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Low photoactive phase temperature all-inorganic, tin–lead mixed perovskite solar cell†

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CsPbI₃ films have recently attracted significant attention as efficient absorbers for thermally stable photovoltaic devices. However, their large bandgap and photoactive black phase formation at high temperature impede their use for practical applications. Using the concept of lattice contraction, we demonstrate a low bandgap (≤ 1.44 eV) cesium-based inorganic perovskite CsPb_xSn_{1-x}I₃ that can be solution processed at low temperature for photovoltaic devices. The results from systematic measurements imply that the partial substitution of lead (Pb) with tin (Sn) results in crystal lattice contraction, which is essential for realizing photoactive phase formation at 100 °C and stabilizing photoactive phase at room temperature. These findings demonstrate the potential of using cesium-based inorganic perovskite as viable alternatives to MA- or FA-based perovskite photovoltaic materials.

Introduction

Inorganic–organic hybrid perovskites, with the general formula of ABX₃ (A is a small organic cation, such as methylammonium [MA⁺]¹ or formamidinium [FA⁺],² B is typically Pb²⁺ or Sn²⁺,^{4,6} and X is a halide),³ have recently gathered tremendous attention as promising photovoltaic^{45,47} materials with combined high-performance and low-cost. Benefitting from their high optical absorption coefficients,⁴ low exciton binding energies,^{5,6} long carrier diffusion lengths,^{7,8} and high tolerance of trap states,⁹ hybrid organic–inorganic perovskite solar cells have seen the meteoric rise of PCEs from 3.8% to over 25.2% in merely 8 years.^{10,11,44} Despite the rapid progress, these materials are found to be sensitive to temperature,^{12,13} radiation,^{14,15} oxygen,¹⁶ moisture,^{17,18} and electric field.^{19,20} Among these factors, thermal stability is one of the more challenging issues. For instance, MAPbI₃ film undergoes a reversible phase transition at 55 °C (ref. 21) and degrades above 85 °C.^{22,23} The incorporation of FA will improve thermal stability over the MA-based perovskites. Incorporating additional inorganic cations,^{13,24,25} bromide,²⁶ or chloride²⁷ can further enhance the thermal stability of resulting perovskites. Nonetheless, the intrinsic volatility of organic components in perovskite remains a liability that may limit these materials' usage under broad operating temperatures. According to the calculation of the

tolerance factor, we found that the size of a cesium ion in the periodic table is enough to occupy the position of an A⁺ ion in perovskite and has a suitable tolerance factor to match the morphology of ABX₃. To avoid the decomposition of organic cations FA⁺, MA⁺ induced by light, Cs is considered an important element that presents the opportunity to replace organic cations. In recent years, the main inorganic perovskite solar cells have proven to have good light and temperature stability.²⁸

In the inorganic perovskite, CsPbI₃ perovskite solar cell as the main research object have also achieved good efficiency. The active phase (cubic) in CsPbI₃ can only be formed and stabilized at high temperature (~ 330 °C). The black phase of the inorganic perovskite has unsuitable tolerance factor and can easily transform to the nonperovskite yellow phase at room temperature, because the smaller Cs atoms are rattling in the oversized perovskite cages.²⁹

In this regard, inorganic cesium lead triiodide (CsPbI₃) is a very attractive material because of its high thermal stability as a photovoltaic absorber. However, its relatively large optical band gap (1.73 eV) and spontaneous phase transition from black phase (perovskite phase) to photo inactive yellow phase (nonperovskite phase) at room temperature limit its potential application.²⁹ Although the stable photocatalytic activity of CsPbI₃ has been demonstrated in quantum dots, the materials and solar cell manufacturing process are rather cumbersome.³⁰ Recently, studies of cesium lead mixed-halide perovskites, CsPbI_{3-x}Br_x, found that the photoactive phase could be stabilized by partial incorporation of Br.^{31–34} However, the incorporation of Br further reduces photocurrent generation due to its increased bandgap over 1.9 eV. Consequently, CsPbI₂Br can only afford a relatively low short-circuit current density (J_{sc}) of 16.3 mA cm⁻² based on the calculation from the Shockley–Queisser limit.³¹ In addition, these inorganic perovskite films

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still require $>200\text{ }^{\circ}\text{C}$ thermal annealing to form the perovskite phase, which further limits their practical application compared to other inorganic–organic hybrid perovskites.

In our work, compositional engineering was employed to overcome the phase transition hurdle to realize inorganic perovskites with low-temperature ($\sim 100\text{ }^{\circ}\text{C}$) solution-processed photoactive phase formation and stable Cs-based photoactive phase at room temperature. Because photoactive CsPbI_3 requires a high amount of energy ($\sim 330\text{ }^{\circ}\text{C}$)²⁹ to take shape, we turned to engineering the B-site cation to mitigate phase transition instead of changing the X-site anion as reported previously. Partially substituting Pb^{2+} with smaller Sn^{2+} not only results in altered internal bonding that influences photoactive phase formation energy but also reduces bandgap, enabling it to be closer to the ideal bandgap ($\sim 1.4\text{ eV}$) for single-junction cells.³⁵ By introducing Sn into CsPbI_3 , we have demonstrated a low-temperature, solution-processed inorganic perovskite with stable photoactive phase at room temperature and tunable bandgap from 1.44 eV to 1.28 eV . In our study, through the calculation of the tolerance factor, we found that in the octahedral perovskite lattice, the mixed lead–tin ratio is $0.75 : 0.25$,

$0.5 : 0.5$, and $0.25 : 0.75$, respectively. Tin can gradually replace the position of lead in the lattice and produce lattice shrinkage effect. Finally, by means of lattice shrinkage effect, we have successfully fabricated three kinds of novel lead-and-tin-mixed Perovskite solar cells with $\text{CsPb}_{0.75}\text{Sn}_{0.25}\text{I}_3$, $\text{CsPb}_{0.5}\text{Sn}_{0.5}\text{I}_3$, and $\text{CsPb}_{0.25}\text{Sn}_{0.75}\text{I}_3$.

Results and discussion

Inorganic cesium lead triiodide perovskite starts to stand out due to its excellent thermal and photo stabilities.^{28,30,36–39} However, its relatively large optical bandgaps and high requirement $>200\text{ }^{\circ}\text{C}$ of thermal annealing temperature (1.73 eV for CsPbI_3 , 1.90 eV for CsPbI_2Br) make it nonideal for solar energy conversion. According to the S–Q theory, the optimal band gap of a single junction solar cell under AM 1.5G solar spectrum is around 1.4 eV ,⁴⁰ Here we develop a new inorganic perovskite semiconductor system, which has a narrower band gap than the traditional lead-based inorganic perovskite, as shown in Fig. 1a. We engineer the B-site cation instead of the widely studied X-site anion. Encouragingly, the partial

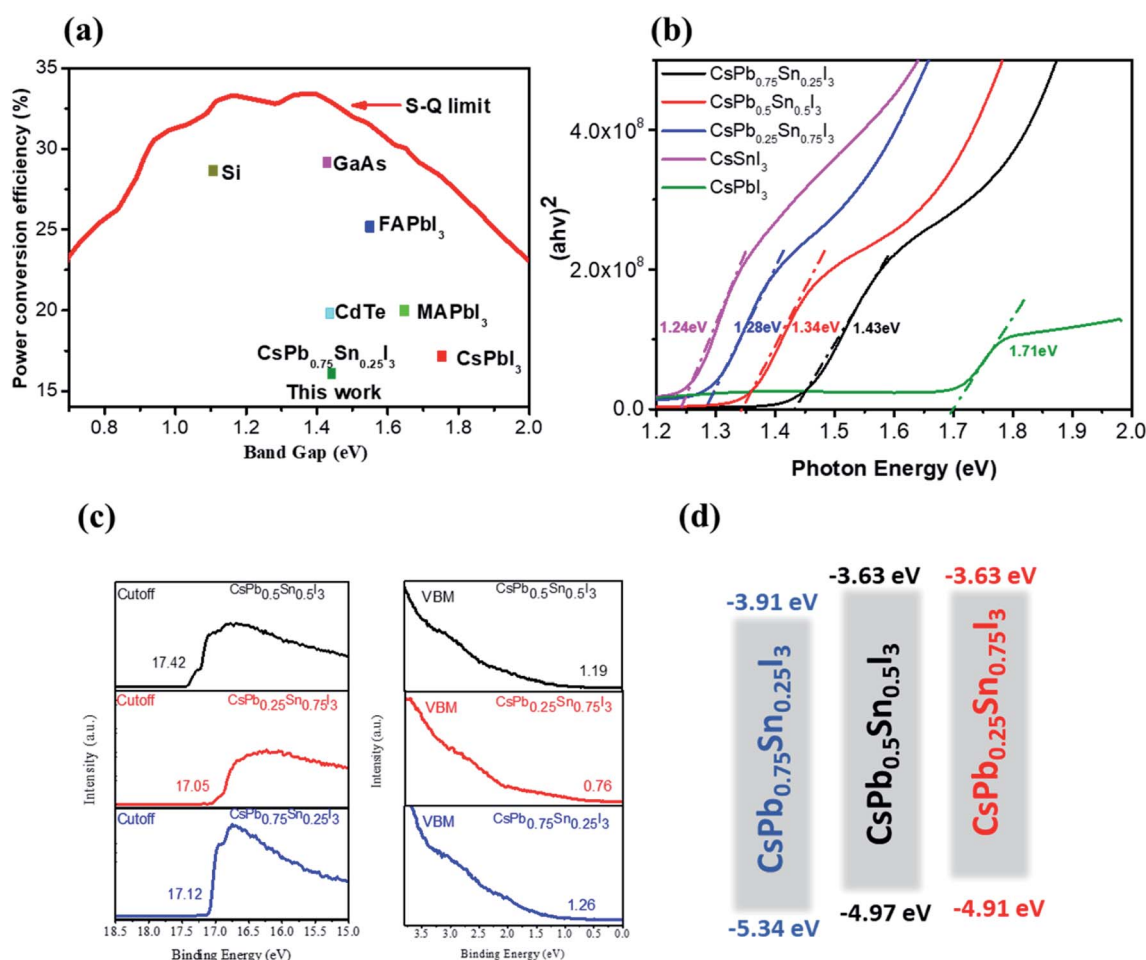


Fig. 1 Basic properties of alloyed $\text{CsPb}_x\text{Sn}_{1-x}\text{I}_3$ inorganic perovskites. (a) Shockley–Queisser limit of $\text{CsPb}_{0.75}\text{Sn}_{0.25}\text{I}_3$ perovskite and other commonly studied photovoltaic materials. (b) Tauc plot of $\text{CsPb}_x\text{Sn}_{1-x}\text{I}_3$ inorganic perovskites. (c) Ultraviolet photoelectron spectroscopy of $\text{CsPb}_x\text{Sn}_{1-x}\text{I}_3$ inorganic perovskites. (d) Band levels of $\text{CsPb}_x\text{Sn}_{1-x}\text{I}_3$ inorganic perovskites (valence band is estimated from ultraviolet photoelectron spectroscopy and the conduction band is calculated by subtracting the band gap estimated from Tauc plot of absorption).

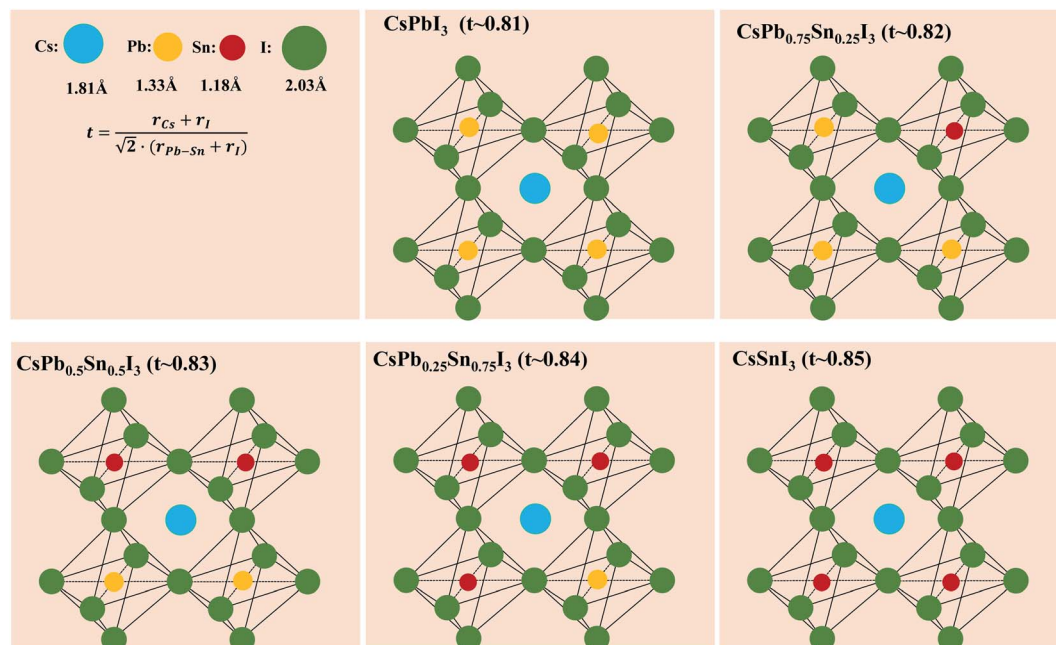


Fig. 2 Possible structures for $\text{CsPb}_x\text{Sn}_{1-x}\text{I}_3$ and their corresponding calculated Goldschmidt tolerance factor. $r_{(\text{Pb-Sn})}$ is defined as the efficient radius of Pb–Sn binary based on its molar ratio.

substitution of Pb^{2+} cation with smaller Sn^{2+} cation not only alters the tolerance factor, but also reduces the bandgap of the inorganic perovskite semiconductor. By introducing Sn into CsPbI_3 , we have demonstrated a series of low-temperature-processed inorganic perovskites with more stable black phases and tunable bandgaps ranging from 1.73 eV to 1.24 eV, as shown in Fig. 1b. A red-shifted absorption onset is observed as the Sn content increases, and the optical bandgaps estimated from the Tauc-plots are 1.43 eV, 1.34 eV, and 1.28 eV for $\text{CsPb}_{0.7}\text{Sn}_{0.3}\text{I}_3$, $\text{CsPb}_{0.5}\text{Sn}_{0.5}\text{I}_3$, and $\text{CsPb}_{0.3}\text{Sn}_{0.7}\text{I}_3$, respectively. Many successful photovoltaic materials have band gaps within this region as predicted by S–Q theory; for example, GaAs (1.43 eV) has a theoretical efficiency limit of 33% and it holds the record of 29.1% as the most efficient single junction photovoltaic ever. The alloyed $\text{CsPb}_{0.7}\text{Sn}_{0.3}\text{I}$ perovskite has the same inorganic semiconductor nature and the same band gap of the GaAs, which appears to be very promising and hence is the focus of this manuscript. Fig. 1c and d show the ultraviolet photoelectron-spectroscopy measurement results and the calculated energy levels of the Sn–Pb-alloyed inorganic perovskites. In addition to the band gap tunability, the lattice contraction induced by smaller sized Sn^{2+} cation also benefits the stabilization of photoactive black phase with low formation energy.

Table S1† gives the annealing temperatures and optical band gaps of inorganic perovskites reported in the literature, revealing the advantages of our work.

Fig. 2 shows the proposed perovskite structure and its corresponding Goldschmidt tolerance factor. In the angle of $[\text{Pb}_x\text{Sn}_{1-x}\text{I}_6]^{4-}$ octahedron, the metal sites are randomly occupied by Sn or Pb atoms.⁴¹ The ionic radius of the components used here can be found in previous literature.⁴² As shown in Fig. 2,

when the Sn content increases, the tolerance coefficient increases from 0.81 in CsPbI_3 to 0.85 in CsSnI_3 . This result predicts that a stable Cs-based triiodide perovskite phase can be obtained due to the partial substitution of lead and tin. The increase of the empirical index indicates that the formation energy of the photoactive symmetric lattice is lower.⁴² For example, CsSnI_3 can obtain stable photoactive phase (orthorhombic phase) at low temperature.

We fabricated inorganic perovskite films with lead–tin ratios of 1 : 0, 0.75 : 0.25, 0.5 : 0.5, 0.25 : 0.75, and 0 : 1. Fig. 3a shows 2-D X-ray diffraction (2-D XRD) patterns of $\text{CsPb}_x\text{Sn}_{1-x}\text{I}_3$ films with x being 0, 0.25, 0.5, 0.75, and 1. In the CsPbI_3 films, no peaks associated with the perovskite phase can be detected in room temperature air. However, after Sn incorporation (>0.25), the characteristic peaks related to photoactive perovskite phase could be clearly observed. Interestingly, the minor features (e.g., peaks at 20.3° , 21.3° , 22.9° , 24.0° , 25.1° , and 27.1°) observed in XRD patterns of perovskite phase are in good agreement with the reported orthorhombic structure (black γ -phase) of CsSnI_3 .²⁹ This result also coincides with the previous work, in which the Cs-based perovskite has small tolerance factor and lower formation energy with orthorhombic structure. To investigate further the influence of replacing with Sn^{2+} on crystal lattice, the zoomed-in region from 14.10° to 14.50° is displayed in Fig. 3b. As expected, with the increased Sn to Pb ratio, the diffraction peak gradually shifted to larger 2θ values, from 14.31° in $\text{CsPb}_{0.75}\text{Sn}_{0.25}\text{I}_3$ to 14.40° in CsSnI_3 . This slight shift of peaks to larger angles is due to lattice contraction from partial replacement of Pb with Sn. To provide a visual comparison of as formed $\text{CsPb}_x\text{Sn}_{1-x}\text{I}_3$ films, atomic force microscopy was used to investigate the morphology. As shown in Fig. 3c, the grain size first decreases significantly then



increases when the Sn content exceeds 50%. Moreover, the films have many pinholes, while the photoactive phase formation is observed in all films (except for CsPbI₃). This result implies the smooth morphology and grain size are not the main factors for forming photoactive phase with low-temperature processing. Thus, we believe that lattice contraction is the reason for forming low-temperature photoactive phase with Sn incorporation. Fig. 3d shows the absorption spectra of CsPb_xSn_{1-x}I₃ films. For CsPbI₃ film under yellow phase from our low-temperature process, almost no absorption occurs beyond 600 nm. For Sn-incorporated films, a red-shifted absorption onset is observed as the Sn content (1 - x) increases. The insert in Fig. 3d is the absorption onset of the films, and the tunable optical bandgap is estimated to be 1.44 eV, 1.35 eV, and 1.28 eV for CsPb_{0.75}Sn_{0.25}I₃, CsPb_{0.5}Sn_{0.5}I₃, and CsPb_{0.25}Sn_{0.75}I₃, respectively, showing their great potential as photovoltaic materials.

To study the thermal stability of the films obtained, a picture of such films annealed at 150 °C in a glove box (N₂) is shown in Fig. 4. We also compared the MAPbI₃ films prepared under the same conditions. As shown in Fig. 4, MAPbI₃ completely decomposed after 30 min, whereas CsPb_{0.75}Sn_{0.25}I₃, CsPb_{0.5}Sn_{0.5}I₃, CsPb_{0.25}Sn_{0.75}I₃, and CsSnI₃ did not change after 1 h. This result confirms that our CsPb_xSn_{1-x}I₃ is much better than MAPbI₃ in thermal stability.

To examine the performance of devices based on these materials, devices containing CsPb_{0.75}Sn_{0.25}I₃, CsPb_{0.5}Sn_{0.5}I₃, and CsPb_{0.25}Sn_{0.75}I₃ were fabricated with the configuration of ITO/PEDOT/PVSK/PCBM/bis-C₆₀/Ag. Fig. 5a shows the corresponding photocurrent density-voltage (*J*-*V*) characteristics of devices under AM 1.5 condition. The photovoltaic parameters, open circuit voltage (*V*_{OC}), short-circuit current density (*J*_{SC}), fill factor (FF), and PCE are tabulated in Table 1. In the CsPb_{0.75}Sn_{0.25}I₃ device, a PCE of 5.78% was achieved with a *J*_{SC} of 15.4 mA cm⁻², a *V*_{OC} of 0.67 V, and a FF of 0.56 under reverse scan.

However, lower PCE of 2.72% and 1.1% were observed in CsPb_{0.5}Sn_{0.5}I₃ and CsPb_{0.25}Sn_{0.75}I₃ devices, respectively. One possible factor for the dropped PCE is the increased recombination at sub-bandgap states due to defects caused by the oxidation of Sn²⁺ when the Sn content is increased.⁴³ Nevertheless, it is interesting to point out that there is negligible hysteresis in these inorganic perovskite devices, unlike that found in FA- and MA-based Pb-Sn binary perovskite devices.⁴³ This phenomenon coincides well with our previous finding⁹ that the motion of cation at A-site contributes noticeably to the hysteresis.

The incident photon-to-current conversion efficiency spectra for these devices is shown in Fig. 5b. This spectrum shows broad absorption for light harvesting, with the edge extending from 860 nm to 950 nm when the Sn content increases from

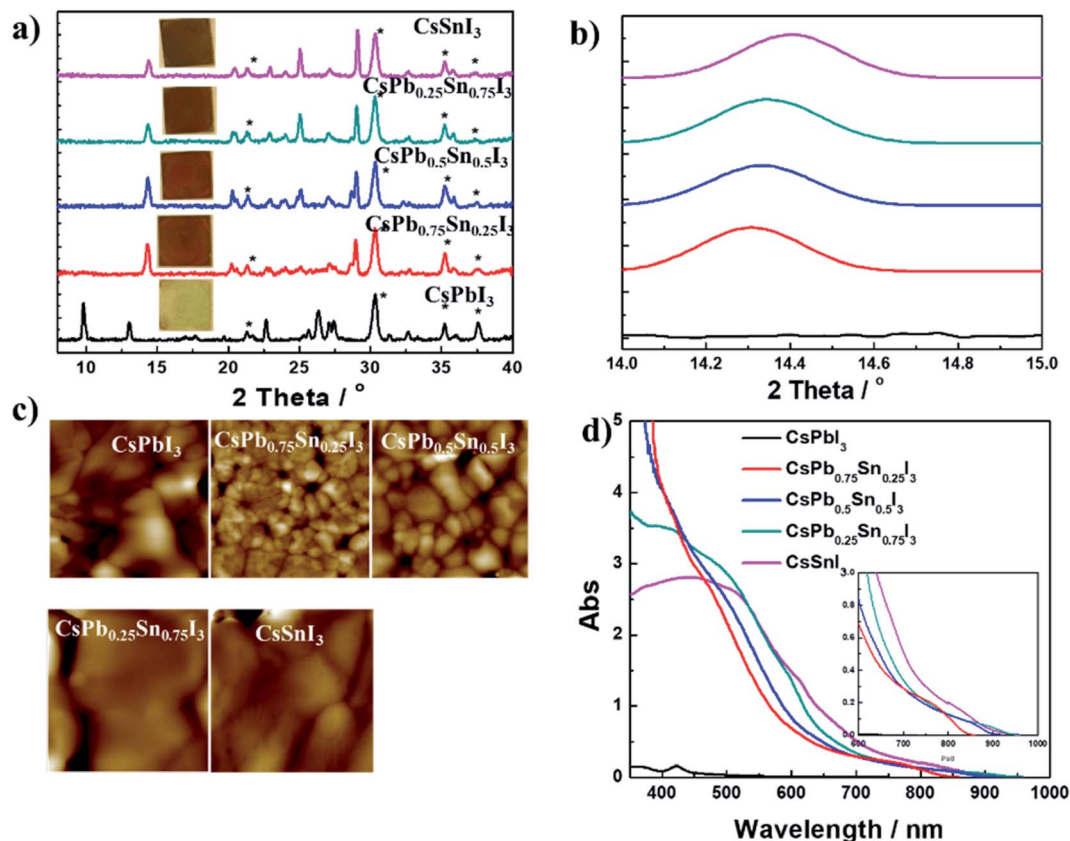


Fig. 3 (a) Photograph of obtained CsPb_xSn_{1-x}I₃ films and their 2D XRD patterns. The peaks associated with ITO are marked with asterisk. (b) Close-up 2-D XRD image of characteristic peak (~14.3°). (c) 1 μm × 1 μm atomic force microscopy images of films. (d) The absorption spectra of CsPb_xSn_{1-x}I₃ obtained by the antisolvent washing method (inserted is the absorption onset). All samples were tested in air without extra protection or encapsulation.

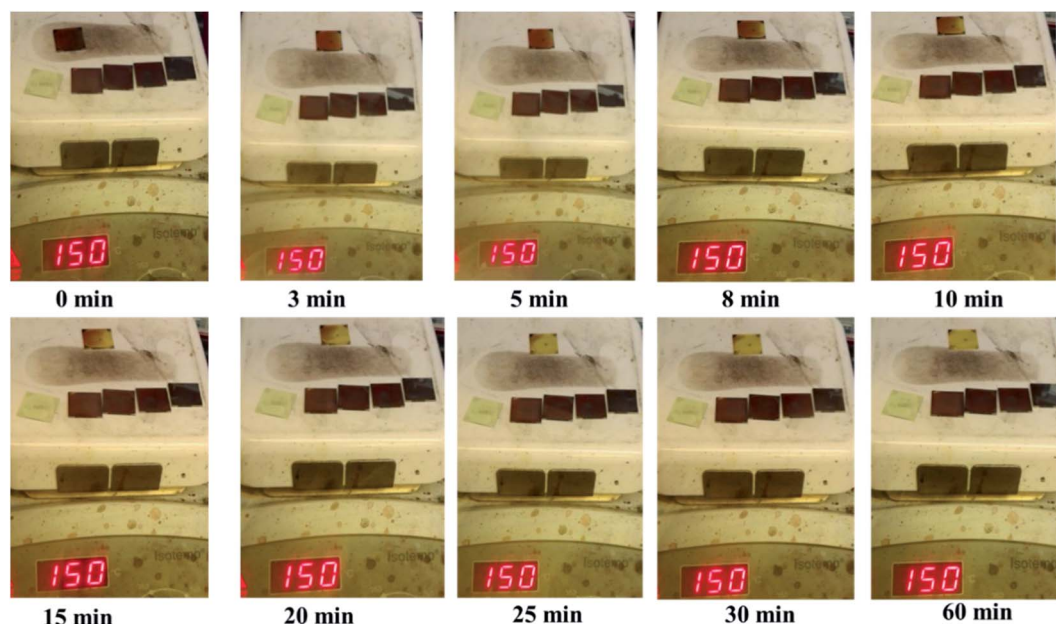


Fig. 4 Images of perovskite films annealed at 150 °C in glove box (N_2). The front row shows only $MAPbI_3$. The second row (from left to right) shows $CsPbI_3$, $CsPb_{0.75}Sn_{0.25}I_3$, $CsPb_{0.5}Sn_{0.5}I_3$, $CsPb_{0.25}Sn_{0.75}I_3$, and $CsSnI_3$.

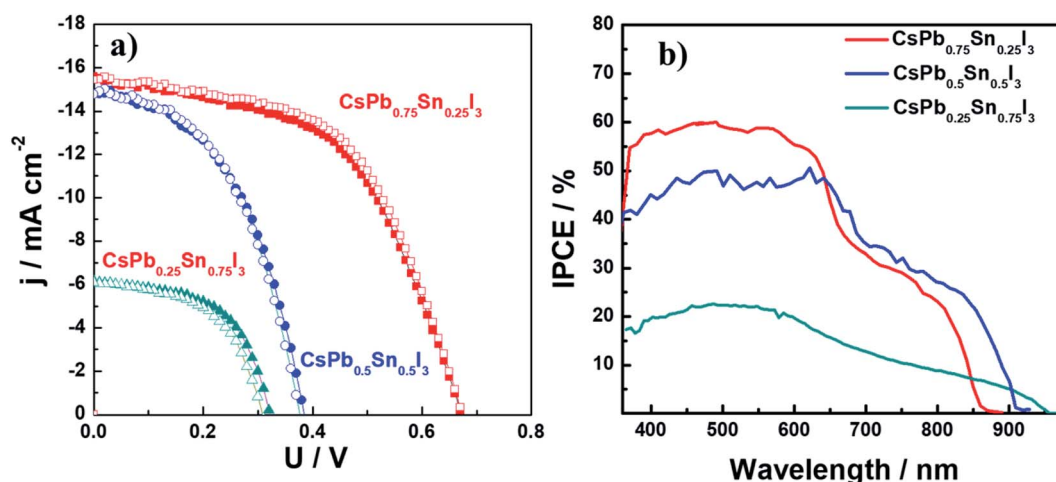


Fig. 5 (a) The J - V curves obtained under full sun illumination for devices based on $CsPb_{0.75}Sn_{0.25}I_3$, $CsPb_{0.5}Sn_{0.5}I_3$, and $CsPb_{0.25}Sn_{0.75}I_3$ as photoactive layers with ITO/PEDOT/PVK/PCBM/bis- C_{60} /Ag configuration. (b) Their incident photon-to-current conversion efficiency spectra. All the samples were tested without extra protection or encapsulation.

Table 1 Detailed photovoltaic parameters of the devices containing different photoactive layers with different scan directions under AM 1.5 G illumination (100 mW cm^{-2} light intensity)

Photoactive layer	Scan direction	V_{oc}/V	$J_{sc}/\text{mA cm}^{-2}$	FF	PCE/%
$CsPb_{0.75}Sn_{0.25}I_3$	Forward scan	0.67	15.6	0.53	5.52
	Reverse scan	0.67	15.4	0.56	5.78
$CsPb_{0.5}Sn_{0.5}I_3$	Forward scan	0.38	14.8	0.48	2.72
	Reverse scan	0.39	14.9	0.47	2.72
$CsPb_{0.25}Sn_{0.75}I_3$	Forward scan	0.31	6.10	0.53	1.00
	Reverse scan	0.32	6.20	0.55	1.10

25% to 75%, which is in good agreement with the corresponding UV-vis absorption.

Conclusion

Organic-inorganic hybrid perovskite solar cells have proven to have good photoelectric properties, but large-scale production requires considerable light and thermal stability. The presence of MA^+ or FA^+ in organic perovskite is easy to decompose, whereas cesium-based, all-inorganic perovskite has good thermal stability. Through the co-doping of lead and tin, the photoactive phase of perovskite successfully reduced the



annealing temperature to 100 °C. To sum up, a low-temperature annealing, Cs-based, all-inorganic perovskite was developed to display stable black phase. Systematic measurements show that the formation of shrinkage lattice cells by partial substitution of Sn for Pb is helpful to the realization and stabilization of photoactive perovskite phase. Through the calculation of the tolerance factor, we have successfully produced all-inorganic, lead-tin-mixed perovskite solar cells with lead-to-tin ratios of 0.75 : 0.25, 0.5 : 0.5, and 0.25 : 0.75. However, when the ratio of lead to tin is 0.75 : 0.25, the initial efficiency of all inorganic perovskite solar cells is 5.78%.

Fabrication of solar cells

ITO glass was first cleaned using detergent, followed by 10 min ultrasonication with detergent, DI water, acetone, and isopropanol, respectively. Then, the substrates were dried by nitrogen gas flow and treated with ultraviolet-ozone for 15 min before depositing PEDOT. The PEDOT/PSS solution was spin-coated onto the cleaned ITO glass substrate at 4000 rpm for 30 s and annealed at 150 °C for 15 min in air. To prepare the precursor solution of $\text{CsPb}_x\text{Sn}_{1-x}\text{I}_3$ perovskite, the CsPbI_3 and CsSnI_3 (with 10% ascorbic acid) precursor was made first in a mixed solvent of DMF and DMSO (4 : 1). Then, the 2 precursor solutions were mixed.

Afterward, a 35 μL perovskite precursor (kept at room temperature) was spin-coated onto the PEDOT layer at 4000 rpm for 35 s in the glovebox. Before spin-coating, the PEDOT substrate was kept at room temperature. Toluene (800 μL) was dripped *in situ* onto the substrate at the 10th s. Afterward, the perovskite films were annealed at 60 °C for 1 min and 100 °C for 4 min. PCBM solution (20 mg ml^{-1} in chlorobenzene) was then spin-coated onto the perovskite film at 2000 rpm for 30 s, followed by the deposition of bis- C_{60} layer (2 mg ml^{-1} in isopropyl alcohol, spin-coated at 3000 rpm for 35 s). Finally, Ag electrode with a thickness of 120 nm was evaporated under high vacuum ($<2 \times 10^{-6}$ torr).

Characterization

XRD was measured by X-ray diffractometer (Bruker F8 Focus Powder for regular XRD and Bruker D8 Discover 2-D XRD). SEM was measured by using FEI Sirion and operated at 5 kV. The absorption spectra were measured using a Varian Cary 5000 UV-Vis spectrometer. The J - V curves were recorded using a Keithley 2400 Source measurement unit, a 450 W xenon lamp was used to produce light, and a standard Si photodiode detector calibrated the light intensity (100 mW cm^{-2}). The EQE spectra were measured by a joint system of monochromated 450 W xenon lamp (Oriol) and a source meter (Keithley 2400), which were calculated using a calibrated Si photodiode (OSI-Optoelectronics).

Time-resolved PL decay traces were acquired using a time-correlated single photon counting system (Fluo Time 100, PicoQuant GmbH). Samples were photoexcited using a 470 nm laser beam (LDH-P-C-470, PicoQuant GmbH) pulsed at

frequencies between 0.5–10 MHz, with a pulse duration of 60 ps and fluence of $\sim 10 \text{ nJ cm}^{-2}$.

Conflicts of interest

There are no conflicts to declare.

Acknowledgements

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