perpendicularly polarized radiation. The small variation in the electron kinetic energy for a particular doubly ionized final state is due to the fitted value extracted from the individual spectra.

An examination of the eigenvectors obtained by quantum chemistry molecular orbital calculations (see Table S1, population analysis in ESI[†]) confirms, unsurprisingly, that the orbitals we label as 3d are essentially pure atomic Xe 3d in character, sharing no population with the F atoms. Moreover, the 3d eigenvalues (reported as Koopmans ionization energies in Table S1, ESI[†]) show that symmetry splitting by the molecular field is completely negligible (≤ 26 meV). Evidently, these deep-lying Xe 3d levels are insensitive to the molecular field. This suggests that the Auger electron angular anisotropy may be more profitably considered as an atomic-like process.

This behaviour parallels our findings concerning the feasibility of using calculated values of β_{ion} to interpret the non-resonant Auger electron angular distributions resulting from I 3d ionization in CH_3I .¹⁴ We can note that the Auger electron bands due to the xenon $\text{M}_{45}\text{N}_1\text{N}_{45}$, $\text{M}_{45}\text{N}_{23}\text{N}_{45}$ and $\text{M}_{45}\text{N}_{45}\text{N}_{45}$ transitions in XeF_2 closely resemble those due to the analogous transitions in atomic xenon, and hence appear atomic-like. We also note that our experimentally derived β_{A} value for a particular final state associated with the $\text{M}_{45}\text{N}_{45}\text{N}_{45}$ transition in XeF_2 is similar to that for the corresponding state in atomic xenon.⁴⁰ Hence, we continue by assuming the atomic limit model, represented by eqn (4), for the $\text{M}_{45}\text{N}_{45}\text{N}_{45}$ Auger electron angular distributions in XeF_2 .

The calculated total, and l -resolved, Xe 3d in XeF_2 photoionization cross sections are plotted in Fig. 8(b). Apart from a small energy range close to threshold, the $l = 3$ partial wave dominates all the others by 1–2 orders of magnitude. Hence, $R_{3\text{d} \rightarrow \text{ep}}^2 / R_{3\text{d} \rightarrow \text{ef}}^2 \rightarrow 0$. In this limit,³⁰ one obtains the energy-independent alignments:

$$A_{20}(\text{D}_{5/2}) = -0.21 \text{ and } A_{20}(\text{D}_{3/2}) = -0.20 \quad (7)$$

where these values of the alignment parameter refer to linearly polarized radiation with the quantization axis parallel to the

electric vector of the radiation, and would be relevant for the non-resonant Auger emission. Full numerical calculations, spanning from threshold up to electron kinetic energies of at least 200 eV, confirm that these limiting alignments, A_{20} , for $3d_{5/2}$ or $3d_{3/2}$ vacancies in atomic Xe^+ , are achieved for all energies more than 10 eV above the ionization threshold.^{30,48}

For photoexcitation by linearly polarized radiation from an atomic ground state having $J = 0$ to an excited state with $J = 1$, A_{20} has an energy independent value of $-\sqrt{2}$.^{52,53} We assume this value of the alignment parameter is valid for the resonantly excited $\text{Xe } 3d \rightarrow \sigma^*$ spectra.

The β_A parameters for the $M_{45}N_{45}N_{45}$ Auger transition, derived from the fits of our non-resonant spectra, are plotted in Fig. 9. For a particular final doubly ionized state, the experimental β_A -values exhibit little variation as a function of photon energy, as already noted. This observation is consistent with the essentially constant alignment terms, A_{20} , theoretically predicted across the energy range over which the measurements were made. Our experimental values for the α_2 parameters, obtained by averaging the β_A parameters for a specific final state over several photon energies, and using the values for A_{20} given in eqn (7), are listed in Table 1.

Several theoretical methods have been employed to calculate the α_2 parameters for the $M_{45}N_{45}N_{45}$ transitions in atomic xenon.^{54–58} The theoretical values calculated by Chen⁵⁶ and Tulkki *et al.*⁵⁷ are given in Table 1. Apart from the value associated with the 1S_0 state belonging to the $M_{5}N_{45}N_{45}$ transition, the agreement between experiment and theory appears reasonable, although not all of the final states give rise to resolvable features in the experimental spectrum. The discrepancy in the calculated and measured values for the 1S_0 state can probably be attributed to the low intensity of the associated peak in the experimental spectrum.

Although our α_2 parameters for the $M_{45}N_{45}N_{45}$ transition in XeF_2 are in fair accord with those derived by Karvonen *et al.*⁴⁰ from the non-resonant $M_{45}N_{45}N_{45}$ spectrum in atomic xenon, obtained through photoionization, some significant discrepancies are apparent. In fact, better agreement is found with the results obtained by Hiltunen *et al.*³⁹ using proton impact, as seen in Table 1.

The alignment of the intermediate neutral state, following resonant excitation ($\text{Xe } 3d_{5/2,3/2} \rightarrow \sigma^*$) is much higher than that of the non-resonant core-ionized state. This higher degree of alignment leads to the Auger electron angular distributions measured in the resonantly excited Auger spectra (Table 2) being more anisotropic than those associated with non-resonant ionization. Unfortunately, no theoretical predictions for these resonantly excited α_2 parameters are available for comparison with the current experimental results.

6. Summary

Polarization dependent electron spectra of XeF_2 , encompassing the bands due to the xenon $M_{45}N_1N_{45}$, $M_{45}N_{23}N_{45}$, $M_{45}N_{45}N_{45}$ and $M_{45}N_{45}V$ and fluorine KVV Auger transitions have been recorded using linearly polarized synchrotron radiation.

Non-resonantly excited spectra have been measured at several photon energies between 687 and 800 eV, and resonantly excited spectra have been measured at energies coinciding with the $\text{Xe } 3d_{5/2} \rightarrow \sigma^*$ and the overlapping $\text{Xe } 3d_{3/2} \rightarrow \sigma^*$ and $\text{F } 1s \rightarrow \sigma^*$ transitions in XeF_2 . The polarization dependent spectra have allowed the Auger electron angular distributions to be determined. The Auger bands associated with the xenon $M_{45}N_1N_{45}$, $M_{45}N_{23}N_{45}$ and $M_{45}N_{45}N_{45}$ transitions in XeF_2 are similar to those for the analogous transitions in atomic xenon, thereby allowing assignments to be proposed for some of the observed structure. The Auger electron spectrum due to the $M_{45}N_{45}V$ transition differs from that due to the $M_{45}N_{45}O$ transition in atomic xenon, indicating the involvement of delocalized molecular orbitals. Moreover, our spectra indicate that peaks due to the $M_{45}N_{45}V$ transitions in XeF_2 occur in a similar kinetic energy range as those resulting from the F KVV decay in XeF_2 .

The interpretation of the F KVV Auger band in XeF_2 has been discussed in relation to results from a theoretical study of the F KLL spectrum in SiF_4 . The theoretical study showed that the 2 h configurations possessing a high one-site population (both holes in the same fluorine atom) were those associated with the strongest peaks in the Auger spectrum. In a similar manner, doubly ionized states of XeF_2 having a high calculated one-site population have been shown to be those predominantly populated in the decay of a F 1s vacancy in XeF_2 . The experimental spectra show that the ionization energies of the doubly charged states populated in the decay of a hole in the Xe 4d orbital in XeF_2 tend to be lower than those populated in the decay of a hole in the F 1s orbital in XeF_2 . This finding is consistent with the calculated one-site population analysis, which reflects the greater electron density on the F atoms (more screening) and the lower electron density on the Xe atom (less screening). Tentative assignments for some of the peaks observed in the F 1s Auger spectrum in XeF_2 have been suggested, based upon predictions for the analogous structure in the F 1s Auger spectrum of SiF_4 . However, a theoretical study of the Xe 4d and F 1s Auger spectra in XeF_2 is needed to confirm these proposed assignments and to identify several additional features.

The Auger electron intrinsic anisotropy parameters for the Xe $M_{45}N_{45}N_{45}$ transition in XeF_2 have been obtained from the measured angular distributions parameters by using a theoretical, limiting value, of the alignment parameter A_{20} , in an atomic approximation. This limiting value should be valid in the photon energy range over which the measurements were taken, where the Xe 3d photoionization cross section is dominated by transitions into the ϵf continuum channel. The values of the derived α_2 parameters for the Xe $M_{45}N_{45}N_{45}$ transition in XeF_2 are similar to those measured and calculated for the analogous transitions in atomic xenon.

Author contributions

The manuscript was written through contributions of all the authors. All the authors have given approval to the final version of the manuscript.

Conflicts of interest

There are no conflicts to declare.

Acknowledgements

We would like to thank C. Buth and L. Cederbaum for making the digital data from their calculations on the double ionization potentials of XeF₂ available to us, and for helping us with their interpretation. D. M. P. H. is grateful to the Science and Technology Facilities Council (United Kingdom) for financial support. S. T. P. is supported by the U.S. Department of Energy, Office of Science, Office of Basic Energy Sciences, Division of Chemical Sciences, Geosciences, and Biosciences under contract No. DE-AC02-06CH11357. We are grateful to the SOLEIL staff for running the facility and providing beamtime under Project No. 20190046.

References

- 1 C. Buth, R. Santra and L. S. Cederbaum, *J. Chem. Phys.*, 2003, **119**, 7763.
- 2 C. Buth, R. Santra and L. S. Cederbaum, *J. Chem. Phys.*, 2003, **119**, 10575.
- 3 M. Pernpointner and L. S. Cederbaum, *J. Chem. Phys.*, 2005, **122**, 214302.
- 4 F. Tarantelli, A. Sgamellotti and L. S. Cederbaum, *J. Chem. Phys.*, 1991, **94**, 523.
- 5 F. O. Gottfried, L. S. Cederbaum and F. Tarantelli, *Phys. Rev. A: At., Mol., Opt. Phys.*, 1996, **53**, 2118.
- 6 R. Forbes, P. Hockett, I. Powis, J. D. Bozek, D. M. P. Holland and S. T. Pratt, *J. Chem. Phys.*, 2021, **155**, 194301.
- 7 A. Picón, C. S. Lehmann, C. Bostedt, A. Rudenko, A. Marinelli, T. Osipov, D. Rolles, N. Berrah, C. Bomme, M. Bucher, G. Doumy, B. Erk, K. R. Ferguson, T. Gorkhover, P. J. Ho, E. P. Kanter, B. Krässig, J. Krzywinski, A. A. Lutman, A. M. March, D. Moonshiram, D. Ray, L. Young, S. T. Pratt and S. H. Southworth, *Nat. Commun.*, 2016, **7**, 11652.
- 8 T. X. Carroll, R. W. Shaw, T. D. Thomas, C. Kindle and N. Bartlett, *J. Am. Chem. Soc.*, 1974, **96**, 1989.
- 9 G. M. Bancroft, P.-Å. Malmquist, S. Svensson, E. Basilier, U. Gelius and K. Siegbahn, *Inorg. Chem.*, 1978, **17**, 1595.
- 10 J. S. Tse, D. J. Bristow, G. M. Bancroft and G. Schrobilgen, *Inorg. Chem.*, 1979, **18**, 1766.
- 11 J. N. Cutler, G. M. Bancroft, J. D. Bozek, K. H. Tan and G. J. Schrobilgen, *J. Am. Chem. Soc.*, 1991, **113**, 9125.
- 12 S. H. Southworth, R. Wehlitz, A. Picón, C. S. Lehmann, L. Cheng and J. F. Stanton, *J. Chem. Phys.*, 2015, **142**, 224302.
- 13 S. Aksela, G. M. Bancroft, D. J. Bristow, H. Aksela and G. J. Schrobilgen, *J. Chem. Phys.*, 1985, **82**, 4809.
- 14 R. Forbes, A. De Fanis, C. Bomme, D. Rolles, S. T. Pratt, I. Powis, N. A. Besley, S. Nandi, A. R. Milosavljević, C. Nicolas, J. D. Bozek, J. G. Underwood and D. M. P. Holland, *J. Chem. Phys.*, 2018, **149**, 094304.
- 15 R. Forbes, A. De Fanis, D. Rolles, S. T. Pratt, I. Powis, N. A. Besley, A. R. Milosavljević, C. Nicolas, J. D. Bozek and D. M. P. Holland, *J. Phys. B: At., Mol. Opt. Phys.*, 2020, **53**, 155101.
- 16 I. Powis, *Chem. Phys.*, 1995, **201**, 189.
- 17 J. L. Dehmer and D. Dill, *Phys. Rev. A: At., Mol., Opt. Phys.*, 1978, **18**, 164.
- 18 D. Dill, J. R. Swanson, S. Wallace and J. L. Dehmer, *Phys. Rev. Lett.*, 1980, **45**, 1393.
- 19 C. R. Brundle, M. B. Robin and G. R. Jones, *J. Chem. Phys.*, 1970, **52**, 3383.
- 20 C. R. Brundle, G. R. Jones and H. Basch, *J. Chem. Phys.*, 1971, **55**, 1098.
- 21 B. W. Yates, K. H. Tan, G. M. Bancroft, L. L. Coatsworth, J. S. Tse and G. J. Schrobilgen, *J. Chem. Phys.*, 1986, **84**, 3603.
- 22 G. M. Bancroft, S. Aksela, H. Aksela, K. Gürtler, K. H. Tan, B. W. Yates and J. S. Tse, *J. Phys. B: At. Mol. Phys.*, 1987, **20**, 3057.
- 23 H. Basch, J. W. Moskowitz, C. Hollister and D. Hankin, *J. Chem. Phys.*, 1971, **55**, 1922.
- 24 F. J. Comes, R. Haensel, U. Nielsen and W. H. E. Schwarz, *J. Chem. Phys.*, 1973, **58**, 516.
- 25 U. Nielsen and W. H. E. Schwarz, *Chem. Phys.*, 1976, **13**, 195.
- 26 I. Powis, D. M. P. Holland, E. Antonsson, M. Patanen, C. Nicolas, C. Miron, M. Schneider, D. Y. Soshnikov, A. Dreuw and A. B. Trofimov, *J. Chem. Phys.*, 2015, **143**, 144304.
- 27 R. Forbes, A. De Fanis, C. Bomme, D. Rolles, S. T. Pratt, I. Powis, N. A. Besley, M. Simon, S. Nandi, A. R. Milosavljević, C. Nicolas, J. D. Bozek, J. G. Underwood and D. M. P. Holland, *J. Chem. Phys.*, 2018, **149**, 144302.
- 28 S. Flügge, W. Mehlhorn and V. Schmidt, *Phys. Rev. Lett.*, 1972, **29**, 7.
- 29 E. G. Berezko and N. M. Kabachnik, *J. Phys. B: At. Mol. Phys.*, 1977, **10**, 2467.
- 30 E. G. Berezko, N. M. Kabachnik and V. S. Rostovsky, *J. Phys. B: At., Mol. Opt. Phys.*, 1978, **11**, 1749.
- 31 J. Jauhiainen, A. Ausmees, A. Kivimäki, S. J. Osborne, A. Naves de Brito, S. Aksela, S. Svensson and H. Aksela, *J. Electron Spectrosc. Relat. Phenom.*, 1994, **69**, 181.
- 32 M. Jurvansuu, A. Kivimäki and S. Aksela, *Phys. Rev. A: At., Mol., Opt. Phys.*, 2001, **64**, 012502.
- 33 D. Dill and J. L. Dehmer, *J. Chem. Phys.*, 1974, **61**, 692.
- 34 J. W. Davenport, PhD thesis, University of Pennsylvania, 1976.
- 35 S. M. Bellm, J. A. Davies, P. T. Whiteside, J. Guo, I. Powis and K. L. Reid, *J. Chem. Phys.*, 2005, **122**, 224306.
- 36 L. O. Werme, T. Bergmark and K. Siegbahn, *Phys. Scr.*, 1972, **6**, 141.
- 37 S. Aksela, H. Aksela and T. D. Thomas, *Phys. Rev. A: At., Mol., Opt. Phys.*, 1979, **19**, 721.
- 38 H. Pulkkinen, H. Aksela and S. Aksela, *Phys. Rev. A: At., Mol., Opt. Phys.*, 1986, **34**, 1195.
- 39 A. Hiltunen, H. Aksela, S. Aksela, G. Víkor, S. Ricz and B. Sulik, *J. Phys. B: At., Mol. Opt. Phys.*, 1998, **31**, L793.
- 40 J. Karvonen, A. Kivimäki, H. Aksela, S. Aksela, R. Camilloni, L. Avaldi, M. Coreno, M. de Simone and K. C. Prince, *Phys. Rev. A: At., Mol., Opt. Phys.*, 1999, **59**, 315.

- 41 L. Partanen, M. Huttula, S.-M. Huttula, H. Aksela and S. Aksela, *J. Phys. B: At., Mol. Opt. Phys.*, 2006, **39**, 4515.
- 42 M. Ohno, *J. Electron Spectrosc. Relat. Phenom.*, 2003, **130**, 19.
- 43 A. Kivimäki, H. Aksela, J. Jauhiainen, M. Kivilompolo, E. Nömmiste and S. Aksela, *J. Electron Spectrosc. Relat. Phenom.*, 1998, **93**, 89.
- 44 R. W. Shaw and T. D. Thomas, *Phys. Rev. A: At., Mol., Opt. Phys.*, 1975, **11**, 1491.
- 45 H. Siegbahn, L. Asplund and P. Kelfve, *Chem. Phys. Lett.*, 1975, **35**, 330.
- 46 W. Mehlhorn, in *Atomic Inner-Shell Physics*, ed. B. Crasemann, Plenum Press, New York, 1985, p.119.
- 47 H. Aksela, S. Aksela and N. Kabachnik, in *VUV and Soft X-Ray Photoionization*, ed. U. Becker and D. A. Shirley, Plenum Press, New York, 1996, p. 401.
- 48 U. Kleiman and B. Lohmann, *J. Electron Spectrosc. Relat. Phenom.*, 2003, **131–132**, 29.
- 49 O. Hemmers, F. Heiser, J. Eiben, R. Wehlitz and U. Becker, *Phys. Rev. Lett.*, 1993, **71**, 987.
- 50 R. F. Fink, M. N. Piancastelli, A. N. Grum-Grzhimailo and K. Ueda, *J. Chem. Phys.*, 2009, **130**, 014306.
- 51 E. Antonsson, M. Patanen, C. Nicolas, S. Benkoulou, J. J. Neville, V. L. Sukhorukov, J. D. Bozek, P. V. Demekhin and C. Miron, *Phys. Rev. A: At., Mol., Opt. Phys.*, 2015, **92**, 042506.
- 52 N. M. Kabachnik, S. Fritzsche, A. N. Grum-Grzhimailo, M. Meyer and K. Ueda, *Phys. Rep.*, 2007, **451**, 155.
- 53 U. Hergenhahn, N. M. Kabachnik and B. Lohmann, *J. Phys. B: At., Mol. Opt. Phys.*, 1991, **24**, 4759.
- 54 N. M. Kabachnik and I. P. Sazhina, *J. Phys. B: At., Mol. Opt. Phys.*, 1988, **21**, 267.
- 55 N. M. Kabachnik, I. P. Sazhina, I. S. Lee and O. V. Lee, *J. Phys. B: At., Mol. Opt. Phys.*, 1988, **21**, 3695.
- 56 M. H. Chen, *Phys. Rev. A: At., Mol., Opt. Phys.*, 1992, **45**, 1684.
- 57 J. Tulkki, N. M. Kabachnik and H. Aksela, *Phys. Rev. A: At., Mol., Opt. Phys.*, 1993, **48**, 1277.
- 58 A. Yu. Elizarov and I. I. Tupitsyn, *J. Exp. Theor. Phys.*, 2003, **97**, 658.