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# Grouping and aggregation of ligand protected Au<sub>9</sub> clusters on TiO<sub>2</sub> nanosheets†

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Atomically precise chemically synthesised Au clusters, in the form of [Au<sub>9</sub>(PPh<sub>3</sub>)<sub>8</sub>](NO<sub>3</sub>)<sub>3</sub>, were deposited onto titania nanosheets after UV pre-treatment of the substrate and examined with scanning tunneling microscopy (STM), atomic force microscopy (AFM) and synchrotron X-ray photoelectron spectroscopy (XPS) before and after heat treatment. The STM, AFM and XPS results complement each other. AFM was performed to determine the height of the deposited species and their dispersion on titania nanosheets. STM shows groups of clusters that at least partially consist of individual clusters both before and after annealing. STM cannot exclude the existence of individual clusters on the titania nanosheets outside the groups. XPS shows that before annealing the Au clusters are attached to the titania surface as individual clusters thus as clusters with non-agglomerated cluster cores. After annealing, both individual and agglomerated clusters are found on the surface. The combination of AFM, STM and XPS shows that the groups formed by the clusters consist of individual and agglomerated clusters.

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## Introduction

Nanoclusters (NCs) are of particular interest because of their unique physical and chemical properties, which are different to the properties of the respective bulk materials.<sup>1–11</sup> Small clusters containing only a few metal atoms show fluxionality leading to a change in conformation of the clusters at room temperature.<sup>12</sup> The fluxionality decreases with the increasing size of the cluster.<sup>13–15</sup> Small clusters show discrete energy levels for the valence electrons. With the increasing size of the cluster the energy levels move closer together. At a size in the order of a few hundred atoms, the metal NCs start to exhibit metallic character.<sup>16–19</sup>

Metal NCs can be generated either in the gas-phase and subsequently deposited onto surfaces under UHV<sup>15,20</sup> Alternatively, they can be synthesised chemically as size-specific ligand-protected clusters, which can be deposited onto surfaces from solutions.<sup>21,22</sup> Such supported clusters can be used in a wide range of applications, such as sensors and catalysis. Metal NCs deposited onto surfaces modify the electronic structure of the

surfaces. For understanding the metal NCs properties deposited onto surfaces it is important to study their size and their dispersion on the surfaces as well as their electronic properties.<sup>23,24</sup> Subsequent to deposition on a surface, the clusters can agglomerate and lose the properties of the small clusters. It is thus of interest to investigate the size of the clusters after application of various post-treatments such as heating or chemical washing in an attempt to remove ligands for activating the deposited clusters as catalytic sites. Among all the metal NCs available today, specifically Au NCs have attracted interest in catalysis for CO oxidation because Au NCs have unique electronic properties.<sup>25,26</sup> The catalytic activity of Au clusters deposited onto surfaces has been shown to depend strongly on the number of atoms forming the cluster core.<sup>21,22,27–33</sup>

The geometrical and electronic structure of the clusters depends on the size of the Au NCs. and are thus affected by the distribution of the clusters over the surface, *e.g.* whether or not clusters have agglomerated to form larger clusters or even nanoparticles (NPs) (>100 atoms). Microscopy techniques such as atomic force microscopy (AFM)<sup>34</sup> and scanning tunneling microscopy (STM)<sup>19,35–39</sup> are important tools for determining the geometrical and electronic properties of clusters on surfaces, especially their size and distribution on the surfaces. Scanning probe microscopy techniques (SPM) such as AFM and STM are less damaging, milder techniques than electron microscopy such as transmission electron microscopy (TEM). Recently, we have determined the geometrical structure of [Au<sub>9</sub>(PPh<sub>3</sub>)<sub>8</sub>](NO<sub>3</sub>)<sub>3</sub> clusters (hereafter abbreviated as Au<sub>9</sub>) clusters deposited on titania nanosheets with atomic

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resolution using high-angle annular dark field scanning transmission electron microscopy (HAADF STEM).<sup>12</sup> In electron microscopy, electron irradiation can easily affect the samples due to the energy transferred from the electron beam to the sample which can cause serious damage to the samples<sup>40,41</sup> by breaking up clusters or inducing agglomeration. In contrast, SPMs are in principle capable of acquiring high-resolution images without affecting the size and distributions of clusters across surfaces. The chemical compositions of the samples can be identified by high intensity electron spectroscopy such as synchrotron X-Ray Photoelectron Spectroscopy (XPS).<sup>21,22,42</sup> In previous work we have characterized Au<sub>9</sub> clusters (hereafter abbreviated as Au<sub>9</sub>) with XPS subsequently to the deposition of the clusters on titania.<sup>21,22</sup> The final state effect of protected Au NCs in XPS offers the possibility to estimate the size of clusters. The XPS data can also be used to determine whether or not ligands are removed from the metallic cluster core and whether or not the Au NCs agglomerate.<sup>21,22</sup> However, XPS does not have the lateral resolution required to determine the distribution of clusters on a surface – a challenge which motivated this study.

Titania is a particularly interesting support for Au NP-based heterogeneous catalysis.<sup>43</sup> On reducible transition metal oxides such as TiO<sub>2</sub>, the presence of O vacancies associated with extra electrons result in a stronger bond with a metal cluster.<sup>44–46</sup> Titania nanosheets are used in the present work as support for Au<sub>9</sub> clusters. Ultrathin oxide films with a thickness of 20 Å or less are considered as an interesting support material because their properties differ from the respective thicker bulk material.<sup>47</sup> Examples are the dielectric constant, which is higher for TiO<sub>2</sub> nanosheets compared to rutile or anatase the density of defect on the surfaces of these two titania materials and the band gaps of the materials.<sup>48</sup> Thus, it can be expected that the interaction of the Au NCs with titania nanosheets is expected to be different to other titania substrates. Other materials used for Au-metaloxide nanohybrids are CuO<sup>49</sup> and ZnO.<sup>50</sup>

In the present work, we have applied STM and AFM to determine the size and the distribution of chemically-synthesised, atomically precise Au<sub>9</sub> clusters deposited onto titania nanosheets. XPS is used to monitor the change in the size of Au<sub>9</sub> clusters before and after annealing. The aim of the work is to determine the degree of aggregation of the Au<sub>9</sub> clusters after deposition and after annealing of the clusters.

## Experimental

The Au<sub>9</sub> clusters used in the present work were synthesized according to a well-established, previously reported method.<sup>51</sup> The structure of Au<sub>9</sub> clusters was confirmed by nuclear magnetic resonance spectroscopy (NMR), mass spectrometry (MS) and X-ray crystallography. Further, electron diffraction using TEM has been applied previously.<sup>52</sup> The Au<sub>9</sub> clusters were deposited onto titania (Ti<sub>0.87</sub>O<sub>2</sub>) nanosheets that have a thickness of 1.1 nm (ref. 53 and 54) and were prepared by delamination from a parent layered K<sub>0.8</sub>Ti<sub>1.73</sub>Li<sub>0.27</sub>O<sub>4</sub> crystal using a soft chemical procedure.<sup>55,56</sup> The thickness of the nanosheets is determined by the crystal structure of the parent crystal and

thus intrinsic to the compound.<sup>52</sup> Further information about the thickness of the nanosheets is provided in the ESI.†

For sample preparation, the titania-nanosheets were deposited on a cleaned 10 × 10 mm<sup>2</sup> silicon wafer and exposed to ultraviolet (UV) light irradiation in air to photocatalytically decompose the tetrabutylammonium ions surrounding the nanosheets in a process as described in.<sup>57</sup> Finally, the silicon/titania nanosheet samples were immersed into a 0.02 mM Au<sub>9</sub> methanoic (ultrapure grade) solution for 30 minutes, followed by brief rinsing in ultrapure methanol (Methanol Infinity Pure, Wako company-Japan) with subsequent drying in a room-temperature stream of pure N<sub>2</sub> (generated from liquid N<sub>2</sub>). Annealing of the samples studied in this paper was carried out at 200 °C for 20 minutes under high vacuum. Annealing was applied in order to remove the PPh<sub>3</sub> ligands and was performed in a very similar way to that used in our previous work.<sup>21,22</sup> The concentration of 0.02 mM has been chosen based on earlier published results.<sup>10,12</sup> The thickness of the nanosheets has been checked with AFM and STEM and is described in the ESI.†

The AFM images were acquired in AC tapping mode on an Asylum MFP3D SPM (Asylum Research, USA) at room temperature in air. Tapping mode can reduce the lateral forces applied to the sample and thus results in a lower degree of damage to samples than contact mode. Au<sub>9</sub> clusters adsorbed from solutions were found to be stable against the tapping AFM mode. AFM cantilevers with a suitable spring constant were chosen to provide stable images under tapping mode but also so that the tip did not move the Au<sub>9</sub> clusters during the imaging process. Olympus OMCL-AC160TS-C<sub>3</sub> probes with resonant frequency of 300 kHz (spring constant 26.2 N m<sup>−1</sup>) were used for this work.

We also performed STM measurements on the same samples previously investigated with AFM. STM observations were carried out with a constant current mode at room temperature under UHV conditions (~10<sup>−9</sup> Pa) with electrochemically etched tungsten tips. STM could stably image the sample surface with conditions of a tunneling current of 25 pA and sample bias voltages (Vs) of 1.5–6 V. Samples were also investigated with STM after heating the sample to 200 °C for 20 minutes after at least 3 to 4 hours under UHV. It should be noted that individual clusters could be moved when applying the STM scanning procedure, which could result in either removal of clusters from the scanned area or lead to drift of clusters in an image.

In order to estimate the probability of tunneling of electrons between the STM tip and the substrate, DFT calculations were employed to calculate the energy levels of the molecular orbitals in the Au<sub>9</sub> cluster. Detailed information about the method employed for the DFT calculations can be found elsewhere.<sup>12</sup>

The photoelectron spectra of Au<sub>9</sub> clusters as deposited on titania nanosheets and after annealing at 200 °C were recorded at the Soft X-ray Beamline at the Australian Synchrotron (AS) by using a SPECS Phoibos 150 hemispherical electron analyser. The photon energy used was 690 eV. The spot size of the irradiation beam was adjusted to ~600 × 600 μm<sup>2</sup> with an X-ray photon flux of approximately 10<sup>12</sup> photons mm<sup>−2</sup> s<sup>−1</sup>. These conditions have been previously shown not to induce damage to samples as those investigated here.<sup>58</sup> High resolution XP spectra of C, O, Si, P, Ti and Au were recorded at a pass energy of 10 eV



with an instrumental resolution of 295 meV.<sup>59</sup> The stability of the X-ray energy was monitored using a bulk Au reference and the scans were repeated several times to ensure that the X-ray irradiation did not affect the samples. For all XP spectra, a Shirley background was applied to remove the electron-scattering background. A sum of Gaussian (30%) and Lorentzian (70%) functions was used to fit all peaks including the peak positions as described previously.<sup>22</sup>

## Results and discussion

AFM study of the height distribution of Au<sub>9</sub> clusters on titania nanosheet.

Fig. 1a–c show the AFM 3D images of the titania nanosheet before and after adsorption of Au<sub>9</sub> clusters and also after annealing, respectively. These 3D images correspond to the magnified area from topographical images shown in Fig. S3.† Fig. 1d–f show AFM 3D images of the titania nanosheets before and after exposure to methanol only, and after annealing, respectively. Methanol was used as solvent in the deposition of the Au<sub>9</sub> clusters and so these images provide a “solvent only” reference of the effect of processing conditions in the absence of the Au<sub>9</sub> clusters. The “solvent only” reference images show that the roughness of the surface increases only marginally due to the exposure to the pure solvent. It can be noted that the random distribution of spikes observed in Fig. 1b and c are not seen in the image of the samples that were exposed to solvent only (Fig. 1e and f). Therefore, the spikes in Fig. 1b and c must be attributed to Au<sub>9</sub> clusters.

AFM imaging allows for height analysis of the Au NCs. Due to the size of the curvature of the AFM tip and the range of the interaction force between the surface and the tip it is not possible to resolve individual Au<sub>9</sub> clusters *i.e.* individual features with a lateral size smaller than the resolution of AFM

will appear almost the same in the image. The measurement of the height of the Au<sub>9</sub> clusters, however, is not affected by the curvature of the tip and can be evaluated. The resulting course of the AFM tip is illustrated in Fig. 2. In Fig. 2a, various possible model configurations of the Au<sub>9</sub> clusters on the titania surface are sketched. On the left a group of Au<sub>9</sub> clusters in close proximity are shown to form a monolayer of clusters. In the centre image a group of Au<sub>9</sub> clusters in close proximity are shown forming a double layer of clusters; such layering is certainly not limited to only double layers. However, it needs to be emphasised that the results shown in the present work do not provide experimental evidence for the formation of double or multi-layers. The possibility of the formation of a double layer is shown in Fig. 2 only because such layer formation cannot be excluded and any layer below the top layer would have a structure that could not be determined with the methods used in the present work. On the right, a single cluster on the titania surface is illustrated. The AFM tip scanning over each of the three groups is indicated by the green line and only the height of the groups or the single cluster can be identified. Thus the left and the right case will appear to be similar in the AFM image. The possible scenarios after annealing are sketched in Fig. 2b. The Au<sub>9</sub> clusters have lost ligands and form groups or individual clusters containing only the cluster core. Similar to the case of as deposited clusters (*i.e.* no heating), the AFM tip can only identify the height of Au species (medium and larger aggregates from NCs), but could not distinguish groups of several closely located clusters from the individual clusters.

Cluster heights were evaluated from the AFM images using the particle analysis tool of the Asylum software. For evaluating the AFM images, a large number (several hundreds) of features in the AFM images were analysed and the height of the features is shown as a statistical distribution. The roughness of the substrate is taken into account for this procedure in order to

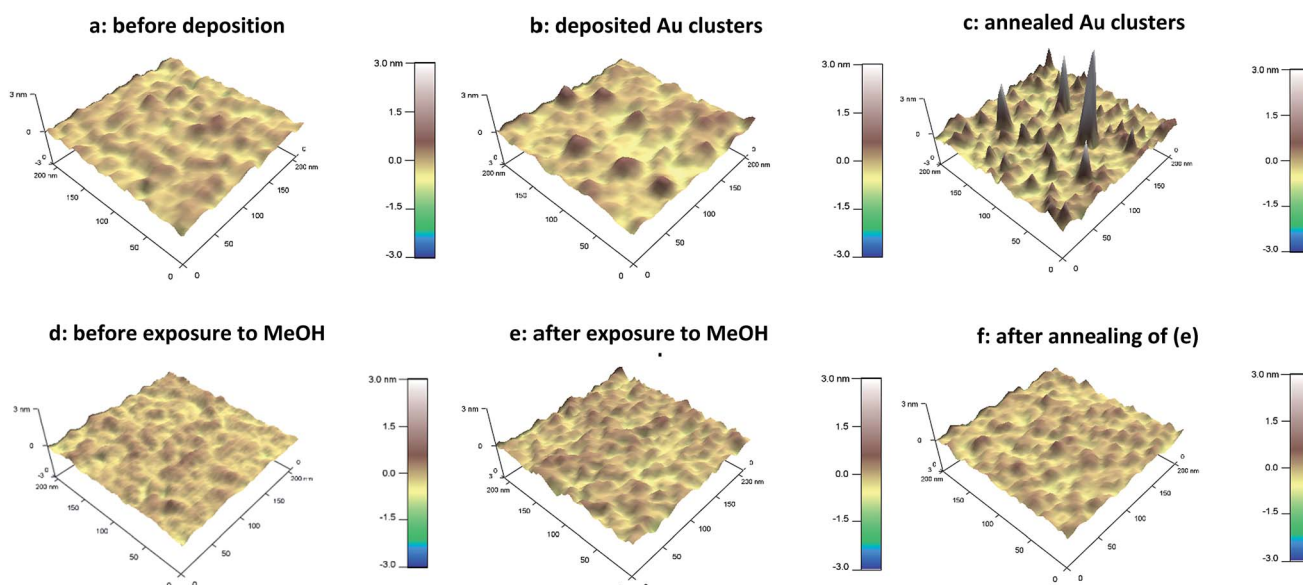


Fig. 1 AFM 3D images of (a and d) pure titania nanosheets, (b) after deposition of Au<sub>9</sub> clusters, (c) after annealing of (b), (e) titania nanosheet exposed to MeOH, and (f) after annealing of (e). Image sizes are 200 × 200 nm<sup>2</sup>.



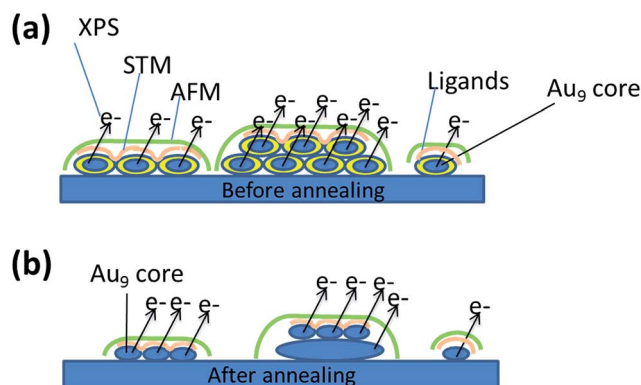


Fig. 2 Schematic illustration of gold clusters deposited on titania nanosheet: (a) before annealing and (b) after annealing. The green and orange lines represent the results that can be obtained from AFM and STM, respectively. The lines indicating the emission of electrons represent the XPS results.

differentiate features related to  $\text{Au}_9$  clusters. It was found that the vast majority of features in the AFM images of bare titania nanosheets with and without exposure to the solvent are less than 0.5 nm in height. Thus, all features in the AFM images of titania nanosheets with adsorbed  $\text{Au}_9$  clusters that were less than 0.5 nm in height were not considered in the statistical analysis.  $\text{Au}_9$  clusters are larger than 0.5 nm in diameter<sup>51</sup> and thus the discrimination of 0.5 nm in the height analysis does not exclude  $\text{Au}_9$  clusters in the evaluation procedure. In Fig. 3a and b, the height distributions of the features identified as  $\text{Au}_9$  clusters on titania nanosheets before and after annealing are shown. The fraction of the features found for a given height interval are calculated by

$$\text{Fraction}_i = \frac{N_i}{\sum_i N_i}$$

where  $\text{fraction}_i$  is the fraction of features found in a given height interval and  $N_i$  the number of features found in a given height interval. Both graphs show one clear peak each in the height distribution. Both distributions are fitted with log-normal distributions using the Levenberg–Marquardt algorithm. Previous studies have shown that the particle size distributions can be well approximated by a log-normal distribution.<sup>60–62</sup> The fitted distributions show a mean height of  $0.75 \pm 0.2$  nm for the non-annealed samples and  $1.0 \pm 0.4$  nm for samples that have been subjected to annealing. The distribution of the features on the annealed samples becomes broader and the maximum has shifted to a larger value. For interpretation of these results, a few aspects must first be considered. Based on the DFT calculation of isolated  $\text{Au}_9$  clusters, the overall shape of the  $\text{Au}_9$  cluster is best described as oblate spheroid and the mean axes of the cluster, including the  $\text{PPh}_3$  ligands, have lengths of 1.75, 1.66 and 1.47 nm.<sup>11</sup> It is likely that the clusters deposit on the surface such that the longest mean axis (1.75 nm) orients parallel to the surface because this would lead to the closest contact and thus strongest interaction between the cluster and the substrate. Experimentally, the average height of the  $\text{Au}_9$  clusters of the non-

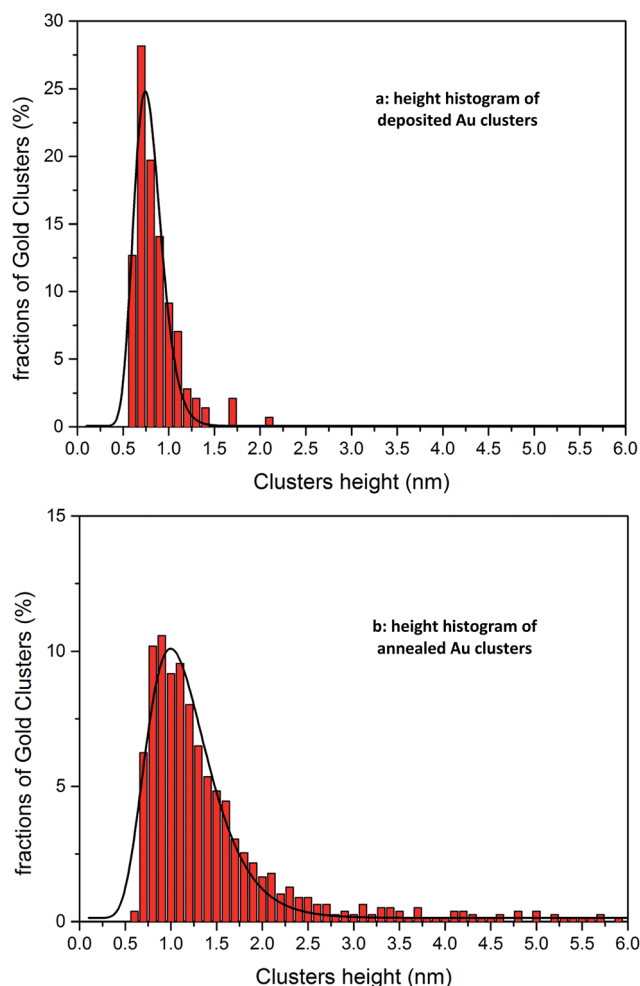


Fig. 3 Height distribution of features in the AFM images: (a) as deposited Au clusters, and (b) after annealing to 200 °C. The cluster mean height before annealing is  $0.75 \pm 0.2$  nm while after annealing it is  $1.0 \pm 0.4$  nm.

annealed samples is smaller than the calculated height. The reason for this discrepancy could be due to the deformation of the  $\text{Au}_9$  clusters when adsorbing onto the surface due to the interaction between the clusters and the surface. Such interaction would lead to a closer contact, which could reduce the height of the adsorbed cluster. A second reason could be that the  $\text{Au}_9$  clusters lose some ligands during the deposition process and therefore decrease their average size.  $\text{Au}_9$  clusters after ligand removal show a calculated size of about 0.5 nm which is twice the average distance between two Au atoms forming the cluster.<sup>63</sup> The DFT calculations also do not take into account that the  $\text{Au}_9$  clusters could change their shape after deposition. After annealing, AFM images show that the small clusters are still present on the surface but their number has decreased. The percentage of small clusters has dropped to about half of the fraction of the non-annealed sample and the fraction of features with larger heights has increased. The change in distribution is interpreted as agglomeration of the smaller clusters to form larger clusters or particles. The reason for this observation is likely that the cluster cores have aggregated or that Ostwald ripening has occurred.<sup>64</sup>





### STM observation and width distribution of Au<sub>9</sub> clusters on titania nanosheet

Fig. 4a and c show STM images of titania nanosheets with deposited Au<sub>9</sub> clusters before heating. The STM images of blank samples before and after annealing (UV-irradiated nanosheets and also MeOH-treated nanosheets on the Si wafers) are shown in the ESI (Fig. S5†). A black arrow in Fig. 4c indicates an example of hemispherically shaped feature, which were not observed in the blank sample before heating (Fig. S5c†). The statistical distribution of the height and width of the features in the STM images are shown in Fig. 5. In the statistical analysis we only considered features which are approximately symmetric spots (typical examples are shown as open blue circles in Fig. 4) to measure the width and height, and excluded asymmetric spots because these could be drifting clusters. The width distribution of the Au NCs before heating varies from 2 nm to 3 nm, with the average found at  $2.4 \pm 0.4$  nm (Fig. 5a). The height distribution varies from 0.2 nm to 0.7 nm with the average found at  $0.5 \pm 0.2$  nm (Fig. 5c). Based on the width of the bright hemispherical features shown in Fig. 4c, we can attribute these to individual Au<sub>9</sub> clusters. The average height of the features, however, is smaller than the theoretical height of the Au<sub>9</sub> clusters. We assume that the reason for this finding is that the STM tip during the scanning process does not get into

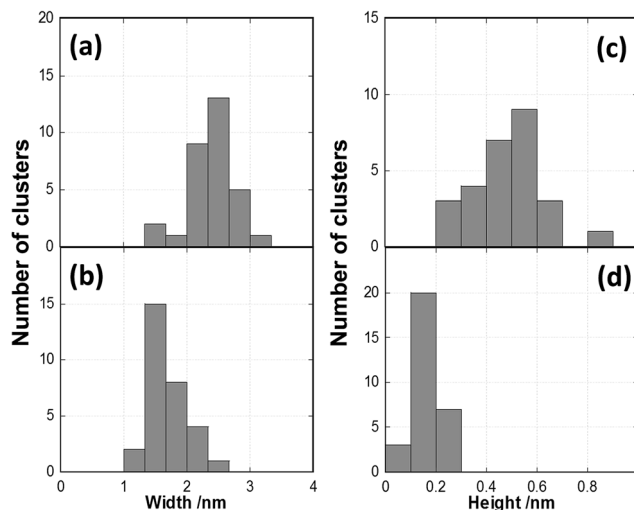


Fig. 5 Histograms of the features in the STM images attributed to the Au clusters. The width is shown in (a) for as-deposited samples and (b) after annealing. The height is shown in (c) for as-deposited samples and (d) after annealing.

contact with the nanosheet layer underlying the Au<sub>9</sub> clusters and thus the height variation found in the STM images is less than the height of individual clusters. The interpretation of the

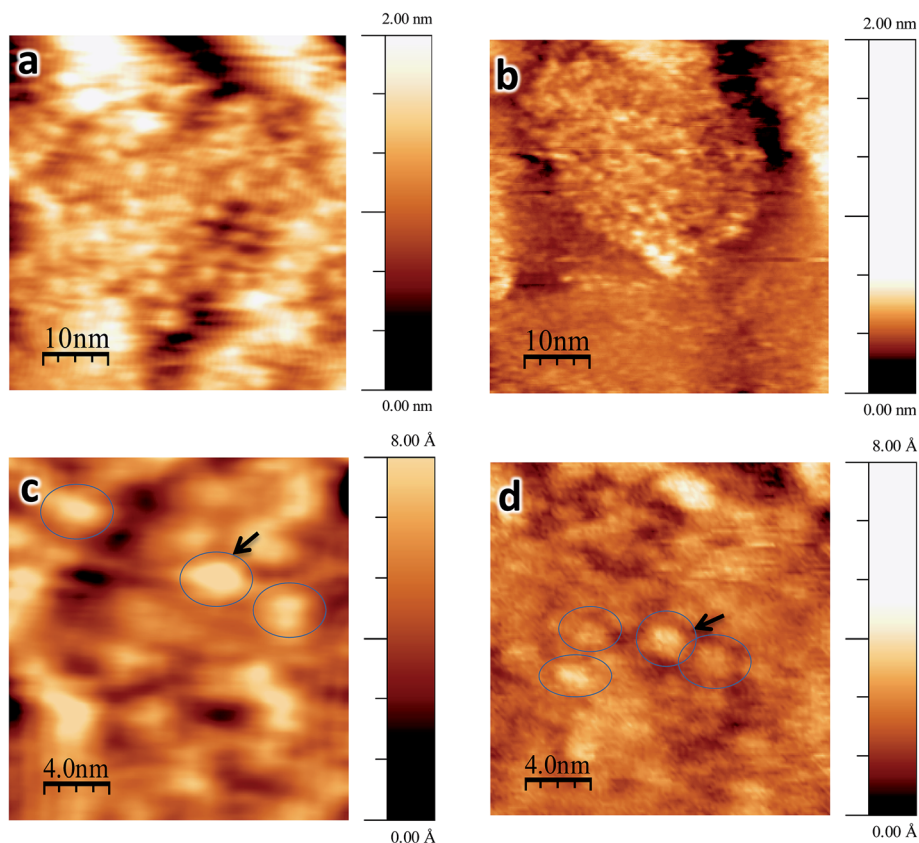


Fig. 4 STM images of Au<sub>9</sub> clusters deposited on titania nanosheets: (a) as-deposited, (b) after annealing to 200 °C. Images (c) and (d) are magnifications of (a) and (b). The image size is  $50 \times 50$  nm<sup>2</sup> for (a) and (b) and  $20 \times 20$  nm<sup>2</sup> for (c) and (d). The STM operation conditions were (a) +1.5 V, 20 pA; (b) +2.0 V, 18 pA; (c) +1.5 V, 20 pA; (d) +2.5 V, 25 pA. Heights and widths of the clusters which are surrounded by open blue circles were used for the histogram analysis (Fig. 5).



STM images is given in Fig. 2a where the orange line illustrates that the STM tip does not reach the titania nanosheet substrate. However, the lateral resolution is better than in AFM and individual Au<sub>9</sub> clusters can be identified.

After annealing at 200 °C for 20 minutes under UHV, STM measurements were repeated for the same sample. Fig. 4b and d show STM images of annealed Au clusters. The brighter spots can be identified in the same way as in the case of non-annealed samples (Fig. 4a and c); note that the bright spots are not observed in the heated blank sample (Fig. S5d†). A black arrow in Fig. 4d indicates one of the individual Au NCs. In Fig. 5b and d the width and height distributions of the features after annealing are shown. In comparison with the non-annealed samples, the width distribution has decreased to 1.3–2.3 nm with an average of  $1.7 \pm 0.3$  nm, and the height distribution to 0.1–0.2 nm with the average  $0.17 \pm 0.05$  nm. The significant decrease in cluster width is compatible with the evidence from the XPS data shown below that the ligands have been removed through the annealing process. The height, however, is less than what is expected from the cluster core. However, it is important to mention that the 2D cluster reported in our previous paper should have a height around 0.2 nm,<sup>12</sup> which is in good agreement with the measured height in this work. The scanning process of the STM tip is illustrated in Fig. 2b. As explained for the non-heated samples, the STM tip does not interact with the layer underlying the top of the Au<sub>9</sub> NCs and thus the height variation is less than the height of an individual cluster core.

Fig. 6a shows the energy levels of ligated Au<sub>9</sub> before heating while Fig. 6b shows the Au<sub>9</sub> cluster core without ligands (corresponding to Au<sub>9</sub> clusters after removing the ligands through heating). These isomers have been observed as preferable configurations for deposited Au<sub>9</sub> clusters on titania nanosheet.<sup>12</sup> The molecular orbital visualisations in Fig. 6a show that

the ligands are not electronically transparent to STM at the chosen bias voltages, thus electrons could tunnel into the respective energy levels. For estimating whether electrons could tunnel into specific orbitals of the clusters, two considerations have to be taken into account; (i) that the Fermi level is the reference potential when applying the bias voltage in STM and that it lies approximately half-way between the occupied and unoccupied orbital and (ii) that the energy of the molecular orbitals of a cluster are broadened owing to the interaction with the adjacent cluster or with the substrate. At room temperature, the molecular orbitals within an energy range of 0.5 to 1 eV overlap due to the energetic broadening of each orbital. Therefore, we consider in Fig. 6 that not only is a single molecular orbital involved (*i.e.* the LUMO) but that combinations of orbitals can be accessed (*i.e.* LUMO, LUMO+1 and LUMO+2). It will be shown below that in all STM images the Au clusters appear approximately hemispherically shaped both before and after heating. This implies that electrons are tunneling into a range of orbitals and not into one specific orbital.

### XPS results of Au<sub>9</sub> clusters on titania nanosheet

High resolution XP spectra of the Au clusters deposited onto titania nanosheet have been measured for the Au 4f, P 2p, Si 2p, Ti 2p, C 1s and O 1s regions. The main C 1s peak is assigned to 285 eV and used to calibrate peak position for all other elements. The Au shows the Au 4f doublet with the Au 4f<sub>7/2</sub> peak found at 84.8 eV (Fig. 7a) for the non-heated sample, which is close to the binding energy previously found for as-deposited, non-agglomerated Au<sub>9</sub> clusters.<sup>9,21,22</sup> No signal is observed at  $84 \pm 0.2$  eV, the binding energy of bulk Au. The shift from bulk gold for the Au 4f<sub>7/2</sub> peak position of isolated Au<sub>9</sub> clusters is attributed to the final state effect as has been described in detail

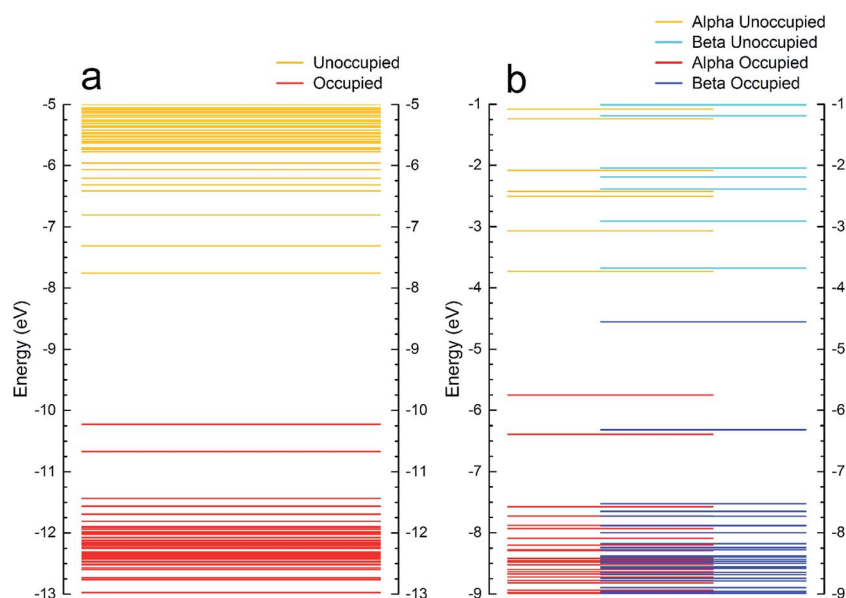


Fig. 6 The energies levels of (a) ligands Au<sub>9</sub> before heating and (b) the naked gold clusters after heating. The molecular orbital visualisations of combined molecular orbitals such as (LUMO, LUMO+1) and (LUMO, LUMO+1, LUMO+2) are also shown in this figure.



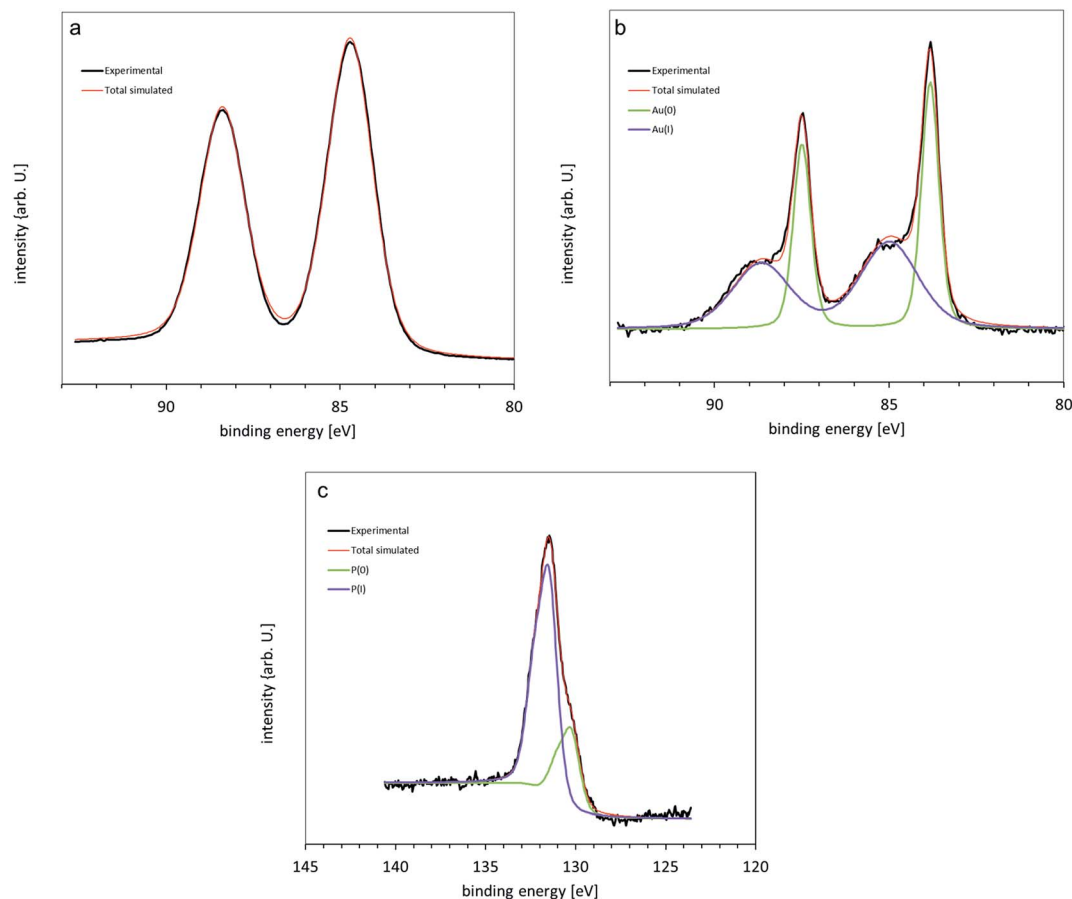


Fig. 7 XPS of  $[\text{Au}_9(\text{PPh}_3)_8](\text{NO}_3)_3$  clusters on titania nanosheet (a) 4f Au spectrum of as-deposited samples, (b) after annealing and (c) the 2p P peak of as-deposited samples.

in our previous work.<sup>21,22</sup> The Au XP spectrum of the annealed titania nanosheet/ $\text{Au}_9$  sample is shown in Fig. 7b and shows two Au 4f doublets. A narrow peak is found at  $83.8 \pm 0.2$  eV and is consistent with the value for the binding energy for bulk Au. This is attributed to aggregated nanoparticles much larger than  $\text{Au}_9$  clusters that arise from the heating procedure.<sup>22</sup> The second, broader Au 4f<sub>7/2</sub> peak is found at higher binding energy,  $85.3 \pm 0.2$  eV. This peak is attributed to non-agglomerated  $\text{Au}_9$  clusters that have lost ligands and are in close, electronic contact with the titania substrate.<sup>22,65</sup> The P 2p<sub>3/2</sub> peak is found at  $131.6 \pm 0.2$  eV and is attributed to ligands attached to the Au NCs (Fig. 7c).<sup>21,22</sup> After annealing, no P signal was found which means that the ligands have been detached from the  $\text{Au}_9$  cluster core through the heating procedure and removed from the sample under UHV. Titanium signal 2p<sub>3/2</sub> is found at 459 eV for all samples (Fig. S6†). There is a small shoulder at 457.5 eV that can be seen in the Ti spectra, owing to the presence of reduced titania surfaces ( $\text{Ti}^{3+}$ ) as an effect of the UV or X-ray irradiation, inducing O vacancy in the titania nanosheet.<sup>8,43,66</sup> After vacuum annealing the  $\text{Ti}^{3+}$  peak increases and is due to the increase in the number of the defects.<sup>8</sup>

The interpretation of the XPS results for the non-annealed samples is illustrated in Fig. 2. It is important to note that the electron mean free path for the samples investigated here is

about 1.5 nm and thus XPS is sensitive for a depth around 4 to 5 nm. Irrespective of whether the clusters form larger groups of  $\text{Au}_9$  clusters or are isolated on the surface, each individual cluster has its own cluster core and the emission of electrons from a given cluster is influenced by the final state effect as illustrated in Fig. 2a. Thus all  $\text{Au}_9$  clusters appear as individual and isolated clusters irrespective whether another  $\text{Au}_9$  cluster is located in close proximity or not. After removing the ligands through annealing, the clusters form groups in which the cluster cores have either merged into a larger particle, similar to the larger block in the middle of Fig. 2b, or they form groups in which the individual clusters are in close proximity but the cluster cores are retained as individual clusters. The latter option is illustrated on the right and left in Fig. 2b, and the top layer in the middle of the figure.

### Integrating AFM, STM and XPS results

AFM, STM and XPS measurements reveal complementary information about the properties of the  $\text{Au}_9$  clusters deposited onto titania nanosheet and can be integrated into a single interpretation. The interpretation is schematically illustrated in Fig. 2. The AFM measurements allow height analysis of the features attributed to  $\text{Au}_9$  clusters. Before heating, the majority of features found in the AFM images show a height



corresponding to the smallest diameter of individual Au<sub>9</sub> clusters. Identifying individual Au NCs is not possible with AFM due to the large tip curvature and the large interaction range between the AFM tip and the sample. In contrast, STM is used to investigate small areas with the size of a few nanometers with high lateral resolution sufficient to resolve individual Au NCs. It needs to be taken into account that individual clusters outside groups were not seen in STM most likely because they were moved out of the image area when scanning the STM tip across the desired area. The STM images identify individual Au<sub>9</sub> clusters even when they form groups of clusters within close proximity to each other. In contrast to AFM, it is very difficult to determine the total height of a group of clusters with STM. Based on the final state effect found in the XPS results, it can be concluded that the Au<sub>9</sub> clusters are individual and separate clusters.

After annealing of the samples, the Au<sub>9</sub> clusters have lost the PPh<sub>3</sub> ligands. The cluster cores could either retain their identity as Au<sub>9</sub> cluster cores or merge into larger particles, as schematically illustrated in Fig. 2b. AFM measurements show that partial agglomeration has occurred leading to an increase of the heights of the features found. However, a significant number of features have retained the height of single clusters. The STM results cannot be used to derive information about the overall size of the groups formed by the clusters but can shed light on the fine structure of the top layer of a group. The STM results allow identifying individual clusters but do not allow us to draw conclusions about whether these individual clusters in the top layer have merged electronically or retained their individual structure. XPS shows that at least a fraction of the Au<sub>9</sub> clusters have not merged to form a bulk electronic structure and have retained their individual electronic structure. This conclusion is based on the XPS data, which shows species with Au 4f<sub>7/2</sub> electron binding energy of  $85.3 \pm 0.2$  eV that can only be due to small clusters. STM also cannot be used to reveal information about the layer below the top Au cluster layer. The layer below could be the substrate, or it could be a layer of agglomerated Au clusters as sketched in the middle section of Fig. 2b, or it could be composed of the individual Au clusters. The layer below the top layer could be formed by agglomerated clusters, which could explain the XPS Au 4f<sub>7/2</sub> peak found at  $83.8 \pm 0.2$  eV which is assigned to larger bulk Au particles. In this context, it is important to note that the depth for which XPS is sensitive is much larger than in the case of STM. Summarising the STM, AFM and XPS data it can be concluded, that agglomeration does not necessarily include all clusters within a group. At least in the top layer of these groups, individual clusters could be identified that most likely have retained the electronic structure of individual clusters.

In a previous study, we investigated the deposition of Au<sub>9</sub> clusters onto plasma treated ALD titania.<sup>10</sup> The coverage of ALD titania with Au in the earlier study was approximately 40 times lower compared to the present case (the Au/Ti ratio was less than 0.1 in ref. 10 and it is 3.5 here). We assume that the higher coverage of the titania nanosheet surface with non-agglomerated Au<sub>9</sub> clusters is due to the larger number of defects on the titania nanosheet surface compared to the surface of ALD titania.

## Conclusions

Au<sub>9</sub> clusters deposited on titania nanosheets have been investigated with AFM, STM and XPS. Each method has allowed the determination of specific information about the structure of the Au<sub>9</sub> clusters before and after annealing which have been integrated into an overall interpretation of the Au<sub>9</sub>/titania surface. After deposition, the Au<sub>9</sub> clusters attached to the surface with a range of configurations from individual clusters to groups of clusters. All clusters have retained their individual electronic structure. Heating of the samples to 200 °C leads to partial agglomeration of the Au NCs that lead to the formation of larger Au particles. Some clusters, however, are not involved in agglomeration and are left as individual isolated clusters on the surface. The electronic structure of the agglomerated Au particles is different to that of the isolated individual clusters. Agglomeration does not necessarily include all clusters within a group. At least in the top layer of these groups, individual clusters could be identified that most likely have retained the electronic structure of individual clusters.

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## References

- 1 T. Imaoka, H. Kitazawa, W. J. Chun, S. Omura, K. Albrecht and K. Yamamoto, *J. Am. Chem. Soc.*, 2013, **135**, 13089–13095.
- 2 F. Baletto and R. Ferrando, *Rev. Mod. Phys.*, 2005, **77**, 371–423.
- 3 W. A. de Heer, *Rev. Mod. Phys.*, 1993, **65**, 611–676.
- 4 G. Schmid, M. Baumle, M. Geerkens, I. Heim, C. Osemann and T. Sawitowski, *Chem. Soc. Rev.*, 1999, **28**, 179–185.
- 5 L. N. Lewis, *Chem. Rev.*, 1993, **93**, 2693–2730.
- 6 M.-C. Daniel and D. Astruc, *Chem. Rev.*, 2004, **104**, 293–346.
- 7 R. Ferrando, J. Jellinek and R. L. Johnston, *Chem. Rev.*, 2008, **108**, 845–910.
- 8 T. Bennett, R. H. Adnan, J. F. Alvino, R. Kler, V. B. Golovko, G. F. Metha and G. G. Andersson, *J. Phys. Chem. C*, 2015, **119**, 11171–11177.
- 9 J.-Y. Ruzicka, F. Abu Bakar, C. Hoeck, R. Adnan, C. McNicoll, T. Kemmitt, B. C. Cowie, G. F. Metha, G. G. Andersson and V. B. Golovko, *J. Phys. Chem. C*, 2015, **119**, 24465–24474.
- 10 G. G. Andersson, V. B. Golovko, J. F. Alvino, T. Bennett, O. Wrede, S. M. Mejia, H. S. Al Qahtani, R. Adnan,





- N. Gunby, D. P. Anderson and G. F. Metha, *J. Chem. Phys.*, 2014, **141**, 014702.
- 11 J. F. Alvino, T. Bennett, D. Anderson, B. Donoeva, D. Ovoshchnikov, R. H. Adnan, D. Appadoo, V. Golovko, G. Andersson and G. F. Metha, *RSC Adv.*, 2013, **3**, 22140–22149.
  - 12 H. S. Al Qahtani, K. Kimoto, T. Bennett, J. F. Alvino, G. G. Andersson, G. F. Metha, V. B. Golovko, T. Sasaki and T. Nakayama, *J. Chem. Phys.*, 2016, **144**, 114703.
  - 13 Z. W. Wang, O. Toikkanen, B. M. Quinn and R. E. Palmer, *Small*, 2011, **7**, 1542–1545.
  - 14 K. J. Batenburg, R. Erni, M. D. Rossell, S. Van Aert and G. Van Tendeloo, *Nature*, 2011, **470**, 374–377.
  - 15 B. Yoon, H. Häkkinen, U. Landman, A. S. Wörz, J.-M. Antonietti, S. Abbet, K. Judai and U. Heiz, *Science*, 2005, **307**, 403–407.
  - 16 D. Lee, R. L. Donkers, G. Wang, A. S. Harper and R. W. Murray, *JACS*, 2004, **126**, 6193–6199.
  - 17 J. Jung, H. Kim and Y.-K. Han, *JACS*, 2011, **133**, 6090–6095.
  - 18 G. Ramakrishna, O. Varnavski, J. Kim, D. Lee and T. Goodson, *JACS*, 2008, **130**, 5032–5033.
  - 19 M. Valden, X. Lai and D. W. Goodman, *Science*, 1998, **281**, 1647–1650.
  - 20 S. Kunz, K. Hartl, M. Nesselberger, F. F. Schweinberger, G. Kwon, M. Hanzlik, K. J. J. Mayrhofer, U. Heiz and M. Arenz, *PCCP*, 2010, **12**, 10288–10291.
  - 21 D. P. Anderson, R. H. Adnan, J. F. Alvino, O. Shipper, B. Donoeva, J.-Y. Ruzicka, H. Al Qahtani, H. H. Harris, B. Cowie, J. B. Aitken, V. B. Golovko, G. F. Metha and G. G. Andersson, *Phys. Chem. Chem. Phys.*, 2013, **15**, 14806–14813.
  - 22 D. P. Anderson, J. F. Alvino, A. Gentleman, H. A. Qahtani, L. Thomsen, M. I. J. Polson, G. F. Metha, V. B. Golovko and G. G. Andersson, *Phys. Chem. Chem. Phys.*, 2013, **15**, 3917–3929.
  - 23 M. Haruta, *Catal. Today*, 1997, **36**, 153–166.
  - 24 K. Judai, S. Abbet, A. S. Wörz, U. Heiz and C. R. Henry, *JACS*, 2004, **126**, 2732–2737.
  - 25 M. S. Chen and D. W. Goodman, *Catal. Today*, 2006, **111**, 22–33.
  - 26 P. Pyykkö, *Angew. Chem., Int. Ed.*, 2004, **43**, 4412–4456.
  - 27 X.-G. Li, X. G. Zhang and H.-P. Cheng, *Nano Lett.*, 2014, **14**, 4476–4479.
  - 28 Y. Zhu, H. Qian and R. Jin, *J. Mater. Chem.*, 2011, **21**, 6793–6799.
  - 29 Y. Zhu, H. Qian, B. A. Drake and R. Jin, *Angew. Chem.*, 2010, **122**, 1317–1320.
  - 30 Y. Liu, H. Tsunoyama, T. Akita, S. Xie and T. Tsukuda, *ACS Catal.*, 2011, **1**, 2–6.
  - 31 M. Haruta, S. Tsubota, T. Kobayashi, H. Kageyama, M. J. Genet and B. Delmon, *J. Catal.*, 1993, **144**, 175–192.
  - 32 R. Meyer, C. Lemire, S. K. Shaikhutdinov and H.-J. Freund, *Gold Bull.*, 2004, **37**, 72–124.
  - 33 R. H. Adnan, G. G. Andersson, M. I. J. Polson, G. F. Metha and V. B. Golovko, *Catal. Sci. Technol.*, 2015, **5**, 1323–1333.
  - 34 P. Rodríguez-Zamora, F. Yin and R. E. Palmer, *J. Phys. Chem. A*, 2014, **118**, 8182–8187.
  - 35 C. C. Chusuei, X. Lai, K. A. Davis, E. K. Bowers, J. P. Fackler and D. W. Goodman, *Langmuir*, 2001, **17**, 4113–4117.
  - 36 A. Kolmakov and D. W. Goodman, *Surf. Sci.*, 2001, **490**, L597–L601.
  - 37 N. Spiridis, J. Haber and J. Korecki, *Vacuum*, 2001, **63**, 99–105.
  - 38 X. Tong, L. Benz, P. Kemper, H. Metiu, M. T. Bowers and S. K. Buratto, *JACS*, 2005, **127**, 13516–13518.
  - 39 E. Wahlström, N. Lopez, R. Schaub, P. Thosttrup, A. Rønnau, C. Africh, E. Lægsgaard, J. K. Nørskov and F. Besenbacher, *Phys. Rev. Lett.*, 2003, **90**, 026101.
  - 40 P. E. Batson, *Microsc. Microanal.*, 2008, **14**, 89–97.
  - 41 Z. Y. Li, N. P. Young, M. Di Vece, S. Palomba, R. E. Palmer, A. L. Bleloch, B. C. Curley, R. L. Johnston, J. Jiang and J. Yuan, *Nature*, 2008, **451**, 46–48.
  - 42 S. Porsgaard, P. Jiang, F. Borondics, S. Wendt, Z. Liu, H. Bluhm, F. Besenbacher and M. Salmeron, *Angew. Chem., Int. Ed.*, 2011, **50**, 2266–2269.
  - 43 M. Chen and D. W. Goodman, *Chem. Soc. Rev.*, 2008, **37**, 1860–1870.
  - 44 A. S. Wörz, U. Heiz, F. Cinquini and G. Pacchioni, *J. Phys. Chem. B*, 2005, **109**, 18418–18426.
  - 45 U. Diebold, *Surf. Sci. Rep.*, 2003, **48**, 53–229.
  - 46 A. Fujishima, X. Zhang and D. A. Tryk, *Surf. Sci. Rep.*, 2008, **63**, 515–582.
  - 47 H.-J. Freund and G. Pacchioni, *Chem. Soc. Rev.*, 2008, **37**, 2224–2242.
  - 48 M. Osada and T. Sasaki, *Adv. Mater.*, 2012, **24**, 210–228.
  - 49 X. Zhang, Y. Yang, W. Que and Y. Du, *RSC Adv.*, 2016, **6**, 81607–81613.
  - 50 N. Guo, R. Lu, S. Liu, G. W. Ho and C. Zhang, *J. Phys. Chem. C*, 2014, **118**, 21038–21041.
  - 51 F. Wen, U. Englert, B. Guttrath and U. Simon, *Eur. J. Inorg. Chem.*, 2008, **2008**, 106–111.
  - 52 T. Sasaki, Y. Ebina, Y. Kitami, M. Watanabe and T. Oikawa, *J. Phys. Chem. B*, 2001, **105**, 6116–6121.
  - 53 T. Shibata, N. Sakai, K. Fukuda, Y. Ebina and T. Sasaki, *PCCP*, 2007, **9**, 2413–2420.
  - 54 T. Sasaki, M. Watanabe, H. Hashizume, H. Yamada and H. Nakazawa, *JACS*, 1996, **118**, 8329–8335.
  - 55 T. Sasaki and M. Watanabe, *JACS*, 1998, **120**, 4682–4689.
  - 56 T. Sasaki, S. Nakano, S. Yamauchi and M. Watanabe, *Chem. Mater.*, 1997, **9**, 602–608.
  - 57 T. Sasaki, Y. Ebina, K. Fukuda, T. Tanaka, M. Harada and M. Watanabe, *Chem. Mater.*, 2002, **14**, 3524–3530.
  - 58 Y.-Y. Fong, B. R. Visser, J. R. Gascooke, B. C. C. Cowie, L. Thomsen, G. F. Metha, M. A. Buntine and H. H. Harris, *Langmuir*, 2011, **27**, 8099–8104.
  - 59 B. C. C. Cowie, A. Tadich and L. Thomsen, *AIP Conf. Proc.*, 2010, **1234**, 307–310.
  - 60 J. Söderlund, L. B. Kiss, G. A. Niklasson and C. G. Granqvist, *Phys. Rev. Lett.*, 1998, **80**, 2386–2388.
  - 61 C. G. Granqvist and R. A. Buhrman, *J. Appl. Phys.*, 1976, **47**, 2200–2219.
  - 62 D. Susan-Resiga, V. Socoliuc, T. Boros, T. Borbáth, O. Marinica, A. Han and L. Vékás, *J. Colloid Interface Sci.*, 2012, **373**, 110–115.



- 63 C. Bosch-Navarro, Z. P. L. Laker, H. R. Thomas, A. J. Marsden, J. Sloan, N. R. Wilson and J. P. Rourke, *Angew. Chem., Int. Ed.*, 2015, **5**, 9560–9563.
- 64 Y. Fukamori, M. König, B. Yoon, B. Wang, F. Esch, U. Heiz and U. Landman, *ChemCatChem*, 2013, **5**, 3330–3341.
- 65 A. Y. Stakheev and L. M. Kustov, *Appl. Catal., A*, 1999, **188**, 3–35.
- 66 M. Ohwada, K. Kimoto, T. Mizoguchi, Y. Ebina and T. Sasaki, *Sci. Rep.*, 2013, **3**, 2801.

