



Cite this: *RSC Adv.*, 2019, 9, 37201

Received 27th May 2019  
 Accepted 30th October 2019

DOI: 10.1039/c9ra03999f

[rsc.li/rsc-advances](http://rsc.li/rsc-advances)

# High responsivity ultraviolet detector based on novel SnO<sub>2</sub> nanoarrays†

Xinhua Pan, \* Tao Zhang, Qiaoqi Lu, Weihao Wang and Zhizhen Ye

In this work, tin oxide (SnO<sub>2</sub>) nanoarrays (NAs) with a novel morphology were fabricated by a hydrothermal method. Each SnO<sub>2</sub> nanorod (NR) consisted of a bunch of nearly ten primary NRs with a width of 4–7 nm. On this basis, gold (Au) nanoparticles (NPs) were used to induce the growth of smaller and more orderly SnO<sub>2</sub> NAs. The prepared ultraviolet (UV) detectors based on SnO<sub>2</sub> NAs have high responsivity and high external quantum efficiency (EQE) under 20 μW cm<sup>-2</sup> UV light with only 1 V external bias voltage.

## 1. Introduction

As a new civil and military technology, ultraviolet (UV) detectors has been widely used in many fields, such as UV communication, forest fire prevention, pollution monitoring and biomedical diagnosis.<sup>1–4</sup> Up to now, the research of UV detectors has gradually developed from vacuum diodes, photomultipliers and silicon photodiodes to the third generation of wide band gap semiconductor based UV detectors. With unique physical and chemical properties, nanostructured metal oxide semiconductors have been proposed for UV detectors with superior responsivity, high quantum efficiency, low-power consumption, and weak signal detection.<sup>3–6</sup>

Tin oxide (SnO<sub>2</sub>), a direct band gap semiconductor with a wide band gap of 3.6–3.8 eV, shows strong exciton binding potential of 130 meV and transparency in visible spectral region.<sup>7</sup> It is considered to be a candidate for visible blind photodetectors.<sup>8</sup> Previous studies on nanostructured SnO<sub>2</sub> mainly focused on morphology and structure. SnO<sub>2</sub> nanospheres, nanorods (NRs), nanobelts, nanotubes and other morphologies have been controllably prepared and widely reported.<sup>9</sup> However, progress in high-performance photodetectors based on nanostructured SnO<sub>2</sub> is limited,<sup>10</sup> compared with other wide band gap semiconductors such as zinc oxide (ZnO) and titanium oxide (TiO<sub>2</sub>). Until recent years, some reports on nanostructured SnO<sub>2</sub> based UV detectors gradually appeared. However, it is studied that the devices constructed by pure nanostructured SnO<sub>2</sub> exhibit poor performance in UV detection, due to the surface defects and trapping effect from oxygen molecules adsorbed on the surface of SnO<sub>2</sub>.<sup>11,12</sup> J. M. Wu *et al.*<sup>13</sup> synthesized SnO<sub>2</sub> nanowires by gas–liquid–solid method at 900 °C, and fabricated UV detectors on Al substrates. The

sensitivity of the devices under 365 nm UV irradiation was only 0.69. Therefore, many efforts have been devoted to improve the performance of SnO<sub>2</sub> based photodetectors. Fabricating complex high-precision devices and utilizing heterojunctions are usually effective methods to improve device performance.<sup>13–19</sup> J. Yan *et al.*<sup>14</sup> constructed SnO<sub>2</sub> nanowires UV detectors with a series of complex optical elements such as spatial light modulators and microscopes. The acquired detectors showed high sensitivity, responsivity and external quantum efficiency (EQE). C. L. Hsu *et al.*<sup>15</sup> prepared Ga<sub>2</sub>O<sub>3</sub>/SnO<sub>2</sub>:Ga core–shell nanowires on sapphire substrates by gas–liquid–solid method at 1000 °C. Under 310 nm UV irradiation, the  $I_{on}/I_{off}$  of Ga-doped devices is 1066.7, but the maximum quantum efficiency is only 0.362%. In addition, S. Abbas *et al.*<sup>16</sup> reported a transparent photodetector governed by a SnO<sub>2</sub>/SnS<sub>2</sub> heterojunction, but they had to introduce an In<sub>2</sub>O<sub>3</sub> nanolayer in the SnO<sub>2</sub>/SnS<sub>2</sub> heterojunction to reduce the dark current. Interestingly, Gan *et al.*<sup>12</sup> reported a novel UV detector based on the kinked SnO<sub>2</sub> nanowires, which shows better performance than the conventional straight SnO<sub>2</sub> nanowire-based photodetectors both in photocurrent and photoresponse speed. The improvement is caused by the enhancement effect of the kinked structure in SnO<sub>2</sub> nanowires to the local electric field, which benefits for the charge separation efficiency and the transmittal time of carriers. However, the exact mechanism of the formation of the kinked SnO<sub>2</sub> nanowires is not clear yet.

Complex synthesis method and device structure may hinder the widespread application of nanostructured SnO<sub>2</sub> based UV detectors. Accordingly, how to improve the comprehensive performance of SnO<sub>2</sub> based detectors by a simple way is the main challenge for SnO<sub>2</sub> nanostructured based UV detectors. In this work, we have successfully prepared SnO<sub>2</sub> nanoarrays (NAs) with peculiar morphology and neat arrangement by a controllable and simple hydrothermal method, and fabricated devices based on SnO<sub>2</sub> NAs with high responsivity. Furthermore, we introduced gold (Au) nanoparticles in order to obtain more compact SnO<sub>2</sub>–Au NAs, and the devices exhibit better performance than pure

State Key Laboratory of Silicon Materials, Cyrus Tang Center for Sensor Materials and Applications, School of Materials Science and Engineering, Zhejiang University, Hangzhou 310027, People's Republic of China. E-mail: [panxinhua@zju.edu.cn](mailto:panxinhua@zju.edu.cn)

† Electronic supplementary information (ESI) available: Surface scan energy spectrum of SnO<sub>2</sub>–Au NAs. See DOI: 10.1039/c9ra03999f



SnO<sub>2</sub> NAs with the help of localized surface plasmon resonances (LSPR). In particular, SnO<sub>2</sub>-Au NAs UV detectors work at low additional bias voltage and show high responsivity, external quantum efficiency and detectivity under weak power UV light (20 μW cm<sup>-2</sup>). More importantly, this work reveals that nano-structured SnO<sub>2</sub> based UV detectors can be developed towards weak signal detection and low-power consumption.

## 2. Experimental section

### 2.1 Synthesis of SnO<sub>2</sub> NAs

In this work, the SnO<sub>2</sub> NAs were prepared by hydrothermal method.<sup>18</sup> 0.1234 g NaBr was dissolved in 3.6 ml deionized water, and 0.1043 g SnCl<sub>4</sub>·5H<sub>2</sub>O was dissolved in 24 ml acetic acid. Under continuous stirring, the two drops were added into 4 ml ethanol to obtain a yellowish liquid. Fluorine-doped tin oxide (FTO) substrates were ultrasonically cleaned in beakers containing acetone, ethanol and deionized water, respectively, and then blown dry with nitrogen. The conductive layer was upward and placed in a 50 ml Teflon-sealed autoclave. The yellowish solution was transferred into the Teflon-sealed autoclave and heated to 200 °C, maintained for 12 hours. When the hydrothermal reaction was completed and the autoclave was cooled to room temperature naturally, samples were taken out with tweezers, washed alternately with deionized water and ethanol, then cleaned by ultrasound in ethanol, and finally dried at 60 °C for 12 hours.

Electron beam evaporation were used to grow 50 nm thick Au NPs on the cleaned FTO conductive layer as the substrate and repeated the above steps to obtain Au-induced SnO<sub>2</sub> NAs. For convenience, Au-induced SnO<sub>2</sub> NAs were named as SnO<sub>2</sub>-Au NAs. UV detectors based on SnO<sub>2</sub> NAs were abbreviated by using electron beam evaporation to coat aluminum (Al) on NAs as electrodes.

### 2.2 Characterization

The surface morphology and crystal structure of the samples were characterized by field-emission scanning electron microscopy (FE-SEM, Hitachi S-4800) and transmission electron microscopy (TEM, PhilipsTecnai F20). The crystallinity of SnO<sub>2</sub> NAs was analyzed by X-ray diffraction (XRD) with a Cu Kα radiation source (λ = 1.54056 Å). The light reflection of the samples was characterized by ultraviolet-visible diffuse reflectance spectroscopy (UV-vis DRS, Shimadzu UV-3600). The electronic structure of SnO<sub>2</sub> NAs was analyzed by ultraviolet photoelectron spectroscopy (UPS, AXIS Supra), and the energy of monochrome UV light source is 21.22 eV. The electrical properties of the devices were studied by current-voltage (*I*-*V*, Agilent E5270B) measurements. The temporal response test of UV detector used 330 nm xenon lamp.

## 3. Results and discussion

### 3.1 Characterization of SnO<sub>2</sub> NAs

The morphological characterization of the prepared samples was analyzed by SEM. Fig. 1a and b are SEM diagrams of pure SnO<sub>2</sub> NAs and SnO<sub>2</sub>-Au NAs, respectively. Comparing the two

diagrams, we can see that each SnO<sub>2</sub> NR of pure SnO<sub>2</sub> NAs has a wider width than SnO<sub>2</sub>-Au NAs, and some of the NRs are dumped to the periphery. This is because the atomic radius of Au is smaller than that of Sn. When Au atoms act as nucleation points of SnO<sub>2</sub> nucleation growth, the distance between each nucleation point is closer, which makes the distance between each SnO<sub>2</sub> NR closer and the size of the NR smaller. Therefore, it is conducive to make the NAs grow more orderly. The insert in Fig. 1b is a high-magnification image. It can be seen that the top of each SnO<sub>2</sub> NR in SnO<sub>2</sub>-Au NAs is rough and exhibits a quadrilateral structure with a side length of about 30 nm. Fig. 1c is a low-magnification SEM image of SnO<sub>2</sub>-Au NAs, which shows that SnO<sub>2</sub>-Au NAs are uniform in a large area. The insert is a cross-section image. It shows that the height of SnO<sub>2</sub>-Au NAs is about 560 nm. Fig. 1d is the XRD image of SnO<sub>2</sub>-Au NAs. It can be seen that SnO<sub>2</sub> NAs are typical tetragonal rutile structure, which matches the JCPDS card of no. 41-1445. The three strong peaks correspond to the (101), (110), (211) crystal planes of SnO<sub>2</sub>.

Because of the low content of Au, there is no obvious characteristic peak of Au in XRD image. So the local sweep energy spectrum of SnO<sub>2</sub>-Au NAs is measured. Fig. S1a† is the energy spectrum diagram of SnO<sub>2</sub>-Au NAs, Fig. S1b-d† are the distribution diagrams of Sn, O, and Au, respectively. It can be seen from the atomic percentage of each element that the oxygen content in the sample exceeds the stoichiometric ratio. This may be due to the small size, large specific surface area and a large number of suspended bonds on the surface of SnO<sub>2</sub> NRs, which make it easy to absorb oxygen.<sup>13,14,20</sup> It is difficult to see Au NPs clearly in SEM because of their small size. However, it can be determined by element distribution diagrams that Au NPs exist in NAs and distribute uniformly.

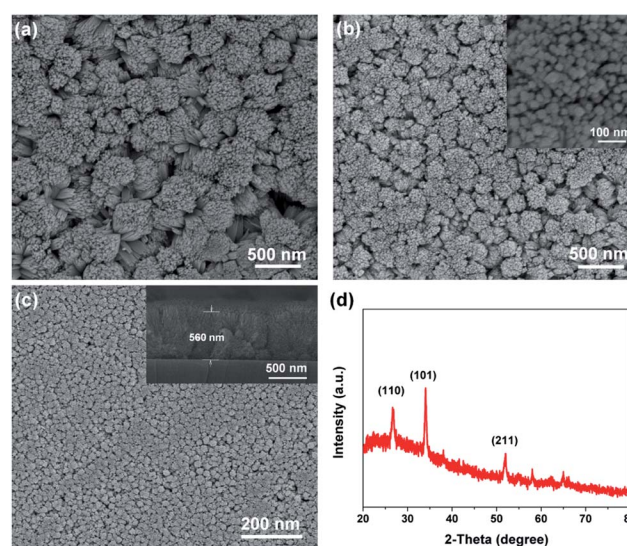


Fig. 1 (a) Surface SEM image of SnO<sub>2</sub> NAs, (b) surface SEM image of SnO<sub>2</sub>-Au NAs, the insert is high-magnification image, (c) low-magnification SEM image of SnO<sub>2</sub>-Au NAs, the insert is cross-section image, (d) XRD pattern of SnO<sub>2</sub>-Au NAs.



In order to further distinguish the rough structure at the top of each SnO<sub>2</sub> NR and determine the location of Au NPs, TEM is needed. Fig. 2a is a TEM image of a single SnO<sub>2</sub> NR. The width of the NR is about 38 nm, which is consistent with the result of SEM. It is surprising to find that a single SnO<sub>2</sub> NR is composed of a bunch of primary NRs with a size of 4–7 nm and the parallel stripes in the middle of the TEM image are formed by the superposition of multi-layer parallel primary NRs, which explain the reason why the top of the NRs is rough. Fig. 2b is a local high-resolution transmission electron microscopy (HRTEM) image of Fig. 2a. It can be seen that the primary NRs in a single SnO<sub>2</sub> NR have the same lattice fringes, indicating that the primary NRs grow in the same direction. The insert is a selected area electron diffraction (SAED) image, which shows that a single SnO<sub>2</sub> NR consisting of a bunch of primary NRs is single crystal phase. Fig. 2c is a TEM image of several SnO<sub>2</sub>-Au NRs, which shows that the lattice fringes of particles on several primary NRs are inconsistent with those of primary NRs, and the local magnification diagram is shown in Fig. 2d. The lattice fringe spacing of SnO<sub>2</sub>-Au primary NRs is 0.265 nm, which corresponds to the (101) crystal plane of rutile SnO<sub>2</sub>, indicating that the primary NRs grow in the direction of [101]. However, the lattice fringe spacing of the top particles is 0.144 nm, which corresponds to the (220) crystal plane of Au with face-centered cubic structure, showing that the Au NPs on the substrate are gradually rising to the top with the growth of SnO<sub>2</sub> NAs. The growth mechanism is similar to that of Au-induced NAs of ZnO and Ga<sub>2</sub>O<sub>3</sub>.<sup>21,22</sup>

The hypothetical mechanism of SnO<sub>2</sub> NAs growth process has been proposed in the literature.<sup>18</sup> In a ternary acidic system containing an appropriate proportion of water, ethanol and acetic acid, controlling the hydrolysis rate can make the supersaturation higher than the critical concentration of

heterogeneous nucleation on the substrate and lower than the critical concentration of homogeneous nucleation in the solution. As a result, SnO<sub>2</sub> can nucleate on the substrate. The organic ligands in solution will adsorb on the surface of the initial core, reduce the surface energy and form the nuclei of primary NRs. With the hydrolysis process, the nuclei increase in width and height, thus forming SnO<sub>2</sub> NRs consisting of a bunch of primary NRs. The NRs continue to grow and form NAs filled with organic ligands in the middle space. The introduction of Au provides more nucleation points for SnO<sub>2</sub> nucleation on the substrate, which makes the NAs more compact and orderly. The growth mechanism of SnO<sub>2</sub> NAs and SnO<sub>2</sub>-Au NAs is shown in Fig. 3.

In order to ensure good ohmic contact between the electrode and the sample in the ultraviolet detector, it is necessary to select the electrode material with appropriate work function. UPS is usually used to analyze the work function of materials. Fig. 4a and b are UPS image of SnO<sub>2</sub> NAs and local enlargement image of cut-off and valence band edges, respectively. The work function of SnO<sub>2</sub> NAs can be obtained from the cut-off energy according to the formula  $W = h\nu - E_{\text{cut-off}}$ , where  $W$  is work function of semiconductor,  $h\nu$  is photon energy and  $E_{\text{cut-off}}$  is cut-off energy, which can be obtained from the intersection point of the fitting line of the cut-off edge and the baseline.<sup>8,23,24</sup> As shown in Fig. 3b,  $h\nu$  used in the instrument is 21.22 eV, the cut-off energy of SnO<sub>2</sub> NAs is 16.68 eV, and the work function is 4.54 eV. Therefore, Al is chosen as the electrode material for ultraviolet detectors. In addition, the intersection point of the fitting line at the edge of the valence band and the baseline is the difference between the valence band and the Fermi level.

As shown in Fig. 4c, UV-vis DRS is used to further characterize the band structure of SnO<sub>2</sub> NAs. The reflectivity of SnO<sub>2</sub> NAs is very strong in the visible band. When the wavelength is less than 350 nm, the reflectivity decreases rapidly, indicating that the samples begin to absorb light. The Kubelka-Munk formula can be used to transform the reflectivity curve, which is shown by the red line in Fig. 4c. The intersection point between the fitting line with the largest slope and the line with zero ordinate is the absorption wavelength of the sample. The absorption wavelength of SnO<sub>2</sub> NAs is 337 nm, and the bandgap of SnO<sub>2</sub> NAs is estimated to be about 3.68 eV. Therefore, we use UV light with a wavelength of 330 nm as excitation light of UV detectors.

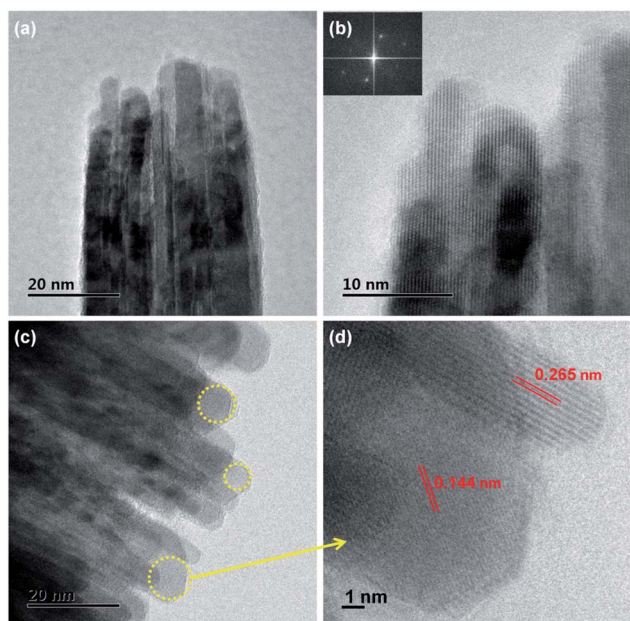


Fig. 2 (a) TEM image and (b) HRTEM image of SnO<sub>2</sub> NAs, the insert is SAED pattern, (c) TEM image and (d) HRTEM image of SnO<sub>2</sub>-Au NAs.

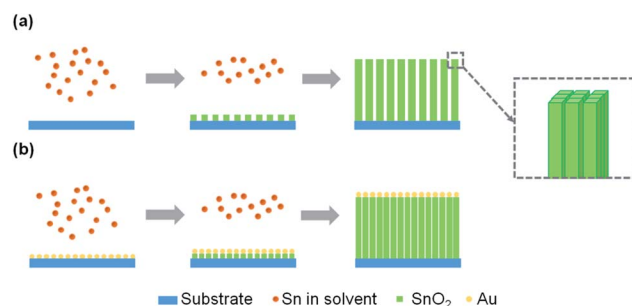


Fig. 3 Growth mechanism diagram of (a) SnO<sub>2</sub> NAs and (b) SnO<sub>2</sub>-Au NAs.





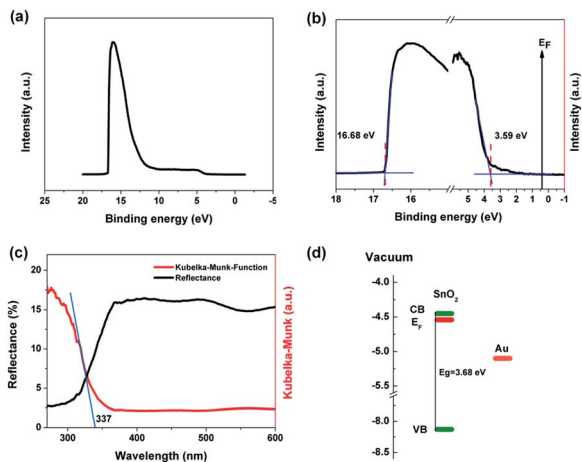


Fig. 4 (a) UPS spectrum, (b) local UPS spectra of the cut-off region and valence band edge region, (c) UV-vis DRS spectra and (d) band structure of SnO<sub>2</sub> NAs.

According to the information obtained by UPS and UV-vis DRS, the band structure of SnO<sub>2</sub> NAs can be drawn, as shown in Fig. 4d. The valence band (VB), Fermi level ( $E_F$ ) and conduction band (CB) of SnO<sub>2</sub> NAs are  $-8.13$  eV,  $-4.54$  eV and  $-4.45$  eV, respectively, relative to the vacuum level. It can be seen that the  $E_F$  of n-type SnO<sub>2</sub> NAs is very high and close to the CB, which indicates that there are many high energy electrons with high concentration.

### 3.2 Photoelectric properties of SnO<sub>2</sub> NAs

The schematic diagram of UV detectors based on SnO<sub>2</sub> and SnO<sub>2</sub>-Au NAs is illustrated in Fig. 5a. In our work, UV detectors are constructed simply and quickly by plating Al electrodes on the top of the arrays through electron beam evaporation. Fig. 5b is the current-voltage ( $I$ - $V$ ) characteristics of UV detectors under dark and light irradiation. UV light with the power of  $20 \mu\text{W cm}^{-2}$  and wavelength of  $330$  nm is used as excitation light. UV irradiation can stimulate valence band electron transitions and generate electron-hole pairs, which makes the photocurrent ( $I_{\text{photo}}$ ) increase sharply. The  $I_{\text{photo}}$  of SnO<sub>2</sub> and SnO<sub>2</sub>-Au NAs can reach  $5.49 \mu\text{A}$  and  $20.03 \mu\text{A}$ , respectively, even if the bias voltage is only  $1$  V. With the increase of bias voltage, the  $I_{\text{photo}}$  of both devices increases linearly, which indicates that good ohmic contact is formed between the Al electrode and SnO<sub>2</sub> NAs. Compared with  $I_{\text{photo}}$ , the dark current ( $I_{\text{dark}}$ ) of the two devices has little difference and almost no gain with the increase of bias voltage.

The sensitivity ( $S$ ) [ $S = (I_{\text{photo}} - I_{\text{dark}})/I_{\text{dark}}$ ] and responsivity ( $R$ ) [ $R = I_{\text{photo}}/PA$ ,  $P$  represents the incident optical power density,  $A$  represents the effective area.] are calculated at bias voltage of  $1$  V, as shown in Table 1. SnO<sub>2</sub> NAs UV detectors have high responsivity, while SnO<sub>2</sub>-Au NAs UV detectors have nearly four times higher responsivity. External quantum efficiency (EQE) is a characterization of the photon utilization rate of detectors.<sup>11</sup> It refers to the number of electron-hole pairs produced by absorbing a photon energy per unit time. The

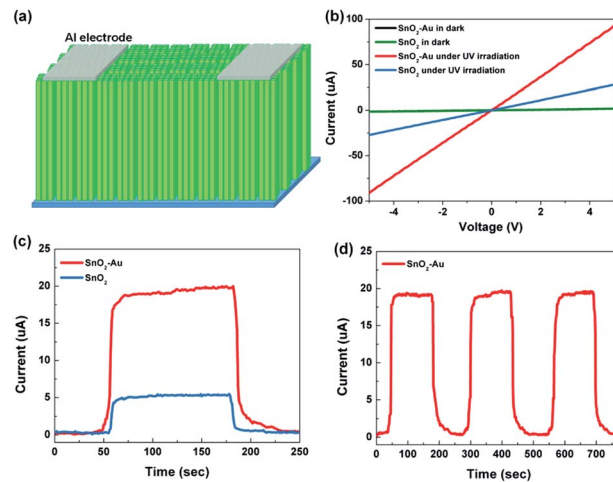


Fig. 5 (a) The schematic diagram of the fabricated UV detectors, (b)  $I$ - $V$  characteristics curves of SnO<sub>2</sub> and SnO<sub>2</sub>-Au based devices under dark and UV irradiation, photocurrent response under on/off radiation UV light in  $250$  s of SnO<sub>2</sub> and SnO<sub>2</sub>-Au NAs based device (c) and in  $780$  s of SnO<sub>2</sub>-Au NAs based device (d) the power density of the UV light used is  $20 \mu\text{W cm}^{-2}$  ( $\lambda = 330$  nm), and the bias voltage is  $1$  V.

formula for calculating EQE can be expressed as  $\text{EQE} = hc(I_{\text{photo}} - I_{\text{dark}})/e\lambda PA$ , where  $h$  is Planck constant,  $c$  is the speed of light,  $e$  is the unit charge, and  $\lambda$  is the excitation wavelength. Specific detectivity ( $D^*$ ) is an important figure in evaluating the performance of photodetectors,<sup>25,26</sup> related to the signal that a detector can distinguish from the background noise. When the shot noise from dark current is considered to be the major contributor, the  $D^*$  can be expressed as  $D^* = RA^{1/2}/(2eI_{\text{dark}})^{1/2}$ . The EQE and  $D^*$  of SnO<sub>2</sub>-Au NAs UV detectors are as high as  $376.321\%$  and  $1.25 \times 10^{12}$  Jones, respectively, which is about four times of pure SnO<sub>2</sub> NAs UV detectors.

In order to study whether the devices can respond to low-power UV light efficiently and stably under low bias voltage, and verify that the devices can be used as high-responsivity and energy-saving UV detectors, the following performance tests set the bias voltage to  $1$  V and the UV power to  $20 \mu\text{W cm}^{-2}$ . Fig. 5c exhibits the time-resolved UV  $I_{\text{photo}}$  on/off measurements of SnO<sub>2</sub> and SnO<sub>2</sub>-Au NAs UV detectors in  $250$  s. When the UV light is irradiated, the  $I_{\text{photo}}$  rapidly rises to saturation state, and maintains a stable state. When the UV light is turned off, the  $I_{\text{photo}}$  falls back rapidly. The response times of the devices listed in Table 1 are obtained by non-linear fitting of the curves,  $\tau_g$  is the rise time constant and  $\tau_d$  is the decay time constant. Because the primary NRs in SnO<sub>2</sub> NAs are small in size and have many grain boundaries, although the primary NRs are in close contact, the scattering of grain boundaries still hinders the

Table 1 The performance parameters of devices

Sample	$S$	$R$ (A/W)	EQE (%)	$D^*$ (Jones)	$\tau_g$ (s)	$\tau_d$ (s)
SnO <sub>2</sub>	30	275	103.145	$3.62 \times 10^{11}$	4.3	5.4
SnO <sub>2</sub> -Au	110	1002	376.321	$1.25 \times 10^{12}$	3.7	5.3



carrier migration, which makes the carrier collection rate not fast enough. Three repeated response cycles of SnO<sub>2</sub> NAs UV detector are displayed in Fig. 5d, in which the  $I_{\text{photo}}$  can keep consistent and repeatable with no significant degeneration during the whole detection process, indicating that our devices have good stability and reliability.

In order to further evaluate the performance of our SnO<sub>2</sub>-Au NAs UV detectors, the performance parameters of different SnO<sub>2</sub> based UV detectors are compared, as shown in Table 2. It can be seen that in the similar nanostructured UV detectors, our detectors have great advantages in responsivity and EQE.

The reason why SnO<sub>2</sub> NAs UV detectors exhibit higher EQE performance and the role of Au introduction will be explained in the following aspects. Compared with ordinary NRs, the SnO<sub>2</sub> NAs composed of several primary NRs have a larger specific surface area, which can increase the scattering and absorption of UV light, thereby increasing the absorption efficiency of UV light and improving the photocurrent. At present, it is generally believed that oxygen molecules will adsorb on the surface of semiconductor oxides and capture free electrons therein,<sup>26–30</sup> resulting in the formation of low conductivity depletion layer or space charge region (SCR) on the surface of materials. When the photon energy irradiated on the material is larger than its bandgap, photogenerated electrons and holes will be excited. The holes will separate from electrons, and then move along the potential gradient caused by the bending of the energy band to the surface of the material and combine with oxygen ions to desorb oxygen molecules. The process of adsorption and desorption of oxygen molecules is expressed as  $\text{O}_2(\text{g}) + \text{e}^- \rightarrow \text{O}_2^-(\text{ad})$ ,  $\text{h}^+ + \text{O}_2^-(\text{ad}) \rightarrow \text{O}_2(\text{g})$ , where  $\text{O}_2(\text{g})$  denotes oxygen molecules,  $\text{O}_2^-(\text{ad})$  denotes adsorbed oxygen ions,  $\text{e}^-$  denotes electrons, and  $\text{h}^+$  denotes holes. Fig. 6a shows the whole process. In addition, nanostructures with large specific surface area usually exhibit a large increase in the probability of oxygen-related hole trapping and the number of surface trapped states due to a large number of suspended bonds on the surface. Therefore, the prepared SnO<sub>2</sub> NAs can desorb more oxygen, separate photogenerated carriers and reduce their recombination, thus greatly prolonging the carrier lifetime. Because EQE can also be expressed as the ratio of carrier lifetime to transition time between electrodes, the reason why SnO<sub>2</sub> NAs UV detectors exhibit higher EQE performance is explained.

As shown in Fig. 4d, considering that the work function of Au is larger than that of SnO<sub>2</sub> NAs, carrier concentration and

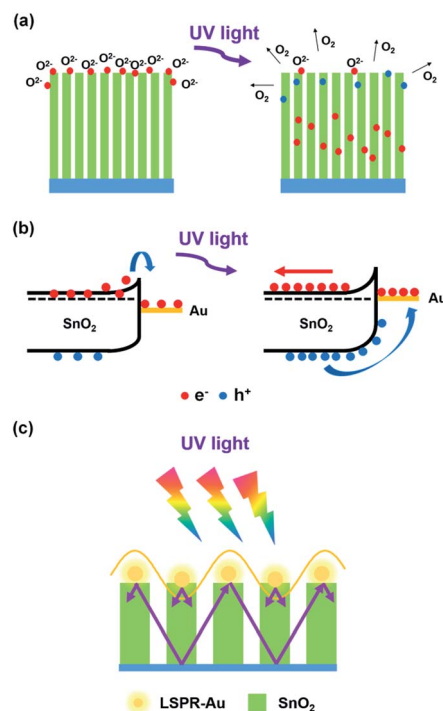


Fig. 6 Schematic diagram of (a) oxygen adsorption and desorption on SnO<sub>2</sub> NAs surface, (b) surface energy band structure of SnO<sub>2</sub>-Au NAs, (c) surface enhanced UV absorption efficiency of Au NPs.

potential redistribution will occur at the interface, forming a Schottky barrier. Electrons transfer from SnO<sub>2</sub> NAs to Au NPs attached to the surface of SnO<sub>2</sub>, and the energy band will bend, which hinders the electron transition and reduces the dark current of devices. When excited by UV light, the photogenerated holes move towards Au NPs along the bend of the energy band, while the photogenerated electrons move backward under the action of the built-in electric field, as shown in Fig. 6b. This effect further reduces the recombination of photogenerated electrons and holes, thus prolonging the carrier lifetime and further improving the EQE performance of devices.

Au NPs at the top of SnO<sub>2</sub> NAs will inevitably increase the scattering and absorption efficiency of UV light. In addition, Au NPs can generate local surface plasmon resonance under UV excitation.<sup>31–35</sup> The near-field amplitude will be enhanced under resonance conditions and localized on the surface of Au NPs. It

Table 2 Comparison of the performance parameters of different SnO<sub>2</sub> based UV detectors

Material	$R$ (A/W)	EQE (%)	$I_{\text{photo}}/I_{\text{dark}}$	$\tau_{\text{g}}$	$\tau_{\text{d}}$	Reference
SnO <sub>2</sub> -Au nanoarrays	1002	376 321	111.3	3.7 s	5.3 s	This work
SnO <sub>2</sub> nanowires	1.18	—	1.67	—	—	13
SnO <sub>2</sub> nanowires	—	8000	5	—	—	27
Au-coated SnO <sub>2</sub> nanowires	—	900	—	—	—	28
Ga <sub>2</sub> O <sub>3</sub> /SnO <sub>2</sub> :Ga core-shell	—	0.362	1066.7	—	—	15
SnO <sub>2</sub> -TiO <sub>2</sub> nanofibers	0.6	—	4500	0.03 s	0.01 s	19
ZnO-SnO <sub>2</sub> nanofibers	—	—	4600	32.2 s	7.8 s	17
ZnS/SnO <sub>2</sub> nanoparticles	$1.2 \times 10^{-8}$	—	13.5	4.2 s	1.5 s	29



can be used as an optical antenna to store UV energy in the near-field, as well as a sub-wavelength scatterer to capture and redistribute UV light onto SnO<sub>2</sub> NAs, as shown in Fig. 6c. This effect can stimulate more photogenerated carriers, which results in high photocurrent and EQE performance of SnO<sub>2</sub>-Au NAs UV detectors.

## 4. Conclusions

In summary, we have fabricated a novel morphology of SnO<sub>2</sub> NAs by a simple and controllable solvothermal method. Oxygen adsorbs on the surface of SnO<sub>2</sub> NRs to form SCRs, which effectively separate space charges. Au NPs induce the growth of smaller and more orderly NRs. At the same time, the band bending of the interface between Au and SnO<sub>2</sub> reduces the recombination of photogenerated electrons and holes, and the LSPR effect of Au improves the absorption efficiency of UV light. Therefore, the prepared UV detectors have high responsivity, EQE and  $D^*$  to UV light with power of only 20  $\mu\text{W cm}^{-2}$  at 1 V bias voltage, which would provide reference value for energy-saving UV detectors.

## Conflicts of interest

There are no conflicts to declare.

## Acknowledgements

This work was supported by the National Natural Science Foundation of China under Grant No. 91833301, and Zhejiang Provincial Natural Science Foundation of China under Grant no. LY17E020005.

## Notes and references

- 1 E. Monroy, F. Calle, J. L. Pau, E. Muñoz, F. Omnès, B. Beaumont and P. Gibart, *J. Cryst. Growth*, 2001, **230**, 537–543.
- 2 E. Muñoz, E. Monroy, J. L. Pau, F. Calle, F. Omnès and P. Gibart, *J. Phys.: Condens. Matter*, 2001, **13**, 7115–7137.
- 3 Q. Zhang, X. Li, Z. He, M. Xu, C. Jin and X. Zhou, *J. Phys. D: Appl. Phys.*, 2019, **52**, 303002.
- 4 Y. Zou, Y. Zhang, Y. Hu and H. Gu, *Sensors*, 2018, **18**, 2072.
- 5 Z. R. Yu and M. Aceves-Mijares, *Appl. Phys. Lett.*, 2009, **95**, 081101.
- 6 E. Monroy, F. Omnès and F. Calle, *Semicond. Sci. Technol.*, 2003, **18**, R33–R51.
- 7 S. Mathur, S. Barth, H. Shen, J. C. Pyun and U. Werner, *Small*, 2005, **1**, 713.
- 8 X. Liu, X. Liu, J. Wang, C. Liao, X. Xiao, S. Guo, C. Jiang, Z. Fan, T. Wang, X. Chen, W. Lu, W. Hu and L. Liao, *Adv. Mater.*, 2014, **26**, 7399–7404.
- 9 Z. R. Dai, Z. W. Pan and Z. L. Wang, *Adv. Funct. Mater.*, 2003, **13**, 9–24.
- 10 J. Zhao, R. Deng, J. Qin, J. Song, D. Jiang, B. Yao and Y. Li, *J. Alloys Compd.*, 2018, **748**, 398–403.
- 11 L. F. Hu, J. Yan, M. Y. Liao, L. M. Wu and X. S. Fang, *Small*, 2011, **7**, 1012–1017.
- 12 L. Gan, M. Liao, H. Li, Y. Ma and T. Zhai, *J. Mater. Chem. C*, 2015, **3**, 8300–8306.
- 13 J. M. Wu and C. H. Kuo, *Thin Solid Films*, 2009, **517**, 3870–3873.
- 14 J. W. Yan, Y. Chen, X. W. Wang, Y. Fu, J. X. Wang, J. Sun, G. Z. Dai, S. H. Tao and Y. L. Gao, *Nanoscale*, 2019, **11**, 2162–2169.
- 15 C. L. Hsu and Y. C. Lu, *Nanoscale*, 2012, **4**, 5710–5717.
- 16 S. Abbas, D.-K. Ban and J. Kim, *Sens. Actuators, A*, 2019, **293**, 215–221.
- 17 W. Tian, T. Zhai, C. Zhang, S. L. Li, X. Wang, F. Liu, D. Q. Liu, X. K. Cai, K. Tsukagoshi, D. Golberg and Y. Bando, *Adv. Mater.*, 2013, **25**, 4625–4630.
- 18 S. Chen, M. Wang, J. F. Ye, J. G. Cai, Y. R. Ma, H. H. Zhou and L. M. Qi, *Nano Res.*, 2013, **6**, 243–252.
- 19 X. D. Li, C. T. Gao, H. G. Duan, B. G. Lu, Y. Q. Wang, L. L. Chen, Z. X. Zhang, X. J. Pan and E. Q. Xie, *Small*, 2013, **9**, 2005–2011.
- 20 X. L. Xu, Y. Chen, G. H. Zhang, S. H. Yan, H. Q. Bian, Q. Chen, Y. Lu and S. Y. Ma, *Mater. Lett.*, 2017, **195**, 159–163.
- 21 X. Chen, K. W. Liu, Z. Z. Zhang, C. R. Wang, B. H. Li, H. F. Zhao, D. X. Zhao and D. Z. Shen, *ACS Appl. Mater. Interfaces*, 2016, **8**, 4185–4191.
- 22 W. I. Park, G. C. Yi, J. W. Kim and S. M. Park, *Appl. Phys. Lett.*, 2003, **82**, 4358–4360.
- 23 Q. Wang, Y. C. Shao, H. P. Xie, L. Lyu, X. L. Liu, Y. L. Gao and J. S. Huang, *Appl. Phys. Lett.*, 2014, **105**, 163508.
- 24 D. Cahen and A. Kahn, *Adv. Mater.*, 2003, **15**, 271–277.
- 25 Y. Wen, Q. S. Wang, L. Yin, Q. Liu, F. Wang, F. M. Wang, Z. X. Wang, K. L. Liu, K. Xu, T. A. Shifa, C. Jiang, J. Xiong and J. He, *Adv. Mater.*, 2016, **28**, 8051–8057.
- 26 J. L. Wang, H. H. Fang, X. D. Wang, X. S. Chen, W. Liu and W. D. Hu, *Small*, 2017, **13**, 1700894.
- 27 C. H. Lin, R. S. Chen, T. T. Chen, H. Y. Chen, Y. F. Chen, K. H. Chen and L. C. Chen, *Appl. Phys. Lett.*, 2008, **93**, 112115.
- 28 C. Y. Chen, J. R. D. Retama, I. W. Wu, D. H. Lien, M. W. Chen, Y. Ding, Y. L. Chueh, C. I. Wu and J. H. He, *ACS Nano*, 2012, **6**, 9366–9372.
- 29 C. Zhang, Y. C. Xie, H. Deng, T. Tumlin, C. Zhang, J. W. Su, P. Yu and J. Lin, *Small*, 2017, **13**, 1604197.
- 30 C. C. Wu, B. W. Du, W. Luo, Y. Liu, T. Y. Li, D. Wang, X. Guo, H. Ting, Z. Y. Fang, S. F. Wang, Z. J. Chen, Y. X. Chen and L. X. Xiao, *Adv. Opt. Mater.*, 2018, **6**, 1800811.
- 31 A. Rajan, H. K. Yadav, V. Gupta and M. Tomar, *Appl. Phys. A*, 2014, **116**, 913–919.
- 32 R. A. Naphade, M. Tathavadekar, J. P. Jog, S. Agarkar and S. Ogale, *J. Mater. Chem. A*, 2014, **2**, 975–984.
- 33 A. Polman, *Science*, 2008, **322**, 868–869.
- 34 H. A. Atwater and A. Polman, *Nat. Mater.*, 2010, **9**, 205–213.
- 35 T. L. Temple and D. M. Bagnall, *J. Appl. Phys.*, 2011, **109**, 084343.

