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**Second Order Jahn-Teller Interactions at Unusually High  
Molecular Orbital Energy  
Separations**

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# 1 **Second Order Jahn-Teller Interactions at Unusually High Molecular Orbital Energy**

## 2 **Separations**

3

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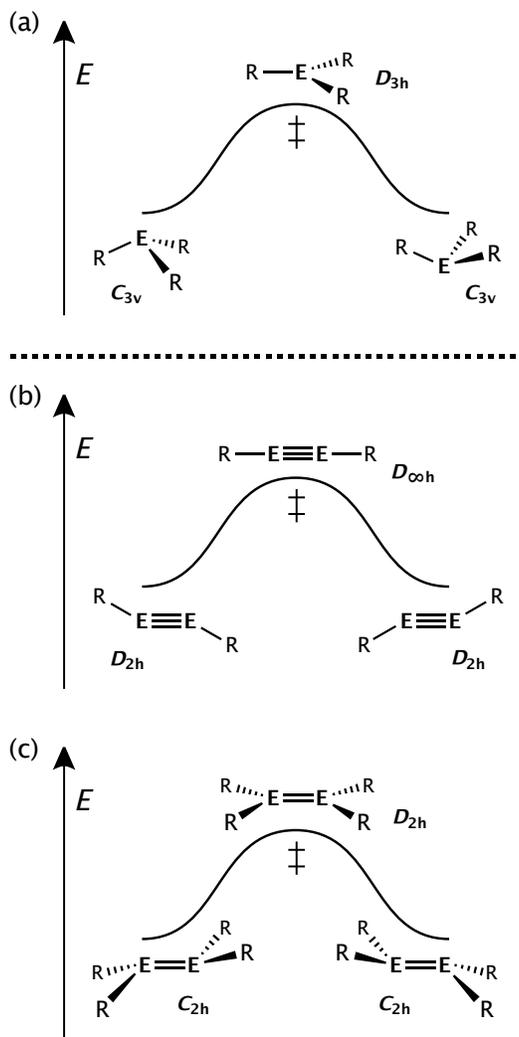
### 10 **Abstract**

11 Second order Jahn-Teller (SOJT) effects arise from interactions between filled and empty molecular  
12 orbitals of like symmetry. These interactions often lead to structural distortions whose extent is inversely  
13 proportional to the energy difference between the interacting orbitals. The main objectives of the work  
14 described here are (1) the calculation (using density functional theory methods) of the energies of the  
15 valence molecular orbitals in the species  $\text{EH}_3$  ( $\text{E} = \text{N}, \text{P}, \text{As}$  or  $\text{Sb}$ ),  $\text{HEEH}$  ( $\text{E} = \text{C}, \text{Si}, \text{Ge}$  or  $\text{Sn}$ ), and  $\text{H}_2\text{EEH}_2$ , ( $\text{E}$   
16  $= \text{C}, \text{Si}, \text{Ge}$  or  $\text{Sn}$ ) and (2) the correlation of these energies with barriers for planarization or linearization.  
17 The calculations suggest an upper limit of about 12 eV energy separation of the interacting levels for SOJT  
18 effects to be significant, which is considerably larger than previously thought and implies that SOJT effects  
19 may be more common than currently realized.

20

## 1 Introduction

2 The process by which pyramidal molecules of the general formula  $ER_3$  ( $E=N-Bi$ ;  $R=H$ , alkyl, aryl, silyl,  
3 halide, etc., or the corresponding anionic group 14 and cationic group 16 species) are transformed into  
4 their inverted conformation (Figure 1) has been studied both experimentally<sup>1</sup> and theoretically<sup>2-8</sup> for many  
5 decades. A detailed understanding of this process is of fundamental interest with respect to bonding in  
6 these molecules and chemical bonding in general. Classically, the inversion process can be regarded as a  
7 (fundamental) vibrational mode that interconverts the two pyramidal conformations via a planar  
8 transition state structure.<sup>9</sup> The energy difference between the pyramidal minimum and the transition  
9 state structure is the so-called inversion barrier. The magnitude of the inversion barrier is related to  
10 several factors.<sup>1</sup> These include the relative electronegativity (EN) of the substituents (R) and central atom  
11 (E), conjugative interactions between E and R, the steric properties of R, the presence of lone pairs on R  
12 and geometrical constraints at E. The influence of electronic factors, e.g., the EN values of the central and  
13 substituent atoms, has been rationalized in terms of the Bent<sup>11</sup>/Walsh<sup>12,13</sup> rules and differences in  
14 hybridization tendencies between the lightest and heavier elements within each group (Kutzelnigg).<sup>14</sup>  
15 Conjugation can lower the relative energy of the planar transition state structure and geometrical  
16 confinement can increase the relative strain of the minimum or transition state structure.<sup>15</sup>



1

2 **Figure 1.** Inversion/linearization processes examined herein.

3 A less prominent view of the inversion process is based on second order Jahn-Teller (SOJT)

4 effects<sup>16-23</sup> involving the extent of HOMO-LUMO interactions and their dependence on symmetries of the

5 frontier orbitals and the electronegativities of E and R. These effects are ubiquitous in chemistry, and in

6 molecular species they generally involve a symmetry governed interaction between a filled and an empty

7 molecular orbital that occurs during a molecular vibration. This interaction can produce a change in the

8 geometry that usually far exceeds those produced by the better known but arguably less important (for

9 chemistry) first order Jahn-Teller effect. The extent of the SOJT distortion is inversely proportional to the

1 energy difference between the interacting levels. It was pointed out recently that there was no  
2 information available on the energy separations between the levels whose interactions produced the SOJT  
3 distortion,<sup>26</sup> and it is generally thought that for SOJT effects to be observable there is an upper limit of ca.  
4 4 eV in the energy separation of the levels.<sup>17,18</sup>

5 The relative energies of the frontier orbitals for the structures in Figure 1 are reflected in the  
6 magnitude of the inversion barrier, yet relatively few calculations<sup>8</sup> that provide energies of the  
7 symmetrically relevant interacting frontier orbitals for a series of ER<sub>3</sub> species with varying E or R groups  
8 have been published to date.<sup>24</sup> Inversion or linearization processes for pyramidal or bent geometries  
9 found for the heavier main group element analogues of alkenes and alkynes are closely related to the ER<sub>3</sub>  
10 inversion process (Figure 1b,c),<sup>19-21</sup> and theoretical studies that relate these processes to SOJT effects are  
11 also scarce.<sup>25,26</sup>

12 Here we describe the results of calculations on the trigonal pyramidal hydrides of group 15 atoms  
13 with emphasis on the relative energies of frontier orbitals and their implications for inversion barriers. We  
14 also describe parallel calculations on the group 14 heavier element ethene and ethyne analogues, which  
15 possess trans-pyramidalized (folded) or trans-bent structures in contrast to their planar or linear carbon  
16 analogues (Figure 1b,c).<sup>27</sup> We emphasize the symmetries and energies of the frontier orbitals during the  
17 inversion or linearization process and show that SOJT interactions occur even when HOMO-LUMO energy  
18 separations are as high as 12 eV, much higher than the previous estimate of 4 eV.<sup>28</sup>

19

## 20 **Methods**

21 All computations were performed using the *Gaussian09* software suite.<sup>29</sup> Calculations were  
22 performed with the  $\omega$ B97X-D functional and the Def2-TZVPP basis set.<sup>30</sup> All calculations utilized an

1 ultrafine quadrature grid (implemented using the “int=ultrafine” command). Some calculations were  
2 carried out at the coupled cluster singles and doubles with perturbative triples (CCSD(T)) level,<sup>31</sup> with the  
3 aug-cc-pVQZ basis set.<sup>32</sup> The results of these calculations, which confirmed, for NH<sub>3</sub> and PH<sub>3</sub>, that the  
4 ωB97X-D/Def2-TZVPP model chemistry satisfactorily predicted orbital energies, are described in the  
5 Supporting Information. *CylView*<sup>33</sup> and *Avogadro*<sup>34</sup> were used to generate ball-and-stick and orbital  
6 (isovalue of 0.02000) images, respectively. To generate molecular orbital diagrams, a custom python script  
7 utilizing *matplotlib* was employed.<sup>35</sup>

8 For the pnictogens, the calculated pyramidal geometries for all equilibrium structures feature C<sub>3v</sub>  
9 point group symmetry. In our case, where R=H, all pnictogen transition state structures have D<sub>3h</sub>  
10 symmetry. Nevertheless, recent calculations by Truhlar and coworkers have shown that when the  
11 substituent is a halogen and E is a heavier pnictogen (e.g., P-Bi), the transition state structures can be T-  
12 shaped with C<sub>2v</sub> symmetry.<sup>9</sup> In addition, open-shell species can also participate in inversion.<sup>9</sup> However, we  
13 focus here on the closed shell C<sub>3v</sub> ⇌ D<sub>3h</sub> ⇌ C<sub>3v</sub> process and the related processes in HEEH and H<sub>2</sub>EEH<sub>2</sub>.

14

## 15 Results and Discussion

### 16 ER<sub>3</sub> Systems

17 A symmetry-directed approach to relate inversion and trans-bending barriers with the relative  
18 energies of the interacting molecular orbitals was used. First, we tested our computational approach by  
19 calculating molecular geometries, energies of valence orbitals, and inversion barriers for simple EH<sub>3</sub> (E =  
20 N, P, As and Sb) molecules.<sup>36</sup> The trigonal pyramidal minimum geometries have C<sub>3v</sub> symmetry and the  
21 planar transition state structures have D<sub>3h</sub> symmetry. In the D<sub>3h</sub> structures, the HOMO-2, HOMO-1, HOMO,  
22 LUMO, and LUMO+1 orbitals have the symmetries a<sub>1</sub>' e', a<sub>2</sub>'', a<sub>1</sub>', and e', respectively (Table 1 and Figure

1 2). Shown in Figure 2 is a representation of each molecular orbital for  $\text{EH}_3$  corresponding to the above-  
2 named MO symmetries. This sequence, which is in order of increasing energy, is observed in the four  $\text{EH}_3$   
3 molecules of the group. The energy spread of the highest and lowest energy levels reflects the E–H bond  
4 strengths, and ranges from 30.62 eV in  $\text{NH}_3$  to 20.96 eV in  $\text{SbH}_3$ . The computed HOMO-LUMO energy  
5 separation for planar,  $D_{3h}$   $\text{NH}_3$  (11.85 eV) is higher by >2 eV than those of its heavier congeners  $\text{PH}_3$  (9.73  
6 eV),  $\text{AsH}_3$  (8.83 eV) and  $\text{SbH}_3$  (7.51 eV). Thus, when the undistorted, planar  $D_{3h}$   $\text{NH}_3$  molecule undergoes a  
7 pyramidalization vibration of  $a_2''$  symmetry, the triple direct product of the HOMO x vibration x LUMO  
8 (i.e.,  $a_2'' \times a_2'' \times a_1' = a_1'$ ) is the totally symmetric irreducible representation of the point group  $D_{3h}$ . In other  
9 words, the mixing becomes symmetry-allowed upon the vibrational distortion from  $D_{3h}$  to  $C_{3v}$  and, in this  
10 case, it produces a stabilization of the HOMO (lone pair) that correlates to the inversion barrier. Due to  
11 the larger HOMO-LUMO energy difference in  $\text{NH}_3$  in comparison to those of its heavier congeners, the  
12 interaction between these orbitals is relatively small and produces a low inversion barrier near 0.18 eV  
13 (4.2 kcal/mol). In other words, since the energy gap in planarized ammonia compared to its heavier  
14 congeners is large, there is less mixing and a smaller interaction between the two levels. In contrast, for  
15 the heavier congeners with smaller HOMO-LUMO energy differences, the HOMO-LUMO interaction is  
16 greater and affords much higher inversion barriers in the range 1.43 to 1.98 eV (33.0 to 45.7 kcal/mol).  
17 The calculated HOMO-LUMO energy differences mirror, but are slightly larger than, those reported earlier  
18 by Schwertfeger, Laakkonen and Pyykkö; in addition, we calculate substantially larger energy differences  
19 (ca. 1.2-2.1 eV) between the unoccupied LUMO and LUMO+1 levels (Table S1) than those (ca. 0.2 eV)  
20 calculated earlier.<sup>8</sup> Figure 3 shows that computed inversion barriers and orbital energy gaps correlate ( $R^2$   
21 = 0.95).

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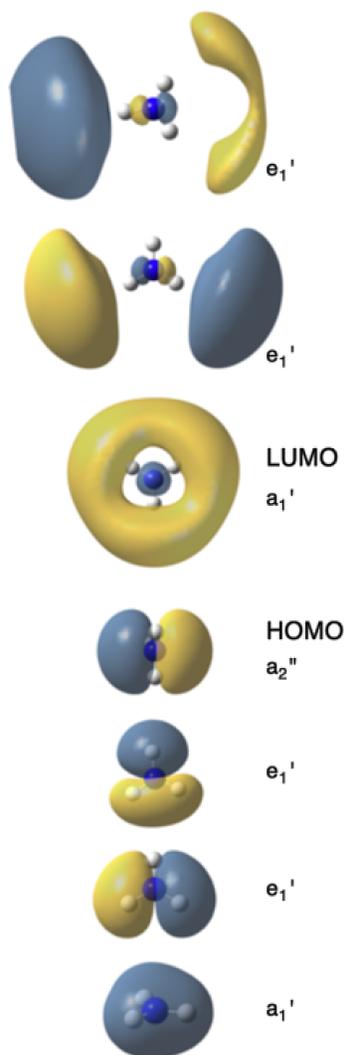
- 1 **Table 1.** The predicted energy differences between the interacting  $a_2''$  HOMO and the  $a_1'$  LUMO in the  
2  $D_{3h}$  planar, geometry  $\text{EH}_3$  systems and predicted barriers for inversion.

<i>Molecule</i>	$\Delta E$ (eV)	<i>Inversion</i> <i>Barrier</i> (kcal/mol)
<b><i>NH<sub>3</sub></i></b>	11.85	4.2
<b><i>PH<sub>3</sub></i></b>	9.73	33.0
<b><i>AsH<sub>3</sub></i></b>	8.83	39.0
<b><i>SbH<sub>3</sub></i></b>	7.51	45.7

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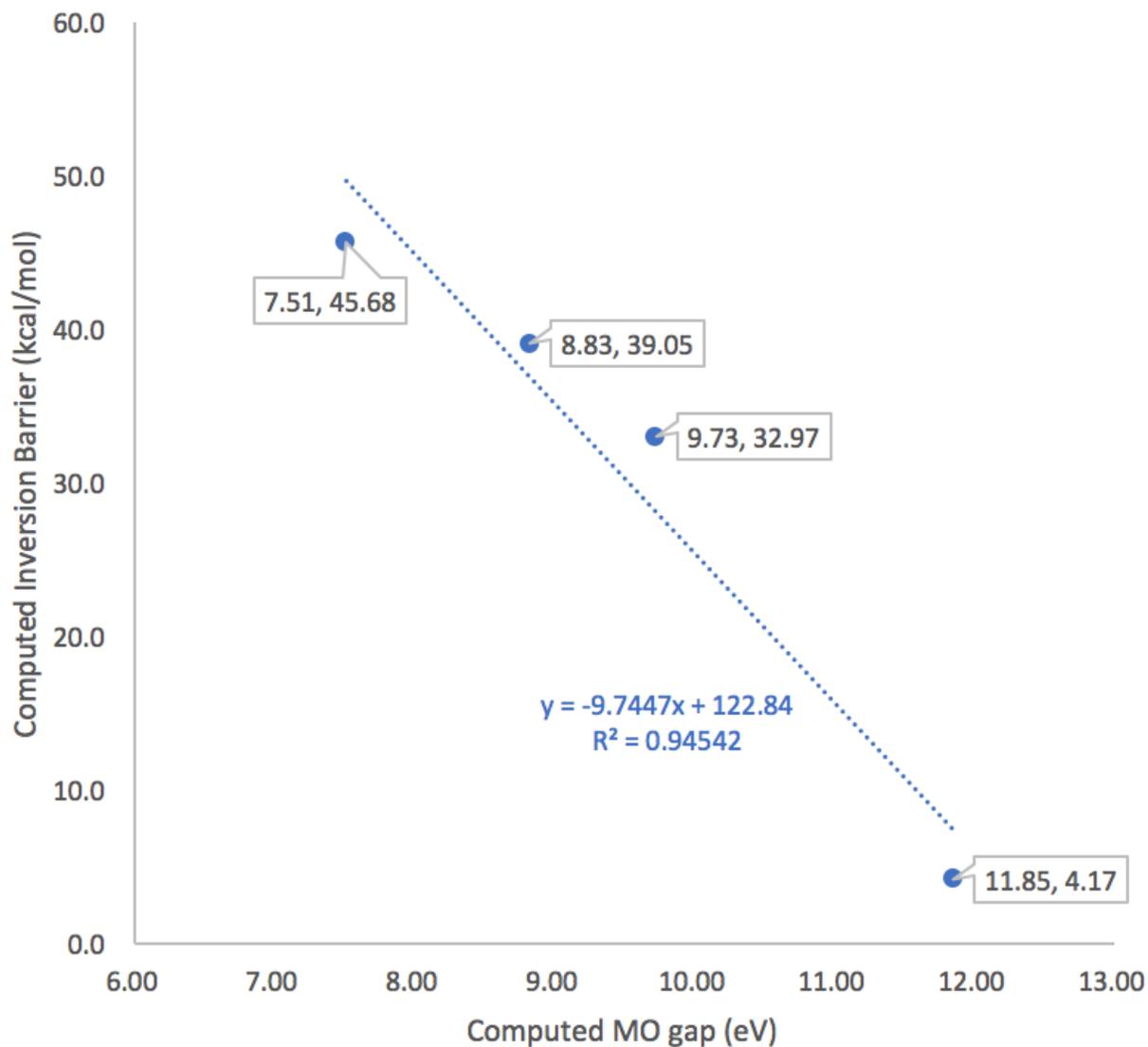


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2 **Figure 2.** Representations of the computed molecular orbitals of ammonia in  $D_{3h}$  symmetry. Note that  
3 different molecular orientations are used for different orbitals to facilitate comprehension of orbital  
4 shapes.

5

1



2

3 **Figure 3.** Correlation between computed inversion barrier and computed energy gap between the  
 4 interacting occupied and unoccupied MOs of  $a_1$  symmetry in  $\text{EH}_3$  molecules.

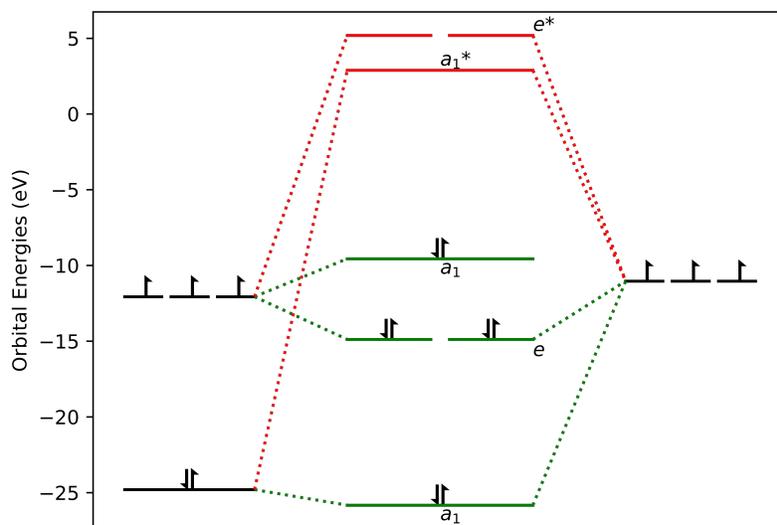
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6 In pyramidal ammonia with  $C_{3v}$  symmetry, the valence orbitals have the symmetries  $a_1, e, a_1, a_1,$   
 7 and  $e$ , in order of increasing energy. This order parallels the sequence given above for  $\text{NH}_3$  in  $D_{3h}$  symmetry  
 8 (Figures 2 and 4). The energy range of the highest and lowest molecular orbital energy levels is 31.02 eV

1 which is just 0.40 eV greater than that of the  $D_{3h}$  structure. In contrast, for the heavier  $\text{PH}_3$ ,  $\text{AsH}_3$ , and  $\text{SbH}_3$   
2 analogues in  $C_{3v}$  symmetry, the ordering of the two highest energy levels changes so that the sequence  
3 becomes  $a_1, e, a_1, e, a_1$ . However, the energy differences between the highest  $e$  and  $a_1$  levels are relatively  
4 low: 0.34 eV ( $\text{PH}_3$ ), 0.22 eV ( $\text{AsH}_3$ ) and 1.02 eV ( $\text{SbH}_3$ ), compared to 2.3 eV in  $\text{NH}_3$ . Nonetheless, we regard  
5 the reversal of the order of the two highest energy orbitals as a consequence of the SOJT mixing of the  
6 HOMO and LUMO ( $a_2''$  and  $a_1'$  in  $D_{3h}$  symmetry), which produces two  $a_1$  levels in  $C_{3v}$  symmetry one of  
7 which (the HOMO) is stabilized. These changes in energy are sufficiently small in  $\text{NH}_3$  that the ordering of  
8 the orbitals remains unchanged. In contrast, in the heavier  $\text{PH}_3$ ,  $\text{AsH}_3$  and  $\text{SbH}_3$ , the SOJT interactions are  
9 significantly larger and the energy of the LUMO is increased sufficiently to become the LUMO+2 level for  
10 the  $C_{3v}$  geometry. Other considerations such as changing orbital overlap and bond strengths can affect the  
11 ordering, but the increased SOJT interactions in the heavier element derivatives offer a simple and  
12 straightforward model for the changed ordering.

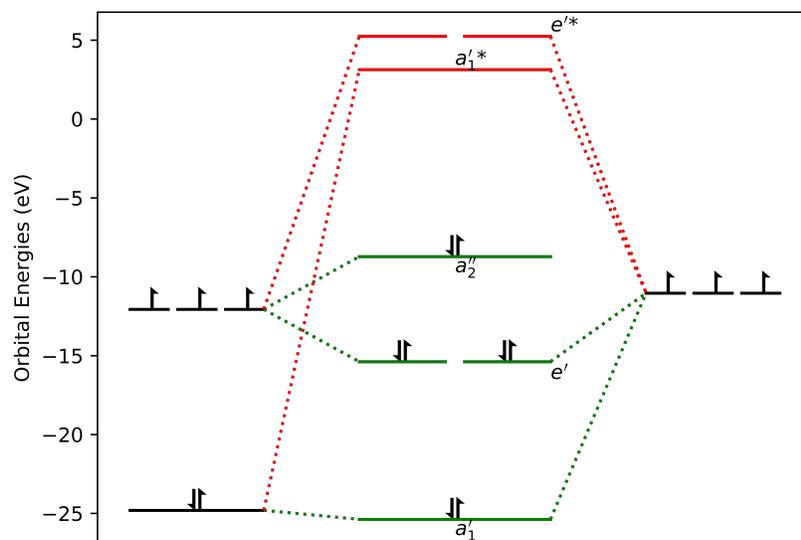
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1



2

(a)



3

(b)

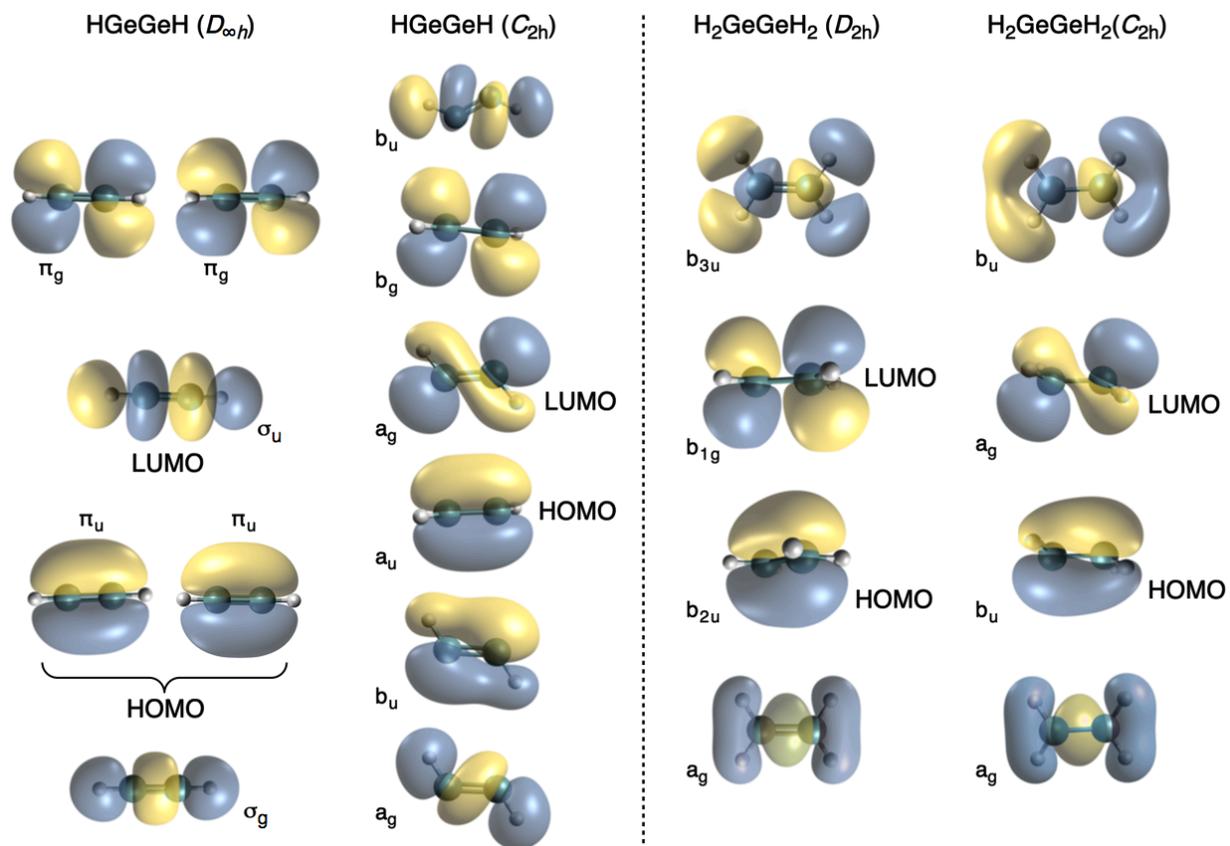
4 **Figure 4.** Molecular orbital diagram for  $\text{NH}_3$  in (a)  $C_{3v}$  symmetry and (b)  $D_{3h}$  symmetry. The ordering of the  
 5 upper two energy levels is reversed for  $E = P, A_s$  and  $S_b$ .

6

1 *REER and R<sub>2</sub>EER<sub>2</sub> Molecules*

2           The bonding in the undistorted linear and planar HEEH alkyne and H<sub>2</sub>EEH<sub>2</sub> alkene and their heavier  
3 congeners (E = C, Si, Ge or Sn), respectively was also investigated.<sup>37</sup> For the silicon, germanium and tin  
4 alkyne analogues, the lowest energy (non-hydrogen bridged) structure was found to have a planar, trans-  
5 bent geometry with C<sub>2h</sub> symmetry (Figure 1b; Table 2). In the case of carbon, the lowest energy structure  
6 was calculated to be linear (D<sub>∞h</sub>), as expected. For the alkene analogues, carbon has the expected planar  
7 geometry with D<sub>2h</sub> symmetry, whereas the trans-bent geometries with C<sub>2h</sub> symmetry (Figure 1c) were  
8 predicted for the heavier element structures. Shown in Figure 5 is a representation of each molecular  
9 orbital for representative heavier element alkyne and alkene analogues corresponding to their respective  
10 MOs.

11



1  
 2 **Figure 5.** Orbital illustrations. Left to right: HGeGeH in  $D_{\infty h}$  symmetry, HGeGeH in  $C_{2h}$  symmetry,  $H_2GeGeH_2$   
 3 in  $D_{2h}$  symmetry, and  $H_2GeGeH_2$  in  $C_{2h}$  symmetry, with illustrations of the calculated orbitals. Note that  
 4 different molecular orientations are used for different orbitals to facilitate comprehension of orbital  
 5 shapes.

6

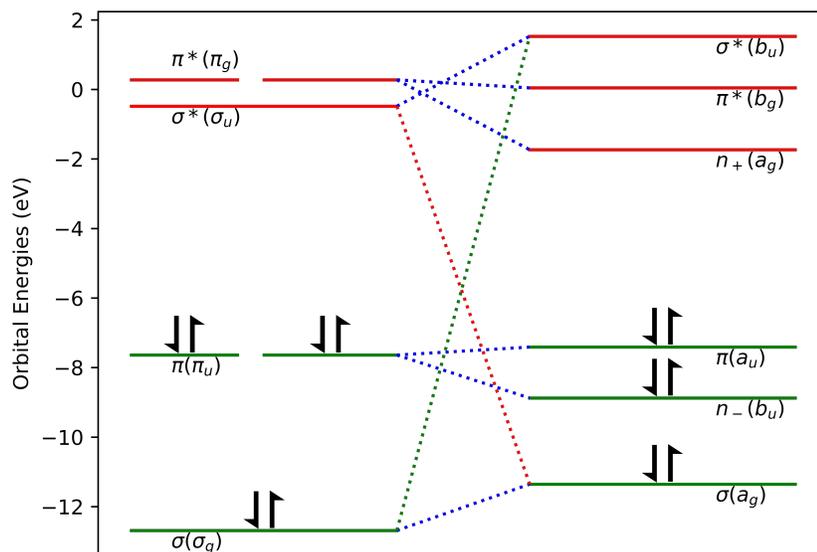
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1 **Table 2.** The predicted energy differences between the interacting  $\pi_u$  HOMO and the  $\sigma_u$  LUMO, which lead  
 2 to a SOJT interaction upon undergoing a trans-bending vibration of  $C_{2h}$  symmetry, and predicted barriers  
 3 for inversion.

<b>Molecule</b>	<b><math>\Delta E</math> (eV)</b>	<b>Linearization Barrier (kcal/mol)</b>
<b>HCCH</b>	14.13	0.0
<b>HSiSiH</b>	7.71	20.2
<b>HGeGeH</b>	7.16	29.5
<b>HSnSnH</b>	5.82	48.9

4

5 In the case of ethyne, HCCH, we start with the  $D_{\infty h}$  point group (linear geometry) where we  
 6 defined the HOMO-1, HOMO, LUMO and LUMO+1 as  $\sigma_g$ ,  $\pi_u$  (doubly degenerate),  $\pi_g$  (doubly degenerate)  
 7 and  $\sigma_u$  levels. However, in the heavier Si, Ge and Sn alkyne analogues,<sup>37</sup> in  $D_{\infty h}$  linear geometry, we find  
 8 the sequence of orbital energies has changed to  $\sigma_g$ ,  $\pi_u$  (doubly degenerate),  $\sigma_u$  and  $\pi_g$  (doubly degenerate)  
 9 (Figure 6). This is due to the large increase in electronic repulsion generated by the heavier cores,  
 10 especially for the valence s electrons.<sup>38</sup> The  $\pi_u$  and  $\pi_g$  orbitals are doubly degenerate in the linear  
 11 geometry, whereas, in the planar trans-bent geometry of  $C_{2h}$  symmetry, the degeneracy of the  $\pi$ -orbitals  
 12 is lifted and one of the original  $\pi$ -orbitals (the in-plane  $\pi$ -orbital) becomes an orbital of  $b_u$  symmetry (the  
 13 out of plane  $\pi$ -orbital has  $a_u$  symmetry) in the trans-bent species with non-bonding character as a result  
 14 of its interaction with a  $\sigma^*$  orbital. In the  $C_{2h}$  structure the energy levels corresponding to HOMO-2,  
 15 HOMO-1, HOMO, LUMO, LUMO+1, and LUMO+2 levels have the symmetries  $a_g$ ,  $b_u$ ,  $a_u$ ,  $a_g$ ,  $b_g$ ,  $b_u$   
 16 respectively. This order holds for all studied alkyne analogues in  $C_{2h}$  trans-bent geometry.

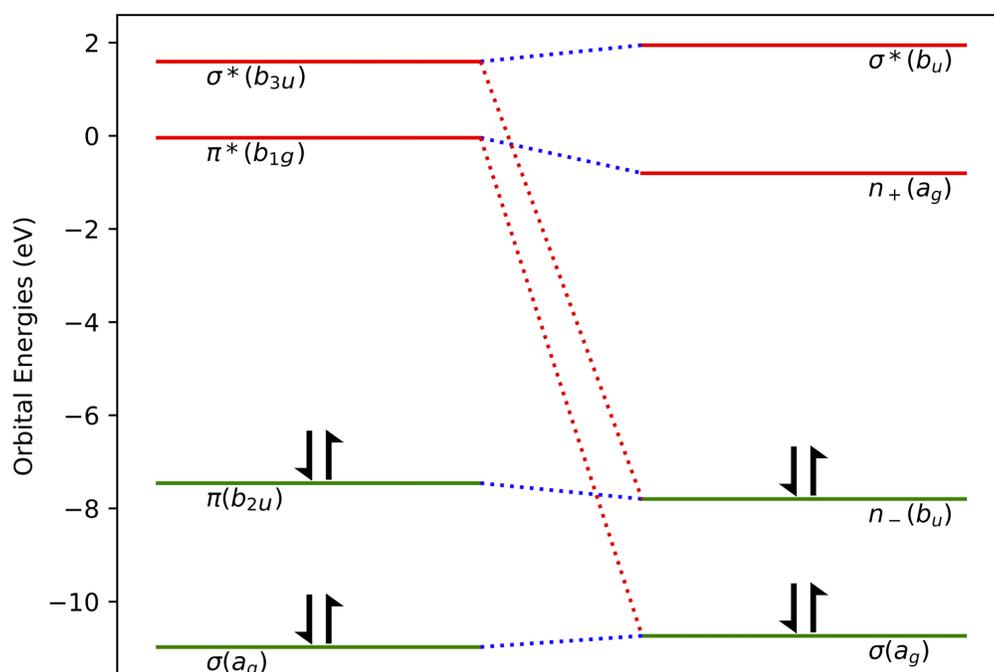


1  
 2 **Figure 6.** Correlation of the interacting valence orbitals of HGeGeH in linear  $D_{\infty h}$  (left) and trans-bent  $C_{2h}$   
 3 (right) geometries.

4  
 5 For linear HCCH in  $D_{\infty h}$  symmetry, LUMO and LUMO+1 are the antibonding  $\pi_g^*$  and  $\sigma_u^*$  orbitals.  
 6 In the corresponding heavier element Si, Ge and Sn species in  $D_{\infty h}$  symmetry, this ordering is changed (see  
 7 above) and the LUMO becomes the  $\sigma_u^*$  orbital and the LUMO+1 becomes the  $\pi_g^*$  orbital. Upon trans-  
 8 bending, the  $\sigma_u^*$  orbital becomes  $b_u$  in  $C_{2h}$  symmetry and one of the original  $\pi$ -bonding orbitals also  
 9 becomes  $b_u$  in symmetry. Thus, as the energy separations between the orbitals decrease in the heavier  
 10 element species, the interaction between the two  $b_u$  orbitals increases and non-bonded electron density  
 11 is generated at the group 14 atom, which favors the trans-bent geometry due to electron/electron  
 12 repulsion.

13 We note that for alkene analogues, the ordering of the orbitals upon changing from E=C to heavier  
 14 elements is far less important than in the case of alkyne analogues (Figure 7). While we see changes in  
 15 relative MO energies as we move through the alkene analogues, the changes mainly involve orbitals that

1 are lower in energy than the HOMO, in contrast to the alkyne analogues where all energy differences  
 2 mentioned involve orbitals that are very close to those involved in the interactions with which we are  
 3 concerned here. While the HOMO is the interacting  $b_{3u}$  orbital for all the alkene analogues, we can see  
 4 that the interacting unoccupied MO  $b_{1u}$ , is not the LUMO or LUMO+1 until we get to E=Ge and Sn. For  
 5 E=Si, it is, in terms of energy, at the LUMO+2 level and for E=C, it is all the way at the LUMO+3 level. This  
 6 also is reflected in the fact that while there is a trans-pyramidalization in the E=Si molecule, it is much  
 7 smaller than the corresponding bends for E=Ge and Sn (Table 3).



8

9 **Figure 7.** Comparison and correlation of  $D_{2h}$  (left) and  $C_{2h}$  (right) symmetries for  $H_2GeGeH_2$ .

10

11

1 **Table 3.** The interacting molecular orbitals ( $b_{2u}$  HOMOs and  $b_{3u}$  LUMOs) and their energies in the alkene  
2 analogues. Inversion barriers are in  $\text{kcal mol}^{-1}$ .

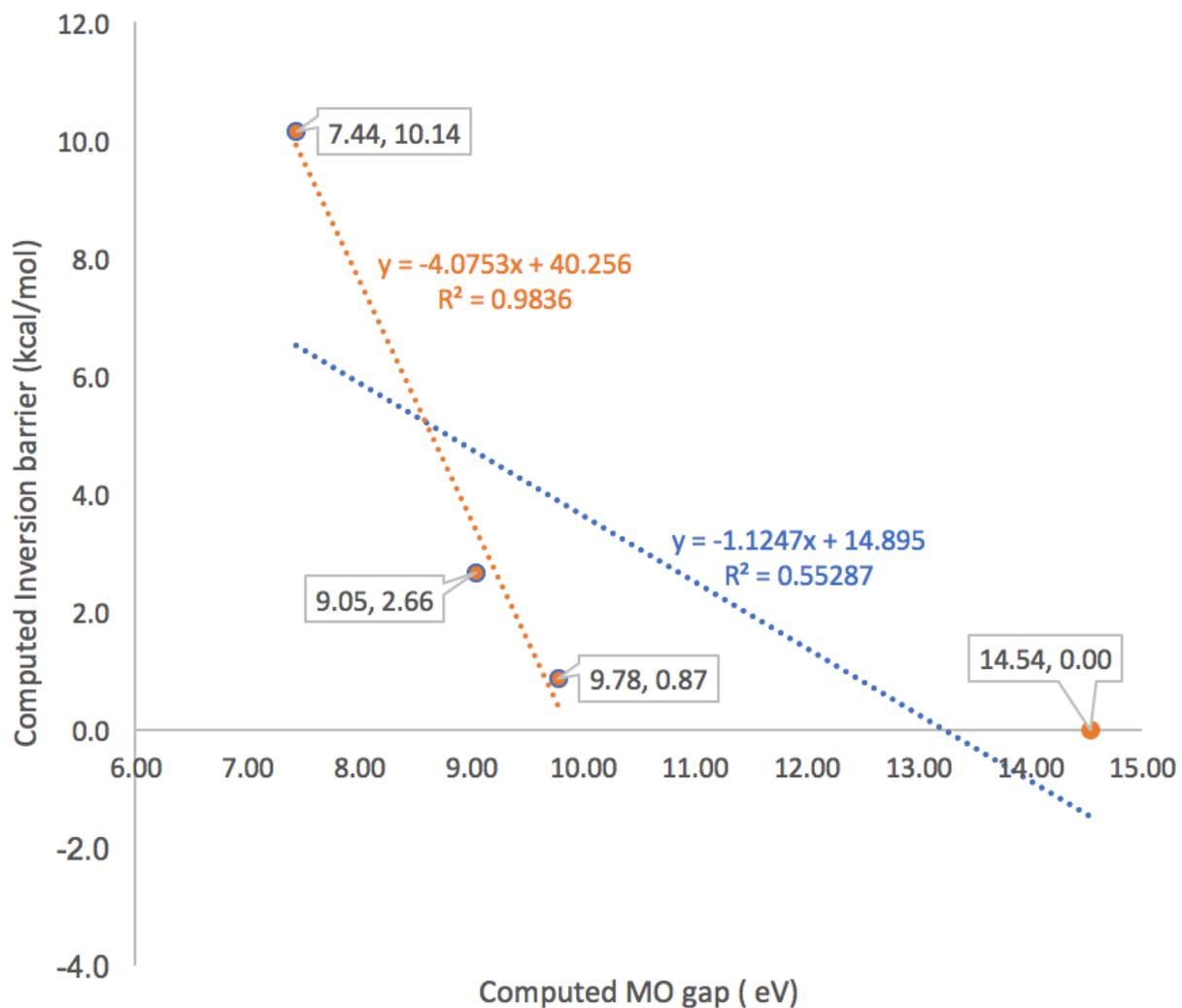
<i>Molecule</i>	<i><math>\Delta E</math></i> <i>(eV)</i>	<i>Planarization</i> <i>Barrier</i> <i>(kcal/mol)</i>
<i><math>H_2CCH_2</math></i>	14.54	0.0
<i><math>H_2SiSiH_2</math></i>	9.78	0.9
<i><math>H_2GeGeH_2</math></i>	9.05	2.7
<i><math>H_2SnSnH_2</math></i>	7.44	10.1

3

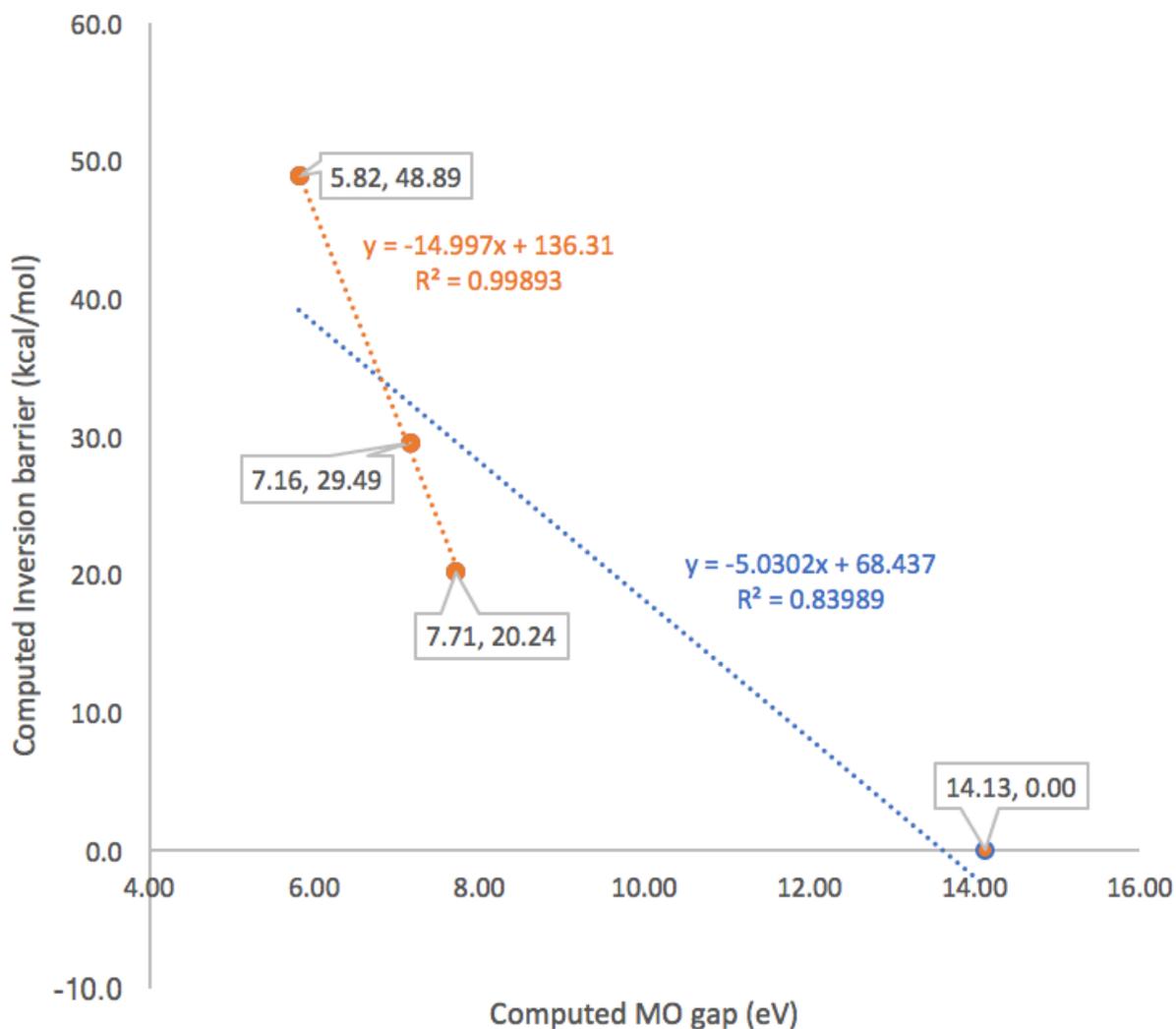
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5 Again, correlations between orbital energy gap and inversion barriers are observed for both  
6 alkene and alkyne analogues (Figures 8 and 9).

7



- 1
- 2 **Figure 8.** Plot of the planarization barriers against the energy separation of the two interacting (occupied
- 3 and unoccupied) MOs for ethylene and its heavier element congeners. Using a best fit line, an  $R^2$  value of
- 4 0.5529 is calculated. When carbon is excluded, an  $R^2$  value of 0.9936 is calculated.



1  
 2 **Figure 9.** Plot of the linearization barriers in HEEH (E = C, Si, Ge or Sn) against the two interacting occupied  
 3 and unoccupied MOs. Using a best-fit line, an  $R^2$  value of 0.8399 is calculated. But when carbon is omitted,  
 4 an  $R^2$  value of 0.9989 is computed.

5  
 6 **Conclusions**

7 Here we examined the relationship between SOJT interactions and planarization/linearization  
 8 barriers. For the  $\text{EH}_3$ , HEEH, and  $\text{H}_2\text{EEH}_2$  species that we examined, SOJT interactions are important even  
 9 when the energy difference between HOMO and the relevant interacting unoccupied orbital is as large as

1 ca. 12 eV, a much larger value than expected. For carbon-containing compounds HCCH, and H<sub>2</sub>CCH<sub>2</sub>, the  
2 predicted occupied/unoccupied orbital energy differences are >14 eV and no SOJT effects are apparent.  
3 In contrast, for the heavier element silicon, germanium, and tin analogues the energy differences tend to  
4 range from 5-10 eV, and these energies are sufficiently low that orbital interactions lead to the  
5 experimentally observed geometrical distortions, which, in some cases, lead to dissociation of the EE  
6 bond.<sup>39</sup> The unexpectedly high magnitude of the energy separations found here implies that SOJT effects  
7 of this type could be much more common than is currently realized.

8

## 9 **Acknowledgements**

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11 (CHE-1565501) for financial support. HBW gratefully acknowledges Sarah Cohen for helpful feedback,  
12 assistance, and advice while performing this work.

13

1 **References**

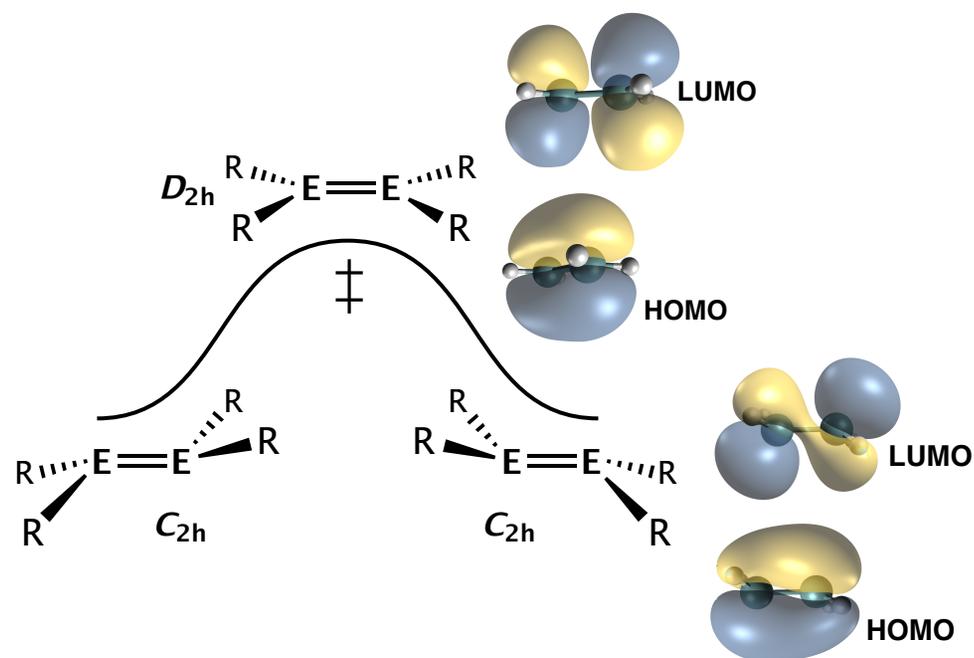
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Calculations suggest an upper limit of approximately 12 eV energy separation of the interacting levels for second order Jahn-Teller effects to be significant, which is considerably larger than previously thought.