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Pulsed-flow growth of polar, semipolar and nonpolar Al-GaN

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The impacts of pulsed-flow growth on aluminium incorporation in polar (0001), semipolar (1013) and (1122), as well as nonpolar (1010) AlGaN layers have been investigated. The layers were grown simultaneously on differently oriented AlN/sapphire templates by metal-organic vapour phase epitaxy. The AlN mole fraction ($0 < x_{AlN} \le 0.85$) of the layers was varied by simply changing the supply time of aluminium precursor while keeping nitrogen and gallium precursors constant. Phase separation has been observed for the (0001) and (1122) layers by x-ray diffraction, which is attributed to their different surface reconstructions during growth. In contrast, no phase separation has been observed for the (1013) layers, attributed to their stable surfaces during growth. The AlN mole fraction of the differently oriented layers generally follows the order: $(1122)<(0001)<(1013)\leq(1010)$, attributed to their different surface measurements, highly carbon-incorporation has quantitatively been found in all the layers.

1 Introduction

Semipolar and nonpolar oriented III-nitrides have attracted much 2 attention due to reduced built-in fields¹, which significantly in-3 crease the wavefunction overlap between electrons and holes re-4 sulting in an increased radiative recombination rate in quantum-5 well structures compared to polar (0001) c-plane counterparts. 6 High-efficiency semipolar and nonpolar InGaN-based visible emit-7 ters have successfully been demonstrated^{2–4}. Additionally, it has 8 been found that due to a very short carrier radiative lifetime, 9 semipolar and nonpolar InGaN-based emitters can be a good can-10 didate for high-speed visible-light-communication systems ^{5,6}. For 11 semipolar and nonpolar AlGaN-based ultraviolet (UV) emitters, 12 together with an increase in the radiative recombination rate, 13 the dominant transverse magnetic polarization ($E \parallel [0001]$) of the 14 light emission over the whole range of composition also can in-15 crease light extraction efficiency^{7–9}. These can increase the per-16 formance of the UV emitters. 17

In contrast to InGaN, there are very limited studies about semipolar and nonpolar Al-containing alloys grown on UVtransparent AlN/sapphire templates, e.g., AlGaN¹⁰⁻¹⁶ and AlInN^{10,17-19}. Additionally, the lack of available non-*c*-plane AlN bulk substrates hinders studies of Al(Ga,In)N on bulk substrates^{9,20-22}. Since 2010, there is only one report about an AlGaN-based UV LED grown on a (1122) AlN/sapphire tem-24 plate operating at 307 nm²³. This is mainly due to very lim-25 ited studies on growth of non-c-plane AlN templates on sap-26 phire $^{10,16,24-27}$. So far, nonpolar (1120) *a*-plane AlGaN 10,15 and 27 (10 $\overline{10}$) *m*-plane AlGaN^{14,15}, as well as semipolar (10 $\overline{13}$)¹⁶ and 28 (1122) AlGaN^{11,12,15,23} have been grown on different planar sap-29 phire substrates by metal-organic vapour phase epitaxy (MOVPE). 30 For these studies, a continuous growth has commonly been em-31 ployed, i.e., Al-, Ga- and N-sources were simultaneously intro-32 duced during AlGaN epitaxy. 33

For polar GaN²⁸ and AlN^{29,30} templates grown on sapphire 34 using MOVPE, to improve the crystalline quality and surface mor-35 phology of the templates, pulsed-flow MOVPE-growth has been 36 employed. (This growth method will be later abbreviated as 37 pulsed-growth.) Such AlN templates have been used to pre-38 pare AlGaN-based UV LEDs operating at 231-261 nm²⁹. Pulsed-39 growth has also been employed for polar quaternary Al(In,Ga)N 40 layers ³¹ and ternary AlGaN quantum wells ³². For MOVPE-grown 41 AlGaN, using pulsed-growth also allows to suppress the gas phase 42 pre-reaction between the group III and group V precursors^{28,31}. 43 Consequently, this suppression can strongly enhance the alu-44 minium incorporation in AlGaN^{12,31}. So far, there is no study 45 about non-c-plane AlGaN epitaxy using pulsed-growth. 46

Recently, it has been found that aluminium incorporation is comparable in (0001), ($10\overline{1}3$), ($11\overline{2}2$), ($11\overline{2}0$) and ($10\overline{1}0$) Al-GaN co-loaded layers grown using continuously growth^{14–16,19}. This has been attributed to a strong Al-N bond that prevents desorption of gallium atoms incorporated on the different surfaces.

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Fig. 1 Schematic illustrations of ideally relaxed Al-polar AlN structures with the (0001) surface (a), ($10\overline{1}0$) surface (b), ($10\overline{1}3$) surface (c), and ($11\overline{2}2$) surface (d) in side view along [$11\overline{2}0$] for (a)-(c) and along [$1\overline{1}00$] for (d).

There is still an open question about the impact of different surface dangling bond densities on aluminium incorporation on differently oriented surfaces. Therefore, to understand this impact, in this paper, we report on pulsed-growth of $Al_xGa_{1-x}N$ layers simultaneously on polar (0001), semipolar (1013) and (1122), as well as nonpolar (1010) AlN/sapphire templates over the entire range of composition.

59 2 Experimental

Growth was performed in an EpiQuest 3×2 -inch close-coupled 60 showerhead MOVPE reactor. Ammonia (NH₃), trimethylgallium 61 (TMGa) and trimethylaluminium (TMAl) were used as precur-62 sors. Differently surface-oriented Al-polar AlN templates grown 63 on sapphire substrates were used to grow AlGaN layers, includ-64 ing (0001) AlN ($d \approx 350$ nm) on *c*-plane sapphire, (11 $\overline{2}2$) AlN 65 $(d \approx 1000 \text{ nm})$ on *m*-plane sapphire, $(10\overline{1}0)$ AlN $(d \approx 550 \text{ nm})$ on 66 *m*-plane sapphire, and untwinned (1013) AlN ($d \approx 350$ nm) on *m*-67 plane sapphire. Growth parameters of these templates are re-68 ported elsewhere^{16,26,27}. The total threading dislocation den-69 sity (TDD) of the (0001) templates is estimated to be about 70 $2-3\times10^9$ cm² by X-ray diffraction (XRD) measurements of the 71 (0002) and (1012) AlN X-ray rocking curves. For the non-c-72 plane templates, due to their broad TDD rocking curves^{16,26}, ac-73 cording to Ref.³³ their TDD is estimated to be in the range of 74 $10^{10} \,\mathrm{cm}^2$. Additionally, the basal-stacking fault density of the 75 non-c-plane templates is estimated to be in the range of 10^5 -76 10⁶ cm^{-1 26}. It should be noted that similar templates had previ-77 ously been used to produce continuously grown $Al_vGa_{1-v}N$ layers 78 (0 < y < 1), which showed clear near-band-edge luminescence at 79 room temperature^{15,16}. Fig. 1 shows schematic illustrations of 80 these differently oriented AlN surfaces in side view. 81

All the 2-inch AlN/sapphire wafers were diced into 1×1 cm² pieces. They were then co-loaded into the reactor chamber for



Fig. 2 Representative growth sequence of a cycle using pulsed-flow growth with fixed $t_{\sf N}=t_{\sf Ga+N}=6\,{\sf s},$ while $t_{\sf Al}=1{\rightarrow}14\,{\sf s}.$

AlGaN epitaxy. Initially, about 50-nm-thick AlN layer was continuously grown on these templates at 1200°C at a reactor pressure of 27 hPa. Afterwards, AlGaN layers were grown on these templates using a pulsed-growth mode at 1100°C at a reactor pressure of 87 100 hPa in hydrogen (H₂) ambient. The total gas flow rate during AlN and AlGaN epitaxy was kept constant at 15 slm (standard litre 89 per minute). The AlN mole fraction (x_{AlN}) of the layers was varied 90 by alternatively supplying TMAl, TMGa and NH₃ precursors into 91 the reactor. Fig. 2 illustrates a cycle of pulsed-growth including 92 (1) NH₃ stabilization ($t_N = 6 \text{ s}$, $P_{NH_3} = 3.33 \text{ kPa}$), (2) only TMAl 93 supply ($t_{Al} = 1 \rightarrow 14$ s, $P_{TMAl} = 0.61$ Pa) and (3) NH₃ + TMGa supply ($t_{Ga+N} = 6 s$, $P_{NH_3} = 3.33 kPa$ and $P_{TMGa} = 0.92 Pa$). The to-95 tal number of cycles was 300. The nominal layer thickness was 96 of 400-500 nm. For comparison, AlGaN layers were also grown 97 on those differently oriented AlN templates using the continu-98 ous growth mode, i.e., all the precursors were simultaneously in-99 troduced during AlGaN epitaxy using the same partial pressures 100 mentioned above (V/III \sim 220). 1 01

The structural properties of the AlGaN/AlN samples were char-102 acterized using a Malvern PANalytical Empyrean triple-axis high-103 resolution XRD system equipped with a hybrid monochromator 1 04 $2 \times \text{Ge}(220)$ Cu asymmetric for CuK_{$\alpha 1$} source ($\lambda = 0.15406 \text{ nm}$). 105 Reciprocal space maps (RSMs) were measured using the detector 106 in a frame-based (1D) mode. To calculate x_{AlN} of the semipolar 107 and nonpolar AlGaN layers, their lattice constants and distortion 108 angles have been calculated from different symmetric and asym-109 metric 2θ - ω diffraction peaks, as previously described in details 110 in Refs.^{14–16}. For the *c*-plane layers, their lattice constants were 111 calculated from the symmetric (0002) and asymmetric ($10\overline{15}$) 112 peaks. These peaks were measured using a PIXcel3D-Medipix3 113 1×1 detector in a scanning mode. The surface morphology of 114 the samples was measured by atomic force microscopy (AFM) in 115 tapping mode (Nanocute, SII NanoTech). The optical properties 116 of the grown samples were investigated by photoluminescence 117 (PL) measurements at room temperature (RT). For RT-PL mea-118 surements, the samples were excited by a Krypton Fluoride (KrF) 119 excimer laser (ExciStar XS-200) with excitation wavelength of 120 248 nm ($E_{ex} = 5 \text{ eV}$), a spot size of $50 \times 500 \,\mu\text{m}^2$ and a power den-121 sity of 5.6 kW/cm². PL signals were recorded by a high-sensitivity 122 Ocean Optics spectrometer (QE65 Pro). 123



Fig. 3 Symmetric $2\theta \cdot \omega$ XRD scans of the (a) (0001), (b) ($10\overline{1}0$), (c) ($10\overline{1}3$) and (d) ($11\overline{2}2$) AlGaN co-loaded layers grown with different $t_N/t_{Al}/t_{Ga+N}$. Scans of the continuously grown layers are also plotted for comparison.

124 3 Results and Discussion

125 3.1 Determine aluminium incorporation by XRD

Fig. 3 shows symmetric $2\theta \cdot \omega$ XRD scans of the (0001), (1013), 126 (1122) and (1010) AlGaN co-loaded layers grown using the con-127 tinuous growth method. From compositional calculations based 128 on XRD measurements $^{14-16}$, estimated x_{AlN} of these layers has 129 been found to be comparable of 0.22 ± 0.02 (see their x_{AlN} val-1 30 ues shown later in Fig. 5). The similar x_{AlN} values of these layers 1 31 confirm results previously reported for differently oriented AlGaN 1 32 layers grown with similar growth conditions^{14–16}. 133

For all the differently oriented AlGaN layers grown using pulsed-growth with different $t_N/t_{Al}/t_{Ga+N}$, it is clearly seen that simply increasing t_{Al} leads to a shift of symmetric 2 θ diffraction peak of AlGaN to a higher value (Fig. 3). This indicates an increase of x_{AlN} in the layers. For the *c*-plane layers grown with $t_{Al} \ge 4$ s (Fig. 3(a)), two diffraction peaks appear, indicating a phase separation in these layers. The dominant (0002) peaks 140 shift to higher 2θ angles with $t_{A1} = 4 \rightarrow 8$ s. For these layers, x_{A1N} 141 of the secondary peak is estimated to be about 0.25 \pm 0.05. For the 142 *c*-plane layer grown with $t_{Al} = 12$ s, its phase separation seems to 143 be suppressed as only one peak has been observed. Interestingly, 144 this peak shifts to a lower angle than those of the dominant peak 145 of the layers grown with $t_{Al} = 6 s$ and 8 s, indicating a reduced 146 x_{AlN} . 147

For the $(11\overline{2}2)$ layers, a similar phase separations have also 148 been observed for the layers grown with $t_{Al} \ge 4 s$ (Fig. 3(d)). 149 However, the dominant (1122) peak monotonically shifts to a 150 higher 2θ angle with increasing t_{Al} . This can be clearly seen in 151 RSMs shown in Fig. 4. For the layers grown with $t_{A1} = 4 \rightarrow 8$ s, the 152 secondary peak appears at an angle higher than that of dominant 153 peak. However, for the layers grown with $t_{Al} > 8 s$, their domi-1 54 nant peaks appears at higher angles with very weak intensity of 155 the secondary peaks. Interestingly, for these layers, besides the 156

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Fig. 4 Symmetric RSM scans of the (1122) AlGaN layers grown on (1122) AlN/sapphire templates with different $t_N/t_{Al}/t_{Ga+N}$.



Fig. 5 XRD-estimated $x_{A|N}$ of the (0001), (1010), (1013), (1122) A|GaN co-loaded layers pulsed-grown with different $t_N/t_{A|}/t_{Ga+N}$. Values (\Rightarrow) estimated for the layers grown using the continuous growth are plotted for comparison.

weak phase separation observed for AlGaN, a very broad peak appears that is close to the (30 $\overline{3}$ 0) sapphire diffraction peak. This can be assigned to the (11 $\overline{2}$ 2) GaN peak (Figs. 3(d) and 4(d)), i.e., some grains of GaN might be formed on top of AlGaN underneath layers. However, such peak is unobservable in the 2θ - ω scans of the other differently oriented co-loaded layers.

In contrast to the *c*-plane and $(11\bar{2}2)$ layers, phase separation is unobservable for the $(10\bar{1}0)$ layers, as only a single peak has been observed (Fig. 3(b)). Additionally, phase separation is very weak for the $(10\bar{1}3)$ layers due to very low intensity of secondary peaks (Fig. 3(c)). Their dominant peak monotonically increases with increasing t_{Al} .

Fig. 5 shows XRD-estimated x_{AIN} of the differently oriented Al-GaN layers grown with $t_N/t_{Al}/t_{Ga+N}$. It should be noted that the shown x_{AIN} values of the *c*-plane and (11 $\overline{2}2$) layers were estimated from the phases with dominant Al-distributions (Figs. 3(a) and (d)). In contrast to comparable x_{AIN} values of 0.22 \pm 0.02 estimated for the continuously grown layers, the pulsed-grown layers have different x_{AIN} with respect to each other. For the pulsed-175 grown layers with $t_{Al} = 1$ s, their x_{AlN} are smaller than that of 176 the continuously grown layers. However, the layers grown with 177 $t_{Al} \ge 2 s$, their x_{AlN} are comparable (as estimated for the *c*-plane 178 and $(10\overline{1}3)$ layers) or even higher (as estimated for the *m*-plane 179 layer) than that of the continuously grown layers, except for the 180 (1122) layer. By increasing t_{Al} , x_{AlN} of the pulsed-grown (1010) 181 and $(10\overline{1}3)$ layers increases considerably, as shown in Figs. 3(b)-182 (c). The maximum x_{AIN} of the (1010) and (1013) AlGaN layers is 183 of $\sim 0.85 \ (t_{\rm Al} = 14 \, {\rm s}).$ 184

For the (11 $\overline{2}2$) layers, the maximum x_{AlN} (of the dominant phase) is ~0.75 ($t_{Al} = 14$ s). As previously mentioned, the *c*-plane layers grown with $t_{Al} \ge 10$ s, their (0002) XRD peaks shift to lower diffraction angles, indicating a reduced x_{AlN} . This can be clearly seen in Fig. 5. The x_{AlN} value of the *c*-plane layers reaches a maximum of about 0.67 \pm 0.02 at $t_{Al} = 6$ and 8 s, then it reduces to a value of about 0.57 \pm 0.02 for the layers grown with longer t_{Al} .

3.2 Surface morphology

Fig. 6 shows $10 \,\mu\text{m} \times 10 \,\mu\text{m}$ AFM images of the (0001), (1010), 193 (1013), (1122) AlGaN co-loaded layers pulsed-grown with dif-1 94 ferent $t_N/t_{Al}/t_{Ga+N}$. The surface morphology of the *c*-plane 195 layers grown with $t_{Al} > 1 s$ shows different areas separated by 196 boundaries, attributed to the occurrence of their phase separation 197 (Fig 3(a)). Root-mean-square roughness values of these *c*-plane 198 layers have been found to decrease from about 43 nm to 13 nm 199 with $t_{Al} = 1 \rightarrow 12$ s. For the layers grown with $t_{Al} > 8$ s, their mor-200 phologies are smoother, likely resulting in their suppressed phase 201 separation. 202

The non-c-plane AlGaN layers show typically undulated 203 surface morphology (with V-shaped patterns), i.e., along 204 $[0001]/[11\overline{2}0]/[1\overline{1}00]$ for the $(10\overline{1}0)/(10\overline{1}3)/(11\overline{2}2)$ layers, re-205 spectively. Such morphology is attributed to anisotropic diffu-206 sion lengths of group-III atoms on these surfaces along two in-207 plane directions^{11-14,16-19}. The morphologies of the pulsed-208 grown (1010) and (1013) layers are similar to that of contin-209 uously grown layers previously reported in Refs.^{14,16}. Com-210 pared to the morphologies of the (1010) and (1013) pulsed-211 grown layers, those of the (1122) layers have much larger V-212 shaped patterns. Additionally, compared to morphology of con-213



Fig. 6 10 μ m × 10 μ m AFM images of the (0001), (1010), (1013), (1122) AlGaN co-loaded layers pulsed-grown with different $t_N/t_{Al}/t_{Ga+N}$. Root-mean square (rms) roughness values of these layers are shown for comparison.

tinuously grown (1122) AlGaN layers previously reported in the 214 literature 11,12 , pattern sizes of the (1122) layers studied here are 215 much larger and less undulated. This can be attributed to the 216 pulsed-growth conditions employed here, which resulted in the 217 different surface diffusions and surface reconstructions (discussed 218 later in section 3.4.1). For continuously grown (1122) AlGaN lay-219 ers on GaN microrods on *m*-plane sapphire³⁴, by using cathodolu-220 minescence hyper-spectral imaging, lower Al incorporations have 221 been found near and around micro-sized V-shaped features com-222 pared to feature-free areas. Given the rough morphology of the 223 layers studied here, one might expect similar alloy fluctuations 224 around their surface features. 225

226 3.3 Photoluminescence

Fig. 7 shows RT-PL spectra ($E_{ex} = 5 \text{ eV}$) of the (0001), (10 $\overline{1}$ 0), (10 $\overline{1}$ 3) and (11 $\overline{2}$ 2) AlGaN co-loaded layers pulsed-grown with $k_N/t_{Al}/t_{Ga+N} = 6/8/6$. For all the continuously grown layers ($x_{AlN} \approx 0.22 \pm 0.02$), only near-band-edge emission energy has been observed at (3.9 \pm 0.1) eV with a full-width at half maximum (Δ_{PL}) of (0.19 \pm 0.03) eV. Similar finding has been found for continuously grown layers with different surface orientations previously reported in Refs.^{14–16}. 234

From the PL spectra shown in Fig. 7, despite different x_{AIN} of 235 the pulsed-grown layers (Fig. 5), all their spectra show domi-236 nant emission energy at about \sim 3.6-3.9 eV. This dominant emis-237 sion mainly originates from substitutional carbons on nitrogen 2 38 sites (C_N), as previously reported for Al(In)N^{19,35-37}. The 239 spectrum of the (1122) layer shows two peaks at about 3.6 240 and 3.8 eV (after Gaussian fitting of the corresponding bands), 241 likely related to their phase separation (Fig. 3(d)) that might 242 incorporate different concentrations of impurities. This dom-243 inant has also been observed for all the other pulsed-grown 244 samples, generally shifting from \sim 3.5 to 3.9 eV with increas-245 ing x_{AIN} ($\Delta_{PL} \approx (0.20 \pm 0.01) \rightarrow (0.42 \pm 0.13)$ eV). For bulk AlN crys-246 tals, it has been found that an increased [C] in AlN shifts 247 the C_N-related emission energy to higher energy regions (e.g., 248 $[C] \approx 2 \times 10^{17} \rightarrow 2 \times 10^{19} \text{ cm}^{-3})^{35}$. This is plausible for the layers 249 studied here, as [C] has previously been found increase with in-250 creasing x_{AlN} in AlGaN³⁸. The dominant C_N emission observed 251 for the pulsed-grown samples indicates highly C-incorporated 252 into them. This can be explained due to the fact that more C-253

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Fig. 7 RT-PL spectra of the (0001), (1010), (1013) and (1122) AlGaN co-loaded layers pulsed-grown with $t_N/t_{Al}/t_{Ga+N} = 6/8/6$.

containing species (e.g., $Al_xC_yH_z$ and $Al_xC_yN_z$ ^{38–40}) will be easily incorporated on the exposed surfaces during TMAl supply without NH₃ (i.e., V/III = 0).

Additionally, the PL shoulders at ~4.2 eV and ~4.7 eV (after Gaussian fitting of the corresponding bands) observed for the (1010) and (1013) layers ($x_{AlN} \approx 0.76$ -0.78) are attributed to acceptor C_N^- and nitrogen vacancies (V_N), respectively³⁷.

261 3.4 Discussion

262 3.4.1 Surface reconstructions

It is well-known that TMAl:NH₃ pre-reactions strongly occur 263 when introducing TMAl and NH3 simultaneously during AlGaN 264 epitaxy in MOVPE, limiting the Al incorporation^{12,31}. Pulsed-265 growth can suppress these pre-reactions, resulting in an enhanced 266 Al-incorporation as observed for the AlGaN layers studied here 267 (Fig. 5). During the pulsed-growth, as TMAl was introduced with 268 different t_{A1} without NH₃, the surface stoichiometry of the lay-269 ers was actually changed from Al-poor to Al-rich conditions. This 270 might lead to different surface reconstructions, causing a forma-271 tion of coexistence areas with different x_{AlN} (i.e., phase separa-272 tion) in one layer. 273

During MOVPE growth of AlGaN in H₂ ambient (i.e., using H₂ 274 carrier gas), the layer surfaces interact with H-rich ambient con-275 ditions. Due to stronger N-H bond energy (\sim 3.5 eV) than those of 276 Al-H (\sim 3 eV) and Ga-H (\sim 2.8 eV) ⁴¹, H atoms easily terminate the 277 topmost N atoms on the Al(Ga)N surfaces. There are few reports 278 for reconstructions of differently oriented AlN and GaN during 279 MOVPE growth, including (0001)^{42,43}, (1122)^{42,44,45}, as well as 280 (1010)^{42,46}. Additionally, for GaN MOVPE-grown at 1100°C (i.e., 281 the temperature used for AlGaN epitaxy in this study), only GaN 282 desorption happens⁴². Therefore, surface reconstructions of the 283 AlGaN layers studied here are mainly governed by the stability of 284 AlN surfaces. 285

For MOVPE-grown (0001) AlN layers grown in H_2 ambient at

For the $(10\overline{1}0)$ GaN and AlN surfaces, by using first-principles 293 calculations, their ideal 1×1 symmetry has been found to be sta-294 ble over the thermodynamically allowed range, irrespective of N-295 rich or H-rich conditions used at $1100^{\circ}C^{42,46}$. The (1013) GaN 296 surface has been found to have a remarkable morphological sta-297 bility, stabilizing a metallic structure (Ga adlayer) over the entire 298 range of the Ga chemical potential in N-rich conditions⁴⁸. So 299 far, no calculations for the $(10\overline{1}3)$ AlN surface have been studied. 300 However, as shown in Figs. 1(c)-(d), the (1013) AlN surface ex-301 hibits a kind of similar atomic arrangement to the $(10\overline{1}0)$ AlN sur-302 face with the topmost N and Al atoms. Thus, one can expect that 303 their surface reconstructions should show the same behaviours 304 over the chemical potential range. Therefore, the $(10\overline{1}3)$ AlN sur-305 face can also be considered as a stable surface, i.e., similar to the 306 GaN surfaces. For the (1122) AlN surface in H-rich conditions, 307 three different reconstructions have been found at 1100°C with 308 increasing Al pressure, including $c(2 \times 2)$ with H atoms [Al_{adatom}-309 H+N-AlH₂], 2×2 8N-6Al and $c(2 \times 2)$ -Al adatom⁴⁵. 310

Conclusively, with the above discussion, for the pulsed-grown 311 $(10\overline{1}0)$ and $(10\overline{1}3)$ layers, their stable surfaces are consistent with 312 the unchanged surface morphology of the AlGaN grown layers as 313 shown in Figs. 6(a)-(h), consistent with the absence of phase sep-314 aration in these layers as shown in Figs. 3(c)-(d). For the pulsed-315 grown (0001) and (1122) AlGaN layers studied here, their dif-316 ferent surface morphologies (Figs. 6(i)-(p)) can be attributed to 317 different surface reconstructions during growth, resulting in their 318 phase separation as observed by XRD (Figs. 3(a)-(b)). Further in-319 vestigations are under going to find out if such phase separation 320 caused by different stacked layers or localized areas. 321

3.4.2 Surface dangling bonds

The differences in x_{AlN} of the pulsed-grown layers can also be 323 attributed to differences in dangling-bond (db) density on their 324 differently oriented surfaces. As shown in Fig. 1, the ideal (0001) 325 and $(11\overline{2}2)$ AlN surfaces both have the topmost Al atoms, while 326 the ideal (1013) and (1010) AlN surfaces are terminated with 327 both N and Al atoms. However, on the $(10\overline{1}3)$ surface, each Al 328 atom has one db, while each N atom has one or two dbs. In con-329 trast, on the $(10\overline{1}0)$ surface, each Al and each N atoms have only 330 one db. Table 1 shows the surface N- and (Al,Ga)-db densities 331 (D_N, D_{Al,Ga}) calculated for different AlN and GaN surfaces. For 332 these calculations, lattice constants of unstrained AlN and GaN 333 have been used 49. 3 34

During the AlGaN pulsed-growth, after introducing TMAl, Al-N 335 bonds will be formed first. The higher number of N-sites exists 336 on a surface, the higher number of Al-N bonds can be formed. 337 Ga-N bonds can only be formed at the residual N-sites after introducing TMGa (with NH₃). Consequently, the higher D_N on the (1010) and (1013) surfaces results in their higher aluminium incorporations than those on the (0001) and (1122) surfaces. 341 $\begin{array}{l} \textbf{Table 1} \text{ Lattice constants of unstrained AIN and GaN 49 used to calculate surface dangling-bond densities (D^{AIN,\,GaN}) of AI, Ga, and N atoms on different ideal AIN and GaN surfaces. \end{array}$

Materials	Lattice constants (angstrom)				
AlN	$a_0 = 3.111\ 97; c_0 = 4.980\ 89$				
GaN	$a_0 = 3.189$ 40; $c_0 = 5.186$ 14				
Surface	Dangling-bond density (nm ⁻²)				
orientations	DAlN	$\mathrm{D}_{\mathrm{N}}^{\mathrm{AlN}}$	$\mathrm{D}_{\mathrm{Ga}}^{\mathrm{GaN}}$	$\mathrm{D}_{\mathrm{N}}^{\mathrm{GaN}}$	
(0001)	11.90	-	11.35	-	
(1010)	6.45	6.45	6.05	6.05	
(1013)	10.15	15.23	9.62	14.43	
(1122)	18.95	-	17.84	-	

As shown in Table 1, $D_N^{Al(Ga)N}$ of the (1013) surface is higher 342 than that of the (1010) surface. Therefore, Ga atoms will 343 have more chance to incorporate on the (1013) surface than 344 on the (1010) surface. This explains why x_{AlN} of the grown 345 $(10\overline{1}3)$ AlGaN layers is lower than that of the $(10\overline{1}0)$ co-loaded 346 layers for $t_{Al} = 1 \rightarrow 6 s$. Further increasing t_{Al} can increase the 347 number of Al-N bonds, resulting in an increase of x_{AlN} . How-348 ever, x_{AlN} values of the (1010) layers grown with $t_{Al} = 8 \rightarrow 14$ s 349 are slightly lower than that of the (1013) layers ($\Delta x \leq 0.04$). As 350 mentioned above, the $(10\overline{1}0)$ AlGaN surface is very stable, i.e., Al 351 adlayers cannot be formed on this surface 42,46. This means that 352 for $t_{Al} \ge 8$ s, instead of forming Al adlayers on the (1010) sur-353 face, abundant Al atoms will be desorpted. In contrast, due to the 354 higher $D_N^{Al(Ga)N}$ of the (1013) surface, the number of Al-N bonds 355 still keeps increasing with t_{Al} , resulting in higher t_{Al} of the (1013) 356 layers with $t_{Al} \ge 8$ s. 357

For the *c*-plane and $(11\overline{2}2)$ layers, due to their phase separa-358 tion, it is difficult to compare their x_{AIN} . However, the trend of 359 lower x_{AIN} with higher db density is evident. Furthermore, these 360 surfaces only have metal dangling bonds, and thus pick up N 361 atoms easily. Since the bond energy of Ga-N $(8.92 \text{ eV}/\text{atom}^{50})$ is 362 lower than that of Al-N (11.52 eV/atom⁵⁰), Al wants to exchange 363 with Ga on the surface until the accumulated strain makes this 364 process unfavourable. Hence, on these two different surfaces, Al 365 atoms are "consuming" the previous GaN layer, and thus form an 366 Al-poor AlGaN layer first until limited by further Al supply and 367 strain. 368

369 4 Conclusions

Growth of $Al_xGa_{1-x}N$ layers simultaneously on polar (0001), 370 semipolar (1013) and (1122), as well as nonpolar (1010) AlN 371 templates has been investigated by metal-organic vapour phase 372 epitaxy. A pulsed-flow growth mode was used to vary the AlN 373 mole fraction ($0 < x_{AlN} \le 0.85$) in the layers. Phase separation 374 has been observed for the (0001) and ($11\overline{2}2$) layers, attributed to 375 their different surface reconstructions during growth. In contrast, 376 no phase separation has been observed for the $(10\overline{1}0)$ and $(10\overline{1}3)$ 377 layers, attributed to their stable surfaces during growth. The AlN 378 mole fraction of the differently oriented layers generally follows 379 the order: $(11\overline{2}2) < (0001) < (10\overline{1}3) \le (10\overline{1}0)$. This has been at-380

tributed to their different surface dangling bonds, from which the 381 surface with higher cation dangling bond density (i.e., $(10\overline{1}3)$ and 382 $(10\overline{1}0))$ incorporated more aluminium during pulsed-growth. For 383 the (0001) and (1122) layers, the trend of lower x_{AIN} with higher 384 anion dangling bond density has also been found, which might be 385 attributed to different bond energies between Ga-N and Al-N, as 386 well as surface relaxation. By means of room-temperature lumi-387 nescence measurements, highly carbon-incorporation has quanti-388 tatively been found in the layers, whose emission energy slightly 389 increases from \sim 3.6 to 3.9 eV with increasing x_{AIN} . 390

Conflicts of interest

There are no conflicts to declare.

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