



Cite this: *Mater. Adv.*, 2024,
5, 8970

Investigating pressure-driven semiconductor-to-metal transition in lead-free perovskites AlGeX_3 ($X = \text{F}$, Cl , and Br): insights from first-principles calculations†

Md. Amran Sarker,^a Md. Mehedi Hasan,^b Md. Rafiqul Islam,^c
Md. Rabbi Talukder,^b Md. Rasidul Islam^d and Ahmed Sharif^{*a}

In the pursuit of commercializing electronic and optoelectronic devices, researchers have turned their attention to non-toxic inorganic cubic metal halide perovskites. This study focuses on novel lead-free compounds—specifically AlGeX_3 (where $X = \text{F}$, Cl , and Br) and examines their structural, electronic, optical, and mechanical properties under the application of hydrostatic pressure through density functional theory (DFT). The mechanical stability of all compounds is rigorously assessed using Born stability criteria and formation energy. The elastic investigations reveal that the materials have anisotropy, ductility, and good Machinability index depending on the halide type and applied pressure. The pressure-dependent electronic band structures are calculated by GGA-PBE functional to demonstrate the intriguing behavior of the compounds. Band structures are also calculated by HSE06 functional without pressure. Further, the substitution of the halide F with Cl/Br leads to an indirect to direct band gap transformation. Additionally, increasing positive hydrostatic pressure results in a tunable band gap with decreasing trends for all the compounds leading them to transit from semiconductor to metallic state. This phenomenon is explained by the partial and total density of states (PDOS and TDOS). The improvement of pressure-dependent optical properties in both visible and UV regions makes them valuable contenders in the quest for efficient solar cells and other electronic and optoelectronic devices.

Received 12th June 2024,
Accepted 21st October 2024

DOI: 10.1039/d4ma00611a

rsc.li/materials-advances

Introduction

The global demand for electrical energy has been rising dramatically over the past few decades as a result of contemporary civilizations and industrialization. Industries rely on abundant,

stable, and affordable energy supplies to construct contemporary civilization. However, a significant portion of the energy used comes from nonrenewable resources like fossil fuels, which are scarce and will eventually run out. Furthermore, the burning of fossil fuels contributes significantly to global

^a Department of Materials and Metallurgical Engineering, Bangladesh University of Engineering & Technology (BUET), Dhaka 1000, Bangladesh.

E-mail: asharif@mme.buet.ac.bd

^b Department of Materials Science and Engineering, Khulna University of Engineering & Technology (KUET), Khulna 9203, Bangladesh

^c Department of Electrical and Electronic Engineering, Khulna University of Engineering & Technology (KUET), Khulna 9203, Bangladesh

^d Department of Electrical and Electronic Engineering, Bangamata Sheikh Fojilatunnesa Mujib Science & Technology University, Jamalpur-2012, Bangladesh

† Electronic supplementary information (ESI) available: Fig. S1 Band structure of AlGeF_3 under hydrostatic pressure. Fig. S2 Band structure of AlGeCl_3 under hydrostatic pressure. Fig. S3 Band structure of AlGeBr_3 under hydrostatic pressure. Fig. S4 The calculated TDOS and PDOS of cubic AlGeF_3 perovskites under pressure. Fig. S5 The calculated TDOS and PDOS of cubic AlGeCl_3 perovskites under pressures. Fig. S6 The calculated TDOS and PDOS of cubic AlGeBr_3 perovskites under pressures. Fig. S7 Anisotropic 3D representation of Young's modulus, shear modulus, and Poisson's ratio of AlGeF_3 at different pressures. Fig. S8 Anisotropic 3D representation of Young's modulus, shear modulus, and Poisson's ratio of AlGeCl_3 at different pressures. Fig. S9 Anisotropic 3D representation of Young's modulus, shear modulus, and Poisson's ratio of AlGeBr_3 at different pressure. Table S1: The calculated lattice parameters (a), cell volume (V), formation energy (ΔH_f), and band gap (E_g) of AlGeX_3 ($X = \text{F}$, Cl , and Br) at different pressures. Table S2: variation in bond length of cubic AlGeX_3 ($X = \text{F}$, Cl , and Br) at different pressure. Table S3 (a): the calculated elastic constants, bulk modulus (B), shear modulus (G), Young's modulus (E), Pugh's ration, Poisson's ratio (ν), hardness (H_v), and machinability index (μ_M) of AlGeF_3 under hydrostatic pressure. (b): changes in anisotropy of cubic AlGeF_3 under hydrostatic pressure. Table S4 (a): the calculated elastic constants, bulk modulus (B), shear modulus (G), Young's modulus (E), Pugh's ration, Poisson's ratio (ν), hardness (H_v), and machinability index (μ_M) of AlGeCl_3 under hydrostatic pressure. (b): changes in anisotropy of cubic AlGeCl_3 under hydrostatic pressure. Table S5 (a): the calculated elastic constants, bulk modulus (B), shear modulus (G), Young's modulus (E), Pugh's ration, Poisson's ratio (ν), hardness (H_v), and machinability index (μ_M) of AlGeBr_3 under hydrostatic pressure. (b): changes in anisotropy of cubic AlGeBr_3 under hydrostatic pressure. See DOI: <https://doi.org/10.1039/d4ma00611a>



warming. Thus, it is essential to research different sustainable and renewable energy sources such as solar energy, which is unlimited and has the potential to alleviate the consequences of global warming. Since photovoltaic cells can generate clean, renewable electricity from light, they have been employed and studied extensively in recent years.^{1–12} Thus, the search for an efficient light-absorbing material capable of absorbing the whole solar spectrum is important for fabricating high-performance solar cells. Perovskites (ABX_3) are promising materials for the absorber layer because they are cheap, easy to make, and have high power conversion efficiency (PCE).^{13,14} MAPbI₃ perovskite solar cells have reached a record PCE of over 25%.¹⁵ However, they have some drawbacks, such as the instability of the organic component under different environmental conditions^{16,17} and the toxicity of the lead element (Pb).^{18,19} These issues can be addressed by replacing the organic cations with inorganic ones²⁰ and the lead element (Pb) with a non-toxic group 14 element, such as Sn or Ge.^{21,22} For example, lead-free germanium-based CsGeX₃ (X is a halogen) perovskite solar cells achieved a PCE of about 4.92%, which increased to 7.11% when tin was added (CsSn_{0.5}Ge_{0.5}I₃).²³ Also, CsSnI₃-based perovskite solar cells reached a PCE of more than 10%.²⁴ However, these PCE are still lower than those of organic–inorganic halide perovskites (MAPbI₃). Therefore, researchers are looking for new lead-free perovskite materials and ways to improve the absorbance of the absorber layer in the whole solar spectrum. Many parameters of the absorbing material, including bandgap, absorbance, carrier mobility, charge carrier recombination rate, and exciton binding energy influence the PCE of photovoltaic cells. These parameters can be adjusted or improved by applying pressure. Therefore, to investigate improved lead-free nontoxic inorganic perovskite-based absorber materials, it is necessary to study their electrical, optical, and mechanical properties under pressure.

Moreover, pressure-induced as well as strain-induced changes can control the material's band gap, improve light absorption, minimize charge carrier recombination, and maximize charge separation.^{25–28} Furthermore, pressure is vital for the exploration of the material's electronic and optical characteristics and structural integrity. To improve the solar energy conversion efficiency of Ge-based perovskite materials applying pressure could be a strategy. Mitro *et al.*, Hasan *et al.*, Sarker *et al.*, and Xiang *et al.* explored pressure-driven structural, electronic, optical, and mechanical properties of RbGeX₃ (X = Cl, Br),²⁹ GaGeX₃ (X = Cl, Br, and I),³⁰ NaGeX₃ (X = F, Cl, Br, and I),³¹ and CsGeX₃ (X = I, Br, or Cl)³² accordingly, to enhance the optoelectronic characteristics, and making them more suitable for practical optoelectronic devices application such as solar cells. Recently, Hasan *et al.* investigated different physical properties of novel AlGeX₃ (X = F, Cl, Br, and I)³³ perovskites using first-principles calculation and whether no previous theoretical studies have been conducted on the pressure-driven Ge-based AlGeX₃ (X = F, Cl, and Br) halide perovskites. So, based on an extensive literature review and leveraging our current understanding, the scarcity of existing research has

driven us to investigate this uncharted area, aiming to overcome existing limitations and harness the untapped potential of these promising materials. Besides, Aluminum-based perovskite minerals are abundant in the Earth's crust and have been used in a wide range of applications. These materials are very resistant to corrosion and have the ideal properties of being lightweight, flexible, and ductile. In the development of functional materials with improved capacities, these attributes are highly valued. After silicon, aluminum is the second most common metallic element in the world that emphasizes its significance, and material science may suggest its wide variety of applications.

This study presents the pressure-induced structural, electronic, mechanical, and optical characteristics of perovskite Ge-halide materials AlGeX₃ (X = F, Cl, and Br) by deliberately altering the pressure circumstances using first-principles computations. The key insights of this study are to investigate the pressure-dependent structural stability of the compounds, energy bandgap engineering, semiconductor–metal transition state, state of optical properties, elastic properties, *etc.*, and hence to explore the compounds' suitability for high-performance solar cells.

Computational details

This research uses the CASTEP module of BIOVIA Materials Studio software³⁴ to perform DFT-based calculations.³⁵ Here the GGA-PBE functional³⁶ to compute the electronic exchange–correlation energy and the bandgap. The existence of the Vanderbilt-type OTFG ultrasoft pseudopotential is assumed³⁷ to handle the electron–ion interactions. The optimal crystal structure is ensured by the BFGS method³⁸ and the plane wave cutoff energy and k-points are set at 500 eV and $6 \times 6 \times 6$ to obtain the optimal structure and the characteristics calculations. The Monkhorst–Pack scheme³⁹ is conducted to sample the k-points in the Brillouin zone and optical properties are calculated at higher k-points. The elastic constants of our models are determined using the stress–strain method⁴⁰ under standard conditions. CASTEP determines the elastic stiffness constants (C_{ij}) based on the finite strain theory. On the other hand, the ELATE software is conducted to visually represent the anisotropic contour plots in three dimensions, specifically for the parameters of Shear modulus, Young's modulus, and Poisson's ratio. The geometry of the unit cells and the point of atomic relaxation optimization in CASTEP were set according to these criteria: convergence of total energy, 5×10^{-6} eV per atom; maximum displacements, 5×10^{-4} Å; maximum force, 0.01 eV per Å; and maximum stress, 0.02 GPa. We set the strain magnitude at 0.003 to achieve the best results and use the same thresholds for convergence.

Structural properties

Fig. 1 shows the supercell of cubic AlGeX₃ crystal structure in 3D schematic form. Al and Ge are at certain Wyckoff



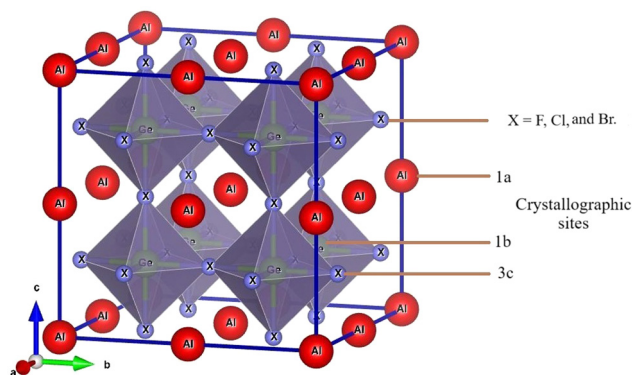


Fig. 1 AlGeX_3 ($X = \text{F}, \text{Cl}, \text{and Br}$) cubic metal halide perovskite structure.

coordinates in this structure: Al is at 1a (0,0,0), Ge is at 1b (0.5,0.5,0.5), and X atoms are at 3c (0.5,0,0.5). This structure belongs to the $pm3m$ space group with number 221. Using the GGA technique, structural characteristics such as the energy of formation (ΔH_f), the lattice constant (a), and the volume (V) were calculated by performing a volume optimization procedure to find the stable ground state unit cell. Furthermore, the energy values as a function of volume were adjusted to the Birch-Murnaghan equation.⁴¹ Using the given fit, we have calculated the lattice parameters under different hydrostatic pressures. The structural stability of those compounds such as AlGeF_3 , AlGeCl_3 , and AlGeBr_3 depends on the applied pressure. The compounds become metallic for the positive hydrostatic pressures (compressive) 15 GPa, 4 GPa, and 2.5 GPa, respectively. The pressure-dependent lattice parameters, cell volumes, and formation energies are listed in Table S1(a–c) (ESI[†]) where the lattice parameters and cell volumes are found to decrease with intensifying the positive pressure for all the compounds. However, for a particular pressure, the lattice parameters as well as cell volumes increase when the halogen size increases (from F to Br) as seen in Table S1 (ESI[†]). The pressure-dependent variations of the lattice parameters for all the compounds are represented in Fig. 2.

To confirm the pressure-dependent structural stability, the formation energy and elastic constants of AlGeX_3 are estimated and listed in Table S1 and S3(a), S4(a), and S5(a) (ESI[†]). The mechanical stability of those compounds is verified by Born stability criteria using the following relations,

$$C_{11} + 2C_{12} > 0, C_{44} > 0, C_{11} - C_{12} > 0$$

The estimated elastic constants satisfy the criteria, so the compounds are mechanically stable. Further, the enthalpies of the compounds are calculated by eqn (1)

$$\Delta H_f(\text{AlGeX}_3) = \frac{[E_{\text{tot}}(\text{AlGeX}_3) - E_s(\text{Al}) - E_s(\text{Ge}) - 3E_s(\text{X})]}{N} \quad (1)$$

here, $E_{\text{tot}}(\text{AlGeX}_3)$ represents AlGeX_3 's total ground state energy, and $E_s(\text{Al})$, $E_s(\text{Ge})$, and $E_s(\text{X})$ are the energies of Al, Ge,

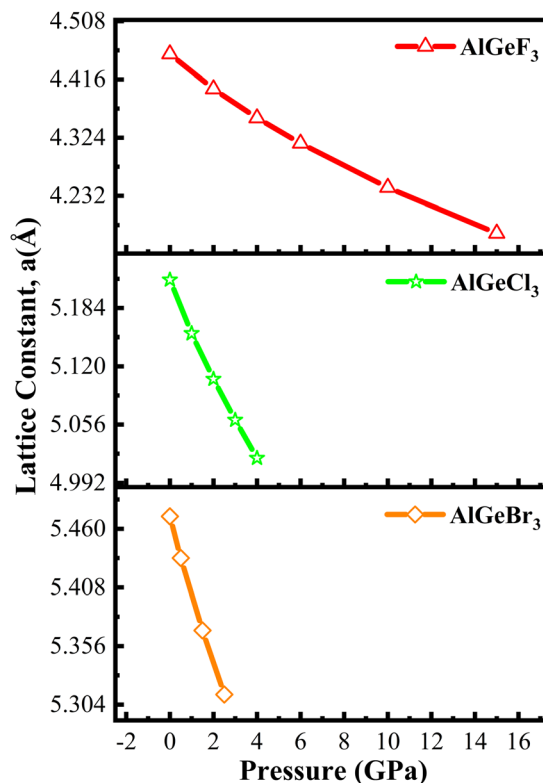


Fig. 2 Compressive pressure-dependent lattice constant of AlGeX_3 ($X = \text{F}, \text{Cl}, \text{and Br}$).

and X (=F, Cl, and Br), respectively, per unit cell. N is the total number of atoms per unit cell.

The results are shown in Fig. 3 where the negative formation energies (ΔH_f) observed up to 15 GPa, 4 GPa, and 2.5 GPa for AlGeF_3 , AlGeCl_3 , and AlGeBr_3 respectively, under compressive pressure. The formation energies of all compounds become less negative with increasing pressure. Thus, the structural stability of AlGeX_3 compounds decreases gradually but they still exhibit excellent stability and tunability in the applied pressure range. The tunable structural and electrical properties of those compounds allow one to realize materials with suitable attributes for various applications. The pressure-dependent bond lengths of those compounds are estimated and listed in Table S2 (ESI[†]). The decrease of bond lengths with increasing pressure is due to the shrinking of interatomic space. This decreasing nature indicates a stronger bond between atoms under pressure. To explore the application areas of those compounds, their pressure-dependent electronic, optical, and mechanical properties are further estimated and presented in the following sections.

Electronic properties

Herein, GGA-PBE approximation is employed to predict the pressure-induced electronic band dispersion profiles of AlGeX_3 . The results are presented in Fig. S1–S3 (ESI[†]), respectively, for the compounds AlGeF_3 , AlGeCl_3 , and AlGeBr_3 . The red dotted lines (0 eV, EF) in Fig. S1–S3 (ESI[†]) denote the Fermi level.





Fig. 3 Pressure-dependent formation energy of AlGeX_3 ($X = \text{F}, \text{Cl}, \text{and Br}$).



Fig. 4 Pressure-dependent band gap energy of AlGeX_3 ($X = \text{F}, \text{Cl}, \text{and Br}$).

Fig. S1 (ESI[†]) elucidates that AlGeF_3 shows a 1.115 eV indirect energy gap at $R-M$ critical points at ambient pressure. With the replacement of F by the larger halogen (Cl and Br) the band gap transformed from indirect to direct at the R point and the energy gap reduced to 0.760 eV and 0.433 eV for AlGeCl_3 and AlGeBr_3 , respectively. AlGeF_3 exhibits an indirect band gap and the rest of the compounds expose a direct band gap for all the applied pressure as seen in Fig. S1–S3 (ESI[†]). Fig. S1 (ESI[†]) shows that the conduction band of AlGeF_3 starts shifting towards the Fermi level up to 10 GPa pressure without changing the valence band position, that's why the energy gap reduces with increasing pressure. This scenario changes, that is, the valence band rises upward and crosses the Fermi level when 15 GPa pressure is applied and the gap turns zero. The compounds AlGeCl_3 and AlGeBr_3 show identical band-shifting characteristics under compressive pressure except band gaps are direct and they become zero up to the applied pressure of 4 GPa and 2.5 GPa, respectively. Thus, by controlling pressure, we can tune the characteristics of compounds for specific applications such as switchable electrical devices that can alternate between metallic and semiconducting states, transistors, memory devices as well as materials with high conductivity. All the calculated band gaps of those compounds are tabulated in Tables S1(a–c) (ESI[†]) and plotted in Fig. 4. Almost linear trends with reduction of energy gaps are observed with intensifying pressure for the compounds AlGeF_3 , AlGeCl_3 , and AlGeBr_3 . Subsequently, semiconductor-to-metallic transition occurs at 15 GPa, 4 GPa, and 2.5 GPa, respectively. Since PBE

underestimates the band gap, HSE06 functional is applied and the calculated values are found to be 2.043 eV, 1.694 eV, and 1.382 eV, respectively, for AlGeF_3 , AlGeCl_3 , and AlGeBr_3 without pressure. A comparison of band dispersion profiles obtained from HSE06 and PBE functionals is shown in Fig. 5.

To assess the contribution of the orbitals in band dispersion profiles, the density of states (DOS) of AlGeX_3 are also estimated and illustrated in Fig. S4–S6 (ESI[†]), respectively, for AlGeF_3 , AlGeCl_3 , and AlGeBr_3 where the vertical line (black colored) at 0 eV designates the Fermi level. The total density of state (TDOS) is critical to understanding the properties of a semiconducting material. The TDOS of all the compounds show zero value near the Fermi level confirming that these are semiconductors at ambient pressure. It is evident from the partial density of states (PDOS) which are illustrated in Fig. S4–S6 (ESI[†]), the valence band maxima (VBM) are mainly derived from the orbital of F-2p and Ge-4s for AlGeF_3 , Cl-3p for AlGeCl_3 , and Br-4p for AlGeBr_3 . While forming the conduction band minima (CBM) are derived from mainly Al-3p for AlGeF_3 , Al-3p & Ge-4p for AlGeCl_3 and Ge-4p for AlGeBr_3 respectively. When compressive pressure is applied, the Al-3p orbital peak in the CB shifts toward the Fermi level for both AlGeF_3 and AlGeCl_3 . But for AlGeBr_3 , the Ge-4p orbital peak in the CB shifts toward the Fermi level up to the pressure of 2.5 GPa. Therefore, the band gap becomes narrower, and semiconductor-to-metallic transition occurs at 15 GPa, 4 GPa, and 2.5 GPa for AlGeF_3 , AlGeCl_3 , and AlGeBr_3 , respectively. The pressure-dependent precise control over the bandgap through mechanical pressure, paving the way for



Fig. 5 A comparison of electronic band structures calculated by GGA-PBE and HSE06 functionals (a) AlGeF₃, (b) AlGeCl₃, and (c) AlGeBr₃.

applications of the compounds in advanced electronic and optoelectronic devices.

Optical properties

Compared to lead-based perovskites, lead-free metal halide cubic perovskites provide intriguing characteristics such as higher absorption, lower reflectance, and increased conductivity.⁴² Assessment of pressure-dependent optical properties is essential for advancing material research in the direction of modern optoelectronic devices, especially solar cell development.⁴³ Here, we conducted a thorough investigation of the dielectric constant, absorption coefficient, conductivity, reflectivity, extinction coefficient, and refractive index of AlGeX₃ (X = F, Cl, and Br) under pressure. It is crucial to assess the dielectric function⁴⁴ first to determine other optical properties. The complex form of the dielectric function can be represented

by $\varepsilon(\omega) = \varepsilon_1(\omega) + i\varepsilon_2(\omega)$ where $\varepsilon_1(\omega)$ is the real part and $\varepsilon_2(\omega)$ is the imaginary.⁴⁵ The real and imaginary parts can be expressed by

$$\varepsilon_1(\omega) = 1 + \frac{2}{\pi} P \int_0^{\infty} \frac{\omega' \varepsilon_2(\omega')}{\omega'^2 - \omega^2} d\omega' \quad (2)$$

and

$$\varepsilon_2(\omega) = \frac{4e^2 \pi^2}{m^2 \omega^2} \sum \int [i|M|j]^2 f_i(1-f_i) \times \delta[E_f - E_i - \omega] d^3k \quad (3)$$

where P represents the integral's principal value. The mathematical expressions of the optical properties like absorption coefficient $\alpha(\omega)$, conductivity $\sigma(\omega)$, refractive index $\eta(\omega)$, and reflectivity (R) are available in.^{44,46}

Here we estimate pressure-dependent dielectric constants and plotted their real, $\varepsilon_1(\omega)$ and imaginary, $\varepsilon_2(\omega)$ parts as a





Fig. 6 Comparison of the real part of the dielectric constant without pressure (a) AlGeX_3 ($X = \text{F}, \text{Cl}$, and Br) and with pressure (b) AlGeF_3 , (c) AlGeCl_3 , and (d) AlGeBr_3 .

function of photon energy. Fig. 6(a) shows a comparison of $\epsilon_1(\omega)$ estimated among the compounds without pressure as a function of photon energy varied from 0 to 30 eV. At 0 GPa, the value of the static dielectric constant $\epsilon_1(0)$ increases when F is substituted by Cl and then to Br as seen in Fig. 6(a). The frequency dependence of $\epsilon_1(\omega)$ shows a dominant peak for the AlGeF_3 , which is further shifted to the lower energy side when F is substituted by Cl and Br. The peak of $\epsilon_1(\omega)$ is resulted due to the resonance between the electric displacement vector and field vector. Afterward, $\epsilon_1(\omega)$ reduces rapidly owing to the reduction of dipole moment resulting in a lack of synchronism between the displacement vector and field vector. The compounds exhibit negative $\epsilon_1(\omega)$ when the electric displacement and field vectors point in opposite directions at frequencies higher than the resonance frequency. The class of negative dielectric materials possesses a negative refractive index and has applications in optical devices. At higher frequencies, the value of $\epsilon_1(\omega)$ is almost zero due to the displacement vector which is unable to respond to the frequency of the field vector. The interaction of light with matter depends on the electronic polarization of the material.⁴⁷ The charge carrier recombination rates are associated with the magnitude of $\epsilon_1(\omega)$ leading to a considerable improvement in the performance of an optoelectronic device. This implies that the compounds AlGeCl_3 and AlGeBr_3 are more appropriate for optoelectronic devices operating in the IR-visible region than the compound AlGeF_3 .

A similar comparison is also shown in Fig. 6(b–d) for the compounds AlGeF_3 , AlGeCl_3 and AlGeBr_3 under pressurized conditions. The values of static dielectric constants $\epsilon_1(0)$ increase with increasing pressure. However, the $\epsilon_1(\omega)$ spectra

show major peaks that shift towards the lower energy side with increasing pressure and then they decrease rapidly and show negative peaks at a particular energy. This may be due to pressure-dependent orientational polarization that may cause the polarization vector stronger and hence the dipole moment. This in turn increases the value of $\epsilon_1(0)$ under pressurized conditions. At higher energy, $\epsilon_1(\omega)$ becomes almost zero displaying small fluctuations as a function of photon energy. It is also found that the major peaks of $\epsilon_1(\omega)$ spectra become sharp and shifted in the lower energy direction after the application of pressure. The density of polarized atoms increases under hydrostatic pressure due to the volumetric compression of the compounds. This may result in sharp peaks due to pressure-dependent orientational change of the displacement vector in the reduced volume. Further, the resonance frequency of the displacement vector gives a redshift with increasing pressure. This demonstrates that the compounds could be more suitable for optoelectronic devices such as solar cells, LEDs, and photodetectors in the visible energy range under pressurized conditions.

The imaginary part of the dielectric function, $\epsilon_2(\omega)$ is directly connected to electron excitation. It signifies the loss of optical energy that is absorption of optical energy in the matter. Fig. 7(a) shows a comparison of the $\epsilon_2(\omega)$ obtained from the compounds AlGeF_3 , AlGeCl_3 , and AlGeBr_3 with the variation of photon energy from 0 to 30 eV without pressure. The significant peak of $\epsilon_2(\omega)$ is observed at the lower energy and it gives a redshift when F is substituted by Cl and Br. Afterward, the magnitude of $\epsilon_2(\omega)$ decreases with photon energy with some arbitrary small peaks. The electric displacement vectors



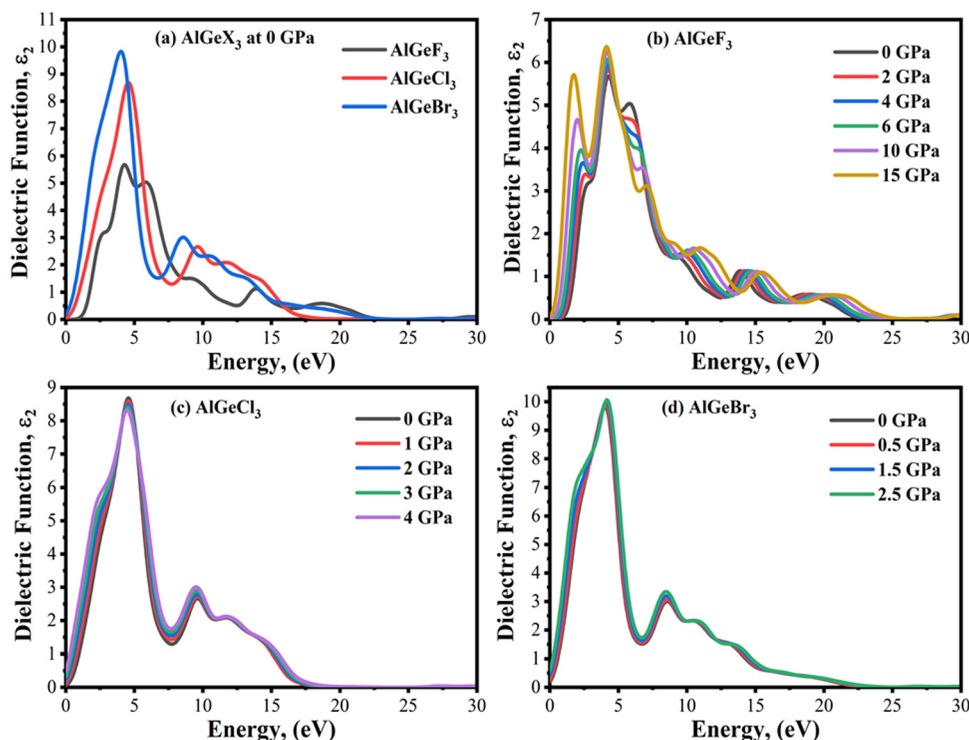


Fig. 7 Comparison of the imaginary part of the dielectric constant without pressure (a) AlGeX_3 ($X = \text{F}, \text{Cl}$, and Br) and with pressure (b) AlGeF_3 , (c) AlGeCl_3 , and (d) AlGeBr_3 .

oscillate with field vectors and resonance condition buildup at different frequencies due to the inherent polarization of the compounds. The pressure-dependent of $\epsilon_2(\omega)$ is represented as a function of photon energy in Fig. 7(b–d). It is found in Fig. 7(b) that $\epsilon_2(\omega)$ shows one strong peak at about 4.85 eV and the second-highest peak at about 2.50 eV for the compound AlGeF_3 at 0 GPa. However, these peaks become very sharp and significant with increasing pressure and they give redshift as well. The causes of sharp peaks and their shifts are already discussed. The influence of pressure is found to be insignificant in $\epsilon_2(\omega)$ spectra for the compounds AlGeCl_3 and AlGeBr_3 with and without pressure as seen in Fig. 7(c and d). Overall, semiconductor materials become more polarizable due to higher dielectric constant (ϵ_1 & ϵ_2). When materials undergo semiconductor-to-metal transitions at higher pressures, these dielectric tend to drop due to metals having low dielectric constants.

Optical absorption is an important factor, especially in visible regions for selecting high-performance photovoltaic materials. The absorption coefficient is a measure of how good a material is to absorb light for a particular wavelength. It depends on the loss of the photon in a material connected to the imaginary part of the dielectric constant, $\epsilon_2(\omega)$. Fig. 8(a) shows a comparison of the absorption coefficient of AlGeX_3 (F, Cl, and Br) at 0 GPa. Here, the absorption spectra give a redshift when F is substituted *via* Cl and Br. A comparison of pressure-dependent absorption spectra of AlGeX_3 ($X = \text{F}, \text{Cl}$, and Br) is shown in Fig. 8(b–d). Pressure-sensitive strong absorption peaks with anomalous shifts are observed in Fig. 8(b) for the

compound AlGeF_3 . However, the influence of pressure in the lower energy of absorption spectra is found to be redshifted with higher intensity when F is replaced by Cl and Br as seen in Fig. 8(c and d). Besides, each compound displays several larger absorption peaks within the higher energy range under pressure, which indicates that the compounds could be suitable as UV sensors and detectors as well. To understand the absorption characteristics in the visible region the extended view of spectra is plotted as a function of wavelength in Fig. 9(a–d). The absorption coefficient decreases with increasing wavelength in the visible range and each compounds show the same pattern. A significant increase in the absorption is observed in the visible area for AlGeX_3 under pressure in Fig. 9(b–d). However, AlGeBr_3 exhibits a higher absorption among the compounds as seen in Fig. 9(a). The absorption in the visible region is consistent with the energy band gap of AlGeX_3 [$E_g(\text{AlGeF}_3) > E_g(\text{AlGeCl}_3) > E_g(\text{AlGeBr}_3)$]. That's why these materials show greater suitability in optoelectronic devices, especially solar cells.

Photoconductivity, $\sigma(\omega)$ is an optical and electrical phenomenon in which a material becomes more electrically conductive when it is exposed to light of sufficient energy. A comparison of $\sigma(\omega)$ is shown in Fig. 10(a–d) estimated for the compounds AlGeX_3 ($X = \text{F}, \text{Cl}$, and Br) without pressure and with pressure. The conductivity spectra follow the same pattern as the imaginary part of the dielectric function $\epsilon_2(\omega)$. When F is substituted by Cl and Br the conductivity shifts towards the lower energy side as seen in Fig. 10(a). The pressure-dependent of $\sigma(\omega)$ is found to be a significant improvement for all





Fig. 8 Comparison of absorption coefficient without pressure (a) AlGeX_3 ($X = \text{F}, \text{Cl}, \text{and Br}$) and with pressure (b) AlGeF_3 , (c) AlGeCl_3 , and (d) AlGeBr_3 .



Fig. 9 Comparison of absorption coefficient in visible wavelength range without pressure (a) AlGeX_3 ($X = \text{F}, \text{Cl}, \text{and Br}$) and with pressure (b) AlGeF_3 , (c) AlGeCl_3 , and (d) AlGeBr_3 .

compounds, especially in the visible region. For the compound AlGeF_3 , the applied pressure causes a shift of the conductivity edge towards the low-energy zone, but the conductivity peaks

provide blueshift at higher energy levels. The extended view of conductivity in the visible area is shown in Fig. 11(a–d) with applied pressure. Without pressure, the compound





Fig. 10 Comparison of the real part of the conductivity without pressure (a) AlGeX_3 ($\text{X} = \text{F}$, Cl , and Br) and with pressure (b) AlGeF_3 , (c) AlGeCl_3 , and (d) AlGeBr_3 .



Fig. 11 Comparison of conductivity in visible wavelength range without pressure (a) AlGeX_3 ($\text{X} = \text{F}$, Cl , and Br) and with pressure (b) AlGeF_3 , (c) AlGeCl_3 , and (d) AlGeBr_3 .

AlGeBr_3 shows the highest conductivity in the visible area as shown in Fig. 11(a). However, the conductivity is found to be more pressure-sensitive for the compound AlGeF_3 than for

the compounds AlGeCl_3 and AlGeBr_3 . So, these tuning conductivity properties make a huge impact on optoelectronic devices.



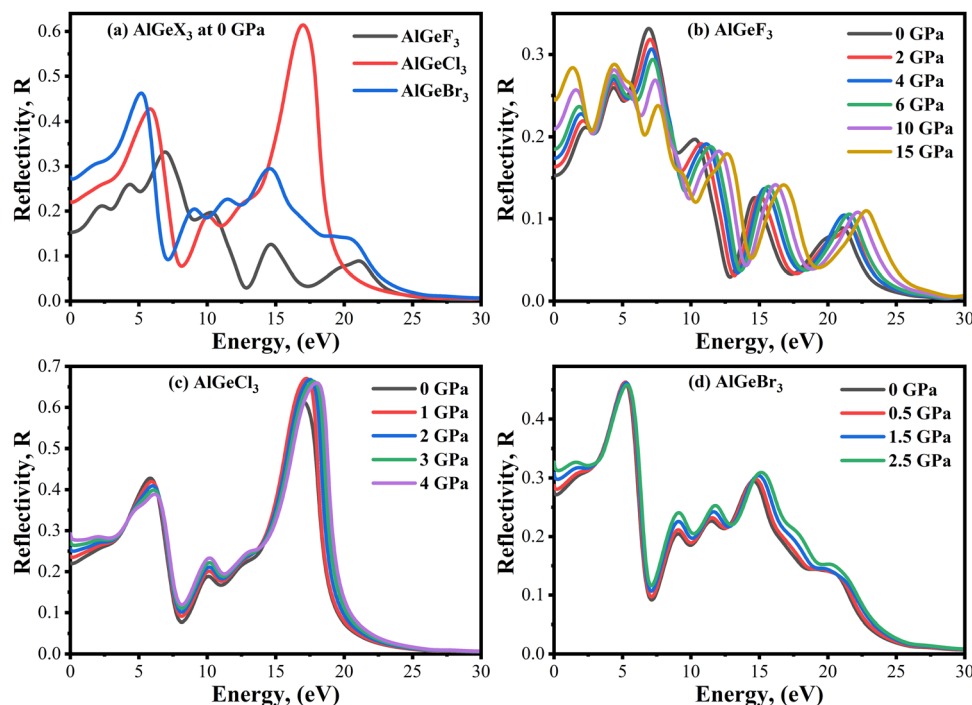


Fig. 12 Comparison of reflectivity without pressure (a) AlGeX_3 ($X = \text{F}$, Cl , and Br) and with pressure (b) AlGeF_3 , (c) AlGeCl_3 , and (d) AlGeBr_3 .

Reflectivity, $R(\omega)$ is an important parameter that determines how much light energy is reflected from a material surface. The $R(\omega)$ spectra of the compounds AlGeX_3 ($X = \text{F}$, Cl , and Br) are shown in Fig. 12(a–d) without and with pressure. The compounds AlGeF_3 and AlGeBr_3 exhibit smaller reflectance over the whole spectrum (IR, Visible, and UV) as seen in Fig. 12(a). However, the compound AlGeCl_3 shows a high reflectance peak at around 17 eV (UV region) indicating its application as a UV reflector. The pressure-dependent reflectivity is shown in Fig. 12(b–d). We observed redshift and blueshift in the reflectivity spectra at the low-energy and higher-energy regions under pressurized conditions. However, the change in the reflectivity value is found to be almost insignificant when pressure is applied.

The refractive index, $\eta(\omega)$ explains the speed of light in a medium. It depends on the real part of $\varepsilon_1(\omega)$.⁴⁸ When the refractive index falls below unity, the traveling photon's group velocity exceeds the speed of light. This property is known as superluminal. At 0 GPa, the variation of $\eta(\omega)$ follows the similar trends of $\varepsilon_1(\omega)$ as shown in Fig. 13(a). The static value of $\eta(\omega)$ is found maximum for AlGeBr_3 and minimum for AlGeF_3 , which decreases with increasing energy. When the energy reaches the UV region, the values of $\eta(\omega)$ go below unity. The values of $\eta(\omega)$ increase with increasing pressure and follow similar trends of $\varepsilon_1(\omega)$ for all the compounds as shown in Fig. 13(b–d). At the lower energy region, the $\eta(\omega)$ peaks give a redshift up to the second peak with increasing pressure for AlGeF_3 as seen in Fig. 13(b). However opposite results (blueshift) are found in the higher energy region, starting from the third peak. But for the compounds AlGeCl_3 and AlGeBr_3 , small changes of $\eta(\omega)$ are observed with pressure in the lower energy region, and no

significant changes are observed in the higher energy region shown in Fig. 13(c) and (d).

Mechanical properties

The response of a material to deformation induced by mechanical stress is quantified by elastic constants (C_{ij}). These basic constants define the connection between stress and strain across different crystallographic orientations and are critical in determining a material's mechanical characteristics. These properties are crucial in determining the anisotropic nature, mechanical stability, and bonding characteristics between adjacent atomic planes of a material.⁴⁹ How a material responds to deformation is greatly aided by the three independent elastic constants in a cubic system: C_{11} , C_{12} , and C_{44} . Elastic constants are calculated using the formulations given in.⁵⁰ C_{11} is the rigidity of the material in the direction of the applied uniaxial stress, C_{12} signifies how stresses are coupled, and C_{44} describes the resistance to shear deformation of a material. Tables S3(a), S4(a), and S5(a) (ESI[†]) listed pressure-dependent elastic constants (C_{ij}), Cauchy pressures ($C_{12} - C_{44}$), and other mechanical properties computed for the compounds AlGeF_3 , AlGeCl_3 , and AlGeBr_3 , respectively. The estimated elastic constants satisfy the Born stability criteria:⁵¹ $C_{11} > 0$; $C_{44} > 0$; $C_{11} - C_{12} > 0$, and $C_{11} + 2C_{12} > 0$ even in pressurized conditions. The values of C_{11} indicate that the compounds have greater resistance to uniaxial compressions compared to shear deformation which further increases under pressure. Also, the values of C_{12} and C_{44} exhibit higher resistance to pressure variations compared to C_{11} as seen in Tables S3(a), S4(a), and S5(a) (ESI[†]). Pressure-dependent elastic constants are illustrated in Fig. 14, where C_{11} increases linearly and C_{12} primarily decreases slightly up to





Fig. 13 Comparison of refractive index without pressure (a) AlGeX_3 ($X = \text{F}$, Cl , and Br) and with pressure (b) AlGeF_3 , (c) AlGeCl_3 , and (d) AlGeBr_3 .

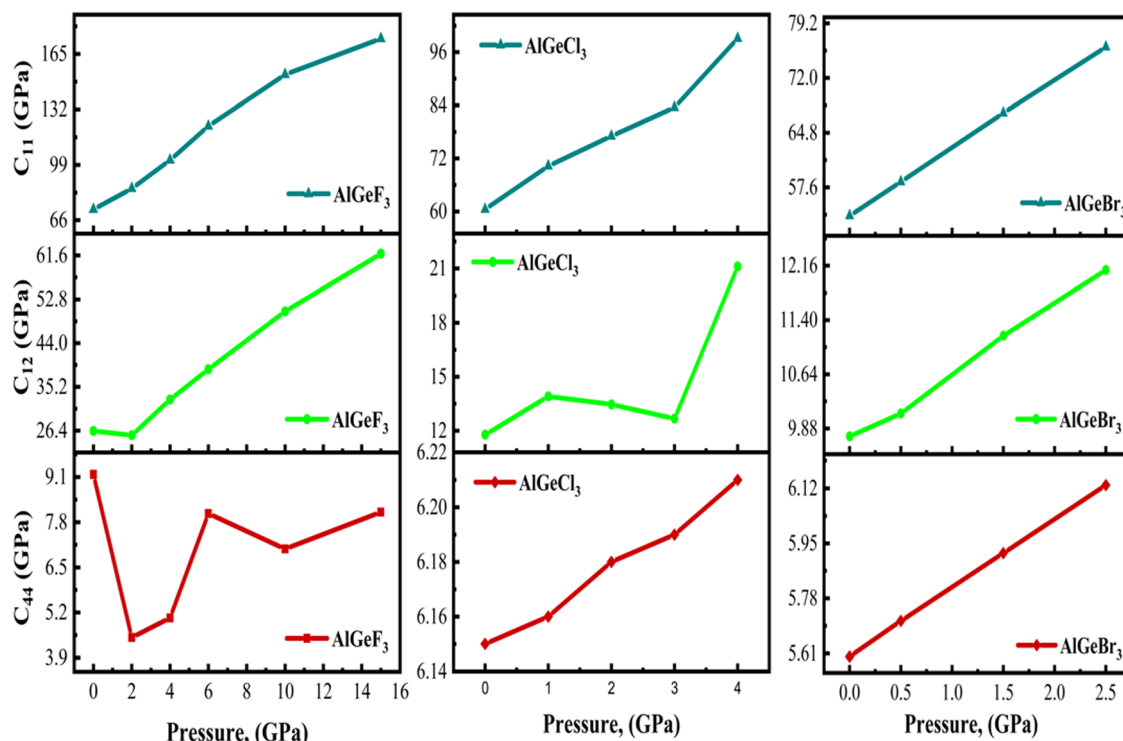


Fig. 14 Comparison of pressure-dependent elastic constants, C_{11} , C_{12} , and C_{44} of AlGeX_3 ($X = \text{F}$, Cl , and Br).

2 GPa, then increases linearly with applied pressure for AlGeF_3 and C_{44} shows non-linear characteristics throughout the pressure. Almost linear variations are found for C_{11} and C_{44} in the

case of AlGeCl_3 under pressure although C_{12} shows non-linearity. Further, linear changes of C_{11} , C_{12} , and C_{44} are observed for AlGeBr_3 up to the entire range of applied pressure.



The observed behavior can be ascribed to the transverse strain's ability to alter the shape of an object instead of changing its volume as a whole.

One of the most widely recognized factors used to assess the ductile and brittle characteristics of materials is the Cauchy pressure, expressed by the equation $C_{12} - C_{44}$. The AlGeX_3 ($X = \text{F, Cl, Br}$) perovskites have positive Cauchy pressure without pressure and it rises with increasing pressure as shown in Tables S3(a), S4(a), and S5(a) (ESI[†]). This demonstrates ductile character which further improved with increasing pressure. The pressure-dependent Cauchy pressure varies almost linearly for the compounds AlGeF_3 and AlGeBr_3 as seen in Fig. 16. But the compound AlGeCl_3 displays a rising trend up to 1 GPa then a decreasing trend at 3 GPa, and finally it increases linearly from 3 GPa to up to metallic. Among the compounds, AlGeF_3 shows the superior ductility. Tables S3(a), S4(a), and S5(a) (ESI[†]) present the calculated mechanical properties including Young's modulus (E), bulk modulus (B), shear modulus (G), Pugh's ratio (B/G), Poisson's ratio (ν), hardness (H_v), and Machinability index (μ_M) for AlGeX_3 ($X = \text{F, Cl, Br}$). To find the B and G , the Voigt–Reuss technique is used. The Voigt and Reuss coefficients stand for the maximum and minimum values of the effective modulus. The Voigt shear modulus (G_V) and Voigt bulk modulus (B_V) are defined for cubic lattices and are given⁵² by:

$$B_V = \frac{1}{3}(C_{11} + 2C_{12}) \quad (4)$$

$$G_V = \frac{1}{5}(C_{11} - C_{12} + 3C_{44}) \quad (5)$$

Furthermore, the Reuss shear modulus (G_R) and the Reuss bulk modulus (B_R) are defined in terms of cubic lattices by the following equations.

$$B_R = B_V = \frac{1}{3}(C_{11} + 2C_{12}) \quad (6)$$

$$G_R = \frac{5C_{44}(C_{11} - C_{12})}{[4C_{44} + 3(C_{11} - C_{12})]} \quad (7)$$

The bulk (B) and shear (G) moduli are calculated from the Voigt–Reuss–Hill approximation:⁵³

$$B = \frac{1}{2}(B_V + B_R) \quad (8)$$

$$G = \frac{1}{2}(G_V + G_R) \quad (9)$$

In addition, Poisson's ratio (ν) and Young's modulus (E) are estimated⁵⁰ by:

$$E = \frac{9BG}{3B + G} \quad (10)$$

$$\nu = \frac{(3B - 2G)}{2(3B + G)} \quad (11)$$

where B represents the fracture resistance, while G corresponds

to plastic deformation. AlGeF_3 shows higher bulk (B) and shear modulus (G) than AlGeCl_3 and AlGeBr_3 . This implies that AlGeF_3 exhibits superior resistance to both fracture and plastic deformation than AlGeCl_3 and AlGeBr_3 . Fig. 15 shows the plots of bulk, shear, and Young's moduli as a function of applied pressure. Under the applied pressure, B increases linearly while G initially shows a decreasing trend after that increases gradually for AlGeF_3 . However, both AlGeCl_3 and AlGeBr_3 show increasing trend linearly under pressure. A higher E indicates greater stiffness, that is, the material requires more stress to induce a given amount of elastic deformation. The E and G show a similar trend for all of the compounds as seen in Fig. 15. With pressure and without pressure, AlGeF_3 shows higher E , that is, it has more stiffness than AlGeCl_3 and AlGeBr_3 .

Pugh's ratio (B/G) is utilized to determine the ductile/brittle characteristic of a material. In general, a higher Pugh ratio greater than the critical value of 1.75 indicates better ductility, while a ratio lower than 1.75 suggests increased brittleness.⁵⁴ Since all the compounds exhibit a Pugh's ratio greater than the critical value at 0 GPa, they are ductile. The pressure-dependent Pugh's ratio of the compounds AlGeX_3 ($X = \text{F, Cl, and Br}$) is shown in Fig. 16. Initially, the Pugh's ratio of the compounds increases with increasing pressure. However, it decreases for the compounds AlGeF_3 and AlGeCl_3 at the pressure of 6 GPa and 3 GPa, respectively, afterward, an increasing trend is observed. In contrast, a linear change in the Pugh's ratio is represented in Fig. 16 for the compound AlGeBr_3 with increasing pressure. Further, the compound AlGeF_3 exhibits higher ductility among the compounds. The stress-induced deformation of a material can be explained by Poisson's ratio. Material with a higher Poisson's ratio exhibits ductile behavior which signifies plastic deformation and energy absorption before failure. Conversely, the lower ratio displays brittle behavior with limited plasticity and a tendency for sudden failure. Frantsevich and Pugh suggested a critical threshold of Poisson's ratio equal to 0.26 to distinguish between ductile and brittle characteristics.⁵⁵ Under pressurized and non-pressurized conditions the studied compounds exhibit higher Poisson's ratio indicating their ductile behavior, which was earlier confirmed by Cauchy pressure. Fig. 16 illustrates the pressure-dependent variations of Cauchy pressure, Pugh's ratio, and Poisson's ratio for the compounds.

The pressure-dependent Hardness (H_v) of the compounds is calculated using Tian's model⁴⁹

$$H_{v,\text{Tian}} = 0.92 K^{1.137} G^{0.708} \quad (12)$$

here, $K = G/B$. The results are illustrated in Fig. 17 as a function of pressure. In the case of AlGeF_3 and AlGeCl_3 , the H_v changes with a decreasing-increasing pattern. But for the AlGeBr_3 , it increases linearly with increasing pressure.

The machinability index (μ_M) indicates the material's lubrication and plasticity. The larger the value of μ_M , the greater the lubrication effectiveness, decreased feed friction, and enhanced plastic strain. The pressure-dependent μ_M is



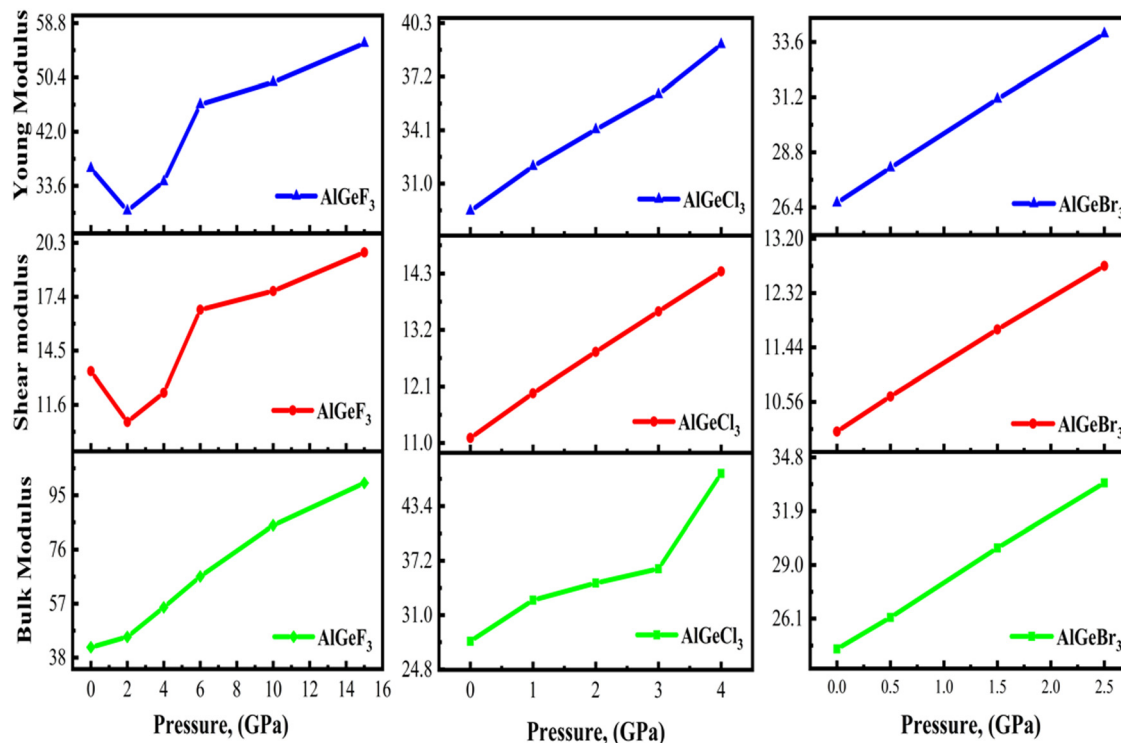


Fig. 15 Comparison of pressure-dependent Young modulus, shear modulus, and bulk modulus of AlGeX_3 ($X = \text{F}, \text{Cl}$, and Br).

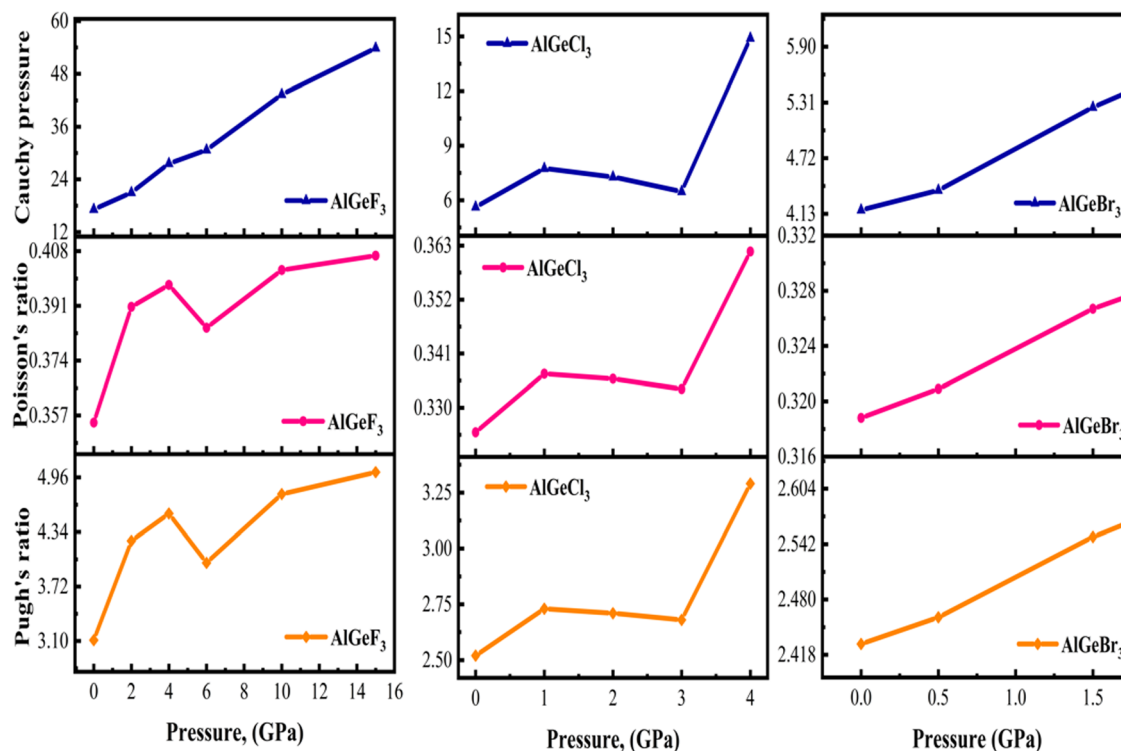


Fig. 16 Comparison of pressure-dependent Cauchy pressure, Poisson's ratio, and Pugh's ratio of cubic AlGeX_3 ($X = \text{F}, \text{Cl}$, and Br).

determined by the following expression:⁵⁶

$$\mu_{\text{M}} = \frac{B}{C_{44}} \quad (13)$$

The results are presented in Fig. 17 for the compounds AlGeX_3 ($X = \text{F}, \text{Cl}$, and Br). It is found that the μ_{M} increases sharply with increasing pressure up to 2 GPa for AlGeF_3 , then it decreases up to 6 GPa, and lastly increases up to metallic



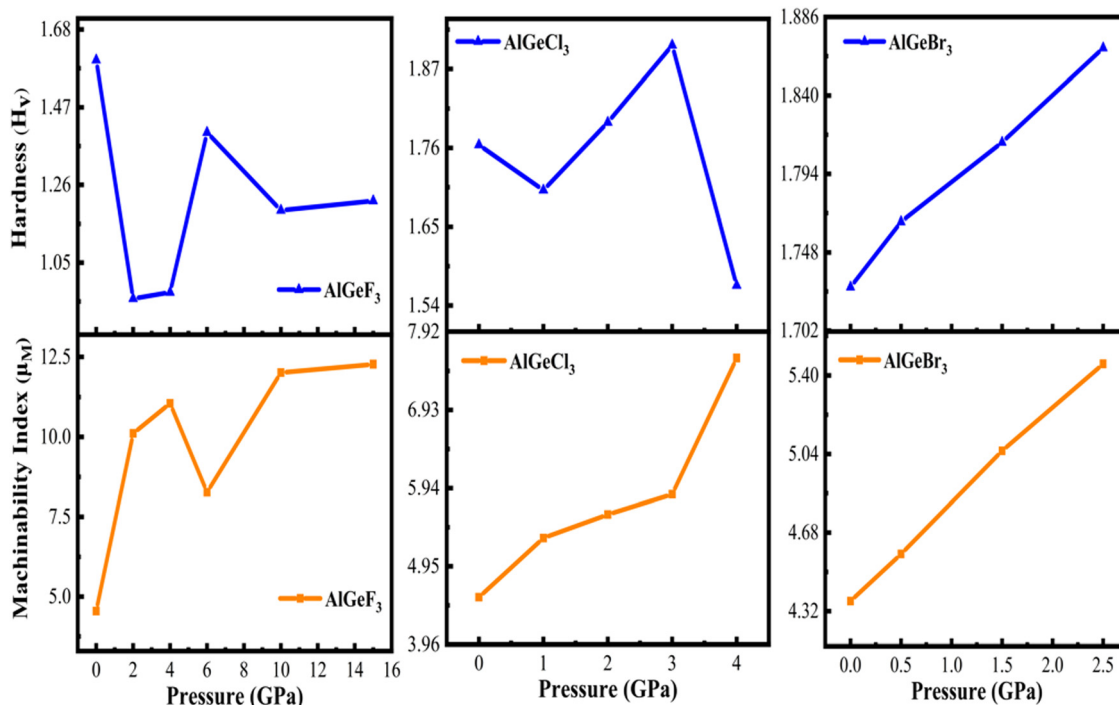


Fig. 17 Comparison of pressure-dependent hardness and machinability index of AlGeX₃ (X = F, Cl, and Br).

pressure (14 GPa). However, for the compounds AlGeCl₃ and AlGeBr₃, the μ_M increases almost linearly with increasing pressure as seen in Fig. 17.

The anisotropic index is utilized to quantify the material properties having directional dependence. The following equation is used to calculate the three shear anisotropy factors (A_1 , A_2 , and A_3) for the {100}, {010}, and {001} planes⁵⁷

$$A_1 = A_2 = A_3 = \frac{4C_{44}}{C_{11} + C_{33} - 2C_{13}} \quad (14)$$

A similar expression for the Zener anisotropy factor (A) can be calculated⁵⁸ by

$$A = \frac{2C_{44}}{C_{11} - C_{12}} \quad (15)$$

The elastic anisotropy increases for the compounds AlGeX₃ (where X represents F, Cl, Br) with increasing pressure as evidenced in Tables S3(b), S4(b), and S5(b) (ESI†). To capture this anisotropy more straightforwardly, a single anisotropy index is preferred over multiple factors for various crystal planes. Moreover, utilizing by Chung and Buessem's empirical formula determines shear (A_G) and bulk (A_B) to give the anisotropy⁵⁹. The calculated values of A_B and A_G demonstrate that all the compounds exhibit anisotropy in both pressurized and unpressurized conditions. The universal anisotropy factor (A^U) can be used for all types of crystal symmetries and its value is zero for isotropic materials.⁶⁰ The equivalent Zener anisotropy (A^{eq}) is determined using the empirical formula which is then used to obtain the proper anisotropy of AlGeX₃ (X = F, Cl, and Br).⁵⁸ The estimated pressure-dependent elastic anisotropies are listed in Tables S3(b), S4(b), and S5(c) (ESI†) for AlGeF₃,

AlGeCl₃, and AlGeBr₃, respectively. 3D anisotropic contour plots for E , G , and ν are presented in Fig. S7–S9 (ESI†) to illustrate the directional dependence of elastic properties in AlGeX₃ (X = F, Cl, Br) perovskites. The plots are generated at pressures of 0 and 15 GPa for AlGeF₃; 0 and 4 GPa for AlGeCl₃; and 0 and 2.5 GPa for AlGeBr₃. The anisotropic character of the perovskites is shown by the deviance that is revealed from the spherical form in these contour plots.⁶¹ The anisotropies of AlGeX₃ (X = F, Cl, Br) halide perovskites increase dramatically when pressure increases from 0 GPa up to metallic, indicating that pressure has a considerable impact on their directional elasticity.

Conclusion

Tuning material properties through hydrostatic pressures is a fascinating avenue of research. This study presents pressure-dependent structural, electronic, optical, and mechanical properties of lead-free AlGeX₃ (X = F, Cl, and Br) perovskites. All properties of our investigated compounds can be improved by substituting a halide component at the X sites, which can be further extended by applying hydrostatic pressure. The material's lattice parameters decrease under pressure, leading to more negative formation energies that confirm the thermodynamic stabilities. Besides, these substances exhibit essential mechanical properties like mechanical stability, anisotropy, and ductility which can be improved further by halide substitution as well as applying pressure. Every compound exhibits semiconducting properties and replacing halide (F to Cl, Br) results in an indirect-to-direct band transition, suggesting their



applications in optical devices and solar cells. Furthermore, positive hydrostatic pressure causes bandgap narrowing and semiconductor-to-metallic transition. This study confirmed that atoms in the Al-3s, F-2p, Cl-3p, Br-4p, and I-5p orbitals mostly contribute to forming the valence bands, while Al-3p and Ge-4p orbitals are mainly involved in the formation of conduction bands. Moreover, these compounds exhibit greater dielectric constant under relaxed conditions which are further improved under pressurized conditions. By precisely tuning the pressure, the absorption coefficient of these compounds can be improved significantly not only in the visible zones but also in the UV regions, demonstrating their application as solar cells, UV sensors, and UV detectors. In addition, pressures affect almost negligible change in the conductivity except AlGeF₃. So, AlGeX₃ perovskites (where X = F, Cl, and Br) are promising candidates for use in optoelectronic devices like solar cells, photosensors, and LEDs with their tunable structural, electronic, optical, and mechanical properties under pressure.

Authorship contribution

Md. Amran Sarker: investigation, methodology, data curation and analysis and interpretation, writing the original draft. Md Mehedi Hasan: investigation, methodology, data curation, graph plotting. Md. Rafiqul Islam: investigation, methodology. Md. Rabbi Talukder: investigation, methodology. Md. Rasidul Islam: bandgap calculation by HSE06 functional, investigation, methodology. Ahmed Sharif: resources allocation, supervision, writing – review & editing.

Data availability

The datasets and computer codes are available upon request from the authors.

Conflicts of interest

The authors declare no conflict of interest.

Acknowledgements

The authors express their sincere gratitude to the Central Computer Center, KUET, Bangladesh for providing the high-performance computing resources used in this research. The authors would like to express their gratitude to the Research and Innovation Centre for Science and Engineering (RISE), Bangladesh University of Engineering and Technology (BUET), Dhaka-1000, Bangladesh for funding through the RISE Internal Research Grant (2022-01-001).

References

- 1 Y. Sun, W. Zhang, H. Chi, Y. Liu, C. L. Hou and D. Fang, *Renewable Sustainable Energy Rev.*, 2015, **43**, 973–980.
- 2 A. M. Humada, M. Hojabri, S. Mekhilef and H. M. Hamada, *Renewable Sustainable Energy Rev.*, 2016, **56**, 494–509.
- 3 P. K. Nayak, S. Mahesh, H. J. Snaith and D. Cahen, *Nat. Rev. Mater.*, 2019, **4**, 269–285.
- 4 M. K. Hossain, S. Bhattacharai, A. A. Arnab, M. K. A. Mohammed, R. Pandey, M. H. Ali, Md. F. Rahman, Md. R. Islam, D. P. Samajdar, J. Madan, H. Bencherif, D. K. Dwivedi and M. Amami, *RSC Adv.*, 2023, **13**, 21044–21062.
- 5 Md. R. Islam, Y. Wu, K. Liu, Z. Wang, S. Qu and Z. Wang, *Adv. Mater. Interfaces*, 2022, **9**(4), 2101144.
- 6 H. S. Jung and N. Park, *Small*, 2015, **11**, 10–25.
- 7 S. Yun, Y. Qin, A. R. Uhl, N. Vlachopoulos, M. Yin, D. Li, X. Han and A. Hagfeldt, *Energy Environ. Sci.*, 2018, **11**, 476–526.
- 8 P. P. Boix, K. Nonomura, N. Mathews and S. G. Mhaisalkar, *Mater. Today*, 2014, **17**, 16–23.
- 9 J.-P. Correa-Baena, M. Saliba, T. Buonassisi, M. Grätzel, A. Abate, W. Tress and A. Hagfeldt, *Science*, 2017, **358**, 739–744.
- 10 I. Hussain, H. P. Tran, J. Jaksik, J. Moore, N. Islam and M. J. Uddin, *Emergent Mater.*, 2018, **1**, 133–154.
- 11 M. A. Green, A. Ho-Baillie and H. J. Snaith, *Nat. Photonics*, 2014, **8**, 506–514.
- 12 M. Solayman, M. R. Islam, M. A. Sarker, R. K. Sharpe, M. Al Momin, M. R. Islam and A. A. Khan, *Phys. Scr.*, 2023, **98**, 125944.
- 13 J. Y. Kim, J.-W. Lee, H. S. Jung, H. Shin and N.-G. Park, *Chem. Rev.*, 2020, **120**, 7867–7918.
- 14 M. K. Hossain, G. F. I. Toki, I. Alam, R. Pandey, D. P. Samajdar, Md. F. Rahman, Md. R. Islam, M. H. K. Rubel, H. Bencherif, J. Madan and M. K. A. Mohammed, *New J. Chem.*, 2023, **47**, 4801–4817.
- 15 M. R. Islam, A. S. M. J. Islam, S. K. Mitro, M. R. Hasan Mojumder, B. K. Moghal, S. Ahmad, S. Roy and M. Z. Islam, *J. Phys. Chem. Solids*, 2022, **170**, 110919.
- 16 J. S. Manser, M. I. Saidaminov, J. A. Christians, O. M. Bakr and P. V. Kamat, *Acc. Chem. Res.*, 2016, **49**, 330–338.
- 17 U.-G. Jong, C.-J. Yu, G.-C. Ri, A. P. McMahon, N. M. Harrison, P. R. F. Barnes and A. Walsh, *J. Mater. Chem. A*, 2018, **6**, 1067–1074.
- 18 R. Ali, Z.-G. Zhu, Q.-B. Yan, Q.-R. Zheng, G. Su, A. Laref, C. S. Saraj and C. Guo, *ACS Appl. Mater. Interfaces*, 2020, **12**, 49636–49647.
- 19 M. K. Hossain, G. F. I. Toki, A. Kuddus, M. H. K. Rubel, M. M. Hossain, H. Bencherif, Md. F. Rahman, Md. R. Islam and M. Mushtaq, *Sci. Rep.*, 2023, **13**, 2521.
- 20 W. S. Yang, J. H. Noh, N. J. Jeon, Y. C. Kim, S. Ryu, J. Seo and S. Il Seok, *Science*, 1979, **205**(348), 1234–1237.
- 21 M. K. Hossain, G. F. I. Toki, D. P. Samajdar, M. Mushtaq, M. H. K. Rubel, R. Pandey, J. Madan, M. K. A. Mohammed, Md. R. Islam, Md. F. Rahman and H. Bencherif, *ACS Omega*, 2023, **8**, 22466–22485.
- 22 Md. A. Sarker, M. M. Hasan, Md Al Momin, A. Irfan, Md. R. Islam and A. Sharif, *RSC Adv.*, 2024, **14**, 9805–9818.
- 23 Md. R. Islam, Md. R. H. Mojumder, R. Moshwan, A. S. M. J. Islam, M. A. Islam, Md. S. Rahman and Md. H. Kabir, *ECS J. Solid State Sci. Technol.*, 2022, **11**, 033001.



- 24 T. Ye, X. Wang, K. Wang, S. Ma, D. Yang, Y. Hou, J. Yoon, K. Wang and S. Priya, *ACS Energy Lett.*, 2021, 1480–1489.
- 25 S. Guo, Y. Zhao, K. Bu, Y. Fu, H. Luo, M. Chen, M. P. Hautzinger, Y. Wang, S. Jin, W. Yang and X. Lü, *Angew. Chem.*, 2020, **132**, 17686–17692.
- 26 Md Solayman, Md. A. Sarker, M. Muntasir, R. K. Sharme and Md. R. Islam, *Opt. Mater.*, 2024, **148**, 114699.
- 27 Md. F. Rahman, Md. A. I. Islam, Md. R. Islam, Md. H. Ali, P. Barman, Md. A. Rahman, Md Harun-Or-Rashid, M. Hasan and M. K. Hossain, *Nano Select*, 2023, **4**, 632–645.
- 28 L.-K. Gao, Y.-L. Tang and X.-F. Diao, *Appl. Sci.*, 2020, **10**, 5055.
- 29 S. K. Mitro, M. Saiduzzaman, T. I. Asif and K. M. Hossain, *J. Mater. Sci.: Mater. Electron.*, 2022, **33**, 13860–13875.
- 30 Md. M. Hasan, Md. A. Sarker, M. B. Mansur, Md. R. Islam and S. Ahmad, *Heliyon*, 2024, **10**, e34824.
- 31 Md. A. Sarker, M. Muntasir, Md Al Momin, Md Solayman and Md. R. Islam, *Adv. Theory Simul.*, 2024, 7(7), 2400112.
- 32 G. Xiang, Y. Wu, M. Zhang, J. Leng, C. Cheng and H. Ma, *Phys. Chem. Chem. Phys.*, 2022, **24**, 5448–5454.
- 33 Md. M. Hasan, Md. A. Sarker, Md. R. Islam and Md. R. Islam, *ACS Omega*, 2024, **9**, 35301–35312.
- 34 S. J. Clark, M. D. Segall, C. J. Pickard, P. J. Hasnip, M. I. J. Probert, K. Refson and M. C. Payne, *Z. Kristallogr. - Cryst. Mater.*, 2005, **220**, 567–570.
- 35 W. Kohn, *Phys. Rev. Lett.*, 1996, **76**, 3168–3171.
- 36 J. P. Perdew, K. Burke and M. Ernzerhof, *Phys. Rev. Lett.*, 1996, **77**, 3865–3868.
- 37 D. Vanderbilt, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 1990, **41**, 7892–7895.
- 38 T. H. Fischer and J. Almlof, *J. Phys. Chem.*, 1992, **96**, 9768–9774.
- 39 H. J. Monkhorst and J. D. Pack, *Phys. Rev. B: Solid State*, 1976, **13**, 5188–5192.
- 40 J. Kang, E.-C. Lee and K. J. Chang, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 2003, **68**, 054106.
- 41 F. D. Murnaghan, *Proc. Natl. Acad. Sci. U. S. A.*, 1944, **30**, 244–247.
- 42 M. Aktary, M. Kamruzzaman and R. Afrose, *Mater. Adv.*, 2023, **4**, 4494–4508.
- 43 Md. F. Rahman, Md. H. Rahman, Md. R. Islam, M. K. Hossain, A. Ghosh, Md. S. Islam, Md. M. Islam, Md Harun-Or-Rashid, H. Albalawi and Q. Mahmood, *J. Mater. Sci.*, 2023, **58**, 13100–13117.
- 44 K. Heidrich, W. Schäfer, M. Schreiber, J. Söchtig, G. Trendel, J. Treusch, T. Grandke and H. J. Stolz, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 1981, **24**, 5642–5649.
- 45 Md. R. Islam, A. A. M. Mazumder, Md. R. H. Mojumder, A. S. M. Z. Shifat and M. K. Hossain, *Jpn. J. Appl. Phys.*, 2023, **62**, 011002.
- 46 M. R. Islam, B. K. Moghal and R. Moshwan, *Phys. Scr.*, 2022, **97**, 065704.
- 47 Md. R. Islam, A. S. M. J. Islam, K. Liu, Z. Wang, S. Qu, C. Zhao, X. Wang and Z. Wang, *Phys. B*, 2022, **638**, 413960.
- 48 A. Ghosh, Md. F. Rahman, Md. R. Islam, Md. S. Islam, M. Khalid Hossain, S. Bhattacharai, R. Pandey, J. Madan, Md. A. Rahman and A. B. Md. Ismail, *Opt. Continuum*, 2023, **2**, 2144.
- 49 Q.-Q. Liang, D.-Y. Hu, T.-Y. Tang, H.-X. Gao, S.-Q. Wu, L. Li and Y.-L. Tang, *J. Mater. Res. Technol.*, 2023, **22**, 3245–3254.
- 50 M. J. Mehl, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 1993, **47**, 2493–2500.
- 51 F. Mouhat and F.-X. Coudert, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 2014, **90**, 224104.
- 52 M. I. Naher and S. H. Naqib, *J. Alloys Compd.*, 2020, **829**, 154509.
- 53 R. Hill, *Proc. Phys. Soc., London, Sect. A*, 1952, **65**, 349–354.
- 54 S. F. Pugh, *Lond. Edinb. Dubl. Phil. Mag. J. Sci.*, 1954, **45**, 823–843.
- 55 R. L. Fleischer, R. S. Gilmore and R. J. Zabala, *Acta Metall.*, 1989, **37**, 2801–2803.
- 56 M. Al-Fahdi, A. Rodriguez, T. Ouyang and M. Hu, *Crystals*, 2021, **11**, 783.
- 57 X. Gao, Y. Jiang, R. Zhou and J. Feng, *J. Alloys Compd.*, 2014, **587**, 819–826.
- 58 C. M. Zener and S. Siegel, *J. Phys. Colloid Chem.*, 1949, **53**, 1468.
- 59 R. K. Sharme, M. R. Islam, M. A. Sarker, M. Solayman, M. Al Momin and M. R. Islam, *Phys. B*, 2024, 415785.
- 60 S. I. Ranganathan and M. Ostoj-Starzewski, *Phys. Rev. Lett.*, 2008, **101**, 055504.
- 61 A. Biswas, Md. S. Alam, A. Sultana, T. Ahmed, M. Saiduzzaman and K. M. Hossain, *Appl. Phys. A*, 2021, **127**, 939.

