


 Cite this: *RSC Adv.*, 2017, **7**, 35175

 Received 1st June 2017  
 Accepted 7th July 2017

DOI: 10.1039/c7ra06130g

[rsc.li/rsc-advances](http://rsc.li/rsc-advances)

## Lead-free and stable antimony–silver–halide double perovskite $(\text{CH}_3\text{NH}_3)_2\text{AgSbI}_6$ †

 Ya-Juan Li,‡<sup>ab</sup> Tao Wu,‡<sup>c</sup> Lei Sun,<sup>a</sup> Rui-Xia Yang,<sup>a</sup> Lei Jiang,<sup>a</sup> Peng-Fei Cheng,<sup>ab</sup> Qun-Qing Hao,<sup>ab</sup> Tian-Jun Wang,<sup>a</sup> Rui-Feng Lu<sup>c</sup> and Wei-Qiao Deng  <sup>id</sup> <sup>\*a</sup>

A mixed metal organic–inorganic perovskite  $(\text{CH}_3\text{NH}_3)_2\text{AgSbI}_6$  was developed in which pairs of the Pb(II) atoms in traditional  $\text{CH}_3\text{NH}_3\text{PbI}_3$  perovskite are replaced by Sb(III)/Ag(I) aliovalent units. We used density functional theory (DFT) calculations (the GLLB-SC method) with spin–orbit coupling corrections to predict the optical band gap of the most stable  $\text{MA}_2\text{AgSbI}_6$  structure (2.00 eV). The results suggest that it is a promising light absorber. We then synthesized the double perovskite  $\text{MA}_2\text{AgSbI}_6$  and confirmed its structure using X-ray diffraction (XRD). The optical band gap (1.93 eV) is in good agreement with the DFT calculations. The band positions of the material (vs. vacuum) are provided for future uses such as achieving the precise energy level alignments for potential photovoltaic applications. This perovskite material exhibits excellent stability in ambient air.

## 1. Introduction

Organic–inorganic hybrid perovskite materials such as  $\text{CH}_3\text{NH}_3\text{PbI}_3$  ( $\text{MAPbI}_3$ ) have attracted broad interest in recent years for thin film solar cells. This is because of their high performance and low cost.<sup>1–4</sup> The power conversion efficiency (PCE) of solar cells with organic–inorganic metal trihalide perovskites has increased from 3.8% (ref. 5) to 22.1%,<sup>6</sup> which is comparable to those of commercial silicon solar cells. However, several drawbacks greatly hamper their use in large-scale applications: (I) the PCE is significantly affected by different measurement delay times, scan direction, light, and voltage bias pre-treatments;<sup>7–10</sup> (II) the PCE decreases sharply if the temperature is above 85 °C and humidity above 80%;<sup>11</sup> and (III) Pb is a bio-accumulative and toxic element. In regard to the first drawback, interface optimization or modification has been used to decrease the hysteresis of perovskite-based solar cells.<sup>12,13</sup> Han and co-workers found that the addition of 5-ammoniumvaleric acid (5-AVA) in a  $\text{TiO}_2$ – $\text{ZrO}_2$ – $\text{MAPbI}_3$ –carbon cell improves the stability of the solar cell.<sup>14</sup> Additionally, the solar cell stability improves upon compositional engineering such as doping with  $\text{Cs}^+$  or  $\text{Rb}^+$  in a mixed-cation lead mixed-halide perovskite absorber,<sup>11,15</sup> or substituting spiro-oMeTAD for an inorganic

hole transporting material.<sup>16–18</sup> Layered two-dimensional perovskites such as  $(\text{PEA})_2(\text{MA})_2[\text{Pb}_3\text{I}_{10}]$  have also been used as absorbers to enhance the moisture stability of the solar cells.<sup>19</sup> While the stability of the devices is enhanced, the Pb pollution is still a concern. Therefore, perovskite crystalline structures containing  $\text{Sn}^{2+}$  or  $\text{Ge}^{2+}$  (same group as Pb), which have a low toxicity, have been used for the light absorption layer. However, both the device performance and stability are not ideal. The PCE of solar cells based on  $\text{MAGEI}_3$  have only reached 0.2%. The inadequate device performance can be attributed to a poor open circuit voltage resulting from Ge oxidation ( $\text{Ge}^{2+} \rightarrow \text{Ge}^{4+}$ ).<sup>20</sup> Additionally,  $\text{Sn}^{2+}$  is also easily prone to be oxidized to the tetravalent state, yielding a maximum efficiency of only 6.4% only under a  $\text{N}_2$  atmosphere.<sup>21–24</sup> These issues, together, have pushed researchers to find new air-stable and non-toxic lead-free halide perovskites for photovoltaic application. Recently, developments of  $\text{Bi}^{3+}$  and  $\text{Sb}^{3+}$  related double perovskites have been explored using both theoretical and experimental methods.<sup>25,26</sup>

$\text{Cs}_3\text{B}_2\text{X}_9$  ( $\text{B} = \text{Bi, Sb, X} = \text{I, Br, Cl}$ ) related to cesium lead halides have been investigated for their luminescence properties in 1970s and 1980s.<sup>27–29</sup> The phase transition of  $\text{MA}_3\text{B}_2\text{I}_9$  ( $\text{B} = \text{Sb, Bi}$ ) at different temperature were also studied to explore their potential applications as ferroelectric and dielectric materials in later 1990s.<sup>30–35</sup> These materials have a hexagonal crystalline phase, which can exhibited either 0-D dimers or 2-D layered structure based on the preparation method.<sup>36</sup> Among them,  $\text{Cs}_3\text{Bi}_2\text{I}_9$  has been reported with a PCE of over 1%,<sup>37</sup> while  $\text{MA}_3\text{Sb}_2\text{I}_9$  and  $\text{MA}_3\text{Bi}_2\text{I}_9$  can have PCEs of ~0.5% and ~0.2%, respectively.<sup>38–40</sup> Furthermore, mixed metal double perovskites with a basic formula  $\text{A}_2\text{B}'\text{B}''\text{X}_6$  ( $\text{A} = \text{Cs}; \text{B}' = \text{Cu, Ag, Au}; \text{B}'' = \text{Bi, Sb}; \text{X} = \text{I, Br, Cl}$ ) exhibit tunable band gaps and lower carrier

<sup>a</sup>State Key Laboratory of Molecular Reaction Dynamics, Dalian Institute of Chemical Physics, Chinese Academy of Science, Dalian 116023, P. R. China. E-mail: dengwq@dicp.ac.cn

<sup>b</sup>University of the Chinese Academy of Sciences, Beijing 100039, P. R. China

<sup>c</sup>Department of Applied Physics, Nanjing University of Science and Technology, Nanjing 210094, P. R. China

† Electronic supplementary information (ESI) available: See DOI: 10.1039/c7ra06130g

‡ These authors contributed equally to this work.



effective masses.<sup>26</sup> It is for these reasons that they have attracted intention as an alternative to  $\text{MAPbI}_3$ .  $\text{Cs}_2\text{AgBiBr}_6$  (2.19 eV) and  $\text{Cs}_2\text{AgBiCl}_6$  (2.77 eV) has been examined to consist of two types of octahedral alternating in rock-salt face-centered cubic structure.<sup>41,42</sup>  $\text{Cs}_2\text{AgBiBr}_6$  is predicted to be a promising photovoltaic material with long carrier recombination lifetime,<sup>43,44</sup> however, it has narrow stable chemical potential region.<sup>45</sup> Deng *et al.* screened hybrid double perovskites  $\text{MA}_2\text{B}'\text{BiX}_6$  ( $\text{B}' = \text{Cu, Ag; X = Cl, Br, I}$ ) with band gaps similar to those of the  $\text{MAPbX}_3$ . However,  $\text{MA}_2\text{CuBiX}_6$  was found to be unstable.<sup>46</sup> Later,  $\text{MA}_2\text{KBiCl}_6$  (3.04 eV)<sup>47</sup> and  $\text{MA}_2\text{AgBiBr}_6$  (2.02 eV)<sup>45</sup> were also synthesized for potential photovoltaic applications. However,  $\text{MA}_2\text{B}'\text{SbX}_6$  materials have not yet been reported.

To address a stable lead-free perovskite material, we replaced the pairs of  $\text{Pb}(\text{II})$  atoms with  $\text{Sb}(\text{III})/\text{Ag}(\text{I})$  aliovalent units, thus resulting in an alternative mixed metal organic-inorganic perovskite  $\text{MA}_2\text{AgSbI}_6$ . In this regard, we first used DFT calculations to predict its properties before synthesizing the material. Its experimental properties are in good agreement with the calculated predictions.  $\text{MA}_2\text{AgSbI}_6$  exhibits efficient optical absorption and has a high stability, therefore making it a promising absorber for lead-free perovskites solar cells.

## 2. Experimental section

### 2.1 Theoretical calculation

All structure optimization were performed by using the generalized gradient approximation (GGA) with density functional theory (DFT) as implemented in the Vienna *ab initio* simulation package (VASP).<sup>48,49</sup> The projector augmented-wave method is applied to describe the interaction between electron and ion.<sup>50</sup> The Perdew–Burke–Ernzerhof (PBE) functional was chosen to describe the electron exchange–correlation.<sup>51</sup> In particular, the kinetic energy cutoff for plane wave basis set was set to 500 eV, and the numerical convergence criteria was  $10^{-4}$  eV for the energy and  $0.02 \text{ eV} \text{ \AA}^{-1}$  for the force. A Monkhorst–Pack  $k$  point mesh of  $6 \times 6 \times 4$  was used for all calculations. The lattice and atom positions of all suspected structures are fully optimized without any symmetry constraints.

### 2.2 The material preparation

The  $\text{MA}_2\text{AgSbI}_6$  was prepared by solid state reactions in a sealed vacuum tube.<sup>52</sup> 502 mg (1 mmol)  $\text{SbI}_3$ , 234 mg (1 mmol)  $\text{AgI}$  and 318 mg (2 mmol)  $\text{MAI}$  were ground uniformly in an agate mortar inside an  $\text{Ar}_2$  gas glove box to ensure a homogenous mixture. This yielded a light red powder, which was sealed in an evacuated (mechanical pump for 40 min)  $1 \text{ cm} \times 10 \text{ cm}$  quartz ampere tube. The powder turned into a crimson color after loading the tube in an oven with constant temperature  $150^\circ \text{C}$  for 2 h.

### 2.3 Characterization

The solid UV-Vis absorption spectra were recorded using a JASCO V-550 UV-Vis spectrophotometer. The XRD spectra were recorded using an Empyrean X-ray diffractometer with  $\text{Cu K}\alpha$  radiation ( $\lambda = 1.54 \text{ \AA}$ ). The UPS measurements were carried out

using a He I ( $h\nu = 21.2 \text{ eV}$ ) source in a UHV chamber. The morphology of the powder was characterized by FESEM (JSM-7800F) and TEM (JEM-2100).

## 3. Results and discussion

The tolerance factor was calculated to evaluate the metal ionic radii mismatch for the  $\text{MA}_2\text{AgSbI}_6$  double perovskite material. Using the approach introduced by V. M. Goldschmidt,<sup>53</sup> the tolerance factor equation can be expressed as follows:

$$\alpha = \frac{(r_A + r_X)}{\sqrt{2}(r_B + r_X)}$$

In our research system,  $r_A$  and  $r_X$  are the effective ionic radii of the MA and iodide, which are 180 pm and 220 pm,<sup>54</sup> respectively.  $r_B$  is the effective ionic radius of the metal with six-coordinates. Considering that there are two metal elements in our system (Sb and Ag) with effective ionic radii corresponding to 76 pm and 115 pm,<sup>55</sup> we set  $r_B$  to their average value (95.5 pm). The calculated tolerance factor of  $\text{MA}_2\text{AgSbI}_6$  is 0.90, which is a little larger than that observed for  $\text{MAPbI}_3$  (0.83). This indicates that  $\text{MA}_2\text{AgSbI}_6$  can form a basic perovskite structure.

The orientation of MA cation in tetragonal phase affects the structural stability and electronic properties of  $\text{MAPbI}_3$ , which has been fully discussed by Claudio *et al.*<sup>56</sup> Herein, we considered ten different orientations of the MA cation in  $\text{MA}_2\text{AgSbI}_6$ . The orientation of the MA cation can be described by the angle of the MA cation projection in the *ab*-plane with respect to the *a*-axis ( $\theta = 45^\circ, 135^\circ, 225^\circ, 315^\circ$ ) and the angle of the MA cation with respect to the *ab* plane ( $\varphi = 30^\circ, -30^\circ$ ). These orientations are illustrated in Scheme S1 (ESI†). After constructing the  $\text{MA}_2\text{AgSbI}_6$  structures with different MA cation orientations (Table S1†), we start the structural optimization of the tetragonal phase. All ten optimized structures (denoted as number 1–10 in Fig. 1), with the exceptions of structure 3 ( $\alpha = 90.5^\circ, \beta = 89.2^\circ, \gamma = 84.5^\circ$ ) and 9 ( $\alpha = 90.1^\circ, \beta = 91.2^\circ, \gamma = 96.6^\circ$ ) clearly exhibit the orthogonal phase. Structures 3 and 9 each have structural distortions, which tend to form hexagonal structures. This serves as evidence that the orientation of the MA cation influences the inorganic framework. The calculated relative energies and band gaps are listed in Table 1. Structures 9 and 3 have lower relative energies than the other structures ( $\sim 0.2\text{--}0.4$  eV), which explains why both of these structures exhibit large distortions from the orthogonal phase and prefer the hexagonal configuration. Additionally, there are three groups in terms of band gap classification: small gaps (0.00 to 0.10 eV), mid-sized gaps (0.60 to 0.80 eV), and large gaps (1.10 to 1.30 eV). The broad gap range indicates that the band gap of  $\text{MA}_2\text{AgSbI}_6$  is heavily affected by the orientation of the MA cation and the inorganic framework. This may be due to different electronic couplings between the MA cation and the metal-iodine framework. Moreover, there are two distinct types of metal-iodine bonds ( $\text{Sb}-\text{I}$  and  $\text{Ag}-\text{I}$ ) in  $\text{MA}_2\text{AgSbI}_6$ . This differs from  $\text{MAPbI}_3$ , which has only one type of metal-iodine bond ( $\text{Pb}-\text{I}$ ). The orientation of the MA cation therefore affects  $\text{MA}_2\text{AgSbI}_6$  more significantly than  $\text{MAPbI}_3$ .<sup>56</sup>



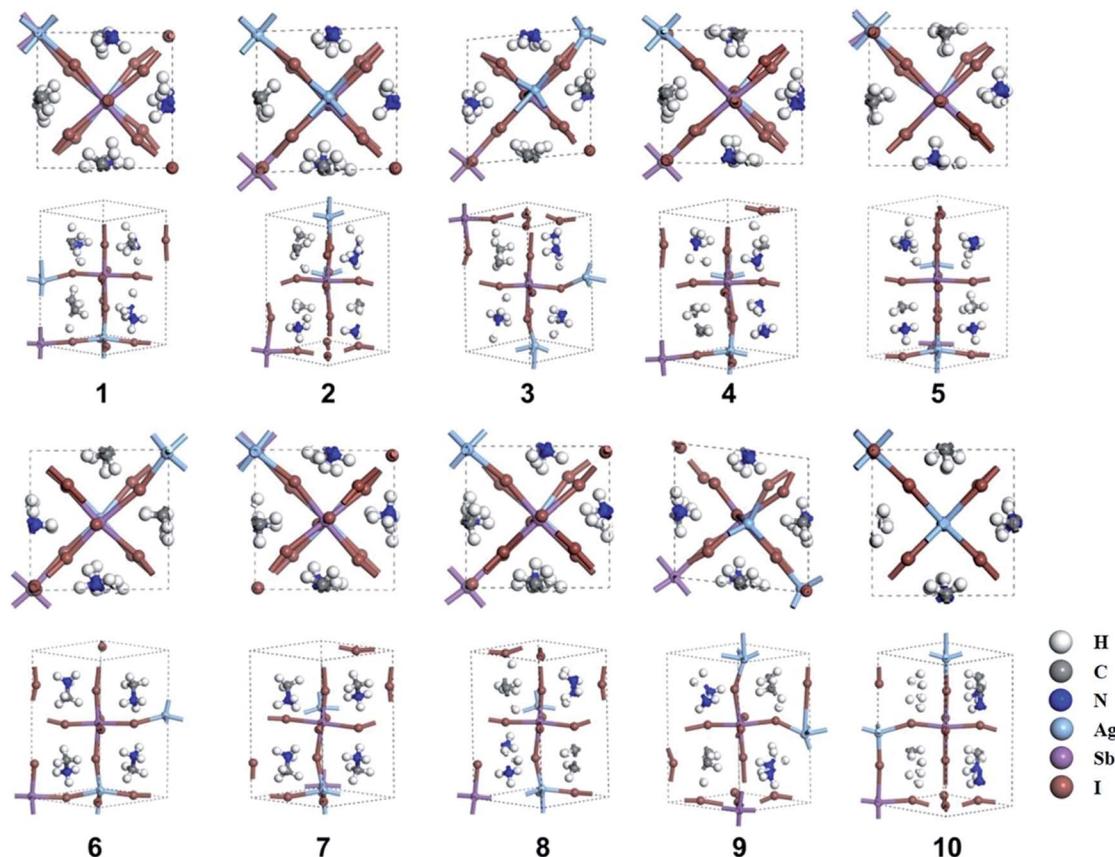


Fig. 1 Ten optimized structures (1–10) of  $\text{MA}_2\text{AgSbI}_6$ . The views along the [001] and [110] planes are shown for each structure.

Table 1 Lattice parameters ( $\text{\AA}$ ), relative energy (eV) and band gap ( $E_g$ , eV) for all investigated structures

| Structure | Lattice parameters ( $a, b, c$ ) | Relative energy | $E_g$ |
|-----------|----------------------------------|-----------------|-------|
| 1         | 8.51, 8.52, 13.02                | 0.19            | 0.61  |
| 2         | 8.46, 8.50, 13.22                | 0.17            | 0.75  |
| 3         | 8.54, 8.59, 13.35                | 0.04            | 1.16  |
| 4         | 8.49, 8.50, 13.08                | 0.27            | 0.65  |
| 5         | 8.45, 8.51, 13.21                | 0.45            | 0.09  |
| 6         | 8.47, 8.52, 13.19                | 0.27            | 0.66  |
| 7         | 8.45, 8.50, 13.27                | 0.44            | 0.02  |
| 8         | 8.45, 8.47, 13.31                | 0.17            | 0.73  |
| 9         | 8.59, 8.55, 13.37                | 0               | 1.25  |
| 10        | 8.48, 8.46, 13.35                | 0.27            | 0.79  |

Structure 9 is the most stable structure as it has the lowest energy based on calculations using the PBE functional. Additionally, the PBE functional shows that its band gap (1.25 eV) is appropriate for photovoltaic application. In light of this, we focus on structure 9 in the following discussion. Considering standard DFT calculations seriously underestimate semiconductor band gaps, the inclusion of relativistic effects is important to calculate more accurate band gaps for hybrid perovskites. The GLLB-SC model potential method<sup>57</sup> implemented in the GPAW code,<sup>58</sup> was previously used to predict the band gaps for hybrid organic-inorganic perovskites. The mean

absolute error was only 0.2 eV compared with experiments.<sup>59</sup> The band gap (2.12 eV) of structure 9 was calculated by using this method. To further improve the accuracy, spin-orbit coupling (SOC) calculations were also taken into account. The SOC effect induces a small decrease in the band gap ( $\sim 0.12$  eV), which results in a band gap of 2.00 eV for structure 9. This proper band gap means  $\text{MA}_2\text{AgSbI}_6$  may exhibit efficient absorption in the visible light region. The band dispersion of structure 9 is depicted in Fig. 2 using the PBE functional. The valence band maximum (VBM) is located at the  $K$  point whereas

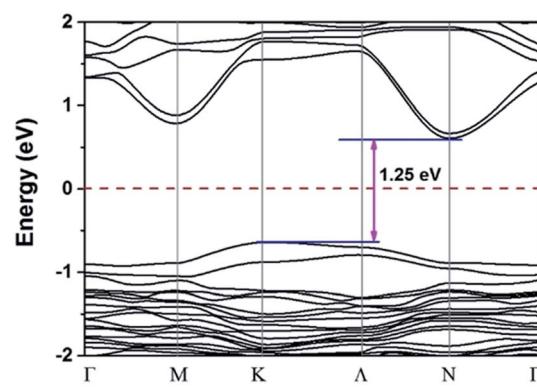


Fig. 2 PBE calculated band structure of 9 in the first Brillouin zone; the Fermi level is set to be zero.

the conduction band minimum (CBM) is at the *N* point, thus indicating that  $\text{MA}_2\text{AgSbI}_6$  has an indirect band gap. Similar to silicon, this means that photoelectric conversion in  $\text{MA}_2\text{AgSbI}_6$  requires the aid of lattice phonons. Further, we calculated the effective mass of  $\text{MA}_2\text{AgSbI}_6$  to evaluate its charge transport properties. Based on the DFT calculations, we obtain a hole effective mass of  $m_h^* = 2.37$  and 4.49 along the *K*-*M* and *K*-*A* directions. However, the electron effective masses are  $m_e^* = 0.39$  and 0.53 along the *N*-*A* and *N*-*T* directions, respectively. These results imply that charge transport is anisotropic and electron transport is much faster than hole transport. Its effective mass of the electron are comparable to  $\text{MAPbI}_3$ ,<sup>60</sup> which indicates that  $\text{MA}_2\text{AgSbI}_6$  exhibits facile electron transport properties. Moreover, the charge densities of the VBM and CBM of structure 9 are plotted in Fig. 3. The charge density of VBM is mainly distributed around I and Ag atoms, while the charge density of CBM is localized mostly on the Sb atoms. This is beneficial for charge separation in photovoltaic application.

To verify our theoretical predictions, we synthesized the  $\text{MA}_2\text{AgSbI}_6$  perovskite by a solid-state reaction (the detailed synthetic method is listed in the Experimental section). Its powder XRD spectrum coincides well with the simulated one (Fig. 4). The XRD spectrum, outside of peak intensity, was the same (Fig. S1†) for each of the three different preparation temperatures 150 °C, 200 °C and 220 °C. This demonstrated the new material is stable in broad temperature range. Notably, the

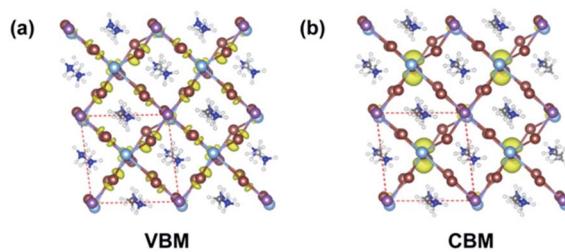


Fig. 3 Partial orbital charge density of VBM (a) and CBM (b), respectively. The isosurface level is 0.005 [purple: Sb; silvery: Ag; brown: I; grey: C; blue: N; white: H atoms].

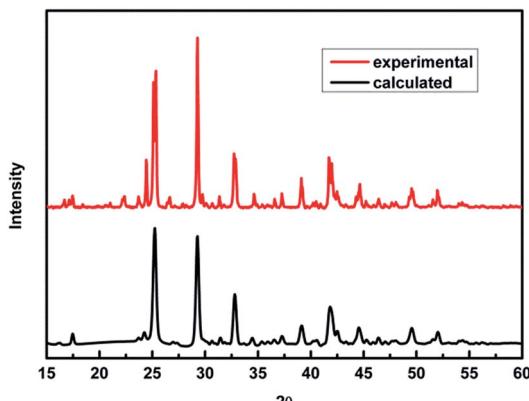


Fig. 4 PXRD patterns of  $\text{MA}_2\text{AgSbI}_6$  predicted by theory and obtained experimentally.

XRD spectrum for the powder is relatively unaltered after being exposed to air for 370 days (Fig. 5). We find the peak intensity increases, which may be attributed to the different crystal sizes. However, there are no additional peaks associated with degradation products. The material is therefore more stable than  $\text{MAPbI}_3$ , which will partially decompose to  $\text{PbI}_2$  yellow powder as shown by XRD spectrum or the naked eye.<sup>43</sup> The thermal decomposition temperature, based on thermogravimetric analysis, can reach up to 260 °C (Fig. S2†) in  $\text{N}_2$  gas flow. This material therefore exhibits excellent air and thermal stability. The synthesized crimson powder efficiently absorbs light in the ultraviolet-visible range of 200–650 nm (Fig. 6), and has an optical band gap of 1.93 eV (based on the Tauc plot shown in the Fig. 6 inset). The band gap is close to the theoretical prediction. Moreover, the powder exhibits a poly-crystalline morphology as shown in the SEM and TEM images, and further confirmed by the TEM selected area electron diffraction (SAED) pattern (Fig. S3†). The Fermi level (4.65 eV) of  $\text{MA}_2\text{AgSbI}_6$  was also obtained from the cut-off energy (16.55 eV) of ultraviolet photo-electron spectrum (Fig. 7) and the energy (21.2 eV) of the He I source. The VBM position is 1.63 eV lower than the Fermi level. The CBM position was calculated according to the optical band gap (1.93 eV), which was obtained from absorption spectrum.

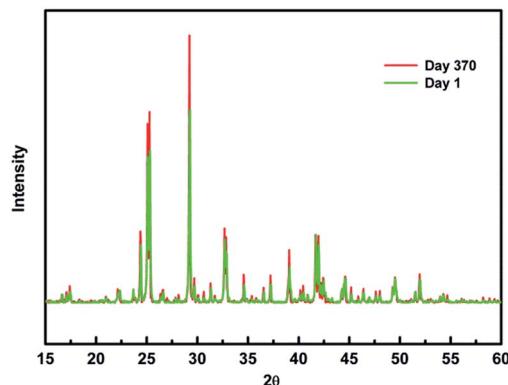


Fig. 5 PXRD patterns of fresh  $\text{MA}_2\text{AgSbI}_6$  compared to  $\text{MA}_2\text{AgSbI}_6$  after 370 days of exposure to air.

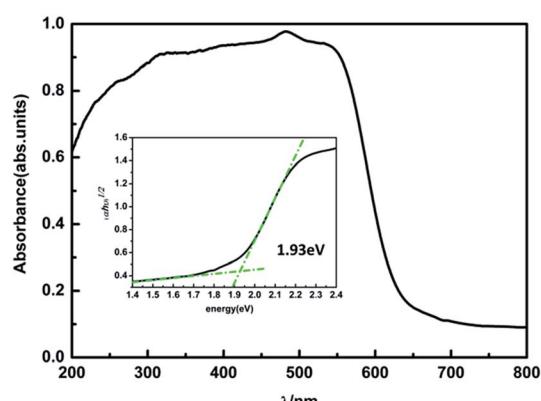


Fig. 6 The UV-vis absorption spectrum and Tauc plot of  $\text{MA}_2\text{AgSbI}_6$  (inset).

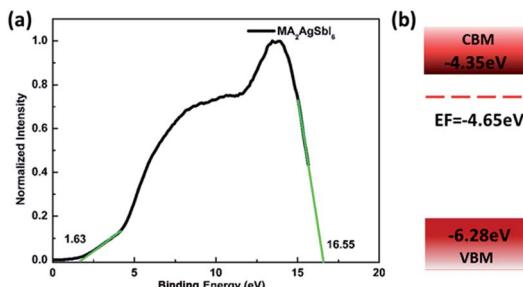


Fig. 7 (a) UPS spectrum of  $\text{MA}_2\text{AgSbI}_6$ ; (b) schematic of energy level alignment.

The Fermi level is closed to the CBM, which indicates that it may be an *N*-type semiconductor (Fig. 7). The energy level alignment is provided as a reference for future endeavors to achieve precise band alignments for high performance photovoltaic devices.

## 4. Conclusions

In summary, we synthesized a new kind of antimony and silver based double perovskite material with a band gap of 1.93 eV. Both DFT calculations and experimental results are in good agreement. The new double perovskite exhibits a high stability in ambient conditions (room temperature, in air at 20–60% humidity) for 370 days. The ideal band gap and excellent stability make this material a promising absorber for photovoltaic applications.

## Acknowledgements

This work was supported by the Ministry of Science and Technology of China (grant number 2017YFA0204800) and National Natural Science Foundation of China (grant number 21525315, 21403211, and 91333116).

## References

- 1 M. M. Lee, J. Teuscher, T. Miyasaka, T. N. Murakami and H. J. Snaith, *Science*, 2012, **338**, 643–647.
- 2 J. Burschka, N. Pellet, S.-J. Moon, R. Humphry-Baker, P. Gao, M. K. Nazeeruddin and M. Graetzel, *Nature*, 2013, **499**, 316.
- 3 X. Li, D. Bi, C. Yi, J.-D. Decoppet, J. Luo, S. M. Zakeeruddin, A. Hagfeldt and M. Gratzel, *Science*, 2016, **353**, 58–62.
- 4 W. S. Yang, J. H. Noh, N. J. Jeon, Y. C. Kim, S. Ryu, J. Seo and S. I. Seok, *Science*, 2015, **348**, 1234–1237.
- 5 A. Kojima, K. Teshima, Y. Shirai and T. Miyasaka, *J. Am. Chem. Soc.*, 2009, **131**, 6050.
- 6 National Renewable Energy Laboratory, *Best Research-Cell Efficiencies*, [http://www.nrel.gov/pv/assets/images/efficiency\\_chart.jpg](http://www.nrel.gov/pv/assets/images/efficiency_chart.jpg), 2016.
- 7 E. L. Unger, E. T. Hoke, C. D. Bailie, W. H. Nguyen, A. R. Bowring, T. Heumueller, M. G. Christoforo and M. D. McGehee, *Energy Environ. Sci.*, 2014, **7**, 3690–3698.
- 8 W. Tress, N. Marinova, T. Moehl, S. M. Zakeeruddin, M. K. Nazeeruddin and M. Grätzel, *Energy Environ. Sci.*, 2015, **8**, 995–1004.
- 9 Z. G. Xiao, Y. B. Yuan, Y. C. Shao, Q. Wang, Q. F. Dong, C. Bi, P. Sharma, A. Gruverman and J. S. Huang, *Nat. Mater.*, 2015, **14**, 193–198.
- 10 Y. H. Zhang, P. Wang, X. G. Yu, J. S. Xie, X. Sun, H. H. Wang, J. B. Huang, L. B. Xu, C. Cui, M. Lei and D. R. Yang, *J. Mater. Chem. A*, 2016, **4**, 18509–18515.
- 11 D. P. McMeekin, G. Sadoughi, W. Rehman, G. E. Eperon, M. Saliba, M. T. Hoerantner, A. Haghaghirad, N. Sakai, L. Korte, B. Rech, M. B. Johnston, L. M. Herz and H. J. Snaith, *Science*, 2016, **351**, 151–155.
- 12 H. P. Zhou, Q. Chen, G. Li, S. Luo, T. B. Song, H. S. Duan, Z. R. Hong, J. B. You, Y. S. Liu and Y. Yang, *Science*, 2014, **345**, 542–546.
- 13 D. Yang, R. Yang, X. Ren, X. Zhu, Z. Yang, C. Li and S. Liu, *Adv. Mater.*, 2016, **28**, 5206.
- 14 A. Mei, X. Li, L. Liu, Z. Ku, T. Liu, Y. Rong, M. Xu, M. Hu, J. Chen, Y. Yang, M. Graetzel and H. Han, *Science*, 2014, **345**, 295–298.
- 15 M. Saliba, T. Matsui, K. Domanski, J. Y. Seo, A. Ummadisingu, S. M. Zakeeruddin, J. P. Correa-Baena, W. R. Tress, A. Abate, A. Hagfeldt and M. Gratzel, *Science*, 2016, **354**, 206–209.
- 16 W. Chen, Y. Wu, Y. Yue, J. Liu, W. Zhang, X. Yang, H. Chen, E. Bi, I. Ashraful, M. Graetzel and L. Han, *Science*, 2015, **350**, 944–948.
- 17 F. Zhang, X. Yang, M. Cheng, W. Wang and L. Sun, *Nano Energy*, 2016, **20**, 108–116.
- 18 R. F. Service, *Science*, 2016, **351**, 113–114.
- 19 I. C. Smith, E. T. Hoke, D. Solis-Ibarra, M. D. McGehee and H. I. Karunadasa, *Angew. Chem., Int. Ed.*, 2014, **53**, 11232–11235.
- 20 T. Krishnamoorthy, H. Ding, C. Yan, W. L. Leong, T. Baikie, Z. Zhang, M. Sherburne, S. Li, M. Asta, N. Mathews and S. G. Mhaisalkar, *J. Mater. Chem. A*, 2015, **3**, 23829–23832.
- 21 F. Hao, C. C. Stoumpos, D. H. Cao, R. P. H. Chang and M. G. Kanatzidis, *Nat. Photonics*, 2014, **8**, 489–494.
- 22 N. K. Noel, S. D. Stranks, A. Abate, C. Wehrenfennig, S. Guarnera, A.-A. Haghaghirad, A. Sadhanala, G. E. Eperon, S. K. Pathak, M. B. Johnston, A. Petrozza, L. M. Herz and H. J. Snaith, *Energy Environ. Sci.*, 2014, **7**, 3061–3068.
- 23 M. H. Kumar, S. Dharani, W. L. Leong, P. P. Boix, R. R. Prabhakar, T. Baikie, C. Shi, H. Ding, R. Ramesh, M. Asta, M. Graetzel, S. G. Mhaisalkar and N. Mathews, *Adv. Mater.*, 2014, **26**, 7122.
- 24 T. B. Song, T. Yokoyama, C. C. Stoumpos, J. Logsdon, D. H. Cao, M. R. Wasielewski, S. Aramaki and M. G. Kanatzidis, *J. Am. Chem. Soc.*, 2017, **139**, 836–842.
- 25 Z. Shi, J. Guo, Y. Chen, Q. Li, Y. Pan, H. Zhang, Y. Xia and W. Huang, *Adv. Mater.*, 2017, **29**, 1605005.
- 26 G. Volonakis, M. R. Filip, A. A. Haghaghirad, N. Sakai, B. Wenger, H. J. Snaith and F. Giustino, *J. Phys. Chem. Lett.*, 2016, **7**, 1254–1259.



27 B. Chabot and E. Parthe, *Acta Crystallogr., Sect. B: Struct. Crystallogr. Cryst. Chem.*, 1978, **34**, 645–648.

28 C. W. M. Timmermans and G. Blasse, *Phys. Status Solidi B*, 1981, **106**, 647–655.

29 C. W. M. Timmermans, S. O. Cholakh and G. Blasse, *J. Solid State Chem.*, 1983, **46**, 222–233.

30 R. Jakubas and L. Sobczyk, *Phase Transitions*, 1990, **20**, 163–193.

31 J. Zaleski, R. Jakubas and L. Sobczyk, *Ferroelectrics*, 1990, **103**, 83–90.

32 R. Jakubas, R. Decressain and J. Lefebvre, *J. Phys. Chem. Solids*, 1992, **53**, 755–759.

33 P. Koziol, Y. Furukawa, D. Nakamura and R. Jakubas, *Bull. Chem. Soc. Jpn.*, 1992, **65**, 1707–1709.

34 G. Bator, R. Jakubas, J. Baran and H. Ratajczak, *Vib. Spectrosc.*, 1995, **8**, 425–433.

35 C. Pawlaczyk, R. Jakubas and H. G. Unruh, *Solid State Commun.*, 1998, **108**, 247–250.

36 K. Yamada, H. Sera, S. Sawada, H. Tada, T. Okuda and H. Tanaka, *J. Solid State Chem.*, 1997, **134**, 319–325.

37 B. Saparov, F. Hong, J. P. Sun, H. S. Duan, W. W. Meng, S. Cameron, I. G. Hill, Y. F. Yan and D. B. Mitzi, *Chem. Mater.*, 2015, **27**, 5622–5632.

38 T. Singh, A. Kulkarni, M. Ikegami and T. Miyasaka, *ACS Appl. Mater. Interfaces*, 2016, **8**, 14542–14547.

39 J.-C. Hebig, I. Kühn, J. Flohre and T. Kirchartz, *ACS Energy Lett.*, 2016, **1**, 309–314.

40 S. Oz, J. C. Hebig, E. W. Jung, T. Singh, A. Lepcha, S. Olthof, J. Flohre, Y. J. Gao, R. German, P. H. M. van Loosdrecht, K. Meerholz, T. Kirchartz and S. Mathur, *Sol. Energy Mater. Sol. Cells*, 2016, **158**, 195–201.

41 E. T. McClure, M. R. Ball, W. Windl and P. M. Woodward, *Chem. Mater.*, 2016, **28**, 1348–1354.

42 M. R. Filip, S. Hillman, A. A. Haghghirad, H. J. Snaith and F. Giustino, *J. Phys. Chem. Lett.*, 2016, **7**, 2579–2585.

43 A. H. Slavney, T. Hu, A. M. Lindenberg and H. I. Karunadasa, *J. Am. Chem. Soc.*, 2016, **138**, 2138–2141.

44 Z. Xiao, W. Meng, J. Wang and Y. Yan, *ChemSusChem*, 2016, **9**, 2628–2633.

45 F. Wei, Z. Deng, S. Sun, F. Zhang, D. M. Evans, G. Kieslich, S. Tominaka, M. A. Carpenter, J. Zhang, P. D. Bristowe and A. K. Cheetham, *Chem. Mater.*, 2017, **29**, 1089–1094.

46 Z. Deng, F. Wei, S. Sun, G. Kieslich, A. K. Cheetham and P. D. Bristowe, *J. Mater. Chem. A*, 2016, **4**, 12025–12029.

47 F. Wei, Z. Deng, S. Sun, F. Xie, G. Kieslich, D. M. Evans, M. A. Carpenter, P. D. Bristowe and A. K. Cheetham, *Mater. Horiz.*, 2016, **3**, 328–332.

48 G. Kresse and J. Furthmüller, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 1996, **54**, 11169–11186.

49 G. Kresse and J. Furthmüller, *Comput. Mater. Sci.*, 1996, **6**, 15–50.

50 P. E. Blochl, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 1994, **50**, 17953–17979.

51 J. P. Perdew, K. Burke and M. Ernzerhof, *Phys. Rev. Lett.*, 1996, **77**, 3865–3868.

52 C. C. Stoumpos, C. D. Malliakas and M. G. Kanatzidis, *Inorg. Chem.*, 2013, **52**, 9019–9038.

53 V. M. Goldschmidt, *Naturwissenschaften*, 1926, **14**, 477–485.

54 N. K. McKinnon, D. C. Reeves and M. H. Akabas, *J. Gen. Physiol.*, 2011, **138**, 453–466.

55 R. D. Shannon, *Acta Crystallogr., Sect. A: Cryst. Phys., Diff., Theor. Gen. Crystallogr.*, 1976, **32**, 751–767.

56 C. Quarti, E. Mosconi and F. De Angelis, *Chem. Mater.*, 2014, **26**, 6557–6569.

57 M. Kuusma, J. Ojanen, J. Enkovaara and T. T. Rantala, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 2010, **82**, 3893–3898.

58 J. J. Mortensen, L. B. Hansen and K. W. Jacobsen, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 2005, **71**, 035109–035119.

59 I. E. Castelli, J. M. Garcia-Lastra, K. S. Thygesen and K. W. Jacobsen, *APL Mater.*, 2014, **2**, 081514–081520.

60 P. Umari, E. Mosconi and F. De Angelis, *Sci. Rep.*, 2014, **4**, 4467–4473.

