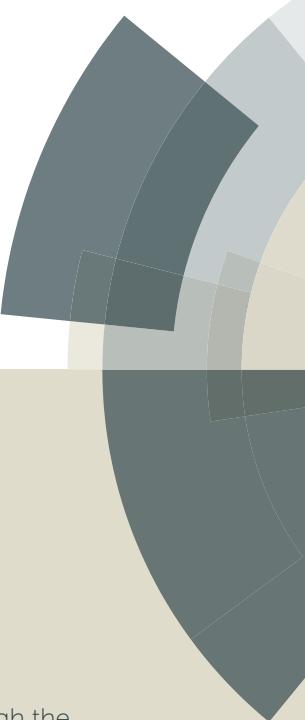


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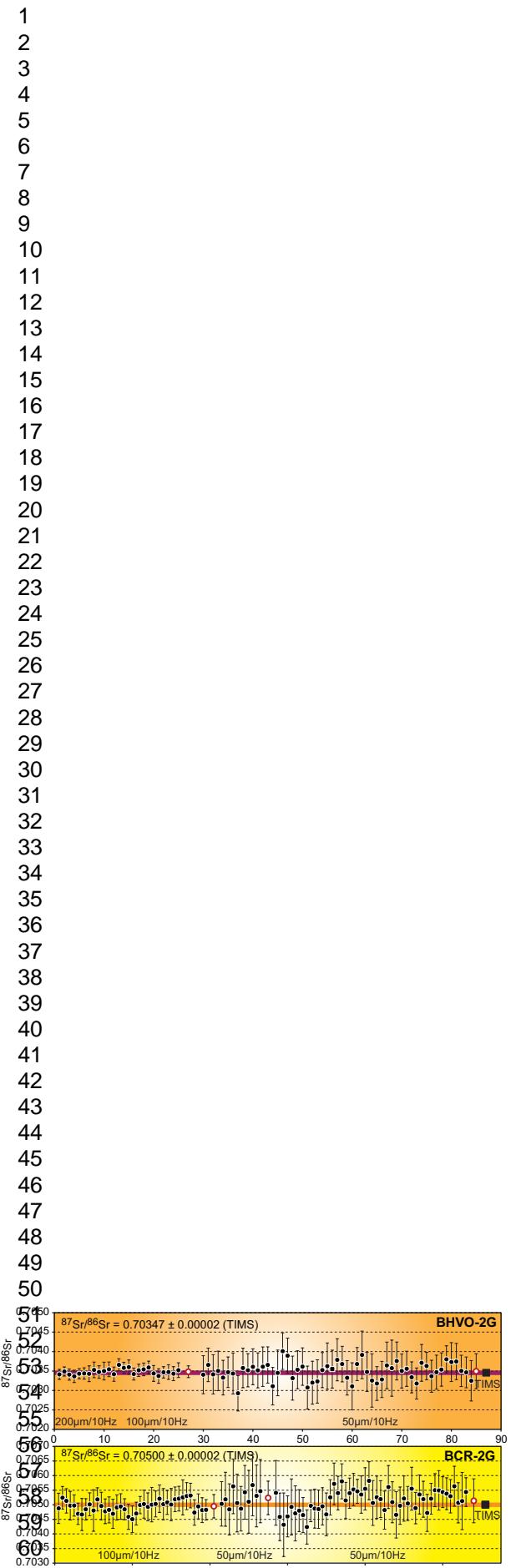
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Journal of Analytical Atomic Spectrometry Accepted Manuscript



## Table of contents entry.

An analytical protocol using LA-MC-ICP-MS for precise Sr isotope ratio analysis from small glass samples is presented.

**TECHNICAL NOTE****In situ Sr isotope measurement of small glass samples using multiple-Faraday collector inductively coupled plasma mass spectrometry with  $10^{12}\text{-}\Omega$  resistor high gain Faraday amplifiers†**

Qing Chang<sup>\*a</sup>, Jun-Ichi Kimura<sup>a</sup>, and Bogdan Stefanov Vaglarov<sup>a</sup>

An analytical protocol was developed for correcting Kr baseline-induced bias and Rb isobaric overlap factors to analyse Sr isotope ratios for small glass samples using excimer laser ablation (LA) with an Aridus II desolvating nebuliser dual-intake system and multiple collector inductively coupled plasma mass spectrometry (MC-ICP-MS). Combined use of a low-oxide interface setup, along with high-gain Faraday amplifiers with a  $10^{12}\text{-}\Omega$  resistor, enabled precise determination of Sr isotope ratios from 50- and 100- $\mu\text{m}$  diameter craters at 10-Hz laser repetition rate. Residual analytical biases of  $^{84}\text{Sr}/^{86}\text{Sr}$  and  $^{87}\text{Sr}/^{86}\text{Sr}$  isotope ratios, obtained from Kr baseline suppressions (Kimura et al., 2013, Journal of Analytical Atomic Spectrometry, 28, 945–957), were found to be nonlinear, but the correction method was applicable to 50–200- $\mu\text{m}/10\text{-Hz}$  craters. We also found that the  $^{85}\text{Rb}/^{87}\text{Rb}$  overlap correction factor changed with time with a change in the surface condition of the sampler–skimmer cones. The correction factor of  $^{85}\text{Rb}/^{87}\text{Rb}$  was thus determined at least once per five unknown measurements using the Aridus solution intake line. We determined  $^{87}\text{Sr}/^{86}\text{Sr}$  isotope ratios from MkAn anorthite ( $\text{Sr} = 305 \text{ ppm}$ ,  $\text{Rb} = 0.07 \text{ ppm}$ ), BHVO-2G, KL2-G, ML3B-G ( $\text{Sr} = 312\text{--}396 \text{ ppm}$ ,  $\text{Rb} = 5.8\text{--}9.2 \text{ ppm}$ ), and BCR-2G ( $\text{Sr} = 337 \text{ ppm}$ ,  $\text{Rb} = 48.5 \text{ ppm}$ ) basalt glasses using a 50–100- $\mu\text{m}/10\text{-Hz}$  crater. The results agree well with their reference values, determined by

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6     25 thermal ionisation mass spectrometry, even with the high Rb/Sr ratio (0.14) in the  
7     26 BCR-2G glass. The internal/intermediate precisions were  $\pm 0.0002$  (two-standard  
8     27 deviation: 2SD) for 100- $\mu\text{m}$  craters and  $\pm 0.0005$  for 50- $\mu\text{m}$  craters. The new instrument  
9     28 settings and analytical protocol improved the precision by a factor of two compared to  
10     29 the previous report using LA-(sector field)-ICP-MS and enables the analysis of sample  
11     30 volumes that are ten times smaller than those used in previous LA-MC-ICP-MS  
12     31 analyses with equal precision.

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14     32 (298 words in abstract, 6836 words in all document)

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26     39 † Electronic supplementary information (ESI) available: Data Table 1: Analytical results  
27     40 of Sr isotope ratios. See DOI:10.1039/c4jaxxxxxg

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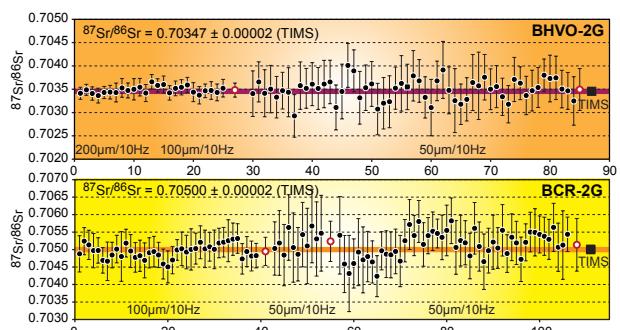


Table of contents entry.

An analytical protocol using LA-MC-ICP-MS for precise Sr isotope ratio analysis from small glass samples is presented.

## 1           43    1. Introduction

### 2           44    1.1. In situ Sr isotope ratio analysis by LA-ICP-MS

3           45    Strontium isotope analysis using in situ laser ablation multiple collector inductively  
4           46    coupled plasma mass spectrometry (LA-MC-ICP-MS) has been a useful geochemical  
5           47    tool for the analysis of plagioclase crystals or carbonate materials and glasses.<sup>1-6</sup>  
6           48    Complex isobaric overlaps from Ar gas blanks and isobaric ions, oxides, and doubly  
7           49    charged ions from the samples are of concern for four naturally occurring Sr isotopes,  
8           50    <sup>84</sup>Sr, <sup>86</sup>Sr, <sup>87</sup>Sr, and <sup>88</sup>Sr. Data correction protocols are therefore very complex,  
9           51    prompting rigorous examination. These problems and their resulting solutions have  
10          52    been summarized by Vroon et al.<sup>7</sup>

11          53    The most commonly employed Sr isotope measurement protocol<sup>7</sup> for the  
12          54    geochemical samples is the use of on-peak background subtraction, which corrects for  
13          55    Kr gas blanks and instrumental memories of Sr and Rb,<sup>1, 2, 5, 6, 8-13</sup> although iterative  
14          56    subtraction of the Kr baselines is also a useful technique.<sup>3, 13</sup> Overlap correction of <sup>87</sup>Rb  
15          57    using <sup>85</sup>Rb is achieved by an external mass bias correction from Sr, determined as  
16          58    <sup>86</sup>Sr/<sup>88</sup>Sr = 0.1194.<sup>1, 2, 5, 6, 8-12</sup> but is successful only with the actual determination of  
17          59    <sup>85</sup>Rb/<sup>87</sup>Rb, either by measurement of a Rb-doped Sr standard solution, using a  
18          60    solution-laser aerosol-dual intake system,<sup>14</sup> or by measurement of both natural and  
19          61    synthetic glasses with known <sup>87</sup>Sr/<sup>86</sup>Sr isotope ratios.<sup>2, 3, 15</sup> Interferences from Ca dimers,  
20          62    CaAr, and <sup>23</sup>Na/<sup>60-64</sup>Ni molecular ions are usually negligible for plagioclase crystals and  
21          63    carbonates.<sup>6, 7, 10</sup> Oxide interferences on <sup>85</sup>Rb and <sup>87</sup>Sr, from <sup>69</sup>Ga/<sup>16</sup>O and <sup>71</sup>Ga/<sup>16</sup>O are  
22          64    significant for samples rich in Al, such as plagioclase and glass, as Ga has a chemical  
23          65    affinity for Al and is highly partitioned in these samples.<sup>14</sup> Therefore, by using particular  
24          66    MC-ICP-MS settings,<sup>16</sup> high oxide yields are avoided.<sup>6, 10</sup> Interference from doubly

charged rare earth element (REE) ions are also obvious,<sup>10, 12, 17</sup> and care must be taken to exclude REE-rich Sr-poor samples (e.g., clinopyroxene) and co-ablation of REE-rich mineral inclusions (e.g., apatite). The REE double charges such as  $^{166}\text{Er}^{++}$ ,  $^{168}\text{Er}^{++}$ , and  $^{170}\text{Er}^{++}$  for  $^{83}\text{Kr}$ ,  $^{84}\text{Sr}$ , and  $^{85}\text{Rb}$ , respectively, and  $^{170}\text{Yb}^{++}$ ,  $^{172}\text{Yb}^{++}$ ,  $^{174}\text{Yb}^{++}$  and  $^{176}\text{Yb}^{++}$  for  $^{85}\text{Rb}$ ,  $^{86}\text{Sr}$ ,  $^{87}\text{Sr}$  and  $^{88}\text{Sr}$ , respectively, can be monitored, or even corrected for, by placing the Faraday collectors at half-mass positions of  $^{83.5}\text{M}$  ( $^{167}\text{Er}^{++}$ ) and  $^{86.5}\text{M}$  ( $^{173}\text{Yb}^{++}$ ).<sup>10, 17</sup> Finally, analytical results are obtained by the internal mass fractionation correction of Sr given that  $^{86}\text{Sr}/^{88}\text{Sr} = 0.1194$ .<sup>7</sup> Additional bias corrections for Kr baseline suppression caused by LA aerosols were recently introduced for  $^{87}\text{Sr}/^{86}\text{Sr}$  measurements to further improve analytical precision and reproducibility for samples low in Sr (e.g., 300 ppm), for which large amounts of LA aerosols needed to be introduced.<sup>14</sup>

## 1.2. Analysis of Sr isotope ratios in glass samples

Analysis of the  $^{87}\text{Sr}/^{86}\text{Sr}$  isotope ratio from a small glass sample, such as olivine-hosted small glass inclusions (diameters of 50–200  $\mu\text{m}$  or less), is particularly useful for the identification of the source processes in the genesis of basalt magmas. These glass inclusions can preserve information regarding the primary magma generated in the deep mantle and is unaffected by chemical modifications during ascent to the surface by encapsulation of tiny drops of the deep, quenched magmas.<sup>3, 18</sup> Attempts to analyse a basalt glass were described by Christensen et al. in the first paper published on in situ  $^{87}\text{Sr}/^{86}\text{Sr}$  isotope analysis using LA-MC-ICP-MS, in which they analysed the groundmass glass of Long Valley basalt from a 300- $\mu\text{m}/20\text{-Hz}$  crater (approximate sample volume:  $V = \sim 2.1 \times 10^{-5} \text{ cm}^3$ ), sampled by an infrared (IR) Nd-YAG laser with

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6 91 an internal/intermediate precision of  $^{87}\text{Sr}/^{86}\text{Sr}$  = approximately  $\pm 0.0005$  (two-standard  
7 deviation: 2SD).<sup>1</sup>  
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11 93 The attempt continued with development in an ultraviolet 213-nm YAG laser  
12 and a third-generation MC-ICP-MS, allowing analysis from a smaller crater of  
13 94 120- $\mu\text{m}$ /20-Hz ( $V = \sim 7.9 \times 10^{-7} \text{ cm}^3$ ) with a similar internal/intermediate precision of  
14 95 approximately  $\pm 0.0005$  2SD.<sup>3</sup> An alternative approach was to use a sector field  
15 96 (SF)-ICP-MS with rapid peak jumping using a single ion counter (IC) connected to a  
16 97 193-nm excimer laser (ExLA). This allowed analysis of a smaller sample volume, a  
17 98 crater from 50  $\mu\text{m}$ /10 Hz ( $V = \sim 9.8 \times 10^{-8} \text{ cm}^3$ ), with an intermediate precision of  
18 99 approximately  $\pm 0.0010$  2SD.<sup>19</sup> Analysis of the small sample volume was achieved  
19 100 through use of the high-sensitivity IC, whereas use of the IC with peak jumping resulted  
20 101 in poor intermediate precision.  
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### 1.3.Focus

In this study, we present an improved analytical technique for Sr isotope analysis of glass samples. We apply a 193-nm ExLA laser sampling method, with simultaneous introduction of Aridus II solution aerosols (dual intake system),<sup>14</sup> connected to an MC-ICP-MS. In order to prevent any potential oxide interferences, the interface of the MC-ICP-MS was set to achieve the lowest oxide molecular yield.<sup>20</sup> Low instrumental sensitivity with the interface was accounted for by the use of a high-efficiency rotary pump at the interface<sup>14, 20-22</sup> and the application of high-gain Faraday amplifiers using a  $10^{12}\text{-}\Omega$  resistor.

Use of these amplifiers increases sensitivity by 10 times by reducing the signal-to-noise ratio to half of what would be obtained from an amplifier with a  $10^{11}\text{-}\Omega$

resistor. This can in turn reduce analytical error generated during the analysis by more than half for low electron currents.<sup>23, 24</sup> We also examine the Kr baseline-induced bias<sup>14</sup> during loading of smaller samples to the ICP and temporal changes in the Rb overlap correction factor for glass samples with high Rb content. The small sample size and the internal/intermediate precision achieved in this study are keys to the successful application of LA-MC-ICP-MS to small glass inclusions.

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## 2. Instrument setting, data acquisition/reduction

### 2.1.Laser ablation (LA) and dual-intake sample introduction

The instrumentation for our LA-multiple correction (MC)-ICP-MS has been reported elsewhere for the in situ Sr isotope analysis of plagioclase crystals<sup>14, 15, 25</sup> In brief, the instrument consists of an LA-Aridus II dual-intake system including a 193-nm excimer LA system with a COMPex 102 laser source (Coherent, Göttingen, Germany) combined with an Aridus II desolvating nebuliser (CETAC Technologies, Omaha, USA). The laser aerosols in the He gas flow and the Aridus II solution aerosols in the Ar gas flow were mixed in a T-piece connector just prior to the ICP torch in the MC-ICP-MS. The excimer laser operated with a fluence of  $20 \text{ J cm}^{-2}$ , a 50–200-μm crater diameter, and, in most cases, a 10-Hz repetition rate, which resulted in craters with depths of ~50 μm after ablation for ~35 s (idle time, 10 s; data acquisition, 25 s), indicating a drilling rate of 0.125-μm per shot.

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### 2.2.MC-ICP-MS

A NEPTUNE MC-ICP-MS was used (Thermo Fisher Scientific, Bremen, Germany) and was modifications at the interface for a high sensitivity as has been reported

elsewhere.<sup>14, 20</sup> The settings of the ion interface were normal (N) Ni sampling and normal (H) Ni skimmer cones with the guard electrode (GE) turned off (electrically disconnected) (**Table 1**). This allowed minimisation of the oxide molecular ion yields ( $^{238}\text{U}^{16}\text{O}^+/\text{U}^+ < 1\%$ , analysed in solution by Aridus II) and thus, any oxide based interference (e.g.,  $\text{GaO}^+$ ) was avoided. The important doubly charged REE ions were monitored by half mass Faradays set at  $^{83.5}\text{M}$  for  $^{167}\text{Er}^{++}$  and  $^{86.5}\text{M}$  for  $^{173}\text{M}^{++}$  (**Table 1**),<sup>20</sup> and these were almost completely absent (grand averages of the per cent  $\text{REE}^{++}$  on the most affected  $^{84}\text{Sr}$  signals were  $^{168}\text{Er}^{++}/^{84}\text{Sr} < 0\%$  and  $^{168}\text{Yb}^{++}/^{84}\text{Sr} = 0.004\%$ ; calculations from **ESI Data Table 1†**) for glass samples.

With the interface designed to maintain low oxide levels, as noted above, the system sensitivity was maintained at 200 V/ ppm Sr (~11 Gcps/ ppm) measured from the Aridus II solution.<sup>3, 10</sup> Analysis of ~300–400 ppm of Sr in glasses yielded a  $^{88}\text{Sr}$  signal of 0.5–0.8 V and ~0.1–0.2 V from 100- $\mu\text{m}/10\text{-Hz}$  and 50- $\mu\text{m}/10\text{-Hz}$  laser craters, respectively (**ESI Data Table 1†**). These values were sufficient for an internal analytical precision of  $^{87}\text{Sr}/^{88}\text{Sr} = \pm 0.0002$  to ~0.0010 (2SE) after ~25 s of data acquisition. This was achieved by the application of high-gain Faraday amplifiers, equipped with  $10^{12}\Omega$  resistors,<sup>23</sup> to all naturally occurring Sr isotopes ( $^{84}\text{Sr}$ ,  $^{86}\text{Sr}$ ,  $^{87}\text{Sr}$ , and  $^{88}\text{Sr}$ ) and  $^{85}\text{Rb}$  (**Table 1**).

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### 2.3. Laser ablation protocol

We especially focused on optimizing the LA analytical protocol for the use of  $10^{12}\Omega$  high-gain amplifiers because of their slow response to transient signals.<sup>23, 26</sup> One cycle of data acquisition consisted of: (1) two 30-s baseline measurements made on-peak before laser firing; (2) subsequent firing of the laser for 10 s without data acquisition,

allowing stabilisation of the LA signal and the current applied to the Faraday amplifiers in order to avoid the acquisition of transient signals during LA signal growth; and (3) data acquisition for ~25 s using a time slice of ~0.524 s for 50 scans, completed while the signal output was flat and stable (**Fig. 1**). The detailed setup parameters of the laser system and the MC-ICP-MS, including the Faraday collector setup in the mass spectrometer, are listed in **Table 1**.

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#### 2.4. Correction of Rb isobaric overlap

Correction of interference from the  $^{87}\text{Rb}$  isobar on  $^{87}\text{Sr}$  is particularly important in the precise analysis of glasses because natural basaltic glasses contain Rb in the range 0.02–0.15 or more times the amount of Sr.<sup>19, 27</sup> Previous studies used both synthetic<sup>2</sup> or natural glasses<sup>3, 19</sup> of known contents of Rb and Sr and of known  $^{87}\text{Sr}/^{86}\text{Sr}$  isotope ratio. This allows determination of the overlap correction factor of Rb by adjusting the  $^{85}\text{Rb}/^{87}\text{Rb}$  isotope ratios to reproduce the known  $^{87}\text{Sr}/^{86}\text{Sr}$  isotope ratio of the glasses.

We have been using a different approach for determination of the Rb overlap correction factor, using a LA–Aridus solution aerosol dual-intake system. This system has been applied to Hf, Nd, and Sr isotope analyses of zircon,<sup>28</sup> monazite/apatite and rutile,<sup>20</sup> and plagioclase,<sup>14</sup> respectively. The advantage of this method is its ability to simultaneous fine-tune the MC-ICP-MS instrument for both the solution and the LA aerosols from the dual-intake lines, by using solution aerosols while LA firing is off and the laser cell gas is on. Precise determination of an overlap correction factor (e.g.,  $^{176}\text{Yb}$  and  $^{176}\text{Lu}$  on  $^{176}\text{Hf}$ ,<sup>28</sup>  $^{144}\text{Sm}$  on  $^{144}\text{Nd}$ ,<sup>20</sup> and  $^{87}\text{Rb}$  on  $^{87}\text{Sr}$ )<sup>14</sup>) is simple by using mixed solutions with known (not necessarily accurate) concentrations of the interfering-interfered elements and known isotope composition.

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6       187       The tuning solution can easily be prepared by mixing standard solutions. The  
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8       188       determined correction factor can be used for the LA aerosols while Aridus II introduces  
9       189       deionised water (DIW). The matrix effect of the concomitant major elements from LA  
10      190       aerosols has been examined by major element-doped solutions<sup>28</sup> or by comparisons to  
11      191       the known TIMS values of the LA analytes.<sup>14, 20, 28</sup> The bias was found to be negligible  
12      192       within the level of intermediate precision of  $\pm 0.00002$  for Hf ( $\sim 90$ -ppm offset in  
13      193       naturally occurring  $^{176}\text{Hf}/^{177}\text{Hf}$ ), Nd ( $\sim 40$ -ppm offset in  $^{143}\text{Nd}/^{144}\text{Nd}$ ), and Sr ( $\sim 30$ -ppm  
14      194       offset in  $^{87}\text{Sr}/^{86}\text{Sr}$ ) isotope analyses.<sup>14, 20, 28</sup> It should be noted that the majority of these  
15      195       analyses were mostly done with a crater size smaller than  $200\ \mu\text{m}/20\ \text{Hz}$  and with the  
16      196       same instruments used in this study. The effectiveness and stability of Rb overlap  
17      197       correction are discussed in **Section 3.2**, particularly for glass samples high in Rb  
18      198       content.

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200     **2.5.Data reduction**

201     First, the peak intensities of  $^{84}\text{Sr}$ ,  $^{86}\text{Sr}$ ,  $^{87}\text{Sr}$ ,  $^{88}\text{Sr}$ , and  $^{85}\text{Rb}$  were measured using on-peak  
202     background subtraction.  $^{87}\text{Sr}$  intensity was corrected for  $^{87}\text{Rb}$  overlap by using the  
203     pre-determined  $^{85}\text{Rb}/^{87}\text{Rb}$  isotope ratio (**Section 2.4**) and the measured  $^{85}\text{Rb}$  intensity.  
204      $^{84}\text{Sr}/^{86}\text{Sr}$  and  $^{87}\text{Sr}/^{86}\text{Sr}$  isotope ratios were calculated from the intensities of  $^{84}\text{Sr}$ ,  $^{86}\text{Sr}$  and  
205      $^{87}\text{Sr}$  (employing the correction for  $^{87}\text{Rb}$  overlap), while  $^{88}\text{Sr}$  was determined by the  
206     exponential law given that  $^{86}\text{Sr}/^{88}\text{Sr} = 0.1194$ . Further Kr baseline bias was corrected by  
207     using  $^{84}\text{Sr}/^{86}\text{Sr}$  and  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios with the following equation:

208     
$$^{87}\text{Sr}/^{86}\text{Sr}_{\text{corrected}} = ^{87}\text{Sr}/^{86}\text{Sr}_{\text{original}} / (-0.14121 \times ^{84}\text{Sr}/^{86}\text{Sr} + 1.0081)$$

209     using the empirical correction coefficients proposed by Kimura et al.<sup>14</sup> The above  
210     calculations were applied off-line to all 50 scans in a given run. An average and twice

the standard error (2SE) were calculated for each spot; in this case, all the errors that propagated from each isotope measurement were included in the calculated 2SE, while errors from the Rb overlap correction factor and the coefficients used for Kr baseline bias were not included because of their minimal effect.

Kr bias correction itself does not improve internal or external precision caused by the short or long term drift in Kr baselines, a feature that is described elsewhere for iterative Kr baseline subtraction.<sup>13</sup> The major role of this technique is to account for Kr baseline suppression due to the mass loading of LA aerosols; the improvement in this bias can be as high 250 ppm, which is far greater than the Kr baseline drift.<sup>14</sup> The effect of the Kr baseline bias with regards to this study is discussed in **Section 3.1**.

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### 3. Results

Although the analytical protocol for Sr isotope measurements using LA-MC-ICP-MS has been rigorously examined and researchers have been largely in consensus regarding the analytical protocol, special care must nevertheless be taken with Kr baseline suppressions and the Rb overlap correction. We identified new features of these factors, as described below.

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#### 3.1.Kr baseline suppression with varied sample loading

##### 3.1.1. Mass loading and Kr baselines: a basic test

We tested the effect of mass loading of LA aerosols by analysing an ~5 ppb SRM 987 Sr standard solution, along with introduction of LA aerosols at different laser crater sizes obtained from an olivine crystal in a picrite lava from St. Helena Island.<sup>29</sup> The SRM 987 is free from any interference from concomitant elements, while Sr and Rb signals from

the tested olivine were both below the detection limits. Kr outgassing during LA was reported from olivine in the JP-1 peridotite<sup>14</sup> but this was not detected from the olivine in the picrite. The olivine sample is therefore suitable for mass loading tests.

The <sup>88</sup>Sr intensity from ~5-ppb SRM 987 was ~0.5 V, which was approximately the same signal size as that from a 100-μm/10-Hz crater on a plagioclase/glass containing ~300-ppm Sr. The SRM 987 solution was first analysed for <sup>87</sup>Sr/<sup>86</sup>Sr without LA firing. Subsequent analyses used 50-μm/10-Hz (rate of LA aerosol loading:  $V_t = 2.5 \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$ ), 100-μm/10-Hz ( $V_t = 9.8 \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$ ), 200-μm/10-Hz ( $V_t = 3.9 \times 10^{-8} \text{ cm}^3 \text{ s}^{-1}$ ), and 200-μm/20-Hz ( $V_t = 7.8 \times 10^{-8} \text{ cm}^3 \text{ s}^{-1}$ ) craters with increasing loaded mass from the olivine aerosols. The Sr signals vary with different LA aerosol loading when analysing a Sr-bearing material with a different crater size; therefore, the absolute role of Kr baseline suppression cannot be tested, particularly for small aerosol loading due to large internal analytical errors in <sup>87</sup>Sr/<sup>86</sup>Sr (see **Section 3.1.2**).

Analytical results are given in **Fig. 2**. The results show a nonlinear relationship, indicating a small suppression of the Kr baseline with 50–100-μm/10-Hz craters, whereas an obvious increase in the uncorrected Sr isotope ratio occurs in 200-μm/10–20-Hz craters. This confirmed the necessity for correction of the Kr baseline suppression. The first report of this suppression was on a 200-μm/5–10-Hz crater for plagioclase analysis,<sup>14</sup> the necessity of which may have been questioned, but our new analysis clearly demonstrates the necessity of the correction when the sample loading of LA is greater than  $3.9 \times 10^{-8} \text{ cm}^3 \text{ s}^{-1}$ . Even with a low mass loading, corrected data are closer to the reference value of SRM 987, indicating the importance of this bias correction (**Fig. 2**).

The baseline bias correction with the heaviest sample loading at  $7.8 \times 10^{-8} \text{ cm}^3$

6       259  $\text{s}^{-1}$  from a 200- $\mu\text{m}/20\text{-Hz}$  crater appears to be invalid, showing a large offset in the  
7       260 corrected Sr isotope ratio (**Fig. 2**). In this particular analysis, we saw  $\sim 10\%$  suppression  
8       261 of the  $^{88}\text{Sr}$  signal for the SRM 987 solution during loading of the olivine LA aerosols,  
9       262 suggesting suppression of Sr in the ICP (not shown). This was not observed with the  
10      263 smaller sample loading. The nonlinear response of the Sr signal might originate from  
11      264 the nonlinear response described by Saha's equation<sup>30, 31</sup> where an increase in the  
12      265 electron number density causes a sigmoidal decrease of the ionisation efficiency.<sup>21, 30-32</sup>  
13      266 Sr signal suppression relative to Kr was not taken into account for the Kr baseline bias  
14      267 correction in the previous paper;<sup>14</sup> therefore, the proposed correction could apply only  
15      268 to the LA aerosol loading smaller than  $7.8 \times 10^{-8} \text{ cm}^3 \text{ s}^{-1}$  with the 200- $\mu\text{m}/20\text{-Hz}$  crater  
16      269 used in this study. The effect of signal suppression is discussed further in **Section 3.2.1**.  
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### 3.1.2. Effect of Kr baseline suppression in plagioclase analysis

We conducted a series of experiments using an extremely homogeneous anorthite plagioclase MkAn.<sup>14, 25</sup> The Sr isotope composition of the crystal is homogeneous,  $^{87}\text{Sr}/^{86}\text{Sr} = 0.70345 \pm 0.00002$  (error in 2SD,  $n = 4$ ) confirmed by TIMS and LA-MC-ICP-MS.<sup>14, 25</sup> This crystal contains minuscule amounts of Rb (0.07 ppm), with Sr = 305 ppm<sup>21</sup> and thus it is the best suited for examination of the Kr baseline suppression without considering the effect of Rb overlap.

Using the MkAn, analyses of Sr isotope ratios were performed with ablation conditions of 200  $\mu\text{m}/20\text{ Hz}$ , 200  $\mu\text{m}/10\text{ Hz}$ , 100  $\mu\text{m}/10\text{ Hz}$ , and 50  $\mu\text{m}/10\text{ Hz}$  (**Fig. 3a** and **ESI Data Table 1†**). We rearranged the data using the Kr baseline bias uncorrected data of  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{84}\text{Sr}/^{86}\text{Sr}$  based on the crater size (**Fig. 3b**). The data from the 200- $\mu\text{m}/10\text{--}20\text{-Hz}$  plot on the same negatively sloped linear regression line, which is

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6 283 sub-parallel to the theoretical fractionation line of Kr baseline suppression proposed by  
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8 284 the previous study.<sup>14</sup> The amount of LA aerosols from a 200- $\mu\text{m}$ /20-Hz crater is about  
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10 285 0.7 times smaller than that from olivine because of the poorer coupling of 193-nm  
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12 286 excimer laser light to the transparent (Fe- and Mg-poor) plagioclase. This would be the  
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14 287 reason for the lack of offset from the Kr baseline bias for this sample amount. In  
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16 288 contrast, the regression lines of the 100  $\mu\text{m}$ /10 Hz and 50  $\mu\text{m}$ /10 Hz data are  
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18 289 sub-parallel, but with shallower slopes than those from 200- $\mu\text{m}$  craters. Kr baseline  
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20 290 corrections, made by the theoretical model of Kimura et al.,<sup>14</sup> resulted in an almost flat  
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22 291 regression line for the data from 200- $\mu\text{m}$  craters (**Fig. 3c**), indicating almost perfect  
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24 292 correction of the bias. The same correction made for the data from 100–50- $\mu\text{m}$  craters  
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26 293 showed slight over-reductions as shown by the positively sloped regression lines in **Fig.**  
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28 294 **3c.**

295 This examination clearly indicates that the Kr baseline suppression is due to  
296 mass loading of the plagioclase aerosols, but the degree of suppression could be  
297 nonlinear against the amount of aerosol loaded. The mass loading from 200- $\mu\text{m}$  craters  
298 clearly showed the bias, whereas 50–100- $\mu\text{m}$  craters showed the same bias, but to an  
299 apparently smaller extent. The systematic errors from the potential nonlinearity,  
300 deduced from the slopes of the regression lines of 50–100- $\mu\text{m}$  craters, are well within  
301 the typical internal precision of  $\pm 0.0002$  for these small craters both before and after the  
302 bias corrections (**Fig. 3b** and **3c**). Therefore, we corrected for the baseline bias even for  
303 the small 50–100- $\mu\text{m}$  craters. As shown, the  $^{87}\text{Sr}/^{86}\text{Sr}$  form SRM 987 solution showed a  
304 closer match to the reference value after the bias correction for a small aerosol loading  
305 (**Fig. 2**), which also supports our approach. The analytical results of MkAn are given in  
306 **Fig. 3a** and **ESI Data Table 1†**, and show an almost perfect match with the isotope

ratios measured by TIMS ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.70345 \pm 0.00002$ ) for all sample sizes;  $0.7035 \pm 0.0002$ , 2SD,  $n = 26$ , 200- $\mu\text{m}$  craters;  $0.7034 \pm 0.0002$ , 2SD,  $n = 10$ , 100- $\mu\text{m}$  craters; and  $0.7035 \pm 0.0008$ , 2SD,  $n = 40$ , 50- $\mu\text{m}$  craters.

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### 311 **3.2. Rb overlap correction factor and its temporal drift**

#### 312 **3.2.1. Determination of the Rb correction factor**

313 We determined the  $^{85}\text{Rb}/^{87}\text{Rb}$  correction factor using the dual-intake system over the  
314 period of analyses of this study and found a temporal change of the value obtained (**Fig.**  
315 **4a** and **4b**). The correction factor was obtained by analysing an ~5-ppb NIST SRM 987  
316 standard solution ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.71024$ )<sup>27</sup> doped with an ~1-ppb Rb, in which Rb/Sr =  
317 ~0.2 is higher than all target glasses. The acquisition was done by the same protocol as  
318 for the LA analysis but with a longer time slice of ~4.25 s for 50 scans in order to obtain  
319 high-precision results. In this measurement, instrumental mass fractionation of Sr was  
320 determined by a stable isotope ratio of  $^{86}\text{Sr}/^{88}\text{Sr} = 0.1194$ .<sup>33</sup> The mass fractionation was  
321 corrected by the exponential law for both Sr and Rb isotope ratios, assuming the same  
322 mass fractionation factor over the mass range  $^{84}\text{M}$ – $^{88}\text{M}$ . This calculation protocol is the  
323 same as that used for the LA analyses.

324 We determined the  $^{87}\text{Sr}/^{86}\text{Sr} = S_1$  ratio of the mixed solution with a provisional  
325  $^{85}\text{Rb}/^{87}\text{Rb}$  value (e.g., 2.5926, the canonical value of  $^{85}\text{Rb}/^{87}\text{Rb}$  from the IUPAC isotopic  
326 abundances of  $^{85}\text{Rb} = 72.165$  and  $^{87}\text{Rb} = 27.835$ ).<sup>33</sup> The obtained  $^{87}\text{Sr}/^{86}\text{Sr}$  was always  
327 incorrect because of the difference between the proper and the given  $^{85}\text{Rb}/^{87}\text{Rb}$  ratio,  
328 due to the different elemental sensitivity between Rb and Sr, which altered the ratio of  
329  $^{87}\text{Rb}/^{87}\text{Sr}$  (see the similar example between Sm and Nd).<sup>20, 34</sup> The second cycle of the  
330 calculation was done using the same analytical data with a different  $^{85}\text{Rb}/^{87}\text{Rb}$  value

(e.g., 2.470) to obtain the second  $^{87}\text{Sr}/^{86}\text{Sr} = S_2$  ratio. The correct  $^{85}\text{Rb}/^{87}\text{Rb} = R_{\text{Rb}}$  ratio should be on the linear regression line passing through the coordinates (X, Y) = ( $S_1$ , 2.5926) and ( $S_2$ , 2.470), obtaining  $R_{\text{Rb}}$  at (0.70124,  $R_{\text{Rb}}$ ) for the Rb doped SRM 987 standard solution.  $R_{\text{Rb}}$  does not represent its natural value but represents the value including the correction factor (see the similar example between Lu, Yb, and Hf).<sup>28</sup>

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### 337 3.2.2. Possible origin of enhanced Rb correction factor

338 The averaged  $R_{\text{Rb}} = ^{85}\text{Rb}/^{87}\text{Rb}$  value on Day 5 was 2.4688 (Fig. 4b), expressed as a  
339 mass fractionation factor of  $\Delta M(\text{Rb}) = 2.417\%$  in comparison to the canonical value of  
340  $^{85}\text{Rb}/^{87}\text{Rb} = 2.5926$ .<sup>33</sup> This fractionation factor is systematically larger than that  
341 measured for  $^{86}\text{Sr}/^{88}\text{Sr}$  which gave  $\Delta M(\text{Sr}) = 2.317\%$  assuming  $^{86}\text{Sr}/^{88}\text{Sr} = 0.1194$  (Fig.  
342 4c). This is due to the difference in elemental sensitivity between Rb and Sr as noted in  
343 Section 3.2.1, showing a 4.35% enhancement of Rb relative to Sr. Note that the  $\Delta M(\text{Rb})$   
344 values in Fig. 4c are derived from the measured and interpolated values shown in Fig.  
345 4b; the values have therefore been corrected for temporal drift (see Section 3.2.3).

346 The enhanced  $\Delta M(\text{Rb})$  values are consistent with the higher ionisation  
347 efficiency of Rb than Sr in the ICP, where the first ionisation energy of Rb (640 kJ  
348 mol<sup>-1</sup>) is lower than Sr (660 kJ mol<sup>-1</sup>), and the boiling and dissociation temperatures of  
349 Rb (312.2 and 961 K, respectively) are both lower than those of Sr (1042 and 1657 K,  
350 respectively) promoting together the higher ionization efficiency of Rb than for Sr.<sup>35</sup>  
351 When the local plasma equilibrium of Saha's equation<sup>30, 31</sup> and electron number density  
352  $ne = 1.47 \times 10^{20} \text{ m}^{-3}$  are assumed, plasma temperature at  $T = 5514 \text{ K}$  is calculated to  
353 yield a 4.35% enhancement in the ionization efficiency of Rb (92.05%) relative to Sr  
354 (88.21%). This temperature is far lower than that of a standard ICP ( $T = 6680 \text{ K}$ ),<sup>36</sup> but

is still similar to the apparent ionization temperature ( $T = \sim 5500$  K) estimated for element fractionation between a refractory SRM 612 glass and a more easily ionized BCR-2G basalt glass (see Fig. 2F in Kimura and Chang).<sup>21</sup> The large particle size of LA aerosols (10–500-nm with ~100-nm median diameter)<sup>21, 37</sup> and short transit time in the ICP may be the source of this apparently low ionization temperature.<sup>21</sup>

It is worth noting that the  $^{85}\text{Rb}/^{87}\text{Rb}$  factor determined by the solution aerosols should differ from that for the LA aerosols owing to the change in the plasma conditions when mass loading of the LA aerosols occurs.<sup>21, 32</sup> This was not the case with the volume of LA aerosols and the internal/intermediate precision being dealt with in this study, or in our previous studies, as shown by the intermediate level of analytical precision of the standard materials (**Sections 3.1.1 and 3.3**). In the case of heavy mass loading, however, determination of the  $^{85}\text{Rb}/^{87}\text{Rb}$  factor by LA aerosols should be recommended.

Finally, the clear correlation between  $\Delta M(\text{Rb})$  and  $\Delta M(\text{Sr})$  (**Fig. 4c**) confirms that the Rb overlap correction factor of  $^{85}\text{Rb}/^{87}\text{Rb}$ , no matter the determination method, should be used in combination with the  $^{86}\text{Sr}/^{88}\text{Sr}$  mass fractionation factor in each analytical run (see **Section 1.1**) Nevertheless, temporal drift in the Rb correction factor occurs regardless (**Fig. 4b**), and should be closely monitored, determined, and applied (**Section 3.2.3**).

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### 375 **3.2.3. Temporal drift of Rb correction factor**

376 The  $^{85}\text{Rb}/^{87}\text{Rb}$  correction factor drifted with time, most likely reflecting a temporal  
377 change in elemental sensitivity between Rb and Sr. The general trend showed a rapid  
378 increase in  $^{85}\text{Rb}/^{87}\text{Rb}$  with time after the plasma was on (**Fig. 4a**). The drift stabilised

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6 379 after 2–3 h but problematic changes could have occurred (see Days 5 and 7 in **Fig. 4a**).  
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8 380 The absolute  $^{85}\text{Rb}/^{87}\text{Rb}$  value changed with different sampler and skimmer cones as  
9 shown for Days 7 and 8 in **Fig. 4a**. In more detail, the change showed a continuous  
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11 381 increase from 2.4669 to 2.4706 and dropped to 2.4694 between runs 39 and 40,  
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13 382 followed by a further increase to 2.4698 on Day 5 (**Fig. 4b**). The sudden drop correlates  
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15 383 to a longer interval of 1 h with no introduction of LA aerosols. We interpreted the  
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17 384 increase in the ratio as being caused by the change in the surface conditions of the  
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19 385 interface cones. This daily trend is common to all examples and is significant when  
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21 386 using a new sampler-skimmer cone for the first time (Day 7 in **Fig. 4b**). It is likely that  
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23 387 the systematic difference between the cones (Day 5–6 and Day 7–8 in **Fig. 4a**) results  
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25 388 from their different geometries.  
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390 The change of the  $^{85}\text{Rb}/^{87}\text{Rb}$  factor from 2.4669 to 2.4706 causes a systematic  
391 bias increase of  $^{87}\text{Sr}/^{86}\text{Sr} = 0.0014$  for BCR-2G (337-ppm Sr, 48.5-ppm Rb, Rb/Sr =  
392 0.144) and  $^{87}\text{Sr}/^{86}\text{Sr} = 0.0002$  for BHVO-2G (396-ppm Sr, 9.2-ppm Rb, Rb/Sr = 0.023).  
393 Considering the temporal change of the factor, frequent measurement of  $^{85}\text{Rb}/^{87}\text{Rb}$  is  
394 crucial for glass samples with a high Rb content. We used the same, ~5-ppb Sr SRM  
395 987 solution doped with ~1-ppb Rb for this purpose and determined the  $^{85}\text{Rb}/^{87}\text{Rb}$   
396 factors for every five unknown LA analyses. The temporal change in  $^{85}\text{Rb}/^{87}\text{Rb}$  was the  
397 linearly interpolated for the five unknown calculations.

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### 399 **3.3. Analyses of basalt glass samples**

400 With the instrument settings and the analytical protocols described in the above sections,  
401 we then analysed the BHVO-2G, KL2-G, ML3B-G, and BCR-2G glasses, which have  
402 also been analysed in the previous LA-ICP-MS studies<sup>18, 19</sup> and TIMS reference values

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6 403 are thus available.<sup>27</sup>  
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9 404 The basalt glasses of BHVO-2G (basalt from the Hawaiian Volcano  
10 Observatory), KL2-G (from Kilauea volcano), and ML3B-G (from Mauna Loa volcano)  
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12 405 are all from the Big Island of Hawaii. As shown in **Table 2** and **ESI Data Table 1†**, the  
13  
14 406 results of BHVO-2G were:  $^{87}\text{Sr}/^{86}\text{Sr} = 0.70345 \pm 0.00020$  ( $n = 15$ ) from 100- $\mu\text{m}/10\text{-Hz}$   
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16 407 craters and  $0.7035 \pm 0.0005$  ( $n = 55$ ) from 50- $\mu\text{m}/10\text{-Hz}$  craters, for the TIMS value of  
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18 408  $^{87}\text{Sr}/^{86}\text{Sr} = 0.703469 \pm 0.000010$  (**Fig. 5a**); those of KL2-G were:  $^{87}\text{Sr}/^{86}\text{Sr} = 0.70352 \pm$   
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20 409  $0.00013$  ( $n = 5$ ) from 100- $\mu\text{m}/10\text{-Hz}$  craters for the TIMS value of  $0.703517 \pm 0.00002$ ;  
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22 410 and those of ML3B-G were:  $^{87}\text{Sr}/^{86}\text{Sr} = 0.70386 \pm 0.00027$  ( $n = 10$ ) from  
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24 411 100- $\mu\text{m}/10\text{-Hz}$  craters for the TIMS value of  $0.703805 \pm 0.000020$ . The LA results are  
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26 412 all in excellent agreement with the TIMS values, indicating that the appropriate  
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28 413 correction factor was used in this study.  
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34 415 BCR-2G (Columbia River Basalt) glass, with Rb/Sr = 0.144, is the most  
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36 416 difficult sample for analysis of the isobaric corrections of the concomitant Rb in this  
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38 417 study. The results of BCR-2G also showed excellent agreement with the TIMS reference  
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41 418 value of  $^{87}\text{Sr}/^{86}\text{Sr} = 0.705000 \pm 0.000020$ ,<sup>27</sup> as shown by  $^{87}\text{Sr}/^{86}\text{Sr} = 0.70495 \pm 0.00040$   
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43 419 ( $n = 39$ ) from 100- $\mu\text{m}/10\text{-Hz}$  craters and  $^{87}\text{Sr}/^{86}\text{Sr} = 0.7051 \pm 0.0008$  ( $n = 50$ ) from  
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45 420 50- $\mu\text{m}/10\text{-Hz}$  craters (**Fig. 5b**). This proves that the method can be applied to many  
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47 421 natural basalt melts of interests.<sup>3</sup>  
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51 422 The intermediate precisions were  $\pm 0.00010$  for 200- $\mu\text{m}/10\text{-Hz}$  craters,  
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53 423  $\pm 0.00013$ – $0.00040$  for 100- $\mu\text{m}/10\text{-Hz}$  craters, and  $\pm 0.00050$ – $0.00080$  for 50- $\mu\text{m}/10\text{-Hz}$   
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55 424 craters (**Table 2** and **Fig. 5**). These internal/intermediate precisions are a factor of  
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57 425 approximately two smaller than those obtained by LA-SF-ICP-MS for a 50- $\mu\text{m}/10\text{-Hz}$   
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59 426 crater in year 2009.<sup>18, 19</sup> The intermediate precisions of  $\pm 0.00050$ – $0.00080$  for a  
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6 427 50- $\mu\text{m}$ /10-Hz crater are comparable to those obtained for a 120- $\mu\text{m}$ /20-Hz crater using  
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8 428 216-nm YAG-LA-MC-ICP-MS in year 2006<sup>3</sup> but from an ~10 times smaller sample  
9 volume. If compared with the first report of 1064 nm by YAG-LA-MC-ICP-MS in year  
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11 429 1995,<sup>1</sup> the necessary sample amount for a  $\pm 0.00010$  internal/intermediate precision is  
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13 430 ~200 times smaller, reflecting significant improvements in both the instruments and the  
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15 431 analytical protocols. These improvements are useful for the analyses of small melt  
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17 432 inclusions.  
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28 435 **4. Conclusions**  
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60 436 We evaluated analytical protocols using LA-MC-ICP-MS for accurate Sr isotope  
analysis of small glass samples. The Kr baseline suppression during LA analysis was  
found to be nonlinear; small for small sample aerosol loading and large for an extreme  
sample loading. We corrected the baseline bias for the all sample amounts from the LA  
crater of 50–200  $\mu\text{m}$ /10 Hz and found it to be applicable over the range. Precise Rb  
overlap correction using a Rb-doped standard solution and temporal variation of the  
correction factor were also examined. The results showed that frequent measurement of  
the correction factor was required for a high Rb content basalt glass. These correction  
methods were applied to the in situ Sr isotope analyses of the basalt glass standards of  
BHVO-2G, BCR-2G, KL2-G, and ML3B-G. The internal/intermediate precisions  
achieved for the glass standards were  $^{87}\text{Sr}/^{86}\text{Sr} = \pm 0.00010$  (200  $\mu\text{m}$ /10 Hz) to 0.00080  
(50  $\mu\text{m}$ /10 Hz), dependent on the crater size and Rb content. The results were a factor of  
two more precise than the previous method using a 50- $\mu\text{m}$  crater with  
ExLA-SF-ICP-MS and were achieved with a ten times smaller sample volume than  
those in 213-nm Nd-YAG-MC-ICP-MS. The results demonstrate the wide application of

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6 451 this method to many naturally occurring melt inclusions in olivine, which typically have  
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8 452 a diameter of 30–200 µm with Sr levels of 200–300 ppm.  
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## 454 Acknowledgments

455 The authors wish to thank Dr Y. Tatsumi of JAMSTEC (currently at Kobe University)  
456 for his generous support of this project. We are grateful to Prof. K. P. Jochum of the  
457 Max Plank Institute for Chemistry for providing the MPI-DING glasses. The technical  
458 and engineering support of Mr K. Ohki of OK Laboratory enabled the manufacture of  
459 the high-quality 193-nm excimer LA imaging optics that helped establish the prolonged  
460 stable ablation of plagioclase. Discussions with Dr D. Bouman of Thermo Scientific  
461 were useful for application of the high-gain amplifiers. Finally, the authors are indebted  
462 to the two anonymous reviewers who provided detailed and constructive comments,  
463 which helped to further improve this paper.  
464

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6 545 **Figure and Table caption**  
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10 547 **Fig. 1.** Signal profiles (a) and isotope ratios (b) acquired from a 50- $\mu\text{m}$  crater in ~25 s  
11 548 on BCR-2G. Stable signals and isotope ratios are obtained after 10 s of idle time with  
12 549 laser firing. No effect from slow response of the Faraday amplifiers was observed.  
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19 551 **Fig. 2.** Variation of  $^{87}\text{Sr}/^{86}\text{Sr}$  of SRM 987 solution (~0.5-V  $^{88}\text{Sr}$ ) measured with various  
20 552 amounts of aerosol loading from a St Helena olivine crystal. The bias from Kr baseline  
21 553 suppression is significant in the crater size 200  $\mu\text{m}$ /20 Hz. Uncorrected values are from  
22 554 isobaric overlap and mass fractionation corrections only. Corrected values are further  
23 555 corrected for Kr baseline-induced bias. Errors are in 2SE. Aerosol loading without LA  
24 556 (w/o LA) is provisionally plotted at  $1 \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$ .  
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35 558 **Fig. 3.** Analytical results of the MkAn anorthite using different crater sizes for different  
36 559 amounts of sample aerosol loading (panel a). Uncorrected (panel b) and corrected  
37 560  $^{87}\text{Sr}/^{86}\text{Sr}$ - $^{84}\text{Sr}/^{86}\text{Sr}$  ratios (panel c) of the Kr baseline-induced biases are compared for  
38 561 the different sample volumes. Data from large volume aerosols (200- $\mu\text{m}$ /10–20-Hz  
39 562 crater) is plotted on the theoretical fractionation line of the Kr baseline suppression,  
40 563 while those from the smaller volume aerosols (100–50- $\mu\text{m}$ /10-Hz craters) show slight  
41 564 over-corrections. Potential biases from these overcorrections are less than  $\pm 0.00020$   
42 565 (thick vertical line), well below the in-run analytical precisions for these data. Thick  
43 566 shaded lines, thin solid lines, and thin dotted lines are regression lines for the results  
44 567 from 200-, 100-, and 50- $\mu\text{m}$  craters, respectively. Thin dashed lines show the reference  
45 568 value of the MkAn as measured by TIMS.  
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570 **Fig. 4.** Daily temporal variations of  $^{85}\text{Rb}/^{87}\text{Rb}$  correction factors during analyses of  
571 basalt glass samples (panel a). The large gap between runs 39 and 40 is due to a long  
572 break for 2 h indicating reverse shift of the drift due to no loading of the laser aerosols  
573 (panel b). Apparent Rb mass bias and actual Sr mass bias, measured on Day 5 (panel c).  
574 Note that open squares show measured values and dots are interpolations in panel b. See  
575 details in the text.

576

577 **Fig. 5.** Analytical results of BHVO-2G and BCR-2G glass standards. Internal precisions  
578 are given by 2SE, whereas those for averages (open circles) are given by two-standard  
579 deviation (2SD). Solid squares and shaded lines are TIMS reference values.

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581 **Table 1.** Laser and mass spectrometer set-up parameters for the LA-MC-ICP-MS Sr  
582 isotope analysis

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584 **Table 2.** Representative LA-MC-ICP-MS analytical results of Sr isotope ratios in  
585 synthetic glass standards

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587 **ESI Data Table 1.** LA-MC-ICP-MS analytical results of Sr isotope ratios

588

589 **Table of Contents Entry: Graphical abstract**

590 An analytical protocol using LA-MC-ICP-MS for precise Sr isotope ratio analysis from  
591 small glass samples is presented

592

**Table 1.** Laser and mass spectrometer set-up parameters for the LA-MC-ICP-MS Sr isotope analysis

Apparatus	Experimental setting
Excimer laser	OK Laboratory Ltd. (in house)
Source wave length/pulse width	193 nm/20 ns
Energy at source	200 mJ
Focusing lens	Imaging optics using field lens and air spaced doublet objective
Spot size	50, 100, and 200 $\mu\text{m}$
Fluence at laser spot	20 J $\text{cm}^{-2}$
Repetition rate	10 and 20 Hz
Aridus II	CETAC Technologies
Spray chamber temperature	110 °C
Membrane temperature	160 °C
Nebuliser flow	1.0 L $\text{min}^{-1}$
Membrane gas flow	~3.5 L $\text{min}^{-1}$
Additional N <sub>2</sub> gas flow	1.5 mL $\text{min}^{-1}$
MC-ICP-MS	Thermo NEPTUNE
RF-power	1400 W
Guard electrode	off (electronically disconnected)
Sampling cone	N-sample cone (Ni)
Skimmer cone	N-skimmer cone (Ni)
Plasma gas (Ar)	1.0 L $\text{min}^{-1}$
Laser carrier gas (He)	0.7 L $\text{min}^{-1}$
Interface vacuum with E2M80	1.7 mbar with He ablation carrier gas
Acquisition time	0.5 s × 50 scans for LA and 4 s × 50 scans for solution
Faraday detector setting	
<sup>83</sup> Kr ( <sup>83</sup> Kr, <sup>43</sup> Ca <sup>40</sup> Ca, <sup>43</sup> Ca <sup>40</sup> Ar)	FC L4 (10 <sup>11</sup> Ω amplifier)
<sup>83.5</sup> M ( <sup>167</sup> Er <sup>++</sup> )	FC L3 (10 <sup>11</sup> Ω amplifier)
<sup>84</sup> Sr ( <sup>84</sup> Kr, <sup>44</sup> Ca <sup>40</sup> Ca, <sup>44</sup> Ca <sup>40</sup> Ar)	FC L2 (10 <sup>12</sup> Ω amplifier)
<sup>85</sup> Rb ( <sup>85</sup> GaO)	FC L1 (10 <sup>12</sup> Ω amplifier)
<sup>86</sup> Sr ( <sup>86</sup> Kr, <sup>46</sup> Ca <sup>40</sup> Ca, <sup>46</sup> Ca <sup>40</sup> Ar)	FC Center (10 <sup>12</sup> Ω amplifier)
<sup>86.5</sup> M ( <sup>173</sup> Yb <sup>++</sup> )	FC H1 (10 <sup>11</sup> Ω amplifier)
<sup>87</sup> Sr ( <sup>87</sup> Rb, <sup>85</sup> GaO)	FC H2 (10 <sup>12</sup> Ω amplifier)
<sup>88</sup> Sr ( <sup>48</sup> Ca <sup>40</sup> Ca, <sup>48</sup> Ca <sup>40</sup> Ar)	FC H3 (10 <sup>12</sup> Ω amplifier)
Background subtraction	on peak background subtraction

FC: Faraday cup; isobaric atomic and molecular ions are shown in parentheses

633      **Table 2.** Representative LA-MC-ICP-MS analytical results of Sr isotope ratios in synthetic glass standards  
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Sample	Sr ppm	Rb ppm	Rb/Sr	$^{87}\text{Sr}/^{86}\text{Sr}$	2SD	n	$^{87}\text{Sr}/^{86}\text{Sr}$	2SD	n	$^{87}\text{Sr}/^{86}\text{Sr}$	2SE	Reference
BHVO-2G	396	9.2	0.023	0.70345	0.00020	15	0.7035	0.0005	55	0.703469	0.00001	This study
	–	–	–	–	–	–	0.7035	0.0010	36	–	–	Jochum et al. (2009)
	–	–	–	–	0.7035	0.0010	–	–	–	–	–	Sovolev et al. (2011)
KL2-G	356	8.7	0.024	0.70352	0.00013	5	–	–	–	0.703517	0.00002	This study
	–	–	–	–	0.7036	0.0008	43	–	–	–	–	Jochum et al. (2009)
	–	–	–	–	0.7035	0.0008	–	–	–	–	–	Sobolev et al. (2011)
ML3B-G	312	5.8	0.019	0.70386	0.00027	10	–	–	–	0.703805	0.00002	This study
	–	–	–	–	0.7038	0.0008	48	–	–	–	–	Jochum et al. (2009)
	–	–	–	–	0.7038	0.0010	–	–	–	–	–	Sobolev et al. (2011)
BCR-2G	337	48.5	0.144	0.70495	0.00040	39	0.7051	0.0008	50	0.705000	0.00002	This study

647 Note: 2SD = two-standard deviation; 2SE = two-standard error; n = number of analyses. TIMS reference values are from GeoREM<sup>27</sup>

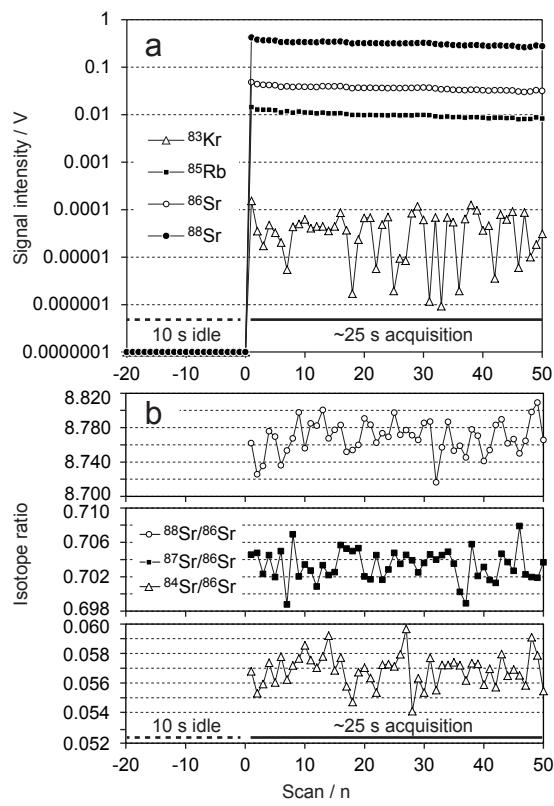


Fig. 1.

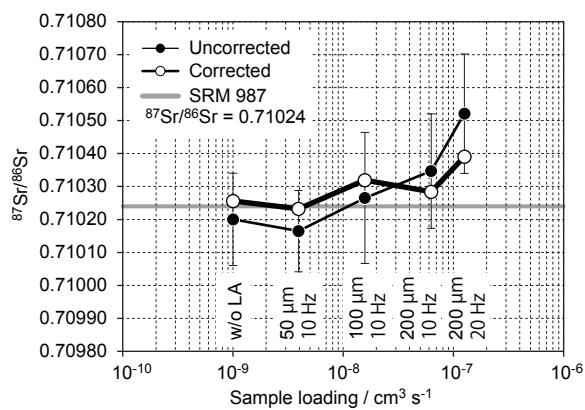


Fig. 2.

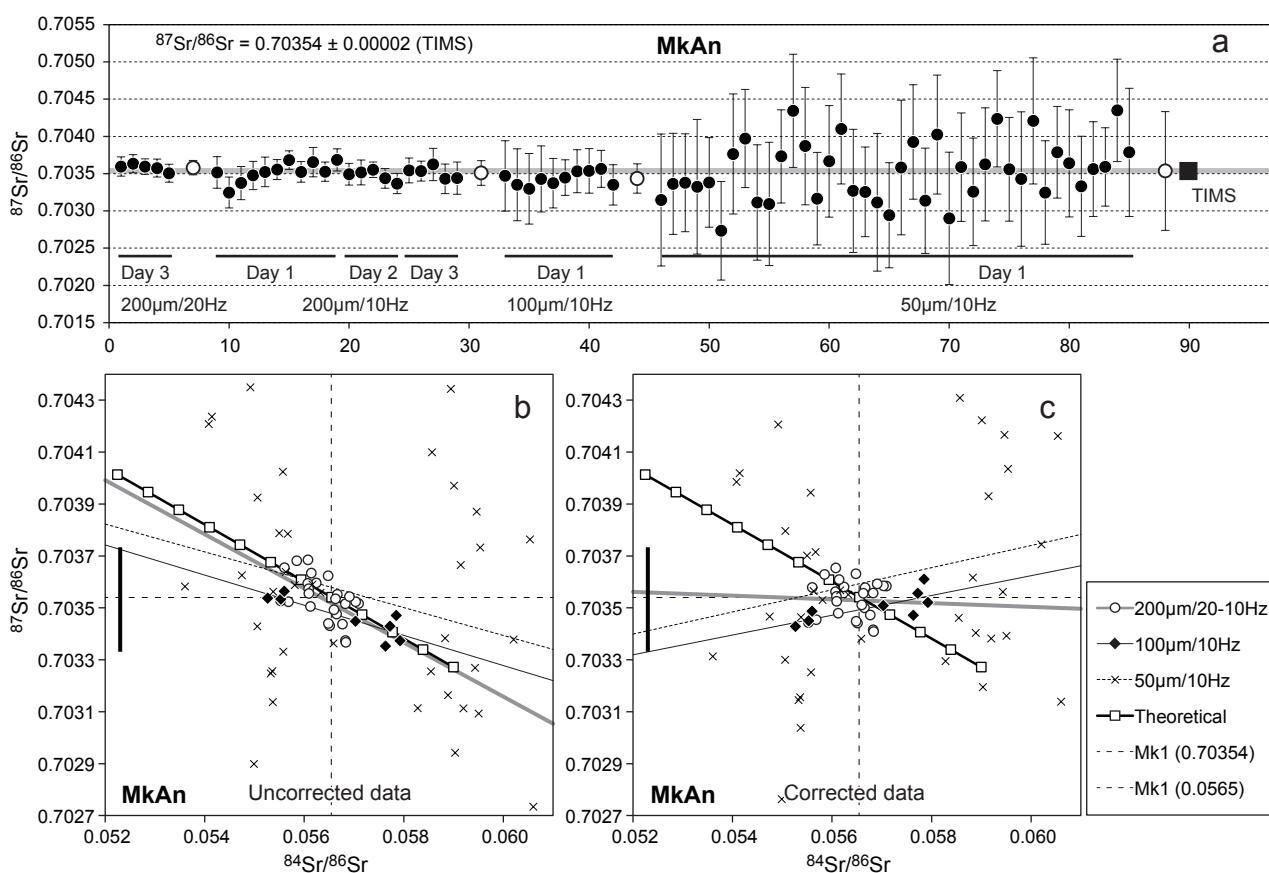


Fig. 3.

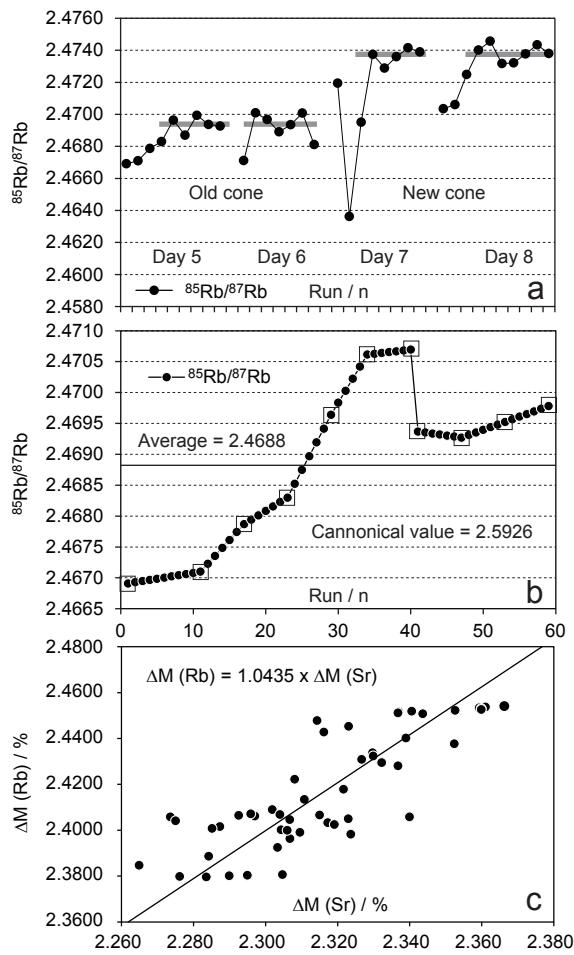


Fig. 4.

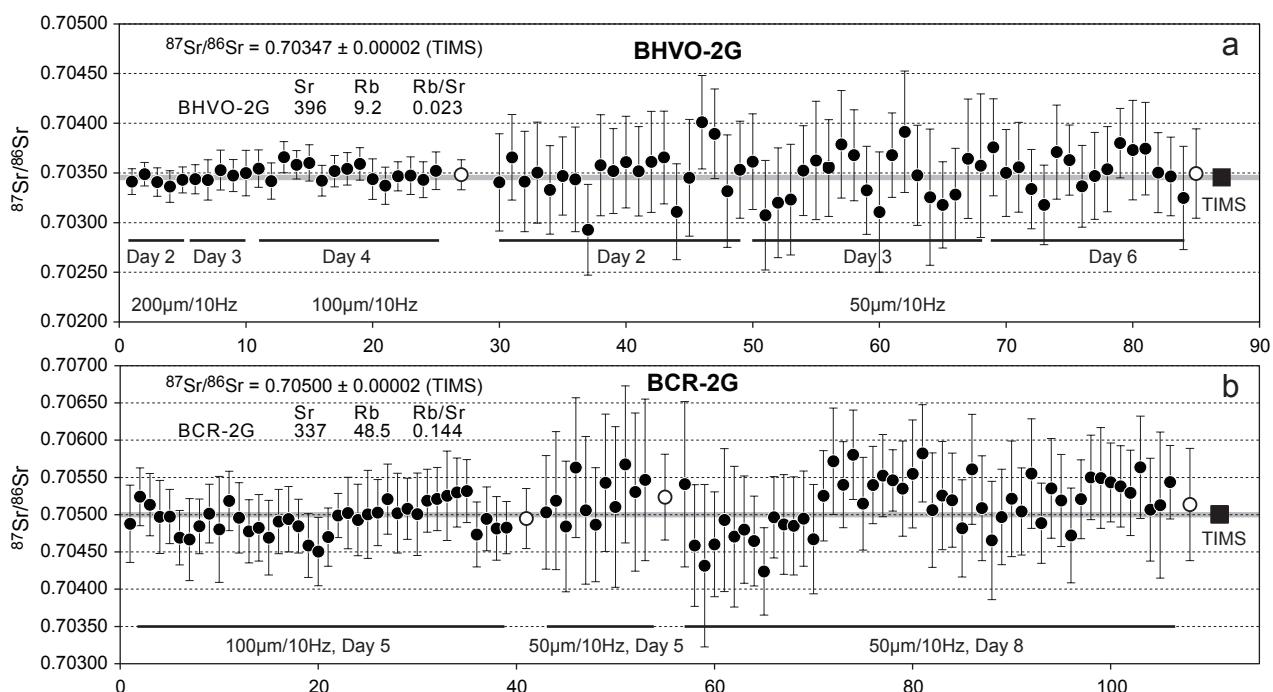


Fig. 5.

**E.S.I. Data Table 1 Analytical results of Sr isotope ratios**

Day	Run ID	Crater	Sample	$^{83}\text{Kr}$ (V)	$^{83.5}\text{M}$ (V)	$^{84}\text{Sr}$ (V)	$^{85}\text{Rb}$ (V)	$^{86}\text{Sr}$ (V)	$^{86.5}\text{M}$ (V)	$^{87}\text{Sr}$ (V)	$^{88}\text{Sr}$ (V)	$^{84}\text{Sr}/^{86}\text{Sr}$	$^{87}\text{Sr}/^{86}\text{Sr}$	$^{88}\text{Sr}/^{86}\text{Sr}$	$^{87}\text{Sr}/^{86}\text{Sr}$ cor	2SE
JUL 26 2014	Run5	200um/20Hz	MkAn1_1	0.00117	-0.00001	0.01178	0.00004	0.21948	-0.00002	0.15805	1.92449	0.05624	0.70360	8.76804	0.70352	0.00013
JUL 26 2014	Run6	200um/20Hz	MkAn1_2	0.00133	0.00001	0.01286	0.00006	0.24015	-0.00001	0.17297	2.10626	0.05614	0.70363	8.77060	0.70355	0.00012
JUL 26 2014	Run7	200um/20Hz	MkAn1_3	0.00154	-0.00001	0.01445	0.00008	0.27020	-0.00001	0.19460	2.36967	0.05607	0.70359	8.76992	0.70350	0.00010
JUL 26 2014	Run8	200um/20Hz	MkAn1_4	0.00180	-0.00001	0.01642	0.00074	0.30698	0.00000	0.22138	2.69283	0.05609	0.70357	8.77176	0.70348	0.00012
JUL 26 2014	Run9	200um/20Hz	MkAn1_5	0.00119	0.00000	0.01175	0.00005	0.21935	0.00000	0.15794	1.92321	0.05613	0.70351	8.76742	0.70342	0.00012
<b>Average and error in 2SD</b>																
JUL 24 2014	Run8	200um/10Hz	MkAn1_1	0.00028	0.00000	0.00565	0.00003	0.10349	0.00000	0.07445	0.90576	0.05710	0.70352	8.75227	0.70352	0.00021
JUL 24 2014	Run9	200um/10Hz	MkAn1_2	0.00026	-0.00001	0.00518	0.00003	0.09405	0.00000	0.06760	0.82241	0.05762	0.70325	8.74408	0.70330	0.00021
JUL 24 2014	Run10	200um/10Hz	MkAn1_3	0.00024	-0.00001	0.00477	0.00003	0.08776	-0.00001	0.06311	0.76786	0.05683	0.70337	8.74932	0.70335	0.00022
JUL 24 2014	Run11	200um/10Hz	MkAn1_4	0.00026	0.00000	0.00518	0.00003	0.09558	-0.00001	0.06875	0.83626	0.05665	0.70348	8.74836	0.70344	0.00019
JUL 24 2014	Run12	200um/10Hz	MkAn1_5	0.00026	0.00000	0.00506	0.00004	0.09280	0.00000	0.06674	0.81166	0.05703	0.70352	8.74643	0.70352	0.00020
JUL 24 2014	Run38	200um/10Hz	MkAn1_31	0.00040	0.00000	0.00835	0.00010	0.15556	0.00000	0.11188	1.36001	0.05609	0.70355	8.74162	0.70346	0.00013
JUL 24 2014	Run39	200um/10Hz	MkAn1_32	0.00040	0.00001	0.00848	0.00005	0.15871	0.00000	0.11416	1.38802	0.05585	0.70368	8.74544	0.70357	0.00012
JUL 24 2014	Run40	200um/10Hz	MkAn1_33	0.00038	-0.00002	0.00827	0.00004	0.15527	0.00000	0.11166	1.35790	0.05568	0.70352	8.74486	0.70339	0.00014
JUL 24 2014	Run41	200um/10Hz	MkAn1_34	0.00042	0.00000	0.00868	0.00005	0.16319	-0.00001	0.11737	1.42700	0.05560	0.70365	8.74416	0.70352	0.00020
JUL 24 2014	Run42	200um/10Hz	MkAn1_35	0.00041	-0.00001	0.00859	0.00005	0.16161	-0.00002	0.11623	1.41360	0.05553	0.70353	8.74682	0.70338	0.00013
JUL 24 2014	Run68	200um/10Hz	MkAn1_61	0.00061	0.00000	0.00863	0.00003	0.16132	-0.00001	0.11618	1.41451	0.05608	0.70368	8.76833	0.70359	0.00015
JUL 25 2014	Run6	200um/10Hz	MkAn1_1	0.00073	0.00000	0.00926	0.00009	0.17127	-0.00001	0.12330	1.50077	0.05660	0.70349	8.76259	0.70345	0.00015
JUL 25 2014	Run7	200um/10Hz	MkAn1_2	0.00064	-0.00001	0.00825	0.00005	0.15199	0.00000	0.10938	1.33119	0.05683	0.70352	8.75837	0.70350	0.00017
JUL 25 2014	Run8	200um/10Hz	MkAn1_3	0.00073	-0.00002	0.00914	0.00005	0.16897	-0.00001	0.12162	1.48030	0.05667	0.70355	8.76053	0.70352	0.00010
JUL 25 2014	Run9	200um/10Hz	MkAn1_4	0.00079	-0.00002	0.00968	0.00002	0.17855	0.00000	0.12849	1.56437	0.05677	0.70344	8.76091	0.70341	0.00013
JUL 25 2014	Run10	200um/10Hz	MkAn1_5	0.00059	0.00000	0.00799	0.00004	0.14725	-0.00001	0.10595	1.28967	0.05684	0.70337	8.75807	0.70335	0.00014
JUL 26 2014	Run10	200um/10Hz	MkAn1_1	0.00054	0.00000	0.00635	0.00003	0.11686	-0.00001	0.08414	1.02456	0.05691	0.70354	8.76722	0.70353	0.00017
JUL 26 2014	Run11	200um/10Hz	MkAn1_2	0.00058	0.00001	0.00682	0.00017	0.12616	0.00000	0.09090	1.10618	0.05663	0.70353	8.76792	0.70349	0.00013
JUL 26 2014	Run12	200um/10Hz	MkAn1_3	0.00045	-0.00001	0.00559	0.00004	0.10370	0.00000	0.07467	0.90882	0.05648	0.70362	8.76379	0.70357	0.00022
JUL 26 2014	Run13	200um/10Hz	MkAn1_4	0.00046	-0.00001	0.00577	0.00012	0.10699	0.00000	0.07706	0.93788	0.05652	0.70343	8.76530	0.70338	0.00020
JUL 26 2014	Run14	200um/10Hz	MkAn1_5	0.00055	-0.00001	0.00672	0.00004	0.12461	-0.00001	0.08971	1.09240	0.05649	0.70344	8.76664	0.70339	0.00022
<b>Average and error in 2SD</b>																
JUL 24 2014	Run13	100um/10Hz	MkAn1_6	0.00011	-0.00001	0.00212	0.00001	0.03820	-0.00002	0.02747	0.33436	0.05785	0.70347	8.75261	0.70355	0.00047
JUL 24 2014	Run14	100um/10Hz	MkAn1_7	0.00009	0.00000	0.00216	0.00001	0.03921	0.00000	0.02820	0.34328	0.05764	0.70335	8.75318	0.70341	0.00048
JUL 24 2014	Run15	100um/10Hz	MkAn1_8	0.00011	0.00000	0.00229	0.00001	0.04164	-0.00001	0.02996	0.36461	0.05761	0.70330	8.75377	0.70335	0.00047
JUL 24 2014	Run16	100um/10Hz	MkAn1_9	0.00012	-0.00001	0.00252	0.00001	0.04584	0.00001	0.03298	0.40156	0.05772	0.70343	8.75992	0.70349	0.00044
JUL 24 2014	Run17	100um/10Hz	MkAn1_10	0.00013	0.00001	0.00270	0.00002	0.04892	0.00000	0.03521	0.42854	0.05793	0.70337	8.75998	0.70346	0.00033
JUL 24 2014	Run43	100um/10Hz	MkAn1_36	0.00011	0.00000	0.00296	0.00003	0.05421	-0.00002	0.03899	0.47416	0.05703	0.70345	8.74680	0.70345	0.00024

1	JUL 25 2014	Run23	50um/10Hz	BHVO-2G_2	0.00004	0.00001	0.00171	0.01014	0.03137	-0.00001	0.02670	0.27526	0.05729	0.7037	8.7731	0.7036	0.0004
2	JUL 25 2014	Run24	50um/10Hz	BHVO-2G_3	0.00003	0.00000	0.00177	0.01046	0.03214	0.00000	0.02737	0.28173	0.05776	0.7034	8.7655	0.7034	0.0005
3	JUL 25 2014	Run25	50um/10Hz	BHVO-2G_4	0.00004	0.00000	0.00183	0.01097	0.03313	0.00000	0.02829	0.29048	0.05799	0.7035	8.7658	0.7036	0.0005
4	JUL 25 2014	Run26	50um/10Hz	BHVO-2G_5	0.00004	0.00001	0.00180	0.01112	0.03335	0.00001	0.02851	0.29244	0.05666	0.7033	8.7664	0.7032	0.0004
5	JUL 25 2014	Run27	50um/10Hz	BHVO-2G_6	0.00004	0.00000	0.00180	0.01109	0.03319	-0.00001	0.02839	0.29111	0.05698	0.7035	8.7684	0.7034	0.0004
6	JUL 25 2014	Run28	50um/10Hz	BHVO-2G_7	0.00004	0.00000	0.00182	0.01108	0.03305	-0.00001	0.02828	0.28988	0.05758	0.7034	8.7694	0.7034	0.0005
7	JUL 25 2014	Run29	50um/10Hz	BHVO-2G_8	0.00004	-0.00001	0.00184	0.01137	0.03399	0.00000	0.02906	0.29807	0.05669	0.7029	8.7679	0.7028	0.0005
8	JUL 25 2014	Run30	50um/10Hz	BHVO-2G_9	0.00002	0.00000	0.00189	0.01144	0.03454	0.00000	0.02950	0.30276	0.05722	0.7036	8.7625	0.7036	0.0005
9	JUL 25 2014	Run31	50um/10Hz	BHVO-2G_10	0.00003	0.00001	0.00191	0.01156	0.03542	0.00000	0.03017	0.31038	0.05662	0.7035	8.7620	0.7034	0.0004
10	JUL 25 2014	Run32	50um/10Hz	BHVO-2G_11	0.00004	0.00002	0.00200	0.01189	0.03654	-0.00002	0.03112	0.32023	0.05731	0.7036	8.7641	0.7036	0.0005
11	JUL 25 2014	Run33	50um/10Hz	BHVO-2G_12	0.00005	0.00002	0.00201	0.01185	0.03657	0.00001	0.03112	0.32060	0.05774	0.7035	8.7671	0.7035	0.0005
12	JUL 25 2014	Run34	50um/10Hz	BHVO-2G_13	0.00003	0.00001	0.00197	0.01177	0.03623	0.00000	0.03085	0.31748	0.05699	0.7036	8.7638	0.7036	0.0005
13	JUL 25 2014	Run35	50um/10Hz	BHVO-2G_14	0.00004	0.00000	0.00197	0.01176	0.03625	0.00000	0.03086	0.31772	0.05682	0.7037	8.7644	0.7036	0.0005
14	JUL 25 2014	Run36	50um/10Hz	BHVO-2G_15	0.00002	0.00001	0.00193	0.01140	0.03568	0.00000	0.03028	0.31249	0.05656	0.7031	8.7587	0.7030	0.0005
15	JUL 25 2014	Run37	50um/10Hz	BHVO-2G_16	0.00005	0.00001	0.00197	0.01179	0.03625	0.00001	0.03086	0.31754	0.05706	0.7035	8.7619	0.7034	0.0006
16	JUL 25 2014	Run38	50um/10Hz	BHVO-2G_17	0.00004	0.00000	0.00197	0.01202	0.03647	-0.00001	0.03113	0.31954	0.05670	0.7040	8.7627	0.7039	0.0005
17	JUL 25 2014	Run39	50um/10Hz	BHVO-2G_18	0.00004	0.00001	0.00198	0.01199	0.03616	0.00000	0.03090	0.31694	0.05736	0.7039	8.7637	0.7039	0.0005
18	JUL 25 2014	Run40	50um/10Hz	BHVO-2G_19	0.00004	0.00000	0.00198	0.01213	0.03649	-0.00001	0.03117	0.31983	0.05693	0.7033	8.7641	0.7033	0.0006
19	JUL 25 2014	Run41	50um/10Hz	BHVO-2G_20	0.00004	0.00002	0.00196	0.01230	0.03630	0.00001	0.03112	0.31829	0.05657	0.7035	8.7669	0.7034	0.0005
20	JUL 26 2014	Run20	50um/10Hz	BHVO-2G_1	0.00004	0.00001	0.00173	0.01025	0.03169	-0.00001	0.02698	0.27806	0.05718	0.7036	8.7748	0.7036	0.0005
21	JUL 26 2014	Run21	50um/10Hz	BHVO-2G_2	0.00001	0.00002	0.00164	0.01011	0.03011	-0.00001	0.02576	0.26406	0.05721	0.7031	8.7688	0.7031	0.0006
22	JUL 26 2014	Run22	50um/10Hz	BHVO-2G_3	0.00003	-0.00001	0.00168	0.01031	0.03109	0.00000	0.02655	0.27264	0.05665	0.7032	8.7696	0.7032	0.0006
23	JUL 26 2014	Run23	50um/10Hz	BHVO-2G_4	0.00006	0.00001	0.00165	0.00976	0.03003	0.00001	0.02557	0.26335	0.05751	0.7032	8.7687	0.7033	0.0006
24	JUL 26 2014	Run24	50um/10Hz	BHVO-2G_5	0.00003	0.00000	0.00180	0.01007	0.03274	-0.00001	0.02766	0.28722	0.05756	0.7035	8.7724	0.7036	0.0005
25	JUL 26 2014	Run25	50um/10Hz	BHVO-2G_6	0.00004	0.00001	0.00185	0.01124	0.03389	0.00000	0.02895	0.29703	0.05724	0.7036	8.7641	0.7036	0.0006
26	JUL 26 2014	Run26	50um/10Hz	BHVO-2G_7	0.00005	0.00001	0.00181	0.01091	0.03292	0.00000	0.02813	0.28872	0.05779	0.7035	8.7687	0.7036	0.0005
27	JUL 26 2014	Run27	50um/10Hz	BHVO-2G_8	0.00004	0.00000	0.00177	0.01066	0.03255	0.00000	0.02776	0.28540	0.05701	0.7038	8.7692	0.7038	0.0005
28	JUL 26 2014	Run28	50um/10Hz	BHVO-2G_9	0.00005	0.00000	0.00180	0.01056	0.03273	0.00000	0.02784	0.28679	0.05747	0.7037	8.7635	0.7037	0.0005
29	JUL 26 2014	Run29	50um/10Hz	BHVO-2G_10	0.00004	0.00001	0.00174	0.01059	0.03213	-0.00001	0.02742	0.28170	0.05679	0.7033	8.7666	0.7033	0.0004
30	JUL 26 2014	Run30	50um/10Hz	BHVO-2G_11	0.00003	0.00000	0.00169	0.01040	0.03132	0.00000	0.02675	0.27462	0.05658	0.7031	8.7678	0.7031	0.0006
31	JUL 26 2014	Run31	50um/10Hz	BHVO-2G_12	0.00003	-0.00001	0.00176	0.01090	0.03241	-0.00001	0.02776	0.28429	0.05711	0.7037	8.7721	0.7037	0.0004
32	JUL 26 2014	Run32	50um/10Hz	BHVO-2G_13	0.00003	0.00002	0.00166	0.00975	0.03039	-0.00001	0.02585	0.26669	0.05744	0.7039	8.7747	0.7040	0.0006
33	JUL 26 2014	Run33	50um/10Hz	BHVO-2G_14	0.00002	0.00000	0.00160	0.00968	0.02946	0.00001	0.02514	0.25853	0.05689	0.7035	8.7743	0.7035	0.0005
34	JUL 26 2014	Run34	50um/10Hz	BHVO-2G_15	0.00005	0.00000	0.00157	0.00939	0.02855	0.00001	0.02436	0.25047	0.05750	0.7033	8.7700	0.7033	0.0007
35	JUL 26 2014	Run35	50um/10Hz	BHVO-2G_16	0.00003	0.00000	0.00172	0.01003	0.03178	-0.00002	0.02694	0.27893	0.05692	0.7032	8		

1	AUG 18 2014 Run05	50um/10Hz	BCR-2G_4	0.00005	0.00000	0.00171	0.05685	0.03161	0.00000	0.04580	0.27715	0.05679	0.7044	8.7663	0.7046	0.0007
2	AUG 18 2014 Run06	50um/10Hz	BCR-2G_5	0.00003	0.00001	0.00173	0.05844	0.03230	-0.00001	0.04695	0.28284	0.05606	0.7049	8.7585	0.7049	0.0010
3	AUG 18 2014 Run08	50um/10Hz	BCR-2G_6	0.00003	0.00001	0.00170	0.05887	0.03265	0.00000	0.04737	0.28590	0.05452	0.7049	8.7569	0.7047	0.0009
4	AUG 18 2014 Run09	50um/10Hz	BCR-2G_7	0.00003	0.00000	0.00174	0.05903	0.03292	0.00001	0.04762	0.28812	0.05525	0.7050	8.7532	0.7048	0.0007
5	AUG 18 2014 Run10	50um/10Hz	BCR-2G_8	0.00004	0.00001	0.00173	0.05955	0.03272	0.00000	0.04769	0.28640	0.05529	0.7048	8.7547	0.7046	0.0006
6	AUG 18 2014 Run11	50um/10Hz	BCR-2G_9	0.00004	0.00002	0.00180	0.06027	0.03293	-0.00001	0.04810	0.28802	0.05727	0.7042	8.7481	0.7042	0.0006
7	AUG 18 2014 Run12	50um/10Hz	BCR-2G_10	0.00007	0.00002	0.00193	0.06047	0.03322	-0.00001	0.04837	0.29021	0.06082	0.7041	8.7383	0.7050	0.0005
8	AUG 18 2014 Run14	50um/10Hz	BCR-2G_11	0.00004	0.00001	0.00159	0.05436	0.03018	0.00000	0.04375	0.26427	0.05519	0.7050	8.7563	0.7049	0.0007
9	AUG 18 2014 Run15	50um/10Hz	BCR-2G_12	0.00005	0.00003	0.00164	0.05442	0.02978	0.00000	0.04347	0.26063	0.05746	0.7048	8.7522	0.7049	0.0007
10	AUG 18 2014 Run16	50um/10Hz	BCR-2G_13	0.00004	0.00000	0.00167	0.05784	0.03153	0.00000	0.04613	0.27604	0.05530	0.7051	8.7556	0.7049	0.0006
11	AUG 18 2014 Run17	50um/10Hz	BCR-2G_14	0.00002	0.00000	0.00163	0.05638	0.03072	0.00000	0.04495	0.26905	0.05558	0.7048	8.7570	0.7047	0.0007
12	AUG 18 2014 Run18	50um/10Hz	BCR-2G_15	0.00002	0.00000	0.00165	0.05704	0.03117	0.00000	0.04556	0.27297	0.05548	0.7054	8.7570	0.7053	0.0006
13	AUG 18 2014 Run20	50um/10Hz	BCR-2G_16	0.00003	0.00002	0.00169	0.06024	0.03288	0.00001	0.04810	0.28811	0.05375	0.7060	8.7631	0.7057	0.0007
14	AUG 18 2014 Run21	50um/10Hz	BCR-2G_17	0.00005	0.00001	0.00178	0.06200	0.03338	0.00000	0.04915	0.29231	0.05579	0.7055	8.7576	0.7054	0.0006
15	AUG 18 2014 Run22	50um/10Hz	BCR-2G_18	0.00004	0.00001	0.00179	0.05980	0.03244	0.00000	0.04758	0.28383	0.05774	0.7057	8.7499	0.7058	0.0006
16	AUG 18 2014 Run23	50um/10Hz	BCR-2G_19	0.00003	0.00002	0.00172	0.05990	0.03289	0.00000	0.04794	0.28796	0.05477	0.7054	8.7561	0.7051	0.0006
17	AUG 18 2014 Run24	50um/10Hz	BCR-2G_20	-0.00001	0.00002	0.00151	0.05992	0.03274	-0.00001	0.04788	0.28708	0.04766	0.7063	8.7714	0.7054	0.0005
18	AUG 18 2014 Run26	50um/10Hz	BCR-2G_21	0.00004	0.00000	0.00181	0.05990	0.03346	0.00000	0.04834	0.29274	0.05646	0.7056	8.7496	0.7055	0.0006
19	AUG 18 2014 Run27	50um/10Hz	BCR-2G_22	0.00006	0.00002	0.00179	0.05920	0.03300	0.00001	0.04772	0.28875	0.05685	0.7055	8.7502	0.7055	0.0004
20	AUG 18 2014 Run28	50um/10Hz	BCR-2G_23	0.00004	0.00000	0.00179	0.06155	0.03309	0.00000	0.04874	0.28958	0.05656	0.7054	8.7508	0.7053	0.0006
21	AUG 18 2014 Run29	50um/10Hz	BCR-2G_24	0.00003	0.00001	0.00140	0.04941	0.02717	0.00000	0.03958	0.23797	0.05406	0.7058	8.7578	0.7055	0.0007
22	AUG 18 2014 Run30	50um/10Hz	BCR-2G_25	0.00005	0.00000	0.00155	0.05109	0.02840	0.00001	0.04114	0.24857	0.05704	0.7058	8.7528	0.7058	0.0007
23	AUG 18 2014 Run32	50um/10Hz	BCR-2G_26	0.00002	0.00002	0.00156	0.05429	0.02983	0.00000	0.04347	0.26122	0.05497	0.7053	8.7581	0.7051	0.0008
24	AUG 18 2014 Run33	50um/10Hz	BCR-2G_27	0.00005	0.00002	0.00149	0.04995	0.02801	-0.00001	0.04039	0.24503	0.05572	0.7054	8.7500	0.7053	0.0007
25	AUG 18 2014 Run34	50um/10Hz	BCR-2G_28	0.00003	0.00001	0.00147	0.05002	0.02799	-0.00001	0.04041	0.24494	0.05502	0.7054	8.7529	0.7052	0.0006
26	AUG 18 2014 Run35	50um/10Hz	BCR-2G_29	0.00002	-0.00001	0.00149	0.05062	0.02845	0.00000	0.04098	0.24906	0.05463	0.7050	8.7553	0.7048	0.0007
27	AUG 18 2014 Run36	50um/10Hz	BCR-2G_30	0.00005	0.00002	0.00157	0.05321	0.02882	0.00000	0.04230	0.25213	0.05713	0.7056	8.7480	0.7056	0.0007
28	AUG 18 2014 Run38	50um/10Hz	BCR-2G_31	0.00004	0.00000	0.00148	0.05012	0.02778	-0.00001	0.04030	0.24328	0.05586	0.7052	8.7566	0.7051	0.0007
29	AUG 18 2014 Run39	50um/10Hz	BCR-2G_32	0.00003	0.00000	0.00142	0.04955	0.02739	0.00001	0.03978	0.23976	0.05409	0.7049	8.7556	0.7047	0.0008
30	AUG 18 2014 Run40	50um/10Hz	BCR-2G_33	0.00004	0.00001	0.00152	0.05214	0.02845	0.00000	0.04159	0.24896	0.05584	0.7051	8.7507	0.7050	0.0006
31	AUG 18 2014 Run41	50um/10Hz	BCR-2G_34	0.00001	0.00001	0.00137	0.05178	0.02867	0.00000	0.04165	0.25133	0.04997	0.7059	8.7651	0.7052	0.0008
32	AUG 18 2014 Run42	50um/10Hz	BCR-2G_35	0.00003	0.00002	0.00163	0.05348	0.02997	0.00000	0.04322	0.26214	0.05689	0.7051	8.7485	0.7050	0.0006
33	AUG 18 2014 Run44	50um/10Hz	BCR-2G_36	0.00003	0.00002	0.00168	0.05801	0.03173	0.00000	0.04635	0.27765	0.05534	0.7057	8.7515	0.7056	0.0007
34	AUG 18 2014 Run45	50um/10Hz	BCR-2G_37	0.00004	0.00001	0.00162	0.05264	0.02985	-0.00002	0.04278	0.26098	0.05691	0.7049	8.7446	0.7049	0.0005
35	AUG 18 2014 Run46	50um/10Hz	BCR-2G_38	0.00005	0.00001	0.00160	0.05318	0.02936	0.00002	0.04267	0.25674	0.05683	0.7054	8.7436	0.7054	0.0007
36	AUG 18 2014 Run47	50um/10Hz	BCR-2G_39	0.00004	0.00001	0.00158	0.05438	0.02924	-0.00001	0.04306</td						