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## Fabrication of earth-abundant Cu<sub>2</sub>ZnSn(S,Se)<sub>4</sub> light absorbers by a sol-gel and selenization route for thin film solar cells

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### 1 Abstract

2  $Cu_2ZnSn(S,Se)_4$  (CZTSSe) absorber for thin film solar cells was fabricated by a nonaqueous thiourea-metal-oxygen sol-gel processing followed by post-selenization. 3 The effects of selenization temperature, selenization time and metal composition on 4 5 the film microstructure and phase were investigated. Under optimized process parameters, the complete CZTSSe solar cell devices produced efficiency values as 6 high as 8.08 % with a short-circuit current density of 32.78 mA cm<sup>-2</sup>, an open-circuit 7 voltage of 487 mV, and a fill factor of 51 %. The factors that limit the device 8 9 performance were discussed. The nonuniform film thickness and nonhomogeneous 10 chalcogen components should be responsible for the low shunt resistance and thereby low open-circuit voltage; and the presence of large number of voids and thick 11  $Mo(S,Se)_2$  interface layer may be the main reasons for high series resistance, 12 detrimental to short-circuit current density and fill factor. 13 14 *Keywords*: CZTSSe; Thin film solar cells; Sol-gel; Selenization

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### 1 **1. Introduction**

2 Kesterite-structured copper zinc tin chalcogenide (sulfoselenide) materials have recently emerged as potential substituents for Cu(In,Ga)Se<sub>2</sub> (CIGSe) because of 3 abundance and lower cost<sup>1</sup>. Cu<sub>2</sub>ZnSn(S,Se)<sub>4</sub> (CZTSSe) exhibits many intrinsic 4 properties, such as a high optical absorption coefficient (>  $10^4$  cm<sup>-1</sup>) and a tunable 5 bandgap that can be varied from 1.0 to 1.5 eV via changing the S/(S+Se) ratio to 6 favorably match the solar spectrum <sup>2-4</sup>. According to current state of the art, the 7 highest power conversion efficiency of 12.7 % has been reported in which the 8 9 prototypical devices were prepared using a hydrazine pure-solution approach by IBM Corporation <sup>5</sup>. 10

A large number of approaches have been reported to synthesize CZTSSe absorber 11 materials, such as thermal evaporation  $^{6}$ , sputtering  $^{7}$ , electrodeposition  $^{8}$ , a 12 solution-particle approach <sup>9</sup>, sol-gel method <sup>10</sup> and so on. Among all of these various 13 14 preparation approaches, sol-gel technique has shown numerous vital advantages including lower cost, non-toxicity, potential for industrial scale-up, and precise 15 16 stoichiometric control. Previously, our group has reported a sol-gel solution approach 17 to synthesize  $Cu_2ZnSnS_4$  (CZTS) precursor films via dissolving metal constituents and excess thiourea in 2-methoxyethanol to form thiourea-metal-oxygen complex 18 followed by thermally decomposing via air annealing to form CZTS xerogel precursor. 19 After post-sulfurization process, a power conversion efficiency equal to 5.1 % of 20 sulfide CZTS device was achieved <sup>10</sup>, which was then improved to 5.7% by low 21 pressure sulfurization <sup>11</sup>. It has been reported that the replacement of S atoms by Se 22 could enhance the grain size, minimize the inner voids and adjust the bandgap<sup>12,13</sup>, 23 yielding higher conversion efficiency. Herein, we studied the fabrication of CZTSSe 24 25 absorbers by post-selenization of the above-mentioned CZTS xerogel precursor and 26 the influence of selenization temperature, selenization time and metal composition in the phase and morphology of CZTSSe thin film. Using the optimized selenization 27 process, conversion efficiency above 8 % without antireflection coating has been 28 29 achieved. In addition, based on the TEM analysis, the limiting factors in solar cell efficiency have been discussed. 30

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### **1 2.** Experimental

### 2 2.1. Preparation of the CZTSSe thin film

The precursor solution was prepared by dissolving Cu(CH<sub>3</sub>COO)<sub>2</sub>·H<sub>2</sub>O (AR), 3  $Zn(CH_3COO)_2 H_2O$  (AR),  $SnCl_2 H_2O$  (AR) and excess  $SC(NH_2)_2$  (2 M, AR) into 4 5 2-methoxyethanol (AR). Total metal ion concentration was 1 M. The final Cu/(Zn+Sn)and Zn/Sn ratios were 0.86 and 1.0, respectively. Triethanolamine with 1 percent in 6 7 volume was added to avoid cracks during the spin coating process. All chemical 8 reagents were purchased from Sinopharm Chemical Reagent Co., ltd. After stirring 9 the solution at 50 °C for 1 hour, a sol solution was prepared. The CZTS precursor film 10 was deposited via spin-coating from the as-prepared sol solution on Mo-coated soda-lime glass substrate at 4000 rpm for 30 s. To get the targeted thickness ( $\sim 1 \ \mu m$ ) 11 of the thin film absorber, several sequent layers were coated following an intermediate 12 heat treatment at 270 °C for 10 min on a hot plate in air, in which the process formed 13 binary metal sulfides and oxides such as Cu<sub>x</sub>S, ZnS, Sn<sub>x</sub>S and SnO<sub>2</sub><sup>10</sup>. Finally, the 14 resultant precursor film was subjected to selenization process at 500  $\sim$  580 °C for 5  $\sim$ 15 16 60 min in 0.04 MPa to achieve desired crystallinity under selenium/Ar gas 17 environment, which allowed the formation of CZTSSe film by substituting part of S by Se. The selenization was performed in a tube furnace with double temperature 18 zone. The temperature of Se zone was programmed to be 400 °C. 19

### 20 *2.2. Device fabrication and characterization*

The CZTSSe-based photovoltaic device was fabricated with the conventional 21 structure of glass/Mo/CZTSSe/CdS/ZnO/ITO. A 70 nm-thick CdS buffer layer was 22 23 deposited on CZTSSe film by chemical bath deposition (CBD) method. More details can be found in reference <sup>14</sup>. Intrinsic ZnO (80 nm) and ITO (~ 300 nm) were 24 deposited by RF and DC magnetron sputtering, respectively. The active area of each 25 device was approximately 0.45 cm<sup>2</sup>. Detailed process of device has been described in 26 references <sup>10, 15</sup>. The device electrical characterization was performed using Xe-based 27 light source solar simulator (Newport, 91160) and a Source Measure Unit 28 29 (KEITHLEY 2400), equipped with a standard Si reference cell, to provide simulated 1 sun AM 1.5G illumination. The external quantum efficiency (EQE) was measured 30

1 using the QEX10 spectral response system from PV measurements, Inc.

2 2.3. Materials characterization

The morphology of the CZTSSe absorber was characterized by scanning electron 3 microscopy (SEM, FEI Quanta-200 and NOVA NanoSEM 230). Elemental analysis 4 5 was determined by energy dispersive spectrometer (EDS, EDAX-GENSIS60S in NOVA NanoSEM 230). Structural characterization was carried out by means of an 6 7 X-ray diffractometer (XRD, Rigaku-TTR III X) and Raman spectroscopy (Raman, 8 Jobin-Yvon LabRAM HR-800). The cross section morphology and composition 9 analysis were measured by scanning transmission electron microscopy (STEM, 10 JEM-2100F).

### 11 **3. Results and discussion**

### 12 *3.1. Effect of the selenization temperature*

Top view and cross-sectional scanning electron microscopy (SEM) images of the 13 CZTSSe thin films selenized under Ar-Se atmosphere at different temperatures are 14 shown in Fig. 1, and the surface micrograph of the precursor film is presented in Fig. 15 16 S1. The CZTS precursor film has a smooth, uniform surface and is consisted of nanocrystalline<sup>10</sup>. After selenization, the crystal size enhances significantly and the 17 selenization temperature has a significant impact on the morphology of the CZTSSe 18 absorber, as shown in Fig. 1. The film formed at 500 °C shows a uniform, crack-free, 19 20 and pinhole-free surface with nanocrystal structure (Fig. 1a). As the selenization temperature increases, the size of grains are enlarged, indicating better crystallinity, 21 while some pores can be observed on the surface (Fig. 1b and c). When the sample is 22 23 selenized at 580 °C, the film is dense, smooth and pinhole-free with micrometer-sized 24 grains, as shown in Fig. 1d. From the cross-sectional SEM images in the insets, it is shown that when the selenization temperature rises to 560 °C or higher, the grains 25 26 span the entire film thickness, which is beneficial to the photocarrier transportation and reduce the bulk recombination. 27

As shown in Fig. 2a, the X-ray diffraction (XRD) patterns of the CZTSSe thin film selenized above 500 °C reveal the formation of CZTSSe without Cu or Sn sulfides/selenides secondary phases. The intensity of the (112) main diffraction peaks

corresponding to the CZTSSe increases gradually with increasing selenization 1 temperature under 560 °C, and shows a significant boost for the sample selenized at 2 580 °C. Furthermore, the (101) characteristic peak from CZTSSe is visible at high 3 selenization temperature, implying that the CZTSSe thin film has good crystallinity <sup>16</sup>. 4 Raman spectroscopy is also applied to detect the phase purity of the CZTSSe thin 5 films because some binary, ternary sulfides/selenides, such as ZnS, Cu<sub>2</sub>SnS<sub>3</sub>, ZnSe 6 and Cu<sub>2</sub>SnSe<sub>3</sub>, have similar XRD patterns with CZTS and Cu<sub>2</sub>ZnSnSe<sub>4</sub> (CZTSe)<sup>17</sup>. 7 8 Fig. 2b shows the Raman spectra of CZTSSe thin films selenized at different temperatures. For all of the CZTSSe films, the main peaks are located at 9 approximately 200 cm<sup>-1</sup> and 330 cm<sup>-1</sup>, which are consistent with the A<sub>1</sub> vibration 10 mode of CZTSe and CZTS, respectively <sup>18</sup>. No other sulfides and/or selenides 11 secondary phases are found in the Raman spectra for the CZTSSe thin films. It is 12 obviously observed that with the increase of selenization temperature, the intensity of 13 14 main peaks increases, indicating that the films obtain better crystallinity. However, the Raman peak intensity significantly decreases at 580 °C due to the degradation of the 15 16 CZTSSe film. Moreover, the atomic ratio image of the CZTSSe films selenized at 17 different temperatures is shown in Fig. S2. The change of Cu/(Zn+Sn) ratio at different temperatures is very little, and the Cu/(Zn+Sn) ratio is a bit lower than that 18 of the precursor solution. The Zn/Sn ratios at 500 °C and 530 °C are 0.9, lower than 19 the ratio of the precursor solution. But the Zn/Sn ratio rises to nearly 1.0 at 560 °C, 20 roughly equal to the ratio of the precursor solution. Besides, the S/(S+Se) ratio 21 slightly increases with the rise in selenization temperature, which can also be reflected 22 23 from Raman spectra in Fig. 2b. For films with intermediate S/Se content, the main peaks from both sulfide and selenide are present, and shift toward one other <sup>9</sup>. When 24 the sample is selenized at above 530 °C, the peak at around 200 cm<sup>-1</sup> shifts to higher 25 wavenumber direction which is closer to the pure sulfide (338 cm<sup>-1</sup>) side compared 26 with thin film selenized at 500 °C. From above results, it is evident that the sol-gel 27 and selenization approach allows kesterite CZTSSe fabrication and easy control to 28 29 phase and microstructure by adjusting selenization temperature. As the temperature increases, the grain size increases gradually. However, when temperature rises to 580 30

°C, it is difficult for the absorber film to stay stable. Hence, we choose 560 °C as the
 optimal selenization temperature in this process.

*3 3.2. Effect of the selenization time* 

The SEM images in Fig. 3 illustrate an improvement in crystallization and film 4 quality with different selenization times. It is shown that the grain size gradually 5 increases with an increase in selenization time. The film selenized for 40 min is 6 7 uniform and pinhole-free with micrometer-sized grains. Besides, as shown in the EDS 8 component atomic ratio image (Fig. S3), metal composition of the thin films selenized 9 for different selenization times have little difference. The ratio of S/(S+Se) keeps steady (about 30%) because the substitution reactions occur spontaneously and fast  $^{13}$ , 10 which demonstrates that with the selenization time increasing, the crystallinity and 11 12 size of grain increase, but the Se incorporation remains the same.

In Fig. 4a, XRD patterns demonstrate kesterite structure for absorbers selenized for 13  $5 \sim 60$  min. A selenium annealing atmosphere makes CZTSSe form as early as 5 min. 14 When selenization time increases, the peak intensities for (112), (220)/(204), 15 (116)/(312) planes become stronger and sharper, and (101) peak belonging to the 16 17 kesterite structure becomes observable, indicating an improvement in crystallization and film quality, consistent with the SEM results (Fig. 3). In order to detect the 18 presence of possible binary and ternary phases, the Raman scattering is applied again. 19 Fig. 4b shows that the selenide-based ( $\sim 200 \text{ cm}^{-1}$ ) and sulfide-based ( $\sim 330 \text{ cm}^{-1}$ ) 20 Raman peaks are the main peaks, which agree with the literature results <sup>18</sup>, and no 21 other impurity phases are observed. Moreover, with the extension of the selenization 22 23 time, the characteristic peaks sharpen, indicating that the film crystallinity gradually 24 gets better, which is in agreement with the XRD results. The peak intensities have 25 little change after 40 min, therefore, 40 min is chosen as the appropriate selenization 26 time.

27 *3.3. Effects of the composition* 

In order to study the influence of the metal component on the morphology and structure of CZTSSe film, the precursor solutions with various Cu/(Zn+Sn) ratios ranging from 0.66 to 0.96 were prepared. Chen et al. reported that acceptors in CZTS

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are due to Cu vacancies  $(V_{Cu})$  and substitution of Cu in Zn site  $(Cu_{Zn})$ . They 1 calculated the acceptor transition energy levels for  $Cu_{Zn}$  and  $V_{Cu}$  to be 0.10 and 0.02 2 eV, respectively, which are both above the valence band maximum  $^{19}$ . Since Cu<sub>Zn</sub> is a 3 relatively deep acceptor, the shallow acceptor of V<sub>Cu</sub> is effective for improving the 4 efficiency of CZTS solar cells. The Cu-poor composition suppresses Cu<sub>Zn</sub> formation 5 and enhances  $V_{Cu}$  formation <sup>20</sup>. As can be seen from Fig. 5, the CZTSSe absorber at 6 Cu/(Zn+Sn) ratio of 0.66 shows uneven morphology with some pores. As the 7 8 Cu/(Zn+Sn) ratio increases, the CZTSSe thin film tends to be uniform with larger 9 grain size. When the ratio of Cu/(Zn+Sn) rises to 0.86, the grains with size above 1 um are obtained. Similar phenomenon was reported by Tanaka et al. for the sol-gel 10 prepared CZTS film <sup>21</sup> and by G. Suresh Babu et al. for the co-evaporated 11 Cu<sub>2</sub>ZnSnSe<sub>4</sub> thin film <sup>22</sup>. However, the film morphology deteriorates obviously with 12 some white clusters on the surface when Cu/(Zn+Sn) ratio is as high as 0.96. In 13 14 summary, the film obtains the optimal morphology at Cu/(Zn+Sn) ratio of 0.86. Besides, no impurity phases are observed at this Cu/(Zn+Sn) ratio, as seen in Fig. S4. 15 Related defect theory study for CZTSSe proves that Zn/Sn ratio has a significant 16

17 impact on intrinsic defects of CZTSSe, which affects the carrier diffusion length and mobility<sup>23</sup>. Under Cu-poor condition, the increase of zinc content can promote the 18 formation of [V<sub>Cu</sub>+Zn<sub>Cu</sub>], which can result in band bending that can promote the 19 separation of photo-generated electron-hole pairs and subsequently improve the 20 photoelectric conversion efficiency of the device <sup>23</sup>. Here, we investigate the influence 21 of Zn/Sn ratio on the crystallinity in the range of 0.8 to 1.4. Fig. 6 shows SEM images 22 23 of CZTSSe films with different Zn/Sn ratios. It can be seen that as the Zn/Sn ratio 24 rises from 0.8 to 1.2, the CZTSSe thin film shows a denser surface morphology with larger crystal size. However, some small white grains, which have been previously 25 assigned as ZnS(Se)<sup>24</sup>, appear on the surface for Zn/Sn ratio of 1.2. When Zn/Sn ratio 26 further increases to 1.4, small white particles almost cover the whole film surface, and 27 the size of grain decreases obviously. Fairbrother et al. interpreted this as ZnSe 28 inhibiting grain growth <sup>24</sup>. Therefore, the most proper Zn/Sn ratio is 1.0 from the 29 viewpoint of morphology, which is lower than those described for previous record 30

8

devices <sup>5, 25</sup>. The Zn/Sn ratio of 1.0 benefits the formation of  $SnO_x$  during the post-annealing in air, which is a common process in device fabrication <sup>26</sup>.  $SnO_x$  at the grain boundaries is found to correlate with high device performance and is proposed to passivate GB recombination sites <sup>27</sup>. Furthermore, no binary and ternary phases are found at such ratio from Fig. S4.

6 *3.4. Device performance and limited factor* 

7 The CZTSSe device synthesized by sol-gel approach under the optimal conditions 8 (as stated above) achieves the highest photoelectric conversion efficiency of 8.08 %, in which the ratios of Cu/(Zn+Sn) and Zn/Sn of CZTSSe thin film are  $\sim 0.86$  and 1.0, 9 respectively (EDS spectrum shown in Fig. S5). The electrical characteristics of the 10 best device in the dark and under illumination AM1.5 are presented in Fig. 7a, 11 without MgF<sub>2</sub> antireflection layer. The device produces remarkably high short-circuit 12 current density (Jsc =  $32.78 \text{ mA cm}^{-2}$ ), moderate open-circuit voltage (Voc = 487 mV) 13 and low fill factor (FF = 51 %). The series resistance (Rs) is 3.8  $\Omega$  cm<sup>2</sup>, which is 14 somewhat larger than those reported by Mitzi et al. for the best CZTSSe solar cells <sup>28</sup>. 15 The low shunt resistance (Rsh =  $181.4 \ \Omega \ cm^2$ ) suggests that there may be some shunt 16 17 paths in the device or a strong voltage-dependent current collection effect. The external quantum efficiency (EQE) curve of the device is shown in Fig. 7b. The 18 optical bandgap of the CZTSSe absorber estimated from the x-axis intercept of [EQE 19  $\times$  E]<sup>2</sup> vs. E plot in the inset of Fig. 7b is 1.20 eV. Recently Chen et al. calculated the 20 band gap bowing of  $Cu_2ZnSn(S_xSe_{1-x})$  to be:  $Eg(x) = x \cdot Eg(CZTS) + (1-x) \cdot Eg(CZTSe)$ 21 -  $b \cdot x \cdot (1-x)$  where b is approximately 0.1 eV <sup>29</sup>. The alloy band gap decreases 22 23 monotonically when the Se content increases, from 1.5 eV at x = 0 to 0.96 eV at x = 1. 24 The decrease is almost linear. From the Fig. S3, under the optimal condition (the CZTSSe absorber selenized at 560 °C for 40 min, with ratios of Cu/(Zn+Sn) = 0.8625 and Zn/Sn= 1.0), the S/(S+Se) ratio is about 0.30. Using x = 0.30, Eg is 1.14 eV, 26 which is very close to the band gap estimated by EQE data. EQE value approaches 27 90% in the visible range, however the EQE decays in the infrared region, especially 28 beyond 800 nm, due to a short minor carrier lifetime <sup>30</sup> and/or a narrow depletion 29 layer <sup>31</sup>. From the analyses of electrical characteristics, the largest room for 30

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improvement compared with chalcopyrite devices with similar bandgap values is in
 open-circuit voltage and fill factor.

To gain insight into the limiting factors in solar cell efficiency, we investigate 3 section feature and compositional distribution of the CZTSSe device. A 4 cross-sectional scanning transmission electron microscopy (STEM) image of the best 5 device in Fig. 8 shows large grains (~1.3-1.9 µm) spanning the entire CZTSSe layer. 6 Typical absorption coefficients reported for CZTSSe thin films are around  $10^5$  cm<sup>-1</sup>, 7 which suggests that a thicker absorber is needed to effectively absorb the solar 8 spectrum and reduce recombination at the back contact <sup>32, 33</sup>. It is evident that the film 9 thickness is not uniform with many thinner regions, which may be considered to be 10 one of the important reasons for low shunt resistance and low Voc. The presence of 11 large pores and voids is also clearly observable in the bottom of CZTSSe film, which 12 is the primary cause of a high series resistance and a low fill factor. In Fig. 9a, an 13 energy-dispersive X-ray spectroscopy (EDX) line scan longitudinally through the 14 device shows uniform elemental distribution. As it is evident in this Fig. and also 15 noted in other reports <sup>34, 35</sup>, a detectable amount of copper is observed in the 16 Mo(S,Se)<sub>2</sub> layer. In addition, a moderate thickness Mo(S,Se)<sub>2</sub> layer (~200 nm) is 17 observed between the Mo substrate and CZTSSe absorber, as thick as devices with 18 efficiency beyond  $10\%^{28,35}$ , but the Mo(S,Se)<sub>2</sub> layer is still thicker than the optimal 19 thickness of 50 nm from numerical analysis  $^{36}$ . A thick Mo(S,Se)<sub>2</sub> can result in the 20 increase in Rs and subsequently decrease in FF<sup>37</sup>. There have been many methods to 21 inhibit  $Mo(S,Se)_2$  formation such as an intermediate layer <sup>38-41</sup>, other alternative back 22 contact materials <sup>42</sup> and a prealloying process <sup>43</sup>. EDX compositional profilings 23 laterally across the absorber are shown in Fig. 9b, which demonstrates that copper, 24 25 zinc and tin concentrations stay basically constant and further confirms that the 26 S/(S+Se) ratio is about 0.30. However, some obvious fluctuations in S/(S+Se) ratio are also observed. Previous research on CIGSe has shown that relatively small 27 fluctuations of band gap can lead to a drastic decrease of the device efficiency, and 28 the higher the voltage at the cell, the smaller are the fluctuations <sup>44</sup>. The low Voc can 29 presumably be caused by the effect of band gap fluctuations resulting from the 30

1 inhomogeneous distribution of Se.

### 2 **4.** Conclusion

Here we synthesized CZTSSe thin film solar cell by sol-gel method which is a 3 simple and facile solution method using commercially available precursors and 4 non-toxic solvents. We investigated the influence of selenization temperature, 5 selenization time and metal compositions on morphology and phase of the films. The 6 CZTSSe absorber selenized at 560 °C for 40 min, with ratios of Cu/(Zn+Sn) = 0.867 and Zn/Sn = 1.0, achieves a uniform and pinhole-free surface morphology. Moreover, 8 9 XRD and Raman analyses reveal that no other impurity phases are observed under 10 such condition. A power conversion efficiency (PCE) of 8.08 % (active cell area) has been achieved. However, the present CZTSSe device still suffers from some problems 11 of voids at the back contact, a thick Mo(S,Se)<sub>2</sub> layer, nonuniform film thickness and 12 inhomogeneous distribution of Se. The PCE could be further enhanced by the 13 optimization of the back contact and homogeneity of chalcogen components. In 14 addition, multivariate chemometrics approach <sup>45, 46</sup> will be applied to take into account 15 all the relevant variables and their interactions, finding the best experimental 16 17 conditions for the proper functioning of the CZTSSe thin films in the future.

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23

### 24 **References**

- J. J. Scragg, P. J. Dale, L. M. Peter, G. Zoppi and I. Forbes, *physica status solidi* (*b*), 2008, 245,
  1772-1778.
- H. Katagiri, K. Jimbo, W. S. Maw, K. Oishi, M. Yamazaki, H. Araki and A. Takeuchi, *Thin Solid Films*, 2009, **517**, 2455-2460.
- 29 3. K. Ito and T. Nakazawa, *Jpn. J. Appl. Phys.*, 1988, **27**, 2094-2097.
- 30 4. J. Krustok, R. Josepson, M. Danilson and D. Meissner, *Solar Energy*, 2010, **84**, 379-383.
- 31 5. J. Kim, H. Hiroi, T. K. Todorov, O. Gunawan, M. Kuwahara, T. Gokmen, D. Nair, M. Hopstaken,
- 32 B. Shin, Y. S. Lee, W. Wang, H. Sugimoto and D. B. Mitzi, Advanced Materials, 2014, 26,

1		7427-7431.
2	6.	K. Wang, O. Gunawan, T. Todorov, B. Shin, S. J. Chey, N. A. Bojarczuk, D. Mitzi and S. Guha,
3		Applied Physics Letters, 2010, <b>97</b> , 143508.
4	7.	T. P. Dhakal, C. Peng, R. Reid Tobias, R. Dasharathy and C. R. Westgate, Solar Energy, 2014,
5		<b>100</b> , 23-30.
6	8.	S. Ahmed, K. B. Reuter, O. Gunawan, L. Guo, L. T. Romankiw and H. Deligianni, Advanced
7		Energy Materials, 2012, <b>2</b> , 253-259.
8	9.	W. Yang, HS. Duan, B. Bob, H. Zhou, B. Lei, CH. Chung, SH. Li, W. W. Hou and Y. Yang,
9		Advanced Materials, 2012, <b>24</b> , 6323-6329.
10	10.	Z. Su, K. Sun, Z. Han, H. Cui, F. Liu, Y. Lai, J. Li, X. Hao, Y. Liu and M. A. Green, Journal of
11		Materials Chemistry A, 2014, <b>2</b> , 500-509.
12	11.	K. Zhang, Z. Su, L. Zhao, C. Yan, F. Liu, H. Cui, X. Hao and Y. Liu, Applied Physics Letters, 2014,
13		104.
14	12.	Y. Wenbing, D. Hsin-Sheng, B. Bob, L. Bao, L. Sheng-Han and Y. Yang, 2012.
15	13.	J. Li, Y. Zhang, H. Wang, L. Wu, J. Wang, W. Liu, Z. Zhou, Q. He and Y. Sun, Solar Energy
16		Materials and Solar Cells, 2015, <b>132</b> , 363-371.
17	14.	F. Liu, Y. Lai, J. Liu, B. Wang, S. Kuang, Z. Zhang, J. Li and Y. Liu, Journal of Alloys and
18		Compounds, 2010, <b>493</b> , 305-308.
19	15.	F. Liu, F. Zeng, N. Song, L. Jiang, Z. Han, Z. Su, C. Yan, X. Wen, X. Hao and Y. Liu, ACS Applied
20		Materials & Interfaces, 2015, 7, 14376-14383.
21	16.	G. Feng, M. Tsuyoshi and W. Takahiro, Japanese Journal of Applied Physics, 2014, 53, 04ER11.
22	17.	W. M. Hlaing Oo, J. L. Johnson, A. Bhatia, E. A. Lund, M. M. Nowell and M. A. Scarpulla,
23		Journal of Elec Materi, 2011, <b>40</b> , 2214-2221.
24	18.	D. B. Mitzi, O. Gunawan, T. K. Todorov, K. Wang and S. Guha, Solar Energy Materials and
25		Solar Cells, 2011, <b>95</b> , 1421-1436.
26	19.	S. Chen, X. G. Gong, A. Walsh and SH. Wei, Applied Physics Letters, 2010, 96, 021902.
27	20.	K. Tanaka, Y. Fukui, N. Moritake and H. Uchiki, Solar Energy Materials and Solar Cells, 2011,
28		<b>95</b> , 838-842.
29	21.	K. Tanaka, Y. Fukui, N. Moritake and H. Uchiki, Solar energy materials and solar cells, 2011,
30		<b>95</b> , 838-842.
31	22.	G. Suresh Babu, Y. B. Kishore Kumar, P. Uday Bhaskar and S. Raja Vanjari, Solar Energy
32		Materials and Solar Cells, 2010, <b>94</b> , 221-226.
33	23.	S. Chen, A. Walsh, XG. Gong and SH. Wei, Advanced Materials, 2013, 25, 1522-1539.
34	24.	A. Fairbrother, X. Fontané, V. Izquierdo - Roca, M. Placidi, D. Sylla, M. Espindola - Rodriguez,
35		S. López - Mariño, F. A. Pulgarín, O. Vigil - Galán and A. Pérez - Rodríguez, Progress in
36		Photovoltaics: Research and Applications, 2014.
37	25.	N. Song, Y. Wang, Y. Hu, Y. Huang, W. Li, S. Huang and X. Hao, Applied Physics Letters, 2014,
38		104.
39	26.	Q. Guo, G. M. Ford, WC. Yang, B. C. Walker, E. A. Stach, H. W. Hillhouse and R. Agrawal,
40		Journal of the American Chemical Society, 2010, <b>132</b> , 17384-17386.
41	27.	K. Sardashti, R. Haight, T. Gokmen, W. Wang, LY. Chang, D. B. Mitzi and A. C. Kummel,
42		Advanced Energy Materials, 2015, DOI: 10.1002/aenm.201402180, 1402180.
43	28.	T. K. Todorov, J. Tang, S. Bag, O. Gunawan, T. Gokmen, Y. Zhu and D. B. Mitzi, Advanced
44		Energy Materials, 2013, <b>3</b> , 34-38.

1	29.	S. Chen, A. Walsh, JH. Yang, X. G. Gong, L. Sun, PX. Yang, JH. Chu and SH. Wei, Physical
2		Review B, 2011, <b>83</b> .
3	30.	W. Ki and H. W. Hillhouse, Advanced Energy Materials, 2011, 1, 732-735.
4	31.	B. Shin, O. Gunawan, Y. Zhu, N. A. Bojarczuk, S. J. Chey and S. Guha, Progress in Photovoltaics:
5		Research and Applications, 2013, <b>21</b> , 72-76.
6	32.	JS. Seol, SY. Lee, JC. Lee, HD. Nam and KH. Kim, Solar Energy Materials and Solar Cells,
7		2003, <b>75</b> , 155-162.
8	33.	Q. Guo, G. M. Ford, WC. Yang, B. C. Walker, E. A. Stach, H. W. Hillhouse and R. Agrawal, J
9		Am Chem Soc, 2010, <b>132</b> , 17384-17386.
10	34.	S. Bag, O. Gunawan, T. Gokmen, Y. Zhu, T. K. Todorov and D. B. Mitzi, Energy &
11		Environmental Science, 2012, <b>5</b> , 7060-7065.
12	35.	W. Wang, M. T. Winkler, O. Gunawan, T. Gokmen, T. K. Todorov, Y. Zhu and D. B. Mitzi,
13		Advanced Energy Materials, 2014, 4, 1301465.
14	36.	P. Chelvanathan, M. I. Hossain, J. Husna, M. Alghoul, K. Sopian and N. Amin, Jpn J Appl Phys,
15		2012, <b>51</b> , 10NC32.
16	37.	KJ. Yang, JH. Sim, B. Jeon, DH. Son, DH. Kim, SJ. Sung, DK. Hwang, S. Song, D. B.
17		Khadka, J. Kim and JK. Kang, Progress in Photovoltaics: Research and Applications, 2014, DOI:
18		10.1002/pip.2500.
19	38.	H. Cui, X. Liu, F. Liu, X. Hao, N. Song and C. Yan, Applied Physics Letters, 2014, 104, 041115.
20	39.	W. Li, J. Chen, H. Cui, F. Liu and X. Hao, Materials Letters, 2014, 130, 87-90.
21	40.	F. Liu, K. Sun, W. Li, C. Yan, H. Cui, L. Jiang, X. Hao and M. A. Green, Applied Physics Letters,
22		2014, <b>104</b> , 051105.
23	41.	F. Zhou, F. Zeng, X. Liu, F. Liu, N. Song, C. Yan, A. Pu, J. Park, K. Sun and X. Hao, ACS Applied
24		Materials & Interfaces, 2015, DOI: 10.1021/acsami.5b05652.
25	42.	G. Altamura, L. Grenet, C. Roger, F. Roux, V. Reita, R. Fillon, H. Fournier, S. Perraud and H.
26		Mariette, Journal of Renewable and Sustainable Energy, 2014, 6, 011401.
27	43.	J. Li, Y. Zhang, W. Zhao, D. Nam, H. Cheong, L. Wu, Z. Zhou and Y. Sun, Advanced Energy
28		Materials, 2015, <b>5</b> , 1402178.
29	44.	J. H. Werner, J. Mattheis and U. Rau, Thin Solid Films, 2005, 480-481, 399-409.
30	45.	F. Bella, M. Imperiyka and A. Ahmad, Journal of Photochemistry and Photobiology A:
31		Chemistry, 2014, <b>289</b> , 73-80.
32	46.	H. Sun and R. Dixon, Journal of Power Sources, 2014, 272, 404-414.
33		

### **Figure Captions**

Fig. 1. Top view SEM images of the CZTSSe thin films annealed under Ar-Se atmosphere at four temperatures of (a) 500 °C, (b) 530 °C, (c) 560 °C and (d) 580 °C for 40 min. Inset: SEM cross-sectional images of the samples. The temperature of Se was always 400 °C under these conditions. The composition of precursor is designed to be Cu/(Zn+Sn) = 0.86 and Zn/Sn = 1.0.

Fig. 2. (a) XRD patterns of CZTSSe annealed at 500 °C, 530 °C, 560 °C and 580 °C for 40 min on molybdenum (Mo). (b) The corresponding Raman spectra of the CZTSSe thin films.

Fig. 3. SEM images of the CZTSSe thin films annealed under Ar-Se atmosphere at 560 °C for (a) 5 min, (b) 20 min, (c) 40 min, and (d) 60 min.

Fig. 4. (a) XRD patterns of CZTSSe annealed at 560 °C for 5 min, 20 min, 40 min and 60 min onMo. (b) The corresponding Raman spectra of the CZTSSe thin films.

Fig. 5. SEM images of the surface of the CZTSSe thin films. The Cu/(Zn+Sn) ratios of precursor solutions were (a) 0.66, (b) 0.76, (c) 0.86 and (d) 0.96, respectively. These four films were annealed under selenium atmosphere at 560 °C for 40 min.

Fig. 6. Four SEM images of CZTSSe films with Zn/Sn ratios of (a) 0.8, (b) 1.0, (c) 1.2, and (d) 1.4. The selenization temperature of 560 °C, the selenization time of 40 min, and the Cu/(Zn+Sn) ratio of 0.86 were utilised in the synthesis.

Fig. 7. (a) I–V characteristics of the 8.08% CZTS solar cell under dark (black solid line) and 1 sun illumination (red dotted line). (b) The quantum efficiency (QE) curve. The inset shows the plot of  $[EQE\times E]^2$  vs. E to evaluate the band gap.

Fig. 8. Cross-sectional STEM image of the 8.08% CZTSSe device.

Fig. 9. EDX analysis on the 8.08% CZTSSe device: (a) across the thickness of the device and (b) laterally across the CZTSSe layer. The red arrow in the inset demonstrats the direction of the EDX scanning. Quantitative analysis for S and Mo in the thin film is not feasible due to the similar spectral peak in the EDX analysis. Scanning range from 2900 to 3800 nm in (a) is considered as Mo layer, in accordance with the sputtering thickness of Mo.



242x210mm (300 x 300 DPI)



164x64mm (300 x 300 DPI)



242x210mm (300 x 300 DPI)



164x64mm (300 x 300 DPI)



242x210mm (300 x 300 DPI)



242x210mm (300 x 300 DPI)



157x56mm (300 x 300 DPI)



258x267mm (300 x 300 DPI)



310x436mm (300 x 300 DPI)



### graphic



CZTSSe thin film solar cell was fabraicated by a sol-gel method with efficiency of 8.08 %.