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Trends in the Adsorption and Reactivity of Hydrogen on Magnesium Silicate Nanoclusters

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Abstract

We study nanoclusters of Mg-rich olivine and pyroxene (having $(\text{MgO})_6(\text{SiO}_2)_3$ and $(\text{MgO})_4(\text{SiO}_2)_4$ compositions) with respect to their reactivity towards hydrogen atoms, using density functional calculations. Ultrasmall silicate particles are fundamental intermediates in cosmic dust grain formation and processing, and are thought to make up a significant mass fraction of the grain population. Due to their nanoscale dimensions and high surface area to bulk ratios, they are likely to also have a disproportionately large influence on surface chemistry in the interstellar medium. This work investigates the potential role of silicate nanoclusters in vital interstellar hydrogen-based chemistry by studying atomic H adsorption and H_2 formation. Our extensive set of calculations confirm the generality of a Brønsted–Evans–Polanyi (BEP) relation between the H_2 reaction barrier and the 2H_{chem} binding energy, suggesting it to be independent of silicate dust grain shape, size, crystallinity and composition. Our results also suggest that amorphous/porous grains with forsteritic composition would tend to dissociate H_2 , but relatively Mg-poor silicate grains (e.g. enstatite composition) and/or more crystalline/compact silicate grains would tend to catalyse H_2 formation. The high structural thermostability of silicate nanoclusters with respect to the heat released during exothermic H_2 formation reactions is also verified.

36 Introduction

37

38 H₂ is the most abundant molecule in the universe and is crucially involved in the formation of more
39 complex molecular species in the interstellar medium (ISM), e.g. the chemistry in interstellar clouds¹.
40 Due to its chemical importance, there is much interest in understanding the mechanisms and
41 formation rates of this vital H₂ feedstock, which needs to be continuously generated from ubiquitous
42 hydrogen atoms. It is widely accepted that the H₂ production reaction most likely occurs on dust
43 grains,^{2,3,4} as all competing gas phase processes are much less efficient at the low interstellar
44 temperatures and gas pressures. The adsorption of atomic hydrogen on small dust particles, can open
45 up reaction pathways which allow for H₂ formation to proceed with relatively reduced barriers. The
46 H₂ formation reaction is also highly exothermic (-434 kJ/mol) with heat that must be dissipated to
47 avoid dissociation of the product. Although significant amounts of energy can go into the internal
48 energy and translational kinetic energy of the formed H₂ molecule,^{5,6,7,8} dust particles can also play the
49 role of a third body to dissipate a significant percentage of the energy released leading to more stable
50 H₂ molecules.

51 Most interstellar dust is thought to be formed in the stellar winds of large dying stars^{9,10} and
52 has a chemical composition that depends on that of the parent star. Two main families of dust types
53 are thus formed in circumstellar environments: (i) carbonaceous and silicon carbide grains in carbon-
54 rich atmospheres, and (ii) magnesium rich silicate grains in the atmospheres of oxygen-rich stars.¹¹
55 After formation the stardust is ejected into the ISM where it is subjected to a multitude of processes
56 (e.g. amorphized by high energy radiation and shockwaves) and eventually incorporated into the
57 relative shelter of interstellar clouds where chemistry can more readily occur.¹² Of the two main
58 families of dust, we focus on silicate grains which are the most abundant and least well understood
59 family of solids in space.¹¹ Specifically, our study concentrates on nanosilicates and their chemical
60 interactions with hydrogen. These nanosilicates are fundamental intermediates in silicate dust
61 formation in stellar outflows, are likely to result from processing of large silicate grains, and are
62 ubiquitous throughout the ISM. Specifically, it has been estimated that 10% of the mass fraction of
63 the interstellar silicate grain population is made up of nanoclusters with ≤ 1.5 Å diameters.¹³ By
64 studying such small species we aim to probe the lower size limit of silicate dust grains with respect to
65 their cosmic chemical importance.

66 The properties of ultra small silicate nanoclusters are extremely difficult to examine in detail
67 through experimental laboratory or observational means, and experimental studies have focussed on
68 bulk silicate surfaces. In particular, H₂ formation has been experimentally studied on surface of
69 polycrystalline¹⁴ and amorphous silicates.^{15,16,17} Quantum chemistry based computational modelling
70 can help to augment our understanding of these systems by providing uniquely detailed insights into
71 many properties of silicates, such as: atomic and electronic structures, thermodynamic stabilities,

72 vibrational spectra, absorption spectra, formation pathways, and chemical reactivity.¹⁸ Previously, via
73 such a computational chemistry approach, the thermodynamic properties of many step-wise additive
74 cluster intermediates based predominately on the abundant Mg, SiO, H₂O species were
75 determined^{19,20}, providing an unprecedented level of detail into nucleation pathways relevant to
76 silicate dust formation in various astronomical environments. Chemical reactivities of silicate
77 nanoclusters towards H₂ and H₂O have also been studied by computational quantum chemical
78 modelling.^{12,21} Addressing the issue of oxygen depletion in dense interstellar clouds it was found that
79 a stable nanopyroxene cluster with a (MgO)₄(SiO₂)₄ composition, could act as an oxygen sink by
80 being significantly overoxidized by water molecules via the formation of surface hydroxyl groups.¹²
81 With respect to the subject of the present work, for one nanosilicate cluster isomer having the
82 (MgO)₆(SiO₂)₃ composition, calculations have also confirmed that nanosilicates could offer several
83 routes to H₂ formation and dissociation.²¹ Similar calculations have also examined H₂ formation on the
84 regular (010) surface of forsterite (i.e. the crystalline bulk silicate with SiO₂(MgO)₂ stoichiometry)
85 using an embedded cluster approach²² and periodic models.^{8,23,24} We note that the majority of silicate
86 dust grains in the ISM are amorphous and modelling using regular crystalline surfaces may not be the
87 most structurally appropriate approach for such systems. On the other hand the nanosilicates that we
88 study herein are so small as to be inherently non-crystalline and can thus be also regarded as a way to
89 simulate the effects of amorphicity within a small model system.

90 In a previous study²¹, using the most energetically stable (MgO)₆(SiO₂)₃ cluster isomer
91 structure, we considered a single mechanism for H₂ formation and dissociation which involved two
92 types of chemisorbed H atoms: (i) O-H⁺_{chem} protonation of an oxygen anion site, together with (ii) H⁻
93_{chem} forming a Mg-H hydride bond. These two sites were found to be associated with one another as
94 the prior formation of O-H⁺_{chem} would entail transfer of negative charge from the oxygen site to a
95 nearby Mg cation site, where formation of the second H⁻_{chem} would be facilitated (i.e. [H⁺_{chem} → H⁻_{chem}]
96 sequential chemisorption of two separate H atoms, 2H_{chem}). For this cluster and this
97 formation/dissociation mechanism for H₂ we also showed that a Brønsted–Evans–Polanyi (BEP)
98 relation (i.e. where the activation energies of a set of reactions of the same family are proportional to
99 their enthalpies of reaction) existed between the transition state energy of H₂ formation and the [H⁺_{chem}
100 → H⁻_{chem}] adsorption energy of the two H atoms.

101 The aim of the present study is first to establish whether the above behaviour is not only
102 general to the most stable olivine (MgO)₆(SiO₂)₃ cluster isomer (i.e. nano-olivine_c in Fig. 1c), but to
103 other structural isomers of olivine clusters of the same size and composition, and to similar sized
104 nanosilicate clusters of different stoichiometry. For all considered nanoclusters (see Fig. 1) we have
105 calculated the energetic pathways associated with: H adsorption, diffusion, recombination and H₂
106 molecular desorption. With our new results we confirm the generality of the H₂ reaction barrier *versus*
107 2H_{chem} binding energy BEP relation for silicates with 20 data points (including six data points from

our previous study²¹ and a data point from a bulk forsterite surface theoretical study²²). We find that the strength of the $[\text{H}_{\text{chem}}^+ \rightarrow \text{H}_{\text{chem}}^-]$ type adsorption process is linked to the distance between the adsorbed polarised H atoms. We further find new highly strongly bound 2H_{chem} adsorption configurations on some silicate clusters that can be interpreted as being of a new $[\text{H}_{\text{chem}}^- \rightarrow \text{H}_{\text{chem}}^+]$ type rather than the $[\text{H}_{\text{chem}}^+ \rightarrow \text{H}_{\text{chem}}^-]$ as has been previously reported.^{8,21} Corresponding data points from this latter type of adsorption are also found to follow the general BEP relation. In all cases the 2H_{chem} binding energy is also found to be proportional to the binding energy of only the H_{chem}^+ species. Comparison with the BEP relation implies that the latter energetic contribution is correlated to the height of the H_2 reaction barrier. Further analysis of this latter trend, suggests that although the $[\text{H}_{\text{chem}}^+ \rightarrow \text{H}_{\text{chem}}^-]$ type of adsorption will be more prevalent for lower total binding energies of both H atoms, the $[\text{H}_{\text{chem}}^- \rightarrow \text{H}_{\text{chem}}^+]$ type of adsorption process will be more common for larger 2H_{chem} binding energies. This in turn suggests that the former type of adsorption will tend to lead to H_2 formation, and the latter mode to H_2 dissociation. From further considering which types of magnesium silicates tend to promote for which types of 2H_{chem} adsorption modes, we further speculate on how structure (e.g. amorphous, porous, crystalline) and Mg-content could be related to the propensity of silicate grains to form H_2 .

Finally, due to the very small size of the nanosilicate clusters we consider, it is conceivable that highly exothermic reactions could significantly perturb, or even destroy, their structural integrity thus potentially changing, or even removing their influence, on subsequent chemical reactions. In order to check this possibility, we explicitly consider the thermodynamic stability of two of our nanosilicate clusters at a temperature estimated to be typically induced through the energy release of the H_2 formation reaction.

130

131 Methodology

132

We consider a number of Mg-rich olivine and pyroxene nanocluster isomers with compositions $(\text{MgO})_6(\text{SiO}_2)_3$ and $(\text{MgO})_4(\text{SiO}_2)_4$ (i.e. having $(\text{MgO})_{2n}(\text{SiO}_2)_n$ forsterite and $(\text{MgO})_n(\text{SiO}_2)_n$ enstatite stoichiometries, respectively), as shown in Fig. 1. We label these clusters as: nano-olivine_a (Fig. 1a), nano-olivine_b (Fig. 1b) and nano-pyroxene (Fig. 1d). For reference with our previous study,²¹ we also label the lowest energy $(\text{MgO})_6(\text{SiO}_2)$ cluster isomer as nano-olivine_c (Fig. 1c). All these nanosilicate clusters were found through global optimization using the methodology detailed in previous works.^{12,19} Due to their very small size all considered clusters do not yet exhibit bulk crystallinity, but all have complete SiO_4 tetrahedra which are either directly joined together through a bridging oxygen (Si-O-Si) linkage, or separated by Mg cations.

For structural optimisations, and for computing harmonic frequencies and transition states, density functional theory (DFT) calculations have been performed using the Gaussian 09 program.²⁵

143

144 The MPWB1K hybrid meta-functional was employed since it has been proven useful to study both
145 systems with weak interactions as well as providing accurate activation barrier energies.^{21,22,26,27,28}
146 This functional has also been employed in previous studies of hydrogenation processes in the
147 interstellar medium^{12,21,22}, thus allowing for comparison with previous data. For Mg and Si atoms,
148 we used a 6-31G(d) basis set and for O and H atoms we used a 6-31+G(d,p) set of basis functions. In
149 all optimisations, all atom positions in the cluster together with any reactants/products were fully
150 relaxed with no symmetry constraints. The structures resulting from the optimisations were checked
151 to have all positive frequencies and thus be true local energy minima. Transition states were obtained
152 using the synchronous transit-guided quasi-Newton method²⁹ and were all verified to have a single
153 imaginary vibrational frequency along the reaction coordinate (see Supplementary Information for
154 lists of all vibrational frequencies). Using this methodology, the nano-olivine_a and nano-olivine_b
155 nanoclusters are found to be metastable with respect to the ground state nanocluster of this
156 composition (i.e nano-olivine_c) by 39 kJ/mol and 51 kJ/mol, per (MgO)₂(SiO₂) formula unit,
157 respectively. The nano-pyroxene cluster studied is the ground state (MgO)₄(SiO₂)₄ nanocluster isomer.

158 The binding energy of a n separated H atoms with a nanocluster is defined as: $E_{\text{bind}} = -$
159 $[E(\text{cluster} + n\text{H}) - E(n\text{H}) - E(\text{cluster})]$, where $E(\text{cluster} + n\text{H})$ is the energy of n H atoms adsorbed
160 on the cluster, $E(\text{H})$ is the energy of the free H atom and $E(\text{cluster})$ is the energy of the cluster. E_{bind}
161 values and other energies are given in kJ/mol and rounded to the nearest kJ/mol to reflect the typical
162 accuracy of DFT methods. When comparing with previous work where Kelvin has been used as an
163 energy unit, the non-rounded kJ/mol value is converted to Kelvin. We note that we report E_{bind} values
164 that are uncorrected for basis set superposition error (BSSE) so that we can compare our results
165 directly with similar data from previous studies. One previous study has indicated that BSSE is most
166 significant for hydrogen physisorption, where it has been shown that it can lead to overestimates of
167 E_{bind} values by between 10-41%, depending on the physisorption type and basis set employed.⁸ Our
168 calculations of BSSE effects for four different physisorption configurations on two clusters (see
169 Supplementary Information) also concur that the uncorrected E_{bind} values are similarly between 18-
170 45% too large. The majority of our results, however, concern chemisorbed H atom configurations
171 where BSSE effects typically lead to systematically smaller average E_{bind} overestimates. For DFT
172 calculations of H chemisorption on the forsterite surface using a moderately large basis set, BSSE was
173 found to account for 8% of E_{bind} .⁸ In the present study, considering nine chemisorption configurations
174 on two nanoclusters (see Supplementary Information), we find on average that BSSE accounts for
175 ~5% of our reported E_{bind} values. As this effect is relatively small, and, moreover, that we are more
176 concerned with trends in reactivity and E_{bind} values (which would not be affected by such effects), we
177 do not apply this correction to our reported chemisorption values.

178 Canonical *ab initio* molecular dynamics (AIMD) calculations of 28 picoseconds (3 ps of
179 equilibration followed by 25 ps production) at a temperature of 800 K were performed on the nano-

180 olivine_c and the nano-pyroxene clusters employing a 0.5 femtosecond time step and the Bussi-
181 Donadio-Parinello thermostat³⁰. The energies and forces during the AIMD runs were calculated using
182 DFT employing the PBE functional³¹ and a light basis as implemented in the FHI-AIMS code³².

183

184 **Results and Discussion**

185

186 **Physisorption of H atoms**

187 The nano-olivine_a and nano-pyroxene clusters offer a variety of sites for impinging gas phase
188 hydrogen atoms. Three physisorption sites for H atoms were found: (i) P1, atop an Mg cation, (ii) P2,
189 above an Mg-O bond, and (iii) P3, over the centre of rings circumscribed by the atoms MgOSiOMgO
190 or OSiOMgOSiO. We remind that reader BSSE is likely to account for ~3-6 kJ/mol of both our
191 reported E_{bind} values, and the E_{bind} values from the literature that we compare our data with.

192 For the nano-pyroxene cluster, only P2 sites were found. Fig. 2a and Fig. 2.b show the
193 physisorption of a H atom with distances of 2.06 Å and 1.94 Å from a Mg cation and 1.93 Å and 1.80
194 Å from an O anion, respectively. These configurations have quite strong binding energies of 11
195 kJ/mol and 19 kJ/mol, respectively.

196 For the the nano-olivine_a cluster, H atoms can adsorb on P1, P2 and P3 sites. The P2 sites
197 are shown in Fig. 2c and Fig. 2d. Here the H atom is physisorbed over an Mg-O bond with distances
198 of 2.06 Å and 2.03 Å from the Mg cation and 1.94 Å and 1.87 Å from the O anion, respectively. The
199 corresponding E_{bind} values are 9 kJ/mol [1045 K] and 16 kJ/mol [1876 K]. For a similar P2 adsorption
200 site on the nano-olivine_c cluster, a binding energy of 11 kJ/mol [1327 K] was previously reported,²¹
201 while for the crystalline forsterite (010) surface, P2 physisorption binding energies have been
202 calculated to be between 10 kJ/mol²² and 17 kJ/mol.⁸

203 Fig. 2e corresponds to the physisorption of H atom on P1 site of the nano-olivine_a cluster at
204 a distance of 2.47 Å from a Mg cation. This configuration has a fairly low E_{bind} value of 7 kJ/mol [864
205 K] and is very similar in magnitude to that reported for the nano-olivine_c cluster (788 K)²¹. For the
206 forsterite (010) surface, Navarro-Ruiz et al.⁸ found that a H atom can physisorb fairly strongly (14
207 kJ/mol) on top of a Mg cation which is surrounded by three oxygen surface anions. In this case the
208 anions are thought to attract the incoming gaseous H atom and then contribute to increase the non-
209 bonded electrostatic interaction.⁸

210 The strongest physisorption was found for a P3 site on the nano-olivine_a cluster whereby a
211 H atom lays above the centre of an MgOSiOMgO ring, as shown in Fig. 2f with the strongest binding
212 energy (17 kJ/mol). Another P3 site (OSiOMgOSiO ring) was also found with the weakest binding
213 energy (5 kJ/mol) of all (not shown). For this weakly bound configuration, the smallest O-H and Mg-
214 H distances were found to be 2.40 Å and 2.65 Å respectively.

215 The present results can be compared with experimental data reported by Katz et al.³³ and
216 Perets et al.¹⁵. In the latter the energy barriers for atomic H desorption were obtained by fitting
217 temperature programmed desorption curves for polycrystalline³³ and amorphous¹⁵ silicate surfaces.
218 The smallest physisorption binding energy of a single H atom on nano-olivine_a (5 kJ/mol [644K]) is
219 close to the experimental values of 4 kJ/mol [513 K]¹⁵ and 3 kJ/mol [373 K]³³. Generally speaking,
220 however, the calculated physisorption E_{bind} values are higher than that found in experiment. We note
221 that the generally higher values reported above are likely to be overestimates due to BSSE effects
222 which may go some way to help to explain the disagreement. The difference between experimental
223 and theory may also be due to the fact that experimental surface is likely to be partially hydroxylated³⁴
224 which would shield the direct interaction with surface ions and thus attract H atoms less strongly.

225 The transition energies between physisorbed states were calculated in order to study the
226 diffusion behaviour of hydrogen atoms on the nano-olivine_a cluster. The barrier for atomic hydrogen
227 to jump between the configuration which has the highest physisorption binding energy (P3 site in 2f),
228 and the configuration (P1 site in 2e) that has the lowest, is about 15 kJ/mol [1753 K]. The calculated
229 activation energy is less than that found in ref [21] (2398 K) for the analogous H diffusion case, which
230 can be explained by the relatively high strength of the P3 binding energy of physisorbed H atom on
231 the MgOSiOMgO ring (2894 K) in this study. For all presently studied sites, H atoms can diffuse over
232 the surface of the clusters with activation barriers ranging between <1 kJ/mol and 15 kJ/mol. All
233 physisorption binding energies are summarised in Table 1.

234

235 **Chemisorption of H atoms**

236 Chemisorption of a single H atom can occur at two different types of site: (i) C1-type, by attachment
237 of H^+_{chem} to an O anion, and (ii) C2-type, by bonding of H^+_{chem} to a MgOMgO ring. On all C1 sites of
238 the nano-pyroxene cluster and the majority of those on the nano-olivine_a cluster, the H atom
239 chemisorbs to an oxygen anion that is bonded to a Si cation. Due to the resulting negative charge
240 transfer to a nearby Mg cation, this type of chemisorption leads to structural rearrangement
241 (stretching/breaking of a nearby Mg-O bond). H^+_{chem} species at this type of C1 chemisorption site on
242 the nano-pyroxene cluster have binding energies between 69-200 kJ/mol [8262-24106 K] (see Fig. 3).
243 These results are in line with previous calculations using a slightly different DFT set-up which
244 reported that a H atoms can chemisorb on the nanopyroxene cluster with a range between 9284 and
245 23000 K¹². The hydrogen adsorption process on a C1 site of the nano-olivine_a cluster gives a range
246 of binding energies between 43-272 kJ/mol [5120-32698 K] which lay above and below those
247 calculated for the nano-olivine_c cluster²¹ (adsorption energy range between 9115 and 22640 K) and
248 for the foresterite (010) surface (68⁸, 101²² and 78²⁴ kJ/mol). The strongest C1-type binding energies
249 of H^+_{chem} , however, were found for configurations corresponding to Fig. 4g in this work and to Fig. 3f

250 in ref [21]. In both these configurations, a MgOHMg group is formed instead of a SiOH group and no
251 Mg-O bond is broken.

252 We performed a Mulliken atomic charge population analysis before and after chemisorption
253 for all C1 configurations in order to identify the Mg cation that receives the negative charge (labelled
254 in Fig. 3 and Fig. 4 by black arrows). We note that although Mulliken partitioned atomic charges are
255 not particularly accurate when considering how they reproduce observables such as molecular dipole
256 moments, as long as a consistent basis set is used throughout, such charges are known to be sufficient
257 to capture trends with respect to changes in the extent of charge transfer. For those C1 configurations
258 corresponding to Figs. 3c-d and 3g (i.e. on the nano-pyroxene cluster) and Figs. 4b, 4d-e, 4g (i.e. on
259 the nano-olivine_a cluster) and 4h (i.e. on the nano-olivine_b cluster), the negative charge localization
260 occurs at the Mg cation bonded to the pre-protonated oxygen anion, resulting in a reduction in the Mg
261 cation positive charge by between +0.34e and +0.45e. Note that the cluster structures in Figs 3 and 4
262 show the optimised final C1 states in which the Mg-O bond has often been broken upon charge
263 transfer. For the remaining C1 configurations, the donation of negative charge from the O anion after
264 the chemisorption of the H atom, is made to a Mg cation that is not the one bonded to the pre-
265 protonated O atom but is located on the periphery of the nanocluster and linked to two O atoms. In
266 these cases a relatively larger decrease in the Mg positive charge is found (between +0.48e and
267 +0.58e) as measured by differences in Mulliken partitioned charges.

268 The conversion of physisorbed H atoms to chemisorbed H atoms was considered through
269 studying the transition from P3 sites to C1 sites on the nano-olivine_a cluster. The largest activation
270 energy barrier for physisorption-to-chemisorption was found to be 8 kJ/mol (1013 K). This
271 corresponds to the transfer of an H atom from the P3 configuration having the strongest physisorption
272 binding energy (Fig. 2f), to the C1 site depicted in Fig. 4d and is slightly higher than that found
273 previously on the nano-olivine_c cluster (855 K).²¹ The transition process from sites of type P2 to
274 sites of type C1 has been the subject of previous studies of H₂ formation on the (010) surface of
275 forsterite.^{22,8} where the reported activation barriers (10-27 kJ/mol) are greater than we find (4
276 kJ/mol).

277 In all our considered cases as the released energy due to physisorption is greater than the
278 activation energy, the barrier to chemisorption can be easily overcome. Therefore, we conclude that
279 hydrogen chemisorption at C1 sites via physisorbed sites on the nano-olivine_a cluster is effectively
280 barrierless. For configurations corresponding to Figs. 4a, 4c, 4f and 4g, impinging gas-phase H atoms
281 can directly chemisorb in a barrierless manner. Conversely, Downing *et al*²⁴ found a barrier of 14
282 kJ/mol for direct chemisorption of a H atom on a C1 site on the (010) forsterite surface. Although this
283 latter study and the present work use different DFT calculation set-ups, the presence or not of a barrier
284 is likely due to the significant difference in structure between the small amorphous-like olivine_a
285 cluster and the extended crystalline silicate surface.

286 H atoms can also directly adsorb at C2 sites on the nano-olivine_a cluster without any barrier
287 above the centre of a MgOMgO ring (see Fig. 5d) and a MgOMgOMgO ring (see Fig. 5e). The
288 binding energies of an H atom at these two sites are 91 kJ/mol and 105 kJ/mol, respectively, which
289 are in a comparable range to those found for the C1 site (see Table. 2). The (MgO)_n rings appear to
290 trap the H atoms as H_{chem} species through strong interactions with at least two Mg atoms, which
291 synergistically attract it. H at the smaller ring C2 site has Mg-H bonds of 1.874 and 2.092 Å, while at
292 the larger ring C2 site the two smallest Mg-H distances are 1.844 Å, 1.951 Å, with one larger distance
293 of 3.18 Å. The analysis of Mulliken atomic partitioned charges shows that the adsorbed H atom is
294 negatively charged (-0.18e and -0.11e for the configurations shown in Figs. 5d and 5e respectively)
295 which confirms that a hydride like Mg-H bond is formed. Downing et al.²⁴ predicted that H atoms
296 could chemisorb above single Mg cation on the forsterite surface (010) with adsorption energy of 43
297 kJ/mol. The fact that this value is significantly lower than that found in this work is probably due to
298 the relative inaccessibility of Mg cations on the (010) forsterite surface where, at most, one Mg cation
299 is available for any one H atom. Conversely, on silicate nanoclusters, where most of the atoms are at
300 the surface, and which structurally do not have the regular dense crystalline packing of forsterite, it is
301 easier to find more than one accessible Mg cation site in close proximity. This rationale would also
302 tend to favour C2 sites occurring more frequently on more Mg-rich silicate nanoclusters (i.e. higher
303 Mg:Si ratios) as is confirmed by our results where C2 sites are only found on the 2:1 (MgO)₆(SiO₂)₃
304 nano-olivine_a cluster and not on the 1:1 (MgO)₄(SiO₂)₄ nano-pyroxene cluster.

305 As only two examples of a C2 site were found on the nano-olivine_a cluster, we examined
306 other possible clusters for this site. Firstly, on the nano-olivine_b cluster (see Fig. 1b) we found that a
307 H atom can be adsorbed close to a MgOMgOSiO ring where it lays most closely to two Mg atoms
308 with Mg-H distances of 2.01 and 2.06 Å (see Fig. 5c). The H atom at this site has a E_{bind} value of 38
309 kJ/mol and the analysis of Mulliken atomic charge on the Mg sites confirms the formation of hydride
310 Mg-H bonding. We thus conclude that this is another example of a C2 site. For the nano-olivine_c
311 cluster, the only configuration with a chemisorbed hydride-like H_{chem} species we could find is
312 illustrated in Fig. 5a. Clearly this H adsorption configuration produces a structural rearrangement (i.e.
313 a broken Mg-O bond) compared to the relaxed structure in Fig. 1c. Our calculations show that the
314 formation of such a C2 site requires energy, resulting in a negative E_{bind} value of -74 kJ/mol with
315 respect to the original relaxed bare cluster. Going to the extreme case of Mg-rich clusters which
316 exhibit no Si atoms, we also attempted to find C2 H_{chem} binding sites on the ground state (MgO)₆
317 cluster. We found that H atoms can chemisorb at a C2 site (see Fig. 5b) but, as in the nano-olivine_c
318 case, this reaction is endothermic (E_{bind} of -25 kJ/mol) since it needs energy to break a Mg-O bond.

319 In both the above cases where the C2 chemisorption is endothermic, the original cluster is the
320 most stable one known for its composition and is thus tend to be structurally compact. Conversely, for
321 the two cases where exothermic C2 adsorption is found, the clusters are slightly metastable with
322 respect to the respective ground state cluster isomer and have accordingly more open structures.

323 Overall, based on these preliminary data, we tentatively suggest that C2-type adsorption could be
324 most energetically favoured on moderately Mg-rich silicate dust grains (e.g. forsteritic composition)
325 with non-compact (e.g. porous and/or amorphous rather than crystalline) surfaces.

326

327

328 **Adsorption of a second H atom**

329 A second gas phase H atom can adsorb on our nanosilicate clusters without a barrier. The binding
330 energies of the second H atom are all considerably larger than that of the first chemisorbed H atom
331 (see Table 2). In all cases of clusters possessing a H atom chemisorbed at a C1 (i.e. H_{chem}^+) site, the
332 transfer of negative charge to a single Mg cation results in a subsequent H atom forming a Mg^+-H^-
333 (C3) hydride species (see Fig. 6 and Fig. 7 for nanopyroxene and nano-olivine_a clusters,
334 respectively) at the Mg site to which the charge is transferred. The charge transfer to a Mg cation due
335 to the formation of a C1 site, and the subsequent energetically favourable adsorption of H in a C3
336 manner at this electron-accepting Mg site has been confirmed in other studies.^{8,21,22} This C1/C3
337 $2H_{\text{chem}}$ adsorption mode is stabilized by favorable electrostatic interactions between the $H_{\text{chem}}^+H_{\text{chem}}^-$
338 surface species. Using our relatively large data set, in Fig. 8, we plot the binding energy of the second
339 H_{chem}^- atom to the respective cluster with one H_{chem}^+ atom already adsorbed at a C1 site, versus the
340 distance between the two adsorbed H atoms. Our results reveal two trends for C1/C3 $2H_{\text{chem}}$
341 adsorption. Firstly, for H-H distances less than approximately 4.8 Å, the binding energy of the second
342 atom is inversely proportional to the H-H bond distance (i.e. the smaller the distance, the greater the
343 adsorption energy of the second H). In this range of H-H separations the electrostatic interaction
344 stabilises the C1/C3 pair, but more weakly with increasing H-H separation (see blue shaded area in
345 Fig. 8). However, when the H-H distance is larger than ~4.8 Å, the binding energies of the second H
346 atom are relatively stable (see red shaded area in Fig. 8). For these larger H-H separations the
347 electrostatic interaction does not appear to play an important role in energetic stabilization.

348 When the first H atom is chemisorbed on a C2 site, this causes a negative charge decrease on a
349 nearby O atom which then becomes the most energetically favourable site for an incoming gas-phase
350 H atom for protonation. The chemisorption of the second H in this case increases the negative charge
351 on the charge-depleted oxygen centre by between 0.45-0.52e. We note that these configurations (see
352 Figs. 9a-d) have the most strongly bound second H atoms (see Table 2). For 9a and 9b (nano-
353 olivine_a) the second H atoms adsorb in a C1 manner at a site which is very similar to that shown in
354 Fig. 4g, which is the site with the highest E_{bind} value for a single H atom. Similarly, the second C1
355 adsorbed H atom in configuration 9c (nano-olivine_b) is almost exactly the same as that found in Fig.
356 4h, which is the H-adsorbed configuration with the second highest E_{bind} value. For these three cases
357 this C2 \rightarrow C1 (i.e. $[H_{\text{chem}}^- \rightarrow H_{\text{chem}}^+]$) adsorption process is also found to lead to the most strongly
358 bound pair of H atoms. In the case of configuration 9d on nano-olivine_c, although the second C1-
359 type H chemisorption has a relatively high E_{bind} value, the first C2-type adsorption is endothermic due

360 to the disruption of the structure of the cluster (Fig. 5a). Thus, in this case the final binding energy of
361 two H atoms is relatively moderated. We note that, although there are several C1 sites that are
362 energetically more favourable for adsorption of a single H atom than the C2 sites discovered herein.
363 However, C1-type H adsorption (e.g. configuration 4g) tends to favour only one Mg atom for
364 subsequent H adsorption (e.g. 4g \rightarrow 7d). Overall such C1 \rightarrow C3 processes yield more weakly bound
365 $2H_{\text{chem}}$ configurations than the $2H_{\text{chem}}$ configurations found to be naturally favoured from C2 \rightarrow C1
366 adsorption (e.g. 9a-c).

367

368 **Formation of H₂**

369 Two chemisorbed H atoms can recombine via a transition state and yield a physisorbed H₂ molecule
370 ($H_{2,\text{phys}}$) that can subsequently desorb to form H₂ in gas phase ($H_{2,\text{g}}$). Several reactions paths were
371 found on the nano-pyroxene and nano-olivine_a clusters. In Table 3, the ($2H_{\text{chem}} \rightarrow H_{2,\text{phys}}$) reaction
372 energy, the transition state energy corresponding to the $\text{TS}[2H_{\text{chem}} \rightarrow H_{2,\text{phys}}]$ reaction barrier, and the
373 $H_{2,\text{phys}} \rightarrow H_{2,\text{g}}$ desorption energy were calculated for reaction paths starting from the configurations of
374 two chemisorbed H atoms shown in Fig. 6, Fig. 7 and Fig. 9. Only for configurations corresponding to
375 Fig. 6c, Fig. 6g and Fig. 7a, and were we unable to locate a $\text{TS}[2H_{\text{chem}} \rightarrow H_{2,\text{phys}}]$ transition state.

376 The most energetically favourable channels for $H_{2,\text{phys}}$ formation correspond to reactions
377 starting from $2H_{\text{chem}}$ configurations shown in Figs. 6d-f for the nano-pyroxene cluster, those shown in
378 Figs. 7a-b, 7f and 7g for the nano-olivine_a cluster. These configurations generally lead to exothermic
379 $H_{2,\text{g}}$ formation, except for those in Figs. 6f and 7f, where the energy required for the desorption of the
380 $H_{2,\text{phys}}$ is higher than the energy released from two chemisorbed H atoms to form physisorbed H₂. The
381 remaining configurations are precursor states to endothermic H₂ formation (see Table. 3).

382 In Fig. 10, we plot the most exothermic and endothermic energy profiles leading to $H_{2,\text{g}}$
383 formation starting from configurations shown in Fig. 6, Fig. 7 and Fig. 9. Following ref 21 we take the
384 sum of the energies of the bare relaxed nano-olivine_c cluster and a $H_{2,\text{g}}$ molecule as a common zero
385 energy. For the nano-pyroxene cluster, reaction paths from configurations 6d and 6e have the highest
386 exothermicities of -94 kJ/mol and -71 kJ/mol, the lowest activation energies of 8 kJ/mol and 45
387 kJ/mol, and the smallest $2H_{\text{chem}}$ binding energies of 380 kJ/mol and 403 kJ/mol, each respectively. For
388 the nano-olivine_a cluster, the most energetically favourable reaction paths to form $H_{2,\text{g}}$ start from
389 configurations 7a and 7g. These two paths have the lowest total $2H_{\text{chem}}$ binding energies of 352 kJ/mol
390 and 426 kJ/mol, respectively, and the highest exothermicities of -123 kJ/mol and -44 kJ/mol,
391 respectively. Configuration 7g also leads to the lowest calculated barrier (22 kJ/mol) of all reaction
392 paths considered. As noted above, a transition state could not be located for reactions starting from
393 configuration 7a, which leads to the most efficient catalytic path for $H_{2,\text{g}}$ formation. $H_{2,\text{g}}$ formation is
394 generally more favoured on the nano-olivine_a cluster compared with the nano-olivine_c cluster, the
395 nano-pyroxene cluster and the foresterite surface.²² For all nanosilicate clusters studied, we have

396 found at least one H_2 formation route that is more energetically favourable than that of the
397 corresponding bulk surface reaction pathway.²²

398 The remaining reaction paths discovered on the clusters are endothermic and they therefore
399 inhibit $H_{2,g}$ formation. They have the strongest total binding energy of two chemisorbed H atoms and,
400 correspondingly, also have a large activation energy (see Table. 3) required to break the Mg^+-H^- bond
401 and to react with a chemisorbed H^+_{chem} . We plot also in Fig. 10 the most and the least reactive profiles
402 relative to $H_{2,g}$ formation on nano-olivine_c (we refer the reader to Fig. 5 in ref [21]). We found that
403 the nano-olivine_a has more sites to catalyse the dissociation of $H_{2,g}$ than the nano-olivine_c.

404 In previous work considering a few reaction paths on the nano-olivine_c cluster,²¹ it was
405 suggested that the H_2 dissociation reaction followed a BEP relation. In order to confirm whether this
406 BEP relation is more general to other bare silicate grains with different structures and compositions,
407 in Fig. 11 we plot the energy of the $TS[2H_{chem} \rightarrow H_{2,phys}]$ transition state barrier height as function of
408 the $2H_{chem}$ binding energy atoms for all calculated reaction paths. Data from Kerkeni and Bromley²¹
409 corresponding to the nano-olivine_c cluster and from Goumans et al.²² corresponding to foresterite
410 surface (010) are also plotted. The results show that all data points follow the same trend whereby the
411 barrier height is linearly dependent on the total $2H_{chem}$ binding energy, clearly extending the BEP
412 relation to all new reactions and clusters studied in the present work. This result is particularly
413 significant as the present work includes: (i) non-olivinic type clusters (e.g. nano-pyroxene), (ii)
414 clusters which are metastable with respect to the ground state cluster isomers for their composition
415 (e.g. nano-olivine_a and nano-olivine_b), and (iii) clusters with particularly strongly chemisorbed H
416 atoms that can be interpreted as being formed through a novel $[H^-_{chem} \rightarrow H^+_{chem}]$ (i.e. C2 \rightarrow C1) type
417 of adsorption process. Taken together, all data in Fig. 11 thus provide more evidence to support the
418 claim that H_2 dissociation on bare magnesium silicates follows a general BEP relation independently
419 of silicate dust grain, shape, size, crystallinity and composition.

420 Of all the reaction energy profiles studied, those corresponding to configurations 7d, 9a and
421 9b are the most energetically favourable for the dissociation of molecular hydrogen. The differences
422 in the exothermicity of these pathways with respect to H_2 dissociation appear to be linked to relatively
423 high $2H_{chem}$ binding energies. Although configuration 7d, exhibits a C1/C3 type of $2H_{chem}$ adsorption,
424 the reaction paths corresponding to configurations 9a and 9b start from C2/C1 H_{chem} adsorption. This
425 observation suggests that reaction paths starting by H adsorption on Mg-O rich areas (C2 sites) on
426 silicate dust grains may be more likely to inhibit $H_{2,g}$ formation. To further investigate this possibility,
427 in Fig. 12, we plot the total binding energy of two chemisorbed H atoms on their host cluster as
428 function of the binding energy of only the H^+_{chem} atom of the corresponding $2H_{chem}$ pair.
429 Corresponding data from references 21 and 22 are also included. We also include data calculated for
430 C2/C1 adsorption of two H atoms on the $(MgO)_6$ cluster, the nano-olivine_b and nano-olivine_c
431 clusters based on the initial C2-type H adsorptions shown in Fig. 5. The plot clearly shows a linear

432 relation between the two energies whereby the greater the binding energy of H^+_{chem} , the stronger the
433 $2H_{\text{chem}}$ binding energy. This trend is clearly similar to the BEP relation in Fig. 10, suggesting that the
434 H^+_{chem} binding energy is a good indicator of the relative $\text{TS}[2H_{\text{chem}} \rightarrow H_{2,\text{phys}}]$ transition state barrier
435 height. (this relation is further confirmed in a plot of the $\text{TS}[2H_{\text{chem}} \rightarrow H_{2,\text{phys}}]$ transition state barrier
436 height versus the H^+_{chem} binding energy – see Supplementary Information). The relatively extended
437 data set in the plot of Fig. 12 allows us to analyse further the trends in H_2 dissociation/formation with
438 respect to $2H_{\text{chem}}$ adsorption. In the inset to Fig. 12 we select ten data points from the main graph: five
439 points corresponding to C1/C3 adsorption, and five points corresponding to C2/C1 adsorption. This
440 selection of points highlights that for smaller H^+_{chem} binding energies the C1/C3 mode of $2H_{\text{chem}}$
441 adsorption yields a higher total binding energy, whereas for higher H^+_{chem} binding energies the C2/C1
442 mode tends to be energetically more favoured. The linear fits to the two sets of five points in the inset
443 to Fig. 12 highlight these trends. In turn, these tendencies indicate that H_2 formation (i.e. smaller
444 $\text{TS}[2H_{\text{chem}} \rightarrow H_{2,\text{phys}}]$ transition state barriers) will tend to be via C1/C3 adsorbed H atoms, and H_2
445 dissociation will tend to proceed more via C2/C1 adsorbed H atom configurations. We further see
446 from our results that strong C2/C1 adsorption can occur on relatively Mg-rich olivinic clusters with
447 less compact structures, but seems not to be common on comparatively Mg-poor pyroxene clusters.
448 Overall, we may speculate that amorphous/porous grains with forsteritic composition are more likely
449 to dissociate H_2 molecules, and that less Mg-rich silicate grains (e.g. enstatite composition) or more
450 crystalline/compact silicate grains in general are more likely to assist in the formation of H_2 from
451 adsorbed H atoms.

452 Finally, we assess the thermal stability of the clusters taking into account the likely
453 temperature increase due to exothermic H_2 formation. Here we assume that one H_2 molecule is formed
454 at a time and that half of the released energy (-217 kJ/mol) goes into heating the dust and the
455 remainder into exciting the degrees of freedom of the H_2 molecule. In order to calculate the
456 temperature rise we use the heat capacity of the cluster in question which can be estimated from a
457 product of the contributions from each harmonic vibrational mode. The constant volume heat
458 capacity, C_v was thus calculated from vibrational frequencies analysis for the nano-olivine_c and the
459 nanopyroxene clusters, and found to be $71.6 \text{ cal mol}^{-1} \text{ K}^{-1}$ and $65.8 \text{ cal mol}^{-1} \text{ K}^{-1}$, respectively. From
460 these data we can calculate that H_2 formation leads to an approximate increase of the temperatures of the
461 nano-olivine_c and nanopyroxene clusters by 725 K and 789 K, respectively. Using a similar
462 approach, Navarro-Ruiz et al.⁸ estimated that H_2 formation on (010) forsterite surface would induce a
463 bulk temperature increase of only 53 K. As our temperature increase for an ultrasmall silicate dust
464 grain is significantly higher than that expected for a large bulk-like dust grain, we have explicitly
465 studied the thermal stability of our clusters at 800 K for 25 ps using AIMD calculations. In the
466 Supplementary Information we show various typical snapshots of the structures of the two clusters
467 from these calculations. Although in each case the clusters are distorted from their 0 K optimised

468 structures the clusters do not break apart or melt. We note that the appearance of bond being broken or
469 formed in the snapshots is largely due to arbitrary bond length cut-offs in the visualisation rather than
470 any true chemical re-ordering. The closeness of the thermally distorted structures to their respective 0
471 K optimised structure was tested by taking a wide selection of snapshots and optimising their
472 structures to the nearest local minimum. In nearly all cases the original 0 K structure, or an energy
473 equivalent symmetry-transformed isomer, was found. Examples of structures leading to the latter can
474 be seen at 22.11 ps for the nano_olivine_c clusters and at 24.47 ps for the nano_pyroxene cluster. In
475 very few cases a slightly more symmetric, though slightly metastable isomer, was also found, having
476 an overall structure which was very similar to the original non-symmetric structure. Snapshots at
477 10.39 ps for nano_olivine_c and at 12.23 ps for nano_pyroxene (see Fig. 14) show cluster structures
478 that yielded these slightly metastable symmetric isomers. In addition, we made a quantitative estimate
479 of the thermally induced structural disorder by calculating the relative root mean square interionic
480 distance fluctuations employing a cluster-adapted Lindemann index (δ), see e.g., Ref. ³⁵. Typically
481 one expects $\delta \geq 0.1$ for systems that are melted, or at least significantly thermally distorted. Although
482 not a rigorous criterion, it is most likely that misleading large values of δ occur through pre-melting
483 effects³⁶ rather than a small δ values incorrectly pointing to non-melting. In Fig. 14 we give the total
484 Lindemann index (δ_{Total}) taking into account all atoms of the respective cluster and also the
485 Lindemann index from only taking into account atom of the same type in a cluster (δ_X , where X=Si, O,
486 Mg). Here we see both $\delta_{\text{Total}} < 0.07$ and $\delta_X < 0.07$ clearly confirming the non-melted character of each
487 cluster and the relatively low degree of structural disorder at 800 K. In each case we also see that $\delta_{\text{Si}} <$
488 δ_{Mg} and $\delta_{\text{Si}} < \delta_{\text{O}}$ showing that the Si atoms are the least perturbed atom type, in line with their 4-
489 coordinated strong directional Si-O bonding which is never seen to be broken during the AIMD runs.
490 The more isotropic ionic interactions between Mg and O ions allows for more structural distortion
491 which allows for the small structural changes observed in each cluster. Following the analysis in ref 8,
492 we expect the radiative dissipation of the heat released into small clusters by H₂ formation to be very
493 rapid.

494

495 Conclusions

496

497 We have performed extensive DFT calculations to investigate the trends in H atom adsorption
498 and H₂ formation and dissociation on nanoscale magnesium silicate clusters as a means to help
499 understand H-based chemistry on silicate dust grains in the ISM. We consider small inherently non-
500 crystalline clusters of both olivine and pyroxene composition and calculate H adsorption energies and
501 barrier heights for the $2\text{H}_{\text{chem}} \rightarrow \text{H}_{2,\text{phys}}$ reaction. We find two 2H_{chem} types: (i) C1/C3 type which
502 comes about through a $\text{H}_{\text{chem}}^+ \rightarrow \text{H}_{\text{chem}}^-$ process, and (ii) C2/C1 type which results from a $\text{H}_{\text{chem}}^- \rightarrow$
503 H_{chem}^+ adsorption sequence. In both cases the resultant 2H_{chem} binding energies with the cluster are

504 found to be proportional to the corresponding $2H_{\text{chem}} \rightarrow H_{2,\text{phys}}$ transition state barrier for all cases
505 considered confirming the generality of a BEP relation for such systems. Further, considering a
506 related relation between the binding energy of only the H^+_{chem} with the bare cluster and the total
507 binding energy of the two H atoms with the cluster, we suggest that there is a crossover between
508 C1/C3 adsorption which promotes H_2 formation for smaller $2H_{\text{chem}}$ binding energies, and C2/C1
509 adsorption which favours H_2 dissociation for relatively higher $2H_{\text{chem}}$ binding energies. Based on
510 where such types of adsorption modes have been found, we speculate that amorphous/porous grains
511 with forsteritic composition are more likely to dissociate H_2 , and that relatively Mg-poor silicate
512 grains (e.g. enstatite composition) or more crystalline/compact silicate grains in general would tend to
513 catalyse H_2 formation. By directly simulating the dynamics of two example nanosilicate clusters at
514 800 K, our ultrasmall grains are assessed to have high thermal structural stability with respect to the
515 heat released during exothermic H_2 formation reactions.

516

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523 We also thank Fedor Goumans for allowing access to some initial structures for C1-type H_{chem}
524 configurations on the nano-pyroxene cluster.

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528 Table 1. Binding energies (kJ/mol) of physisorbed H atoms on the nano-pyroxene and nano-olivine_a
529 clusters. Note that both our own data and cited literature data relate to calculated values that have not
530 been corrected for BSSE effects and are thus typically overestimated by 3-6 kJ/mol (see main text,
531 Supplementary Information and ref. 8). Values in brackets are in Kelvin.
532

Nano-cluster	Physisorption type	Previous Calculations	This work
Nano-pyroxene	P2	-	11 [1316], 19 [2312]
Nano-olivine_a	P1	[788], [1661]	7 [864] (Fig.2e)
	P2	[1327], [2063], [1240]	9 [1045] (Fig.2c), 16 [1876] (Fig.2d)
	P3	[1036-2894]	17 [1999] (Fig.2f), 5 [644]

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534 ^a Kerkeni and Bromley²¹

535 [¶] Navarro-Ruiz et al.⁸

536 ^γ Goumans et al.²²

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Table 2. Binding energies (kJ/mol) of single chemisorbed H atoms corresponding to the configurations shown in Fig. 3 for the nano-pyroxene cluster, and in Fig. 4 and Fig. 5 for the nano-olivine_a, nano-olivine_b and nano-olivine_c clusters, are listed in column 3. Column 5 shows the chemisorption energy of the second H atom (kJ/mol) for each configuration in column 4 with respect to the initial configurations listed in column 2 (following labelling in Figs. 3-5). The $2H_{\text{chem}}$ chemisorption type is listed in column 6. Column 7 lists the total $2H_{\text{chem}}$ binding energy (kJ/mol) with respect to the corresponding bare relaxed cluster.

Nano-cluster	Configuration of the first H_{chem}	E_{bind} of the first H_{chem}	Configuration of the second H_{chem}	E_{bind} of the second H_{chem}	$2H_{\text{chem}}$ chemisorption type	Binding Energy (E_{bind})
Nano-pyroxene	Fig. 3a	69	Fig. 6d	311	C1/C3	380
	Fig. 3b	91	Fig. 6e	311	C1/C3	403
	Fig. 3c	145	Fig. 6f	311	C1/C3	456
	Fig. 3d	162	Fig. 6a	304	C1/C3	466
	Fig. 3e	177	Fig. 6c	311	C1/C3	488
	Fig. 3f	196	Fig. 6g	312	C1/C3	508
	Fig. 3g	200	Fig. 6b	310	C1/C3	511
Nano-olivine_a	Fig. 4a	43	Fig. 7a	309	C1/C3	352
	Fig. 4b	92	Fig. 7g	334	C1/C3	426
	Fig. 4c	121	Fig. 7b	310	C1/C3	431
	Fig. 4d	146	Fig. 7f	324	C1/C3	470
	Fig. 4e	188	Fig. 7c	310	C1/C3	499
	Fig. 4f	208	Fig. 7e	311	C1/C3	519
	Fig. 4g	272	Fig. 7d	311	C1/C3	582
Nano-olivine_b	Fig. 5d	91	Fig. 9b	533	C2/C1	625
	Fig. 5e	105	Fig. 9a	529	C2/C1	634
Nano-olivine_b	Fig. 4h	253	Fig. 7h	334	C1/C3	588
	Fig. 5c	38	Fig. 9c	555	C2/C1	593
Nano-olivine_c	Fig. 3, Ref. 21	180	Fig. 4c, Ref. 21	313	C1/C3	493
	Fig. 5a	-74	Fig. 9d	521	C2/C1	448

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555 Table 3. Reaction barriers, recombination energies and desorption energies (kJ/mol) for $H_{2,g}$ release
556 are given in columns 2, 3 and 4, respectively starting from configurations of two chemisorbed H
557 atoms (following labelling of adsorbed $2H_{chem}$ configurations in Figs. 6, 7 and 9).

558

559

Initial configuration	$2H_{chem} \rightarrow H_{2,phys}$	TS [$2H_{chem} \rightarrow H_{2,phys}$]	$H_{2,phys} \rightarrow H_{2,g}$
Fig. 6a	2	55	9
Fig. 6b	37	102	20
Fig. 6c	14	-	20
Fig. 6d	-94	8	20
Fig. 6e	-71	45	20
Fig. 6f	-18	74	20
Fig. 6g	21	-	33
Fig. 7a	-123	-	20
Fig. 7b	-43	34	20
Fig. 7c	24	90	20
Fig. 7d	114	137	14
Fig. 7e	41	72	24
Fig. 7f	-3	51	20
Fig. 7g	-44	22	16
Fig. 9a	159	188	12
Fig. 9b	163	195	17

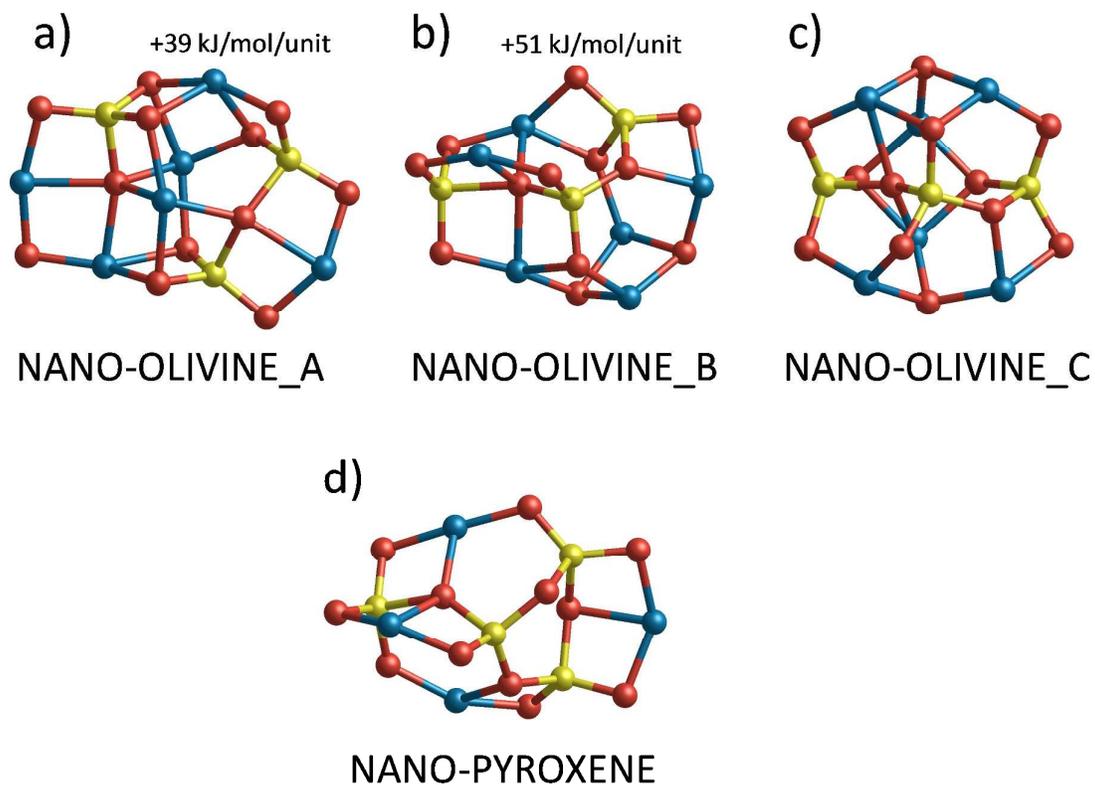
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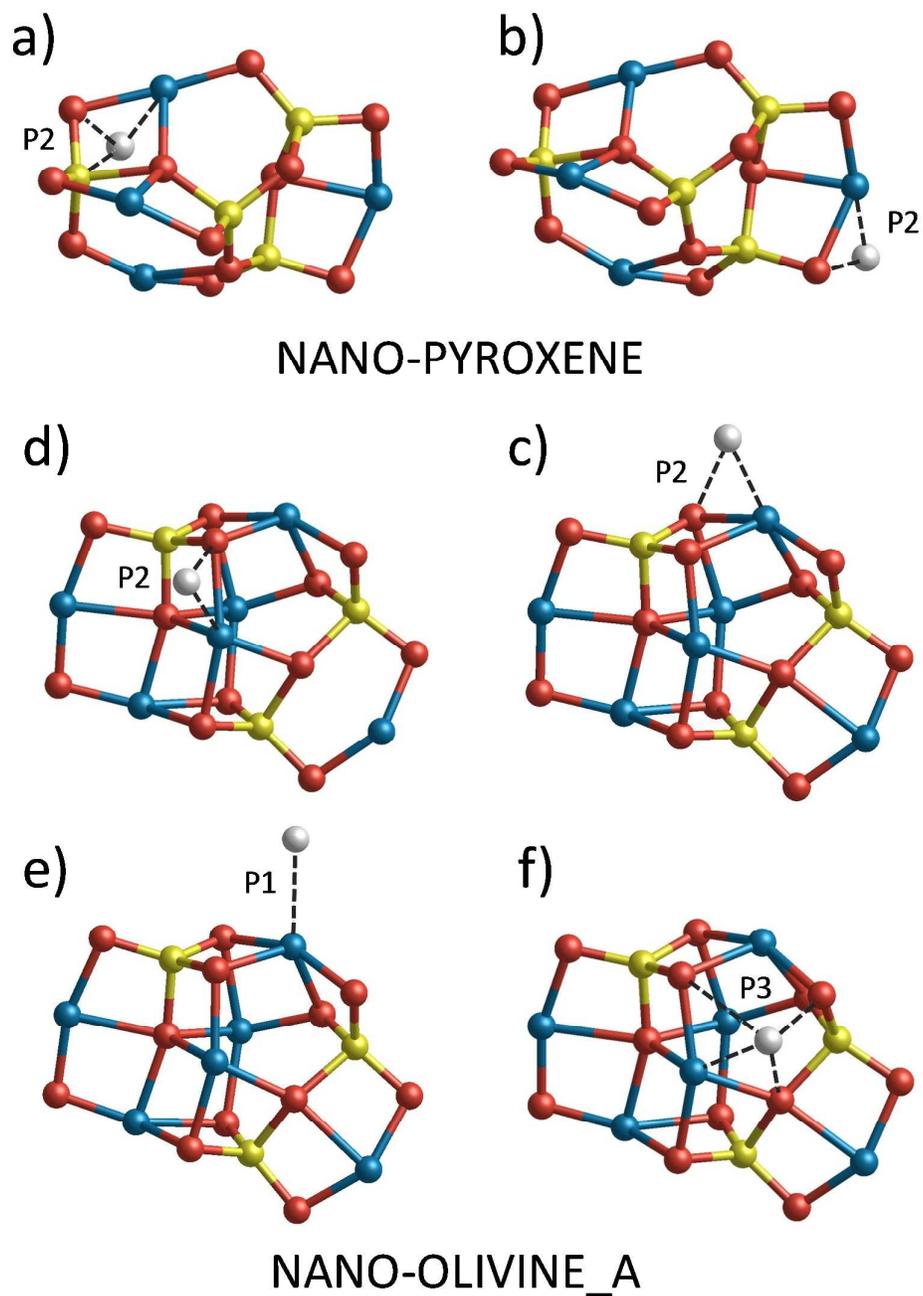
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571 Fig. 1. Optimized structures of: a) nano-olivine_a, b), nano-olivine_b, c) and nano-olivine_c cluster

572 isomers (all with composition $(\text{MgO})_6(\text{SiO}_2)_3$), and d) the nano-pyroxene cluster (composition573 $(\text{MgO})_4(\text{SiO}_2)_4$). The nano-olivine relative energies are per $(\text{MgO})_2(\text{SiO}_2)$ formula unit with respect to

574 the nano-olivine_c ground state. Atom key: Si - yellow, Mg - blue and O - red.

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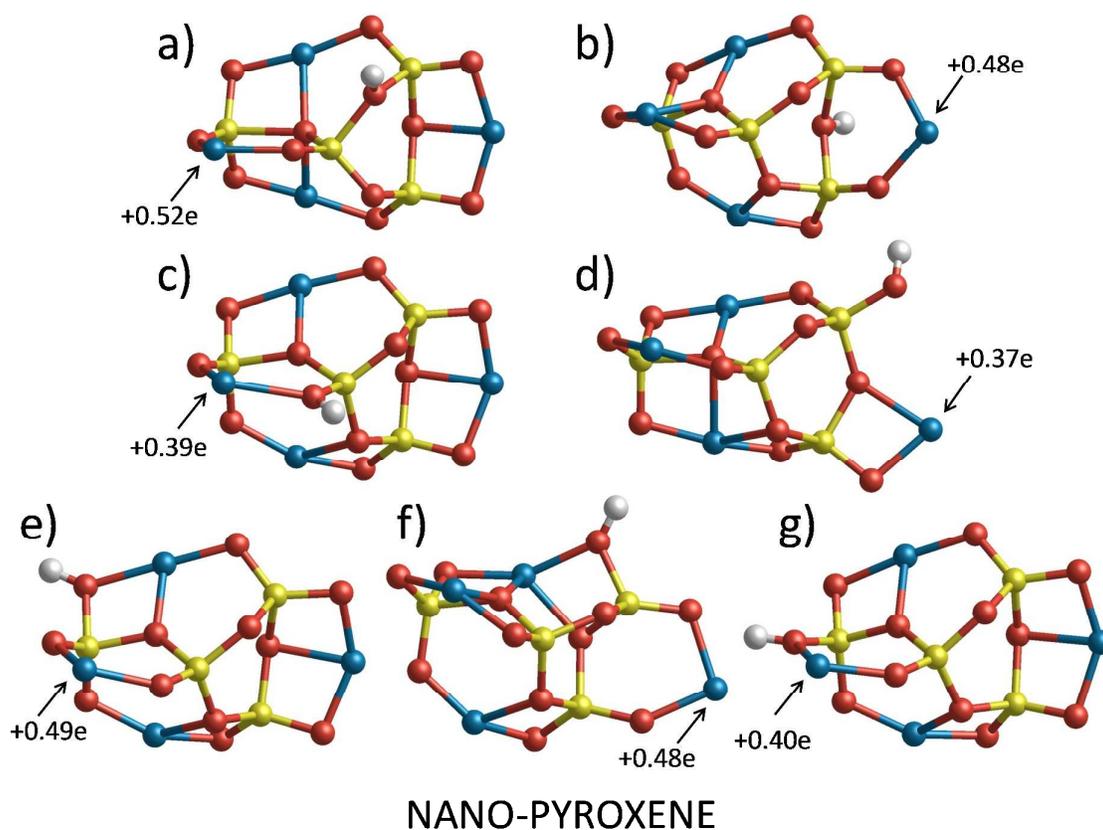
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578 Fig. 2. Physisorption sites (P1, P2 and P3) for a single physisorbed H atom on the $(\text{MgO})_4(\text{SiO}_2)_4$
579 nano-pyroxene cluster (a, b), and on the $(\text{MgO})_6(\text{SiO}_2)_3$ nano-olivine_a cluster (c-f). Grey dashed lines
580 highlight all interatomic distances between the H atom and the respective cluster that are $\leq 2.7\text{\AA}$. Atom
581 key: Si - yellow, Mg - blue and O - red.

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586 Fig. 3. C1-type H chemisorption configurations on the on the $(\text{MgO})_4(\text{SiO}_2)_4$ nanoproxene cluster
587 ordered by increasing E_{bind} values (a-g). The Mg atom that receives the negative charge in each case is
588 indicated by an arrow and the corresponding negative charge increase. Atom key: Si - yellow, Mg -
589 blue and O - red.

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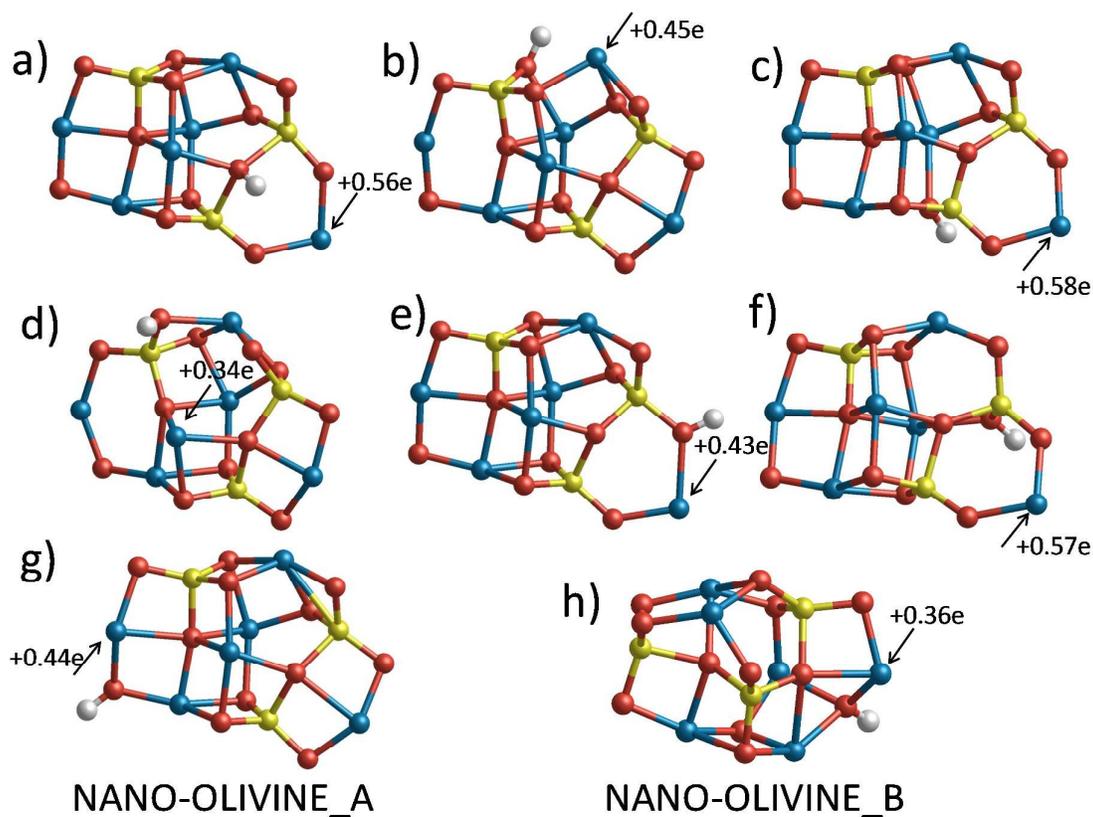
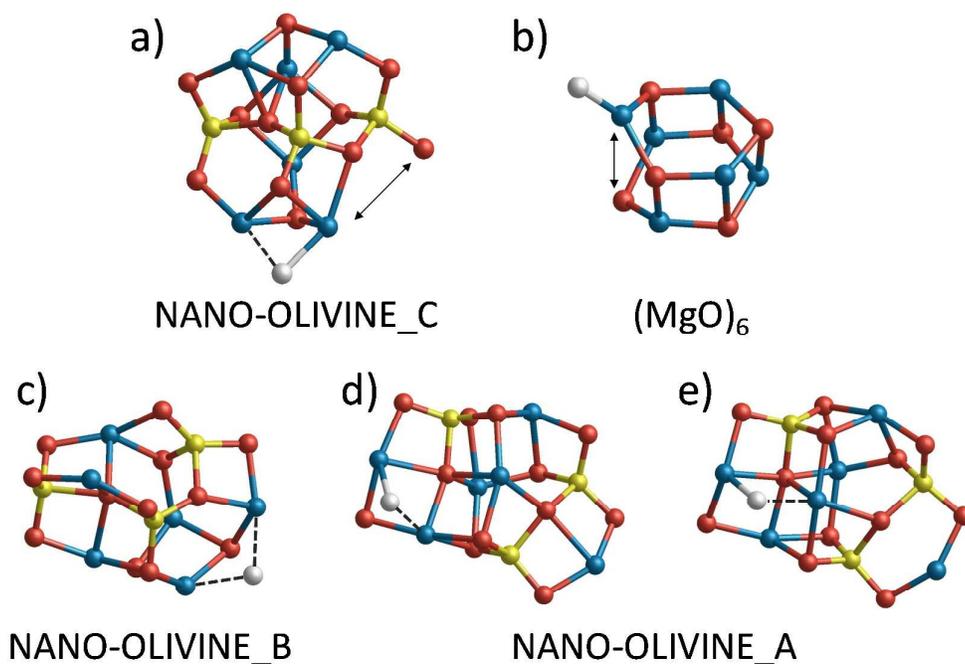
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Fig. 4. C1-type H chemisorption configurations on the $(\text{MgO})_6(\text{SiO}_2)_3$ nano-olivine_a cluster a)-g), ordered by increasing E_{bind} values, and on the $(\text{MgO})_6(\text{SiO}_2)_3$ nano-olivine_b cluster h). The Mg atom that receives the negative charge in each case is indicated by an arrow and the corresponding negative charge increase. Atom key: Si - yellow, Mg - blue and O - red.

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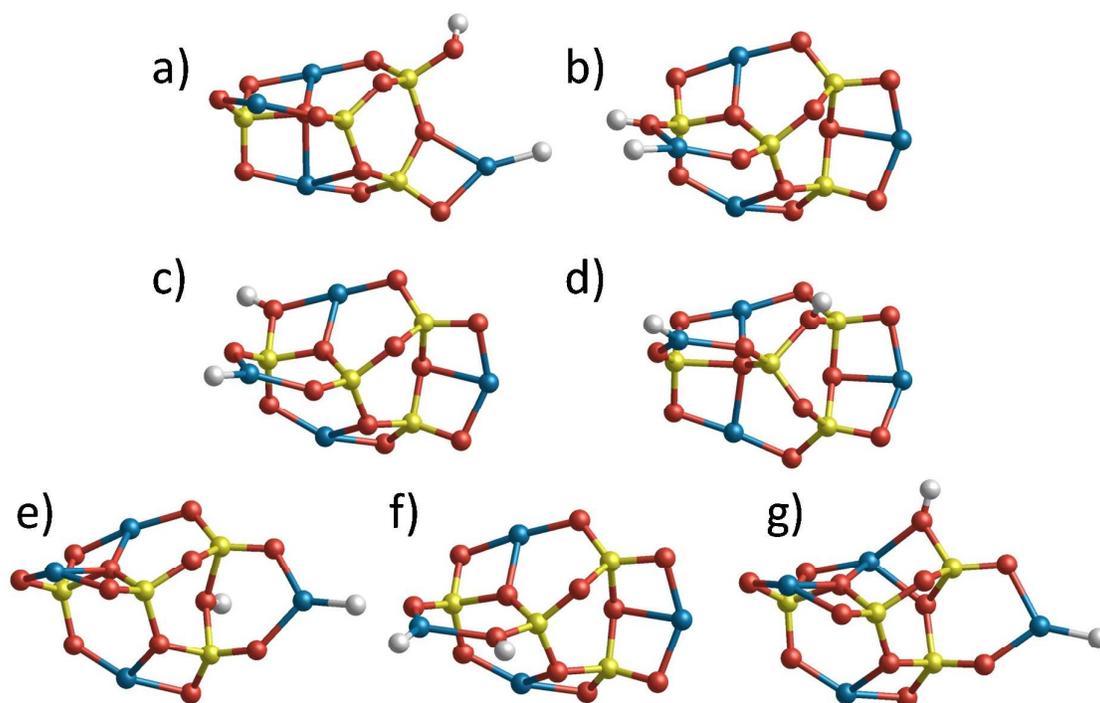


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603 Fig. 5. C2-type H chemisorption configurations on: a), the (MgO)₆(SiO₂)₃ nano-olivine_a cluster, b) a
 604 (MgO)₆ cluster, and c-e) the (MgO)₆(SiO₂)₃ nano-olivine_a clusters. Grey dashed lines highlight all
 605 interatomic distances between the H atom and the respective cluster which are >1.9 Å and ≤2.7 Å. H-
 606 Mg distances which are ≤1.9 Å are drawn as bonds. Double headed arrows indicate Mg-O bond
 607 breakage induced upon H adsorption. Atom key: Si - yellow, Mg - blue and O - red.

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NANO-PYROXENE

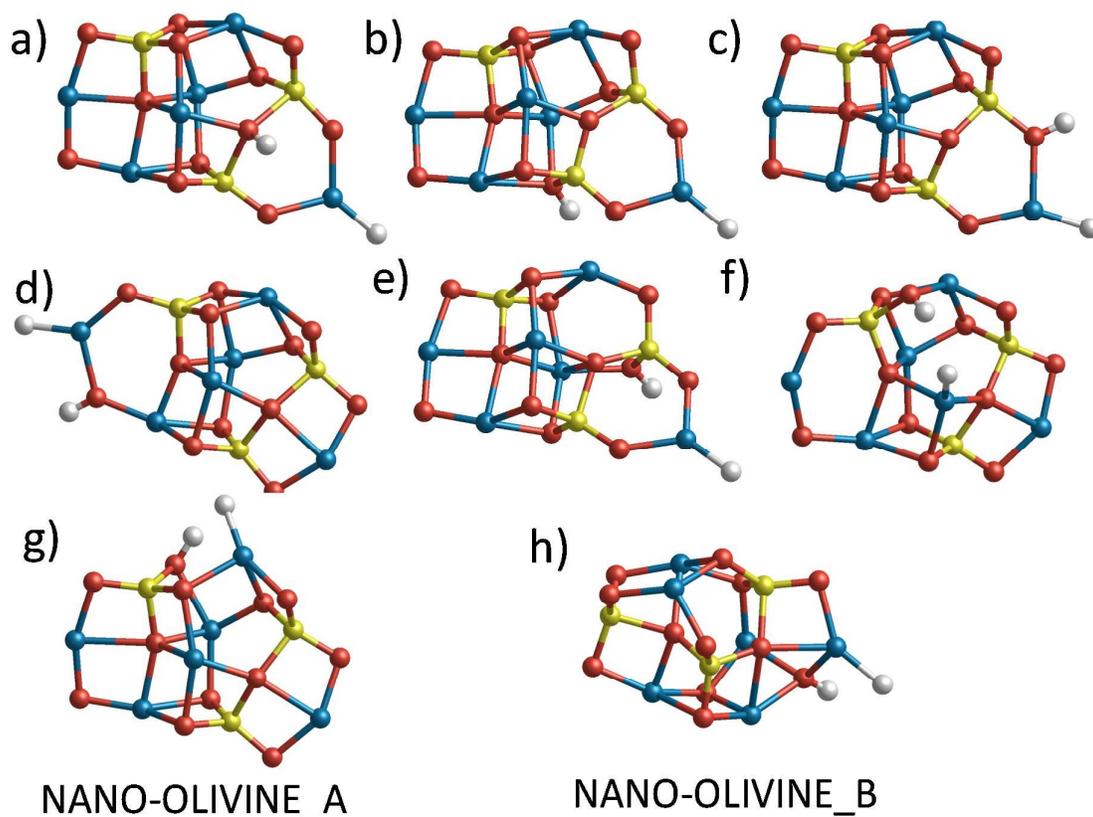
609

610 Fig. 6. C1/C3-type $2H_{\text{chem}}$ configurations on the $(\text{MgO})_6(\text{SiO}_2)_3$ nano-olivine_a cluster ordered by
611 increasing E_{bind} of H atom which is bound to a Mg atom relative to the corresponding cluster with
612 one chemisorbed H atom on a C1 site. Atom key: Si - yellow, Mg - blue and O - red.

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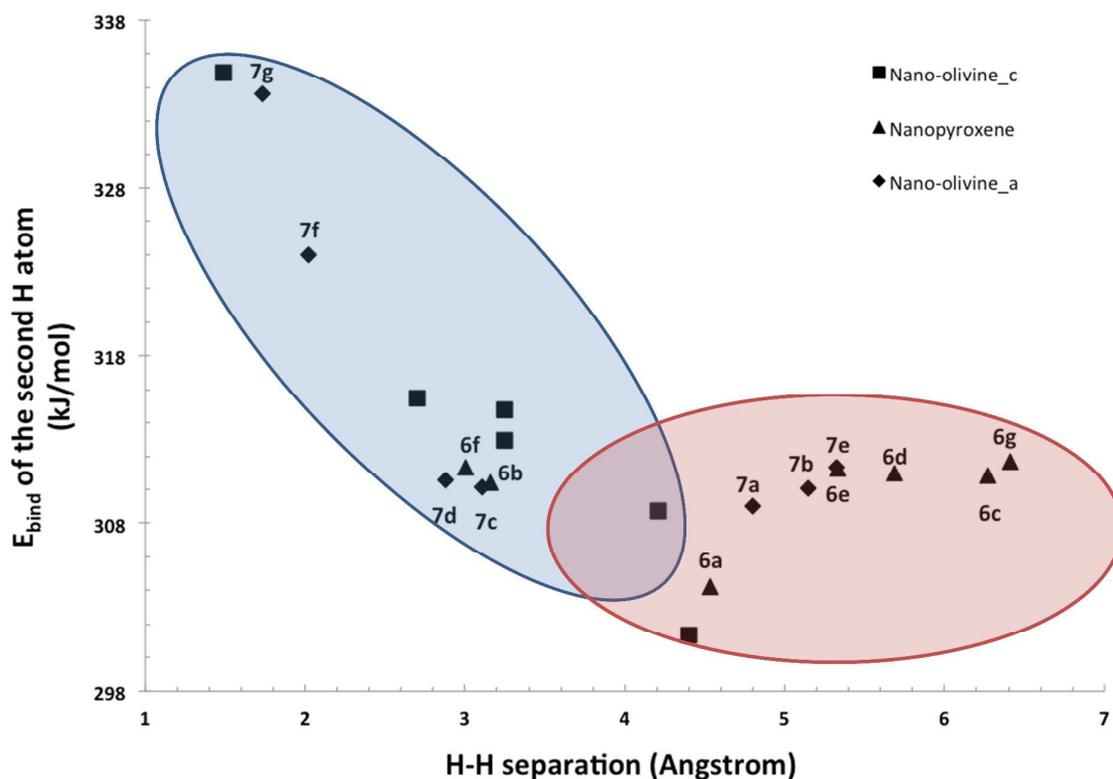


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617 Fig. 7. C1/C3-type $2H_{\text{chem}}$ configurations on the $(\text{MgO})_4(\text{SiO}_2)_4$ nanoproxene cluster ordered by
618 increasing E_{bind} of the H atom which is bound to a Mg atom relative to the corresponding cluster with
619 one chemisorbed H atom on a C1 site. Atom key: Si - yellow, Mg - blue and O - red.

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623 Fig. 8. Plot of binding energy of the second chemisorbed H_{chem}^- atom on the nano-olivine_a and nano-624 pyroxene clusters versus H-H separation for C1/C3 type 2H_{chem} adsorption (i.e. $[\text{H}_{\text{chem}}^+ \rightarrow \text{H}_{\text{chem}}^-]$)

625 Labelling corresponds to the configurations in Fig. 6 for the nano-pyroxene cluster and in Fig.7 for

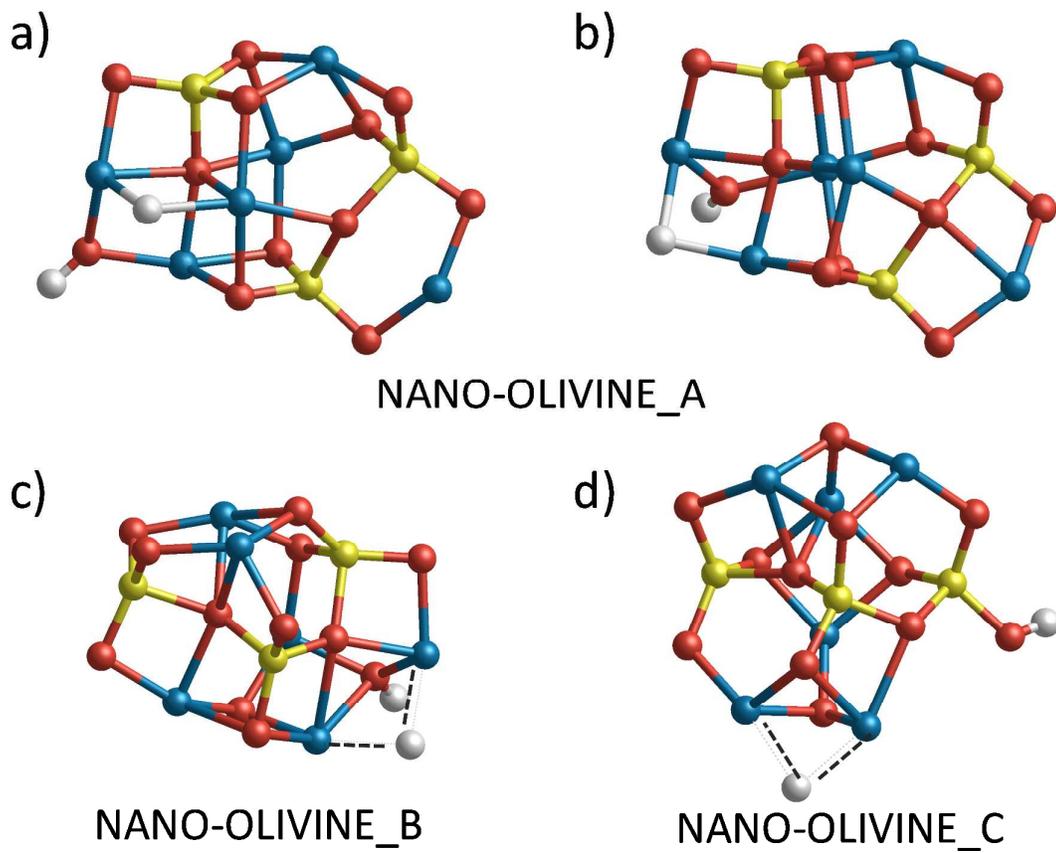
626 the nano-olivine_a cluster. Data from Kerkeni and Bromley²¹ corresponding to the nano-olivine_c

627 cluster is also included. The blue shaded area roughly encircles data points for which the

628 corresponding $\text{H}_{\text{chem}}^+ + \text{H}_{\text{chem}}^-$ pairs are electrostatically stabilised. The red shaded area roughly629 encircles data points for which the corresponding $\text{H}_{\text{chem}}^+ + \text{H}_{\text{chem}}^-$ pairs appear not to be significantly

630 affected by electrostatics.

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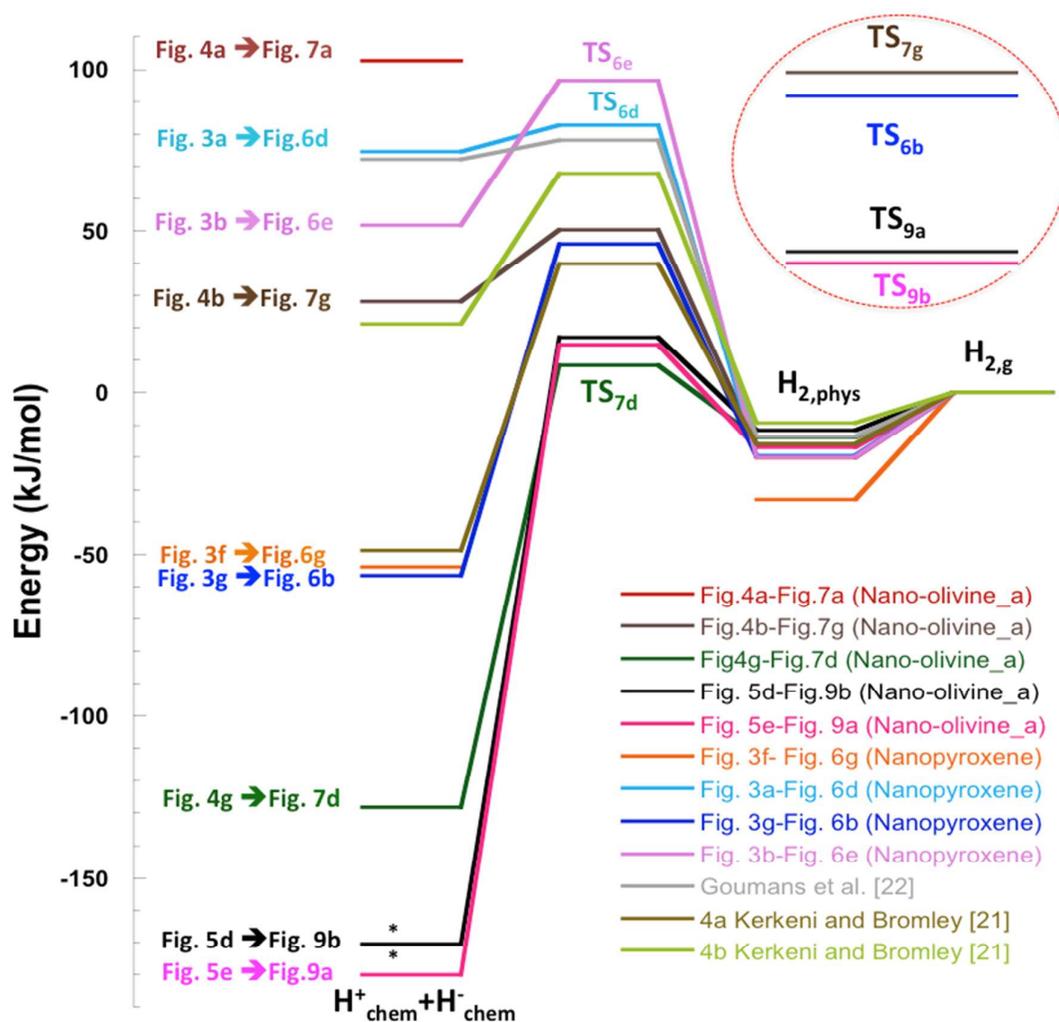
633 Fig. 9. C1/C2-type $2H_{\text{chem}}$ configurations on: a-b) the nano-olivine_a cluster, c) the nano-olivine_b634 cluster and d) the nano-olivine_c cluster, all with $(\text{MgO})_6(\text{SiO}_2)_3$ composition. Atom key: Si - yellow,

635 Mg - blue and O - red.

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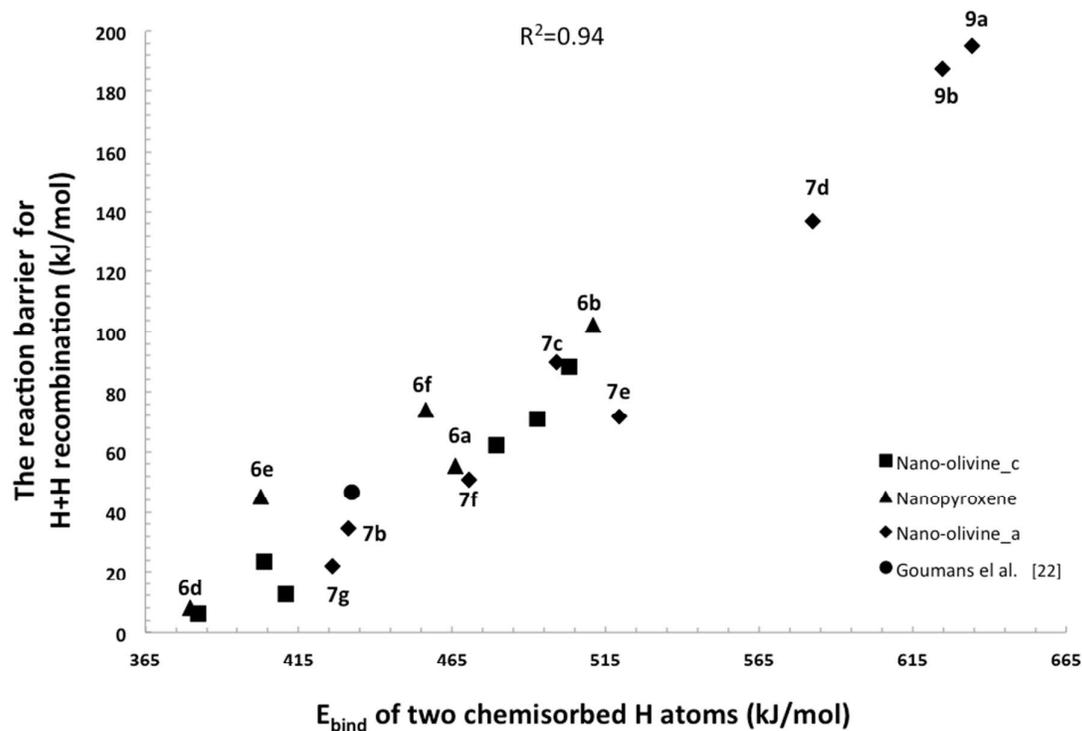
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640 Fig. 10. A selection of pathways for the $2\text{H}_{\text{chem}} \rightarrow \text{H}_{2,\text{g}}$ reaction on the nano-olivine_a cluster and on
 641 the nano-pyroxene cluster. The y-axis refers to energies relative to the sum of the energies of the bare
 642 relaxed nano-olivine_c cluster and a $\text{H}_{2,\text{g}}$ molecule (see also ref 21). The pathway labelling refers to
 643 the structures in the corresponding figures and denote the manner in which the initial state 2H_{chem}
 644 found. For example the pathway label “Fig. 5e → Fig. 9a” signifies that the starting 2H_{chem}
 645 configuration is that of Fig. 9a, which was derived from adding a H atom to the structure found in Fig.
 646 5e. The two “*”s correspond to pathways starting from C2/C1 type 2H_{chem} adsorption. The “TS”
 647 labels refer to transition states for the $2\text{H}_{\text{chem}} \rightarrow \text{H}_{2,\text{phys}}$ reaction found in this study, where the
 648 subscripts relate to the figure showing the corresponding initial 2H_{chem} configuration (all TS structures
 649 can be found in the Supplementary Information). The inset encircled by a dashed line is a magnified
 650 portion of the main plot showing transition state labels. Data for corresponding reaction pathways on
 651 the foresterite (010) surface (ref. 22), and on the nano-silicate_c cluster (ref. 21) are also included.

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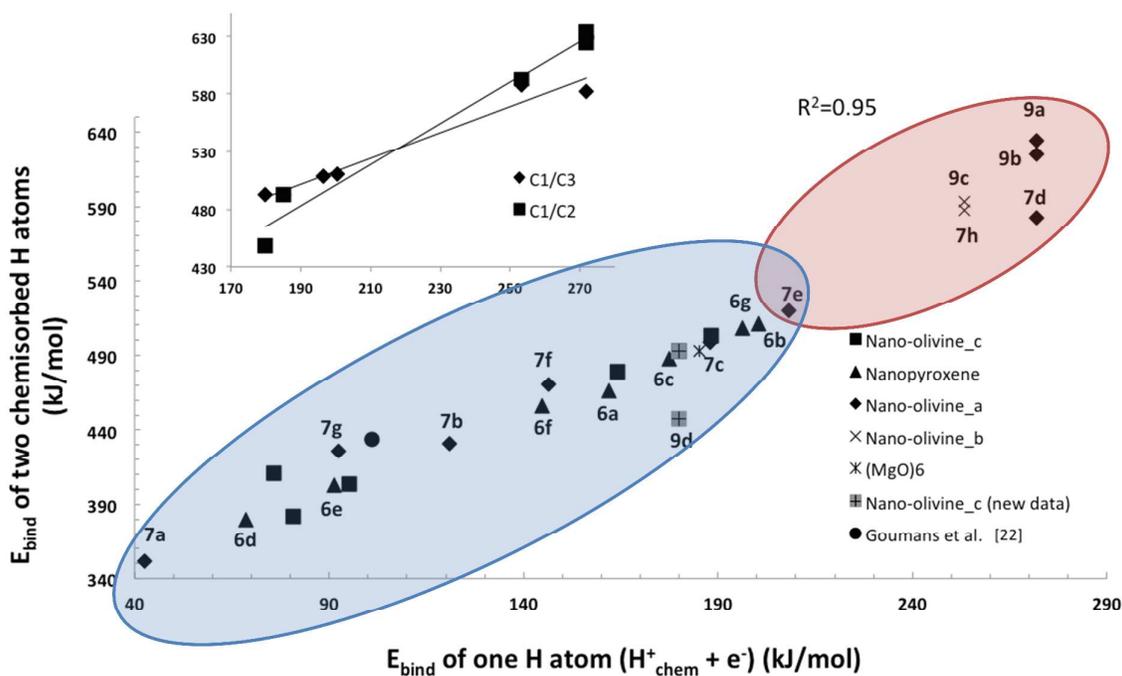
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655 Fig. 11. $2\text{H}_{\text{chem}} \rightarrow \text{H}_2$ recombination barrier height versus the total binding energy of two chemisorbed
 656 H atoms for all considered reaction paths on the nano-olivine_a and nano-pyroxene clusters. The
 657 labelling of the data points follows that in Figs. 6, 7 and 9. Data from Goumans *et al.*²² corresponding
 658 to the foresterite (010) surface and from Kerkeni and Bromley²¹ corresponding to the nano-olivine_c
 659 cluster are also shown. The high coefficient of determination (R^2) confirms the strong linearity of the
 660 relation (fitting line not shown).

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666 Fig. 12. The total binding energy versus binding energy of one H atom chemisorbed to the respective
 667 bare cluster in a C1 mode (i.e. H^+_{chem}) for all considered nano-olivine and nano-pyroxene clusters and
 668 the $(\text{MgO})_6$ cluster in Fig. 5. The labelling of the data points for the nanosilicates follows that in Figs.
 669 6, 7 and 9. Data from Goumans et al.²² corresponding to the foresterite (010) surface and from
 670 Kerkeni and Bromley²¹ corresponding to the nano-olivine_c cluster are also shown (new data for
 671 nano-olivine_c from this study is labelled). Red shading highlights data points with both high $2H_{\text{chem}}$
 672 and high H^+_{chem} binding energies, all of which correspond to nano-olivine clusters, and most of which
 673 to metastable nano-olivine isomers. The blue shading highlights trend of the remaining data points
 674 which correspond to a mix of nano-olivine and nano-pyroxene clusters. The high coefficient of
 675 determination (R^2) confirms the strong linearity of the relation for all data points in the main plot
 676 (fitting line not shown). The inset highlights the separate trends in the strength of C1/C2 *versus* C1/C3
 677 types of 2H chemisorption with respect to binding energy of the C1-bound H_{chem} atom. The fitting
 678 lines indicate a crossover in energetic favourability from C1/C3 type chemisorption to C1/C2
 679 chemisorption with increasing chemisorption strength of the C1-bound H_{chem} atom.

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685 **References**

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- ¹ E. Herbst, *Chem. Soc. Rev.*, 2001, 30, 168-176.
 - ² D. Hollenbach and E. E. Salpeter, *Astrophys. J.*, 1971, 163, 155-160.
 - ³ W. W. Duley and D. A. Williams, 1984, *Interstellar chemistry*, Academic Press, New York.
 - ⁴ D. A. Williams, 2003, in V. Pirronello, J. Krelowski, G. Manico, eds *Solid state astrochemistry*, Kluwer, Dordrecht, 1.
 - ⁵ J. L. Lemaire, G. Vidali, S. Baouche, M. Chehrouni, H. Chaabouni and H. Mokrane, *Astrophys. J.*, 2010, 725, L156-160.
 - ⁶ F. Islam, C. Cecchi-Pestellini, S. Viti and S. Casu, *Astrophys. J.*, 2010, 725, L1111 – 1123.
 - ⁷ J. Le Bourlot, F. Le Petit, C. Pinto, E. Roueff and F. Roy, 2012, *Astronom. Astrophys.*, 541, A76-94.
 - ⁸ J. Navarro-Ruiz, M. Sodupe, P. Ugliengo and A. Rimola, *Phys. Chem. Chem. Phys.*, 2014, **16**, 17447-17457.
 - ⁹ A. S. Ferrarotti, H. –P. Gail, *Astronom. Astrophys.*, 2006, 447, 553-576.
 - ¹⁰ I. Cherchneff, *Astronom. Astrophys.*, 2006, 456, 1001-1012.
 - ¹¹ S. T. Bromley, T. P. M. Goumans, E. Herbst, A.P. Jones, B. Slater, *Phys. Chem. Chem. Phys.*, 2014, **16**, 18623-18643.
 - ¹² T. P. M. Goumans and Stefan T. Bromley, *Mon. Not. R. Astron. Soc.*, 2011, 414, 1285-1291.
 - ¹³ A. Li and B. T. Draine, *Astrophys. J.*, 2001, 550, L213-L217.
 - ¹⁴ V. Pirronello, C. Liu, L. Shem and G. Vidali, *Astrophys. J.*, 1997, 475, L69-L72.
 - ¹⁵ H. B. Perets, A. Lederhendler, O. Biham, G. Vidali, L. Li, S. Swords, E. Congiu, J. Roser, G. Manicó, J. R. Brucato and V. Pirronello, *Astrophys. J.*, 2007, 661, L163-L166.
 - ¹⁶ G. Vidali, V. Pirronello, L. Li, J. Roser, G. Manicó, E. Congiu, H. Mehl, A. Lederhendler, H. B. Perets, J. R. Brucato and O. Biham, *J. Phys. Chem. A.*, 2007, 111, 12611- 12619.
 - ¹⁷ L. Gavilan, J. L. Lemaire and G. Vidali, *Mon. Not. R. Astron. Soc.*, 2012, 424, 2961-2970.
 - ¹⁸ S. T. Bromley, T. P. M. Goumans, E. Herbst, A. P. Jones and B. Slater, *Phys. Chem. Chem. Phys.*, 2014, **16**, 18623-18643
 - ¹⁹ T. P. M. Goumans and S. T. Bromley, *Mon. Not. R. Astron. Soc.*, 2012, 420, 3344-3349.
 - ²⁰ T. P. M. Goumans and S. T. Bromley, *Phil. Trans. R. Soc. A*, 2013, 371, 20110580.
 - ²¹ B. Kerkeni and S.T. Bromley, *Mon. Not. R. Astron. Soc.*, 2013, 435, 1486-1493.
 - ²² T. P. M. Goumans, C. Richard A. Catlow and Wendy A. Brown, *Mon. Not. R. Astron. Soc.*, 2009, 393, 1403-1407.
 - ²³ S. Garcia-Gil, D. Teillet-Billy, N. Rougeau, and V. Sidis, *J. Phys. Chem. C.*, 2013, 117, 12612–12621
 - ²⁴ C. A. Downing, B. Ahmady, C. R. A. Catlow and N. H. de Leeuw, *Phil. Trans. R. Soc. A.*, 2013, 371, 20110592.
 - ²⁵ M. J. Frish et al., 2004, *Gaussian 09*, Revision A.02. Gaussian Inc., Walling-ford, CT, 2009.

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- ²⁶ Y. Zhao and D.G. Truhlar, *J. Phys. Chem. A.*, 2004, 108, 6908-6918.
- ²⁷ Y. Zhao and D.G. Truhlar, *J. Phys. Chem. A.*, 2005, 109, 5656-5667.
- ²⁸ T. P. M. Goumans, C. Richard A. Catlow, W. A. Brown, J. Kastner and P. Sherwood, *Phys. Chem. Chem. Phys.*, 2009, 11, 5431-5436.
- ²⁹ H. B. Schlegel, *J. Comp. Chem.*, 1982, 3, 214-218.
- ³⁰ G. Bussi, D. Donadio and M. Parrinello, *J. Chem. Phys.* 2007, 126, 014101(1-7)
- ³¹ J. P. Perdew, K. Burke, and M. Ernzerhof, *Phys. Rev. Lett.*, 1996, 77, 3865-3868.
- ³² V. Blum, R. Gehrke, F. Hanke, P. Havu, V. Havu, X. Ren, K. Reuter, and M. Scheffler, *Comp. Phys. Commun.*, 2009, 180, 2175-2196.
- ³³ N. Katz, I. Furman, O. Biham, V. Pirronello and G. Vidali, *Astrophys. J.*, 1999, 522, 305-12
- ³⁴ J. Arsic, D. Kaminski, M. Radenovic, P. Poodt, W. S. Graswinckel, H. M. Cuppen and E. Vlieg, *J. Chem. Phys.*, 2004, 120, 9720-9724
- ³⁵ J. Jellinek, T. L. Beck and R. S. Berry, *J. Chem. Phys.*, 1986, 84, 2783-2794.
- ³⁶ F. Viñes, J. Carrasco, and S. T. Bromley, *Phys. Rev. B*, 2012, 85, 195425(1-6).