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# Rotational excitation of $C_2(X^1\Sigma_g^+)$ by *para*- and *ortho*- $H_2$

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A new four dimensional (4D) potential energy surface for the  $C_2(X^1\Sigma_g^+)-H_2$  van der Waals system is generated. The potential was obtained from a multi-reference internally contracted configuration-interaction method including the Davidson correction (MRCI+Q). The four atoms were described using the augmented correlation-consistent quadruple zeta (aug-cc-pVQZ) basis sets. Both molecules were treated as rigid rotors. Close-coupling calculations of the inelastic integral cross sections of  $C_2$  in collisions with *para*- $H_2(j_{H_2} = 0)$  and *ortho*- $H_2(j_{H_2} = 1)$  were also carried out at low energies. After Boltzmann thermal averaging, rate coefficients were obtained for temperatures ranging from 5 to 100 K. The rate coefficients for collisions with *ortho*- $H_2$  are significantly larger than the rate coefficients for collisions with *para*- $H_2$ .

## 1 Introduction

The  $C_2$  molecule is one of the best known constituents of the interstellar medium. It was first detected by Souza and Lutz.<sup>1</sup> Subsequently, many absorption lines have been detected in the near-infrared region for the  $A^1\Pi_u-X^1\Sigma_g^+$  Phillips system,<sup>2-8</sup> and in the UV region for the  $D^1\Sigma_u-X^1\Sigma_g^+$  Mulliken system<sup>9</sup> and the  $F^1\Pi_u-X^1\Sigma_g^+$  band.<sup>10,11</sup> The  $C_2$  molecule is used as a diagnostic probe of the physical condition in diffuse molecular clouds. The populations of its rotational levels are determined by competition between collisional excitation and radiative excitation through absorption into excited electronic states followed by fluorescence to various vibrational levels of  $X^1\Sigma_g^+$ .<sup>2</sup> The analysis of rotationally excited  $C_2$  provides useful information about the density of collision partners, the kinetic temperature and the strength of the radiation field permeating the gas.<sup>12</sup> In order to perform a more accurate analysis of  $C_2$  excitation, rotational rate coefficients of  $C_2$  excited by its principal interstellar collisional partner,  $H_2$ , are required. In ref. 13, the full dimensional PES was calculated by the CCSD(T)-F12a/cc-pCVTZ-F12 method. However, we suppose that there is a lack of static correlation because of the multi configurational character of the PES for interacting  $C_2$  and  $H_2$ . In this paper, we present a new complete four dimensional (4D) PES for the ground electronic state of the  $C_2-H_2$  collisional system calculated at the MRCI+Q level. Then,

this accurate PES was used to perform rotational excitation of  $C_2$  with both *para*- $H_2$  and *ortho*- $H_2$ .

The activation energy of the  $C_2 + H_2 \rightleftharpoons C_2H + H$  reaction being about  $900\text{ cm}^{-1}$ ,<sup>14</sup> inelastic cross sections of  $C_2$  in collisions with  $H_2(j_{H_2} = 0)$  and  $H_2(j_{H_2} = 1)$  are reported for total energies of up to  $1000\text{ cm}^{-1}$  for the first 9 even rotational levels, yielding converged rate coefficients up to 100 K.

Excitation rate coefficients of  $C_2$  excited by *para*- $H_2(j_{H_2} = 0)$  were provided by Najar *et al.*<sup>15</sup> In this study, the PES was derived by averaging only over three specific  $H_2$  orientations. These calculations gave a first estimate of  $C_2$ -*para*- $H_2$  collisional rate coefficients. However, rotational excitation rate coefficients for  $C_2$  in collision with *ortho*- $H_2$  are also needed for a complete description of  $C_2$  excitation.<sup>16</sup>

The paper is organized as follows: Section 2 describes the calculation of the potential energy surface, Section 3 contains a concise description of the scattering calculations, and in Section 4, we present and discuss our results.

## 2 Potential energy surface

As mentioned in previous work,<sup>17</sup> the ground electronic state of the  $C_2-H_2$  system is described by two dominant configurations. Because of this multi configurational character, the potential energy surface (PES) of this van der Waals complex was accomplished by first performing full valence complete active space (CASSCF) calculations<sup>18,19</sup> followed by multi-reference internally contracted configuration-interaction (MRCI) calculations,<sup>20,21</sup> including all single and double excitations from the CASSCF reference function. The contributions of higher-order excitations were estimated using the normalized multi-reference Davidson correction.<sup>22,23</sup> In the CASSCF-MRCI calculations, the active space is defined by distributing 10 electrons

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in 10 molecular orbitals. The  $1s^2$  orbitals of the carbon atoms were kept frozen. To achieve a good description of weak interactions, the calculations were performed using a rather large aVQZ basis set of Dunning.<sup>24–26</sup> All *ab initio* calculations were carried out using the MOLPRO program package.<sup>27</sup>

The body-fixed coordinate system presented in Fig. 1 was used. The center of the coordinates is placed in the middle of the bond of the  $C_2$  molecule. Vector  $\mathbf{R}$  is directed along the z-axis and connects the  $C_2$  and  $H_2$  centers of mass. The rotation of the  $C_2$  molecule is defined by angle  $\theta_{C_2}$  and the rotation of the  $H_2$  molecule is defined by angle  $\theta_{H_2}$ , while  $\phi$  is the dihedral angle.

The *ab initio* calculations were performed using the approximation of rigid interacting molecules with the  $H_2$  bond length corresponding to the ground vibrational state:  $r_{H-H} = 1.449$  Bohr.<sup>28</sup> For  $C_2$ , we used the experimental equilibrium distance  $r_{C-C} = 2.348$  Bohr,<sup>29</sup> which differs by less than 1% from the calculated value obtained by averaging over the  $\nu = 0$  wavefunction. We note that employing state-averaged geometries is a reliable approximation for including zero-point vibrational effects within a rigid-rotor PES.<sup>30–32</sup>

In order to determine the interaction potential  $V(R, \theta_{C_2}, \theta_{H_2}, \phi)$ , the basis set superposition error (BSSE) was corrected for all geometries with the Boys and Bernardi counterpoise scheme:<sup>33</sup>

$$V(R, \theta_{C_2}, \theta_{H_2}, \phi) = E_{C_2-H_2}(R, \theta_{C_2}, \theta_{H_2}, \phi) - E_{C_2}(R, \theta_{C_2}, \theta_{H_2}, \phi) - E_{H_2}(R, \theta_{C_2}, \theta_{H_2}, \phi) \quad (1)$$

where the energies of the  $C_2$  and  $H_2$  subsystems are computed using a full basis set of the complex. The size inconsistency was corrected for all geometries by adding the energy  $-33.75$   $\text{cm}^{-1}$ .

At each point  $R$  the calculations were carried out for 122 angular orientations with  $\theta_{C_2}$  varying from  $0^\circ$  to  $180^\circ$ ,  $\theta_{H_2}$  varying from  $0^\circ$  to  $90^\circ$ , and  $\phi$  varying from  $0^\circ$  to  $90^\circ$ .  $R$ -distances were varied from 4.8 to 20 Bohr, giving 29 grid points.

For the solution of the close-coupling scattering equations, it is most convenient to expand the interaction potential  $V(R, \theta_{C_2}, \theta_{H_2}, \phi)$  into the angular functions at each value of  $R$ . For the scattering of two linear rigid rotors, we used the expansion of Green:<sup>34</sup>

$$V(R, \theta_{C_2}, \theta_{H_2}, \phi) = \sum_{l_1, l_2, l} v_{l_1, l_2, l}(R) A_{l_1, l_2, l}(\theta_{C_2}, \theta_{H_2}, \phi) \quad (2)$$

The basis functions  $A_{l_1, l_2, l}(\theta_{C_2}, \theta_{H_2}, \phi)$  are products of the associated Legendre polynomials  $P_{lm}$ :

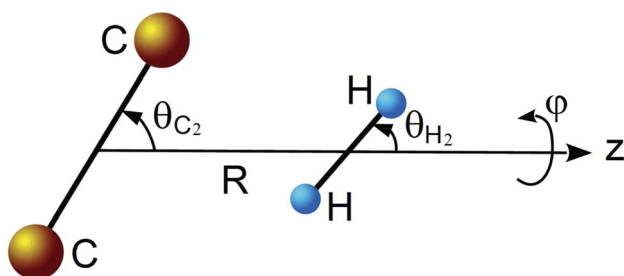


Fig. 1 Geometry describing the  $C_2-H_2$  complex.

$$A_{l_1, l_2, l}(\theta_{C_2}, \theta_{H_2}, \phi) = \sqrt{\frac{2l_1 + 1}{4\pi}} \left\{ \langle l_1 0 l_2 0 | l_1 l_2 l 0 \rangle P_{l_1 0}(\theta_{C_2}) P_{l_2 0}(\theta_{H_2}) + 2 \sum_m (-1)^m \langle l_1 m l_2 -m | l_1 l_2 l 0 \rangle P_{l_1 m}(\theta_{C_2}) P_{l_2 m}(\theta_{H_2}) \cos(m\phi) \right\} \quad (3)$$

where  $\langle \dots | \dots \rangle$  is a Clebsch–Gordan coefficient. The  $P_{lm}$  functions are related to coupled spherical harmonics through  $Y_{lm}(\theta, \phi) = P_{lm}(\theta) e^{im\phi}$ . Here  $l_1$  and  $l_2$  are associated respectively with the rotational motion of  $C_2$  and  $H_2$ , and  $l = |l_1 - l_2|, \dots, |l_1 + l_2|$ . In eqn (2), the homonuclear symmetry of the interacting molecules forces the indexes  $l_1$  and  $l_2$  to be even.

To obtain the expansion described above, at each point of the radial grid  $R$  we fitted the *ab initio* points to expression (2) using a linear least squares method, with  $l_1^{\text{max}} = 6$  and  $l_2^{\text{max}} = 4$ . The choice of such expansion parameters limits the results to a total of 28 radial expansion coefficients  $v_{l_1, l_2, l}(R)$ . The root-mean-square error (rmse) of the fits in the long and (attractive) medium range was of the order of  $10^{-3}$  to  $10^{-2}$   $\text{cm}^{-1}$ . The medium repulsive wall (energies of the order of  $10^2$  to  $10^3$   $\text{cm}^{-1}$ ) was reproduced with rmse of 1–10  $\text{cm}^{-1}$ .

Each radial coefficient was interpolated using cubic splines for  $4.8 < R \leq 16a_0$ . For  $R > 16a_0$  the energy (not the coefficients) was extrapolated using the  $c_5/R^5 + c_6/R^6$  functional form based on its values at  $R = 16a_0$  and  $R = 18a_0$ . For  $R < 4.8a_0$  the energy was extrapolated using the functional form  $A \exp(BR)$  based on its values at  $R = 4.8a_0$  and  $R = 5a_0$ . Thus, our PES covers the whole range of  $R$ . However, we suggest to use it in the range from 3.0 to  $100a_0$ .

The global minimum of the 4D PES corresponds to a T-shaped structure with  $\theta_{C_2} = 0^\circ$ ,  $\theta_{H_2} = 90^\circ$ ,  $\phi = 0^\circ$  and  $R = 6.92a_0$  with  $V = -109.22$   $\text{cm}^{-1}$ . In Fig. 2–4 we present the contour plots of the interaction energy surface for fixed equilibrium parameters. It is seen that there is a weak anisotropy of

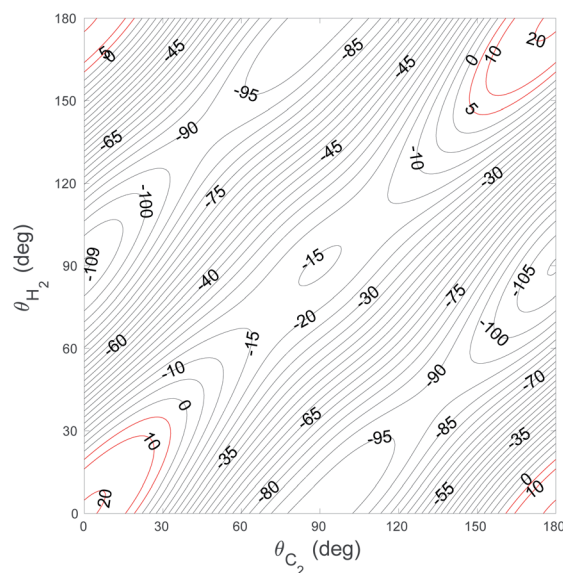


Fig. 2 Contour plot of the cut of the 4D PES for fixed  $\phi = 0^\circ$  and  $R = 6.92a_0$ . Energy is in  $\text{cm}^{-1}$ .



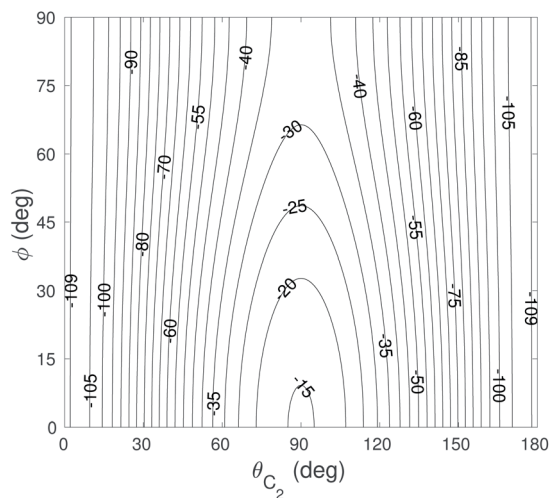


Fig. 3 Contour plot of the cut of the 4D PES for fixed  $\theta_{\text{H}_2} = 90^\circ$  and  $R = 6.92a_0$ . Energy is in  $\text{cm}^{-1}$ .

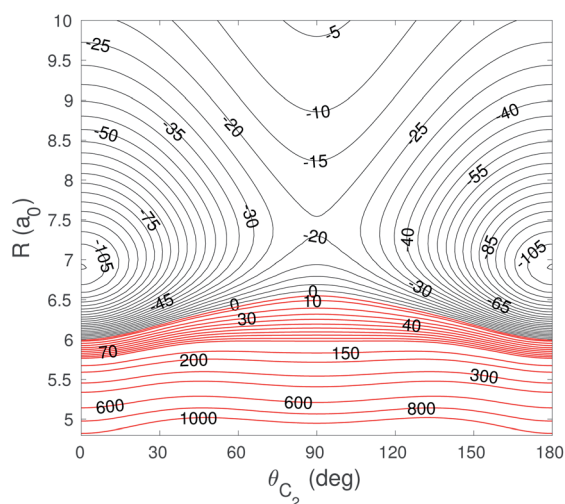


Fig. 4 Contour plot of the cut of the 4D PES for fixed  $\theta_{\text{H}_2} = 90^\circ$  and  $\phi = 0^\circ$ . Energy is in  $\text{cm}^{-1}$ . Red contour lines represent repulsive interaction energies.

the PES for the Jacoby angle  $\phi$  and a strong anisotropy for the angle  $\theta_{\text{C}_2}$ . The local minimum of the PES occurs for another T-shaped structure with  $\theta_{\text{C}_2} = 90^\circ$ ,  $\theta_{\text{H}_2} = 0^\circ$ ,  $\phi = 0^\circ$  and  $R = 6.70a_0$ . The interaction energy at this point is  $V = -101.63 \text{ cm}^{-1}$ .

### 3 Scattering study of excitation of $\text{C}_2$ by *para*- and *ortho*- $\text{H}_2$

In the present work, we consider collisions of  $^{12}\text{C}_2$  with *para*- and *ortho*- $\text{H}_2$ . Only even rotational levels  $j$  were considered here because, in homonuclear molecules with zero nuclear spins like  $\text{C}_2$ , rotational levels with odd values are not allowed. On the other hand, the  $\text{H}_2$  molecule can exist in two forms denoted *para*- and *ortho*- $\text{H}_2$ . In the former, the nuclear spins of the hydrogen atoms are opposed resulting in total  $I = 0$ , whereas, in

the latter, the nuclear spins are aligned and  $I = 1$ . As the nuclear wavefunction must be asymmetric, in *para*- $\text{H}_2$ , only even rotational states,  $j_{\text{H}_2} = 0, 2, \dots$  are allowed, and in *ortho*- $\text{H}_2$ , only odd values,  $j_{\text{H}_2} = 1, 3, \dots$  are allowed.

Full close-coupling calculations<sup>35</sup> were carried out for total energies up to  $1000 \text{ cm}^{-1}$ , using  $\text{C}_2$  and  $\text{H}_2$  as rigid rotors with rotational and centrifugal distortion constants  $B_{\text{C}_2} = 1.820 \text{ cm}^{-1}$  and  $D_{\text{C}_2} = 6.92 \times 10^{-6} \text{ cm}^{-1}$  (ref. 36), and  $B_{\text{H}_2} = 60.853 \text{ cm}^{-1}$  and  $D_{\text{H}_2} = 3.3 \times 10^{-6} \text{ cm}^{-1}$ ,<sup>29</sup> respectively. Transitions among  $\text{C}_2$  rotational levels up to  $j = 16$  were computed for collisions involving *para*- $\text{H}_2$  ( $j_{\text{H}_2} = 0$ ) and *ortho*- $\text{H}_2$  ( $j_{\text{H}_2} = 1$ ). For convergence of the cross sections, the rotational states  $j_{\text{H}_2} = 0, 2$  and  $j_{\text{H}_2} = 1, 3$  were included in the rotational basis sets for *para*- and *ortho*- $\text{H}_2$ , respectively. The inclusion of the  $j_{\text{H}_2} = 4$  channel in *para*- $\text{H}_2$  affects the cross sections by less than 1 percent. For the  $\text{C}_2$  molecule, the rotational basis set was extended up to  $j_{\text{max}} = 24$  at the largest total energy considered ( $1000 \text{ cm}^{-1}$ ). The coupled scattering equations were solved by the MOLSCAT code<sup>37</sup> using the propagator of Manolopoulos.<sup>38</sup> For the  $\text{C}_2$ - $\text{H}_2$  system the reduced mass is  $\mu = 1.8597 \text{ amu}$ . The integration parameters were chosen to ensure convergence of the cross sections in this range. The minimum and maximum integration distances chosen were  $R_{\text{min}} = 3 \text{ Bohr}$  and  $R_{\text{max}} = 50 \text{ Bohr}$ , respectively. The maximum value of the total angular momentum  $J$  used in the calculations was chosen according to a convergence criterion of  $0.01 \text{ \AA}^2$ . The energy grid was adjusted to describe all the details of the resonances of the inelastic cross sections. The thermal rate coefficients at temperature  $T$  are obtained by averaging the rotational inelastic cross sections,  $\sigma_{j \rightarrow j'}$ , over a Boltzmann distribution of collision energy  $E_c$ :

$$k_{j \rightarrow j'}(T) = \left( \frac{8}{\pi \mu \beta} \right)^{1/2} \beta^2 \int_0^\infty E_c \sigma_{j \rightarrow j'} e^{-\beta E_c} dE_c, \quad (4)$$

with  $\beta = \frac{1}{k_{\text{B}}T}$ , where  $k_{\text{B}}$  is the Boltzmann constant.

## 4 Results

First we compare the  $\sigma_{0 \rightarrow 2}$  cross sections for the excitation of  $\text{C}_2$  with *para*- $\text{H}_2$  ( $j_{\text{H}_2} = 0, 2$ ) calculated with the 4D PES and including  $j_{\text{H}_2} = 0, 2$  in the rotational basis set of  $\text{H}_2$  with the cross sections for the excitation of  $\text{C}_2$  with *para*- $\text{H}_2$  calculated previously by Najar *et al.*<sup>15</sup> with a 2D effective PES averaged over three specific  $\text{H}_2$  orientations using equal weights of  $1/3$  for the three geometries considered. As we can see in Fig. 5, the two cross sections are close, having the same form of resonances with a shift and small difference in amplitudes. The difference is appreciable at high energies, approaching 35 percent. To understand where the difference comes from, we present in Table 1 cross sections at different energies for  $\text{C}_2$ -*para*- $\text{H}_2$  calculations with a PES averaged over 3 orientations,  $\text{C}_2$ -*para*- $\text{H}_2$  4D PES calculations with just the basis  $j_{\text{H}_2} = 0$  and  $\text{C}_2$ -*para*- $\text{H}_2$  4D PES calculations with the rotational basis involving  $j_{\text{H}_2} = 0$  and  $j_{\text{H}_2} = 2$  levels ( $j_{\text{H}_2} = 0, 2$ ). We note that at low energies the major difference ( $\sim 15\%$ ) is due to lack of anisotropy in the PES averaged over 3 orientations. At high energies the inaccuracies of the PES have only a minor influence on the cross sections, and



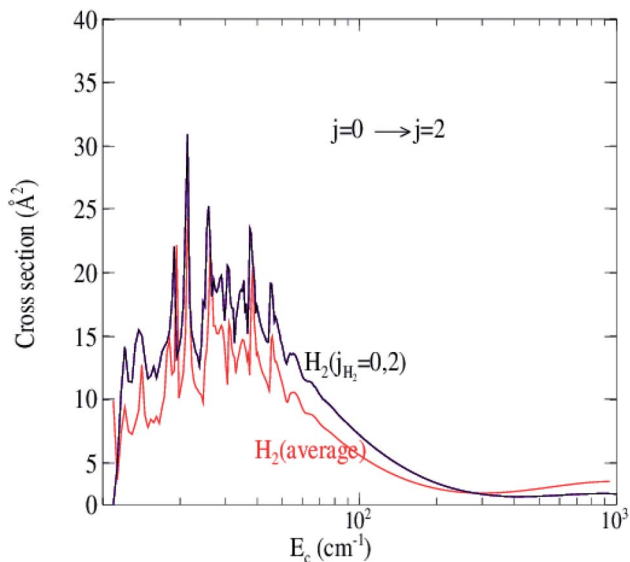


Fig. 5 Cross sections as a function of collision energy for  $0 \rightarrow 2$  rotational excitation of  $C_2$  with *para*- $H_2$  obtained by a 2D effective PES (red curve) and by the 4D PES including ( $j_{H_2} = 2$ ) in the rotational basis set of  $H_2$  (black curve).

**Table 1** Cross sections  $\sigma_{0 \rightarrow 2}$  ( $\text{\AA}^2$ ) for  $C_2$ -*para* $H_2$  obtained using a 2D effective PES, a 4D PES with basis  $j_{H_2} = 0$  and a 4D PES with the basis  $j_{H_2} = 0, 2$

Energy ( $\text{cm}^{-1}$ )	2D effective PES	4D PES, $j_{H_2} = 0$	4D PES, $j_{H_2} = 0, 2$
50	11.249	13.307	14.160
100	5.613	6.580	7.087
300	2.622	2.650	2.515
500	2.978	2.819	2.379
1000	3.540	3.298	2.567

the differences stem mainly from neglect of the  $j_{H_2} = 2$  rotational state. This result was also confirmed by Lique *et al.*<sup>39</sup> in the SiS- $H_2$  system.

Fig. 6 shows the collision energy dependence of the  $0 \rightarrow 2$  inelastic integral cross sections of  $C_2$  in collision with *para*- $H_2$  ( $j_{H_2} = 0$ ) (including  $j_{H_2} = 0, 2$ ) and *ortho*- $H_2$  ( $j_{H_2} = 1$ ) (including  $j_{H_2} = 1, 3$ ). The figure illustrates many resonances in the cross sections for collision energies below  $100 \text{ cm}^{-1}$ , which are a consequence of the temporary trapping of the  $H_2$  molecule in the attractive potential well, involving the formation of quasi-bound states of the van der Waals  $C_2$ - $H_2$  complex. We note a difference in amplitude and shape of the resonances between the two cross sections. The collisions with *ortho*- $H_2$  are more strongly inelastic. This is expected because, unlike with *ortho*- $H_2$ , the long-range interactions with *para*- $H_2$  ( $j_{H_2} = 0$ ) do not involve the permanent quadrupole moment of the  $H_2$  molecule; only induction and dispersion forces participate.<sup>40</sup> This behavior was also found in  $H_2$ - $H_2$  (ref. 41) and  $SO_2$ - $H_2$  (ref. 42) collisions.

Rate coefficients for rotational excitation of  $C_2$  by *para*- $H_2$  ( $j_{H_2} = 0$ ) and *ortho*- $H_2$  ( $j_{H_2} = 1$ ) have been computed from

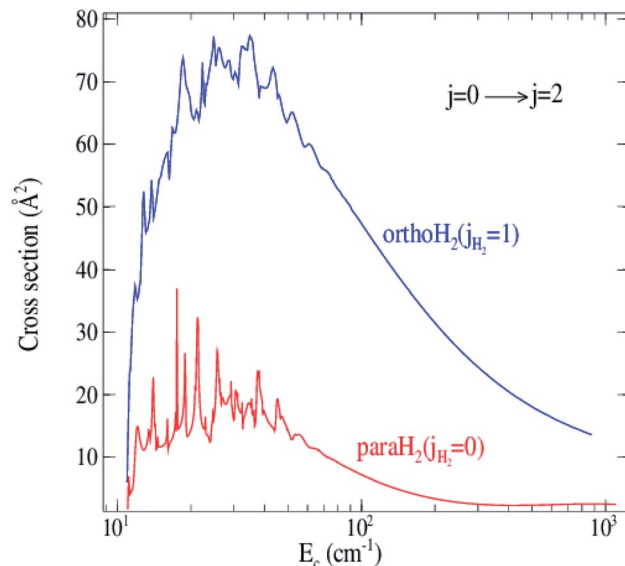


Fig. 6 Cross sections for rotational excitation of  $C_2$  by *para*- $H_2$  ( $j_{H_2} = 0, 2$ ) and *ortho*- $H_2$  ( $j_{H_2} = 1, 3$ ) for the  $0 \rightarrow 2$  transition as a function of collision energy.

cross sections generated from the grid of collision energies. The representative variation with temperature is illustrated in Fig. 7. The curves in Fig. 7 and the results in Table 2 show that the rates for *ortho*- $H_2$  ( $j_{H_2} = 1$ ) are significantly larger by a factor of 5 on average. This result, also found previously for some systems like CO- $H_2$ ,<sup>43</sup>  $H_2O$ - $H_2$  (ref. 44) and SiS- $H_2$ ,<sup>45</sup> confirms the expectation that *ortho*- $H_2$  ( $j_{H_2} = 1$ ) is a more efficient perturbant than *para*- $H_2$  ( $j_{H_2} = 0$ ).

In the paper of Lavendy *et al.*,<sup>46</sup> for construction of the PES the authors considered only 5 relative orientations of  $C_2$  and  $H_2$

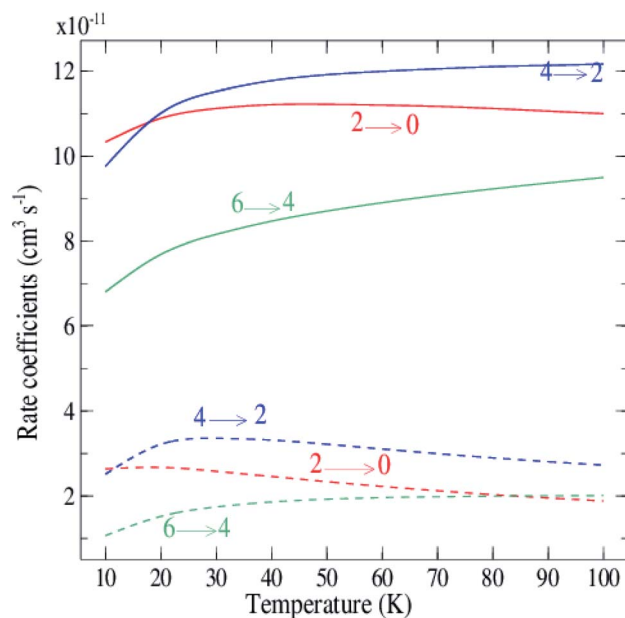


Fig. 7 Temperature dependence of rotational de-excitation rate coefficients of  $C_2$  excited by  $H_2$  for  $\Delta j = 2$  transitions. Dashed lines: *para*- $H_2$  ( $j_{H_2} = 0$ ); solid lines: *ortho*- $H_2$  ( $j_{H_2} = 1$ ).



Table 2 Rate coefficients ( $\text{cm}^3 \text{s}^{-1}$ ) for rotational de-excitation in  $\text{C}_2$  induced by *para*- $\text{H}_2(j_{\text{H}_2} = 0)$  and *ortho*- $\text{H}_2(j_{\text{H}_2} = 1)$ . The upper entry for each transition is for  $T = 50 \text{ K}$ , and the lower entry is for  $T = 100 \text{ K}$

Transition	<i>Para</i> - $\text{H}_2$	<i>Ortho</i> - $\text{H}_2$
2 $\rightarrow$ 0	0.234(-10)	1.122(-10)
	0.188(-10)	1.100(-10)
4 $\rightarrow$ 2	0.322(-10)	1.191(-10)
	0.273(-10)	1.216(-10)
6 $\rightarrow$ 4	0.192(-10)	0.870(-10)
	0.201(-10)	0.949(-10)
8 $\rightarrow$ 6	0.110(-10)	0.610(-10)
	0.141(-10)	0.704(-10)

for each  $R$  value. Later, Phillips<sup>47</sup> used this PES to compute the rate coefficients using the close-coupling approach. In our previous work,<sup>15</sup> we constructed a 2D PES based on nearly 900 *ab initio* points. For each  $(R, \theta)$  we considered 3 angular orientations of the  $\text{H}_2$  molecule, the energies of which were then averaged. In that work,<sup>15</sup> we compared our collision rates for  $\text{C}_2$ -*para*- $\text{H}_2(j_{\text{H}_2} = 0)$  to those of Phillips<sup>47</sup> for the temperature range  $T < 100 \text{ K}$ . We found that the rates calculated by Phillips are greater by a factor of 5. This difference is due to the fact that the interaction potential used by Phillips overestimates the anisotropy of the potential. In addition, Phillips found that *para*- $\text{H}_2(j_{\text{H}_2} = 0)$  and *ortho*- $\text{H}_2(j_{\text{H}_2} = 1)$  are nearly identical colliders, which doesn't agree with our findings that there is a notable difference. This can also be explained by the fact that the potential of Phillips based on the *ab initio* energies of Lavendy *et al.* was constructed by neglecting the terms responsible for the quadrupole-quadrupole interactions.

## 5 Conclusions

We have applied full close-coupling calculations to investigate rotational excitation in the  $\text{C}_2$  molecule for collisions involving *para*- $\text{H}_2(j_{\text{H}_2} = 0)$  and *ortho*- $\text{H}_2(j_{\text{H}_2} = 1)$ . The calculations are performed using a highly accurate 4D potential energy surface. The inclusion of  $j_{\text{H}_2} = 2$  in the rotational basis set of  $\text{H}_2$  makes an appreciable contribution at high energies to the *para*- $\text{H}_2$  cross sections, approaching 35 percent. Rate coefficients for transitions among the levels of  $\text{C}_2$  up to  $j = 16$  were determined for the range of temperatures  $5 \leq T \leq 100 \text{ K}$ . Rate coefficients with *ortho*- $\text{H}_2$  are larger by a factor of 5 on average than rate coefficients with *para*- $\text{H}_2$ .

## Conflicts of interest

There are no conflicts to declare.

## Acknowledgements

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## Notes and references

- S. P. Souza and B. L. Lutz, *Astrophys. J.*, 1977, **216**, L49.
- F. H. Chaffee, B. L. Lutz, J. H. Black, P. A. Vanden Bout and R. L. Snell, *Astrophys. J.*, 1980, **236**, 474.
- L. M. Hobbs and B. Cambell, *Astrophys. J.*, 1982, **254**, 108.
- A. C. Danks and D. L. Lambert, *Astron. Astrophys.*, 1983, **124**, 188.
- E. F. van Dishoeck and J. H. Black, *Astrophys. J.*, 1989, **340**, 273.
- S. R. Federman, C. J. Strom, D. L. Lambert, *et al.*, *Astrophys. J.*, 1994, **424**, 772.
- P. Sonnentrucker, D. E. Welty, J. A. Thorburn and D. G. York, *Astrophys. J., Suppl. Ser.*, 2007, **168**, 58.
- M. Kaźmierczak, M. R. Schmidt, A. Bondar and J. Krelowski, *Mon. Not. R. Astron. Soc.*, 2010, **402**, 2548.
- T. P. Snow, *Astrophys. J.*, 1978, **220**, L93.
- D. J. Lien, *Astrophys. J.*, 1984, **287**, L95.
- D. L. Lambert, Y. Sheffer and S. R. Federman, *Astrophys. J.*, 1995, **438**, 740.
- E. F. van Dishoeck and J. H. Black, *Astrophys. J.*, 1982, **258**, 533.
- H. Han, A. Li and H. Guo, *J. Chem. Phys.*, 2014, **141**, 244312.
- N. Masakazu, M. Akira and M. Akira, *J. Phys. Chem. A*, 2009, **113**, 8963.
- F. Najjar, D. Ben Abdallah, N. Jaidane, Z. Ben Lakhdar, G. Chambaud and M. Hochlaf, *J. Chem. Phys.*, 2009, **130**, 204305.
- R. C. Hupe, Y. Sheffer and S. R. Federman, *Astrophys. J.*, 2012, **761**(1), 38.
- F. Najjar, D. Ben Abdallah, N. Jaidane and Z. Ben Lakhdar, *Chem. Phys. Lett.*, 2008, **460**, 31.
- P. J. Knowles and H.-J. Werner, *Chem. Phys. Lett.*, 1985, **115**, 259.
- H.-J. Werner and P. J. Knowles, *J. Chem. Phys.*, 1985, **82**, 5053.
- P. J. Knowles and H.-J. Werner, *Chem. Phys. Lett.*, 1988, **145**, 514.
- H.-J. Werner and P. J. Knowles, *J. Chem. Phys.*, 1988, **89**, 5803.
- E. R. Davidson and D. W. Silver, *Chem. Phys. Lett.*, 1977, **52**, 403.
- S. R. Langhoff and E. R. Davidson, *Int. J. Quantum Chem.*, 1974, **8**, 61.
- T. H. Dunning Jr, *J. Chem. Phys.*, 1989, **90**, 1007.
- R. A. Kendall, T. H. Dunning Jr and R. J. Harrison, *J. Chem. Phys.*, 1992, **96**, 6796.
- D. E. Woon and T. H. Dunning Jr, *J. Chem. Phys.*, 1994, **100**, 2975.
- H.-J. Werner, P. J. Knowles, *et al.*, *MOLPRO, version 2010.1, a package of ab initio program*, 2010, <http://www.molpro.net>.
- M. P. Hodges, R. J. Wheatley, G. K. Schenter and A. H. Harvey, *J. Chem. Phys.*, 2004, **120**, 710.
- K. P. Huber and G. Herzberg, *Molecular Spectra and Molecular Structure. IV. Constants of Diatomic Molecules*, Van Nostrand Reinhold, New York, 1979.
- P. Valiron, M. Wernli, A. Faure, L. Wiesenfeld, C. Rist, S. Kedzuch and J. Noga, *J. Chem. Phys.*, 2008, **129**, 134306.



## Paper

- 31 Y. Scribano, A. Faure and L. Wiesenfeld, *J. Chem. Phys.*, 2010, **133**, 231105.
- 32 A. Faure, P. Jankowski, T. Stoecklin and K. Szalewicz, *Sci. Rep.*, 2016, **6**, 28449.
- 33 S. F. Boys and F. Bernardi, *Mol. Phys.*, 1970, **19**, 553.
- 34 S. Green, *J. Chem. Phys.*, 1975, **62**, 2271.
- 35 A. M. Arthurs and A. Dalgarno, *Proc. R. Soc. A.*, 1960, **256**, 540.
- 36 M. Martin, *J. Photochem. Photobiol., A*, 1992, **66**, 263.
- 37 J. M. Hutson and S. Green, *MOLSCAT Computer Code, version 14*, Distributed by Collaborative Computational Project No. 6 of the Engineering and Physical Sciences Research Council, UK, 1994.
- 38 D. E. Manolopoulos, *J. Chem. Phys.*, 1986, **85**, 6425.
- 39 F. Lique, R. Tobola, J. Klos, *et al.*, *Astron. Astrophys.*, 2008, **478**, 567.
- 40 D. R. Flower, J. M. Launay, E. Kochanski and J. Prissette, *Chem. Phys.*, 1979, **37**, 355.
- 41 G. Danby, D. R. Flower and T. S. Monteiro, *Mon. Not. R. Astron. Soc.*, 1987, **226**, 739.
- 42 J. Cernicharo, A. Spielfiedel, C. Balança, *et al.*, *Astron. Astrophys.*, 2011, **531**, A103.
- 43 A. R. Offer, M. C. van Hemert and E. F. van Dishoeck, *J. Chem. Phys.*, 1994, **100**, 362.
- 44 M. Wernli, P. Valiron, A. Faure, L. Wiesenfeld, P. Jankowski and K. Szalewicz, *Astron. Astrophys.*, 2006, **446**, 367.
- 45 J. Klos and F. Lique, *Mon. Not. R. Astron. Soc.*, 2008, **390**, 239.
- 46 H. Lavendy, J. M. Robbe, G. Chabaud, B. Levy and E. Roueff, *Astron. Astrophys.*, 1991, **251**, 365.
- 47 T. R. Phillips, *Mon. Not. R. Astron. Soc.*, 1994, **271**, 827.

