

CrossMark
click for updatesCite this: *RSC Adv.*, 2016, 6, 55499

Broadband-sensitive Ni^{2+} – Er^{3+} based upconverters for crystalline silicon solar cells†

Hom N. Luitel,* Shintaro Mizuno, Toshihiko Tani and Yasuhiko Takeda

We have developed Ni^{2+} , Er^{3+} codoped CaZrO_3 broadband-sensitive upconverters that significantly broaden the sensitive range, and hence overcome the shortcomings of conventional Er^{3+} doped upconverters used for crystalline silicon (c-Si) solar cells that utilize only a small fraction of the solar spectrum around 1550 nm. We have designed the combination of sensitizers and host material to utilize photons that are not absorbed by c-Si itself or Er^{3+} ions. Six coordinated Ni^{2+} ions substituted at the Zr^{4+} sites absorb (1060–1450) nm photons and transfer the energies to the Er^{3+} ions, and the Er^{3+} upconverts at 980 nm. Co-doping with monovalent charge compensators such as Li^+ for high Er^{3+} solubilisation at the Ca^{2+} sites and multivalent ions (Nb^{5+}) for stabilization of Ni^{2+} at the Zr^{4+} sites is essential. In addition to 1450–1600 nm ($\approx 2 \times 10^{20} \text{ m}^{-2} \text{ s}^{-1}$) photons directly absorbed by the Er^{3+} ions, we have demonstrated upconversion of 1060–1450 nm ($\approx 6 \times 10^{20} \text{ m}^{-2} \text{ s}^{-1}$) photons in the Ni^{2+} absorption band to 980 nm photons using the $\text{CaZrO}_3:\text{Ni}^{2+}, \text{Er}^{3+}$ upconverters. Compared with the current density gain of present c-Si solar cells ($\sim 40 \text{ mA cm}^{-2}$), the upconverted photons could increase this by $\sim 7.3 \text{ mA cm}^{-2}$, which is about 18% improvement. This architecture for broadband-sensitive upconversion may pave a new direction for the improvement in efficiency of the present c-Si solar cells to surpass the limiting conversion efficiency of single-junction solar cells.

Received 26th April 2016

Accepted 27th May 2016

DOI: 10.1039/c6ra10713c

www.rsc.org/advances

1. Introduction

Incident solar radiation is energetically broad and poses challenges for efficient conversion to electricity using solar cells consisting of finite bandgap semiconductors.¹ Even with an optimized crystalline silicon (c-Si) semiconductor, $\sim 30\%$ of the incident radiation is not absorbed and is simply transmitted through the present c-Si solar cells. Considering the maximum conversion efficiency of single-junction solar cells, $\sim 33\%$ sunlight into electricity in the radiative limit,² the unused near-infrared (NIR) radiation provides an important opportunity for enhancing the efficiency of solar cells. Although the NIR photons have insufficient energy, combining two or more infrared photons to generate a single visible or ultraviolet photon, a process known as upconversion, enables these NIR photons to contribute to solar energy conversion.^{3,4} An upconversion layer can be placed at the back of a solar cell and converts a part of the transmitted photons to wavelengths that can be absorbed; it is straightforward to demonstrate a positive contribution from the upconversion layer.

Upconversion in various materials in which rare-earth ions are doped have been reported.^{5–9} Among them, upconverters

using Er^{3+} ions that convert 1550 nm photons to 980 nm photons are applied to c-Si solar cells,^{10–12} while those using a combination of Er^{3+} and Yb^{3+} , which converts 970 nm to the visible photons, to amorphous silicon (a-Si), organic, and dye-sensitized solar cells.^{13–17} However, the absorption bands of rare-earth ions originating from f–f transitions in the NIR region (700–2100 nm) are narrow, and hence only a small fraction of the solar spectrum can be utilized.^{17–19} Thus, to realize highly efficient solar cells, efficient upconverters with sufficiently wide absorption spectra are desired. One of the approaches is to use a combination of rare earth ions with sensitizers that absorb photons in a broadband range, which are not directly absorbed by the rare-earth ions, and transfer the absorbed energies to the rare-earth ions. There are reports on upconversion of 970 nm photons to the visible (550, 650 nm) using the Er^{3+} – Yb^{3+} pairs where organic dyes function as the sensitizers that absorb photons in a range of 700–850 nm and transfer the energies to the Yb^{3+} (ref. 19) and similarly Cr^{3+} for 600–650 nm.²⁰ On the other hand, upconversion with a rather broadband sensitivity and a high efficiency has been demonstrated using triplet–triplet annihilation of organic molecules but the absorption range is limited up to 800 nm at present and thus can have applications only for a-Si, organic, and dye-sensitized solar cells.^{21–23}

2. Materials design

There are detailed reports on absorption spectra of transition-metal ions that extend up to the NIR range.^{24–31} It has been

Toyota Central Research and Development Laboratories Inc., 41-1, Yokomichi, Nagakute, Aichi 480-1192, Japan. E-mail: e1698@mosk.tytlabs.co.jp; takeda@mosk.tytlabs.co.jp; Tel: +81 956 717134

† Electronic supplementary information (ESI) available. See DOI: 10.1039/c6ra10713c

demonstrated that six coordinated Ni^{2+} ions located at the centers of octahedrons have substantial absorption around 1100–1400 nm that can further be tuned in a wide range by manipulating the crystal field strength around the Ni^{2+} ions.^{26,30,32} Accordingly, its NIR emission band is also wide ranging from 1200 nm to longer than 1700 nm with extremely high quantum yield reaching unity in some crystals such as LiGa_5O_8 .³³ There are other reports of the Ni^{2+} emission bands in the NIR ranges^{25–27,34–37} that sufficiently overlap with the Er^{3+} absorption bands making possibility of $\text{Ni}^{2+} \rightarrow \text{Er}^{3+}$ resonance energy transfer (sensitization).^{32,38,39} Zhang *et al.* has reported that Ni^{2+} to Er^{3+} energy transfer or *vice versa* in glass ceramics depending on the host matrix and Ni^{2+} to Er^{3+} physical separation to achieve NIR broadband Stokes emission for amplifier and lasers. Furthermore, a high concentration of the Ni^{2+} sensitizers is required, because it is equivalent to high excitation intensity for the Er^{3+} emitters. In the similar way, the optimum Er^{3+} concentration should also be high, 10–25% range, to promote energy transfer between two neighbouring excited Er^{3+} ions in addition to excited state absorption in a single Er^{3+} ion.³⁴

Thus, the host materials to realize broadband-sensitive upconversion using Er^{3+} and Ni^{2+} dopants are selected according to the following three criteria: (1) the host material must include cations located at octahedron centers, and should not include tetrahedron or other symmetry sites that can be occupied by Ni^{2+} ions. Four co-ordinated Ni^{2+} ions located at tetrahedron centers and Ni^{3+} have lower energy levels^{26,30} and hence rapid relaxation occur rather than energy transfer to the Er^{3+} emitters after photoexcitation. (2) The ionic radius of the cations located at the octahedron centers should be close to or at least not much smaller than that of Ni^{2+} (0.069 nm) to enable a high Ni^{2+} concentration, otherwise introduced Ni^{2+} ions would change to smaller Ni^{3+} ions (0.056 nm) or segregate. (3) The host material should include bigger ions (favourably isovalent to Er^{3+}) such that sufficient amount of Er^{3+} can be solubilised. Most widely used upconversion materials such as $\text{NaYF}_4:\text{Er}^{3+}$, $\text{Y}_2\text{O}_3:\text{Er}^{3+}$, or $\text{YF}_3:\text{Er}^{3+}$ (ref. 10 and 40) do not contain octahedron centers for Ni^{2+} hosting thus not suitable for the present purpose. We have already found that ABO_3 type perovskite structures such as the LaGaO_3 structure fulfil these criteria, because Ga ions (B sites) are located at the octahedron centers and the ionic radius of Ga^{3+} (0.062 nm) is similar to that of Ni^{2+} . La^{3+} ions (A sites) are isovalent to Er^{3+} ions with very similar ionic radii. Under such materials design, we have successfully demonstrated the broadband-sensitive upconversion utilizing the Ni–Er pairs in $\text{La}(\text{Ga},\text{Sc})\text{O}_3$, where the Ni^{2+} sensitizers absorb the photons around 1100–1400 nm and transfer the absorbed energies to the Er^{3+} ions, followed by Er^{3+} upconversion at around 980 nm, highly suitable for c-Si solar cells.⁴¹ However, the Ni-sensitized Er-upconverter exhibited low efficiency, and La and Sc used in the $\text{La}(\text{Ga},\text{Sc})\text{O}_3$ compounds are extremely expensive. For practical applications, inexpensive materials with high upconversion efficiency are desired.

It has been reported that alkaline earth perovskites such as CaZrO_3 are good hosts for rare earth doped phosphors and can solubilize as high as 20% rare earth ions when proper charge

compensators are utilized.⁴² The crystal structure of CaZrO_3 consists of ZrO_6 octahedra and CaO_{12} polyhedra as depicted in Fig. 1(c). The 6-coordinated Zr^{4+} ions occupy the centres of octahedra. The ionic radius of Zr^{4+} is 0.072 nm (6-coordination) while that of Ca^{2+} is 0.112 nm (12-coordination).⁴³ Thus, it is expected that codoped Ni^{2+} ions (0.069 nm) occupy Zr^{4+} sites while Er^{3+} ions (0.106 nm) occupy the Ca^{2+} sites. When the tetravalent Zr^{4+} ions are substituted with divalent Ni^{2+} ions, oxygen vacancies are created for charge neutrality, which would cause parasitic absorption and nonradiative relaxation of excited Ni^{2+} and Er^{3+} ions. Further, Ni^{2+} ions located at the higher valence sites (Zr^{4+}) would change to Ni^{3+} ions and do not contribute to upconversion as described above.³⁰ Nb^{5+} ions seem suitable as the charge compensators because they are stable and the ionic radius ($\text{Nb}^{5+} \approx 0.064$ nm) is close to that of Zr^{4+} (0.072 nm). Furthermore, it has been reported that the substitution of larger or smaller ions into host cation sites remarkably distorts the crystal field from the ideal symmetry and hence changes the crystal field splitting.⁴⁴ As a result, the f–f transition probabilities improve leading to increased absorption and upconversion. In the $\text{CaZrO}_3:\text{Er}$ system, substitution of

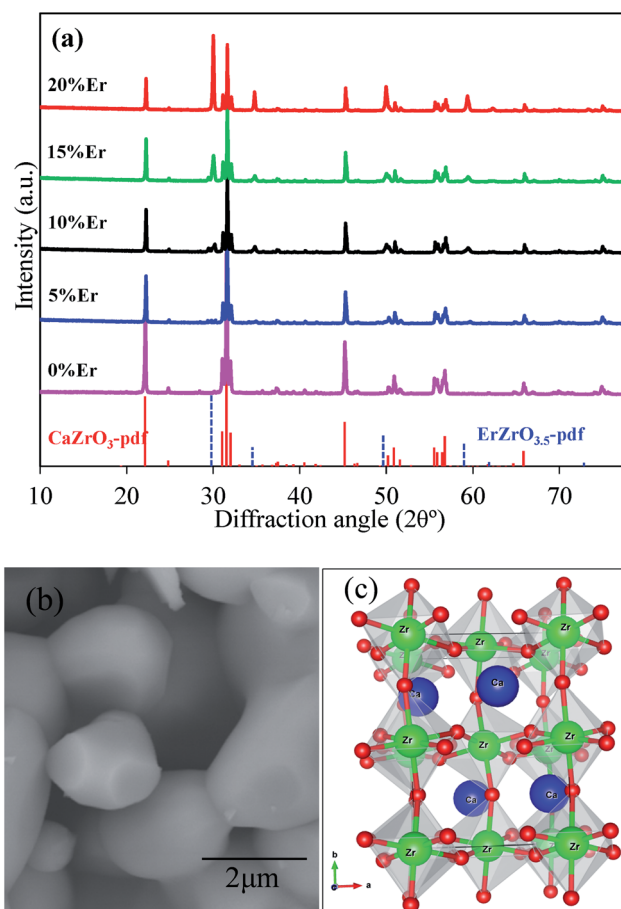


Fig. 1 (a) XRD patterns of CaZrO_3 doped with 0.2 mol% Ni and different mol% Er. For comparison, standard pdf data of CaZrO_3 (red solid lines) and $\text{ErZrO}_{3.5}$ (blue dotted lines) are also shown. (b) SEM image of the typical sample and (c) crystal structure of CaZrO_3 created by VESTA program.



Table 1 Ionic radii of cations at different coordination environments^{31,43}

Metal ions	Er ³⁺	Ca ²⁺	Li ⁺	Na ⁺	K ⁺
Ionic radii (nm) (12 coordination)	0.106	0.112	0.092	0.118	0.151
Metal ions	Ni ²⁺	Ni ³⁺	Zr ⁴⁺	Nb ⁵⁺	
Ionic radii (nm) (6 coordination)	0.069	0.056	0.072	0.064	

a part of the Ca²⁺ ions with smaller Er³⁺ ions distorts the crystal structure and gives the possibility of improved upconversion. In addition, substitution of smaller (Li⁺) or bigger (K⁺) ions for the Ca²⁺ ions further distorts the CaZrO₃ crystal, which can more remarkably increase the upconversion efficiency (Table 1).

Therefore, in the present study we have elaborately investigated the broadband-sensitive upconversion characteristics in the newly designed Ni²⁺–Er³⁺ codoped CaZrO₃ that can significantly improve the conversion efficiency of present c-Si solar cells.

3. Experiments

3.1. Synthesis of Er, Ni-codoped samples

We synthesized powder samples of CaZrO₃ codoped with Er³⁺ and Ni²⁺. Equivalent amount of Li⁺ ions (or Na⁺/K⁺) to that of Er³⁺ ions and double amount of Nb⁵⁺ ions to that of Ni²⁺ ions were also co-doped for charge balance unless specified. Because Er³⁺ occupy Ca²⁺ sites and Ni²⁺, Nb⁵⁺ substitute Zr⁴⁺ sites, the compositions were determined as (Er_xLi_xCa_{1–2x})(Zr_{1–3y}Ni_yNb_{2y})O₃. The compound oxide powders were synthesized using component metal-oxide or carbonate powders by the solid state reaction method. Predetermined amounts of the oxides/carbonates (Kojundo Kagaku, Japan) were well mixed with the help of small amount of ethanol and then dried at RT. Next the dry powders were again mixed well and heat-treated at 1400 °C for 6 h in air for reaction and crystallization. As synthesized powders might contain mixture of Ni²⁺ and Ni³⁺, post annealing of the powder samples was carried out at 800 °C in N₂ gas flow to convert all the Ni³⁺ ions to Ni²⁺ ions.

The crystalline structures were identified by X-ray diffraction (XRD) using Cu-Kα line in a range from 10 to 80 degree. Scanning electron microscope (SEM) observations were carried out using Hitachi SU3500.

3.2. Optical measurements

Each powder sample was sandwiched between two glass plates with a 0.5 mm thick spacer. Absorption spectra were measured using an integrating sphere. To determine the photon flux at the Ni²⁺ or Er³⁺ absorption bands, we have integrated the standard AM1.5 spectral solar irradiation at earth surface⁴⁵ in the absorption range of Ni²⁺ and Er³⁺, respectively. The reference spectrum is available in irradiance (W m^{–2} nm^{–1}) form, that was converted to the photon flux and current density forms. Continuous wave (CW) laser diodes of ~1 mW output power were used as excitation sources to acquire upconversion and Stokes emission spectra. Excitation beam intensity (mW cm^{–2}) at the sample surface was obtained using a power meter and

a knife-edge and it was in the order of 10⁶ mW cm^{–2}. An optical parametric oscillator (OPO) pumped by the third harmonic of a Nd-YAG laser (7 ns pulse duration, 5–20 mW output power depending on output wavelengths) was employed for evaluation of wavelength dependent upconversion sensitivity and time-resolved measurements. Appropriate bandpass filters were used to receive emitted photons of desired wavelengths. Si and InGaAs photodiodes were used to detect the 980 nm and longer wavelength photons, respectively, and the output signal was accumulated using a storage oscilloscope.

4. Results and discussion

Fig. 1(a) shows XRD patterns of the CaZrO₃ doped with different mol% Er. The measured diffraction peaks of the powder samples matched well to those of the standard CaZrO₃ card data (PDF# 01-080-6213), and confirmed that the samples were well crystallized in the orthorhombic phase (oP20). The sharp peaks are indication of good crystallization of the particles. The SEM image presented in Fig. 1(b) with well determined facets further confirmed the highly crystallized particles. The particles dimension varied over a range of 1–5 μm. The lattice constants calculated from the XRD result of the undoped CaZrO₃ sample are *a* = 5.5861 Å, *b* = 8.012 Å and *c* = 5.7568 Å, which is in good agreement with the standard JCPDS card data (PDF# 01-080-6213). With increasing Er-doping concentration, the lattice constants (*a*, *b*, *c*) and cell volume gradually decreased as summarised in the ESI S1† indicating smaller Er³⁺ (0.106 nm) ions occupied the larger Ca²⁺ (0.112 nm) sites. However, at/above 15 mol% Er addition, no further lattice contraction was observed indicating the limiting solid solubility of the Er³⁺ ions in the CaZrO₃ host, which was confirmed by existence of distinct XRD peaks of Er₂O₃ phase (*2θ* = 30 degree) in the XRD diagram in Fig. 1(a). On the other hand, no distinct lattice changes were observed by the Ni substitution at the Zr sites because the amount of Ni was very small (<0.5 mol%) in addition to the fact that the ionic size of Ni²⁺ (0.069 nm) is very similar to that of Zr⁴⁺ (0.072 nm). Further, because trivalent Er³⁺ substituted divalent Ca²⁺ ions, charge unbalance limited the solubility of Er³⁺ in the CaZrO₃ matrix as observed by the existence of Er₂O₃ peaks (*2θ* = 29.4 degree; PDF# 01-073-6274) in the Li free sample as shown in Fig. S2.† When equal amount of monovalent charge compensators such as Li⁺, Na⁺, K⁺ were codoped together with Er³⁺, the solid solubility of Er³⁺ in the CaZrO₃ host increased, and as a result the Er₂O₃ phase disappeared (see ESI Fig. S2†). However, lower formation energy of pyrochlore structures (–3.88 eV) compared to the perovskite structures, CaZrO₃ (–3.66 eV), the ErZrO_{3.5} pyrochlore phase was unavoidable at higher Er-doping concentration.⁴⁶ Reflecting the fact that codoped Li⁺ (Na⁺/K⁺) occupied Ca²⁺ sites, the lattice parameters and cell volume changed according to the ionic radii of these alkali ions (Fig. S2†).

The absorption spectra of the CaZrO₃ doped with 10 mol% Er and different mol% of Ni are presented in Fig. 2 and S3 (refer ESI S3† for visible absorption range). Sharp absorption peaks located at 970, 800, 660, 520, 490, 410, 380 nm and comparatively broad absorption band around (1450–1600) nm



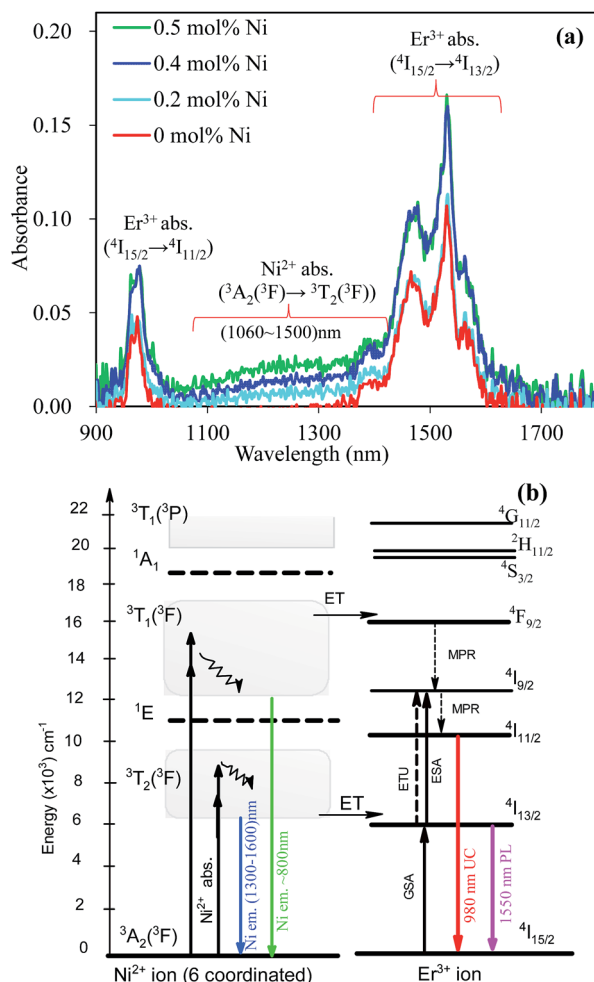


Fig. 2 (a) Absorption spectra of the CaZrO_3 :10 mol% Er codoped with different mol% of Ni exhibiting Er–Ni combined broadband absorption and (b) energy level diagram of the Ni^{2+} and Er^{3+} ions in the CaZrO_3 host with possible transitions and Ni \rightarrow Er energy transfer.

were assigned to the f–f transitions of Er^{3+} ions,^{24,34} while rather broad but comparatively weak absorption bands located around (1050–1500) nm, (680–860) nm and (390–460) nm were originated from the $^3\text{A}_2(^3\text{F})$ to $^3\text{T}_2(^3\text{F})$, $^3\text{T}_1(^3\text{F})$ and $^3\text{T}_1(^3\text{P})$ optically allowed transitions, respectively, of six coordinated Ni^{2+} ions.^{20,32} No distinct Ni^{3+} absorption bands were identified indicating that most of the doped Ni stabilised as Ni^{2+} ions.^{25,33} We used Nb^{5+} ions as charge compensator in double amount to that of doped Ni that might have stabilized the added Ni as Ni^{2+} ions. The relative locations of the energy levels of the Er^{3+} and Ni^{2+} ions are illustrated in Fig. 2(b). Furthermore, the Ni^{2+} absorption bands remarkably increased with increasing doped Ni^{2+} ion concentration. The extra broad Ni^{2+} absorption band (1060–1500 nm) in combination with the Er^{3+} absorption band (1450–1600 nm) covers the most of the NIR solar spectrum (1050–1600 nm) which is not absorbed by c-Si solar cells themselves and will have a great advantage for the spectral conversion purpose.

Fig. 3(a) and (b) show the upconversion spectra of the various mol% of Er (Li) doped CaZrO_3 :0.2 mol% Ni (Nb) samples when

excited at 1490 and 1300 nm, respectively. Under the 1490 nm excitation, corresponding to direct Er^{3+} excitation, clear upconverted emissions at around 980 nm were observed (Fig. 3(a)). In addition, very similar upconverted emissions appeared while excited at 1300 and 1180 nm (for 1180 nm excited UC see Fig. S5†) despite the excitation light was not absorbed by the Er^{3+} ions themselves. The maximal upconversion intensity was achieved at 15 mol% Er doping concentration under both the 1300 and 1490 nm excitations. Further, it has been confirmed that the upconverted emission at 980 nm was proportional to the square of the Er concentrations for both excitation wavelengths at lower concentrations as depicted in the insets of Fig. 3(a) and (b). The power dependent upconversion intensity measurements confirmed that the present upconversion is two photon processes under both the 1490 and 1300 nm excitations (Fig. 3(c)).

Fig. 4 compares the Stokes emissions of the Ni (Nb) only doped and Er, Ni (Li, Nb)-codoped samples excited at 1180 nm. For the Ni only doped (no Er) sample, broad emission band ranging from 1300 nm to 1600 nm originated from the $^3\text{T}_2(^3\text{F}) \rightarrow ^3\text{A}_2(^3\text{F})$ transition of Ni^{2+} was observed.^{26–33} However, by introducing Er^{3+} (Li^+) co-dopants, the Ni^{2+} emission almost disappeared and Er^{3+} emissions at around 1550 nm ($^4\text{I}_{13/2} \rightarrow ^4\text{I}_{15/2}$) and 980 nm ($^4\text{I}_{11/2} \rightarrow ^4\text{I}_{15/2}$) appeared (see Fig. S5† for 1180 nm excited UC at 980 nm). This suggests that the energies absorbed by the Ni^{2+} ions transferred to the Er^{3+} ions, followed by the Er^{3+} emissions *i.e.* the Ni^{2+} ions act as donors while Er^{3+} acceptors. Further, under sufficient Er^{3+} co-doping (15 mol% or higher), the Ni^{2+} emission was almost perfectly quenched and only the Er^{3+} emission appeared, indicating the Ni \rightarrow Er energy transfer efficiency was close to unity. The energy transfer efficiency estimated from the ratio of the Ni^{2+} emission intensities integrated over 1300–1450 nm in the Ni only doped sample (0.2 mol% Ni) and Ni–Er codoped sample (0.2 mol% Ni, 10 mol% Er) was as high as 84%. However, at lower Er^{3+} concentrations, the Er^{3+} ions were far apart from the Ni^{2+} ions and the Ni \rightarrow Er energy transfer was less efficient, as a result, Ni^{2+} emission was detected although it was weak. The energy transfer efficiency and subsequent excited Er^{3+} number should increase with the Er^{3+} concentration, and consequently the Er^{3+} Stokes emission intensity. In fact, the Er^{3+} Stokes emission, which is a one-photon process, was proportional to the Er concentration (see Fig. S4†). On the other hand, Er^{3+} upconversion (980 nm) is a two-photon process (Fig. 3(c)); it means two excited Er^{3+} ions are utilised to emit a single upconverted photon at a shorter wavelength. Thus, the upconversion emission intensity should be proportional to the square of the Er^{3+} ion concentration, which was obeyed at relatively lower Er^{3+} concentrations. However, at higher Er^{3+} -doping concentrations, the Ni \rightarrow Er energy transfer efficiency saturated and concentration quenching processes predominates, consequently the upconversion and Stokes emissions weakened.³⁴

Effect of the Ni^{2+} -doping concentration on the upconversion spectra of CaZrO_3 :15 mol% Er(Li) are presented in Fig. 5. At the beginning, the upconversion intensity increased with increasing Ni-doping concentration that is due to increased Ni absorption as depicted in Fig. 2(a). However, at slightly higher



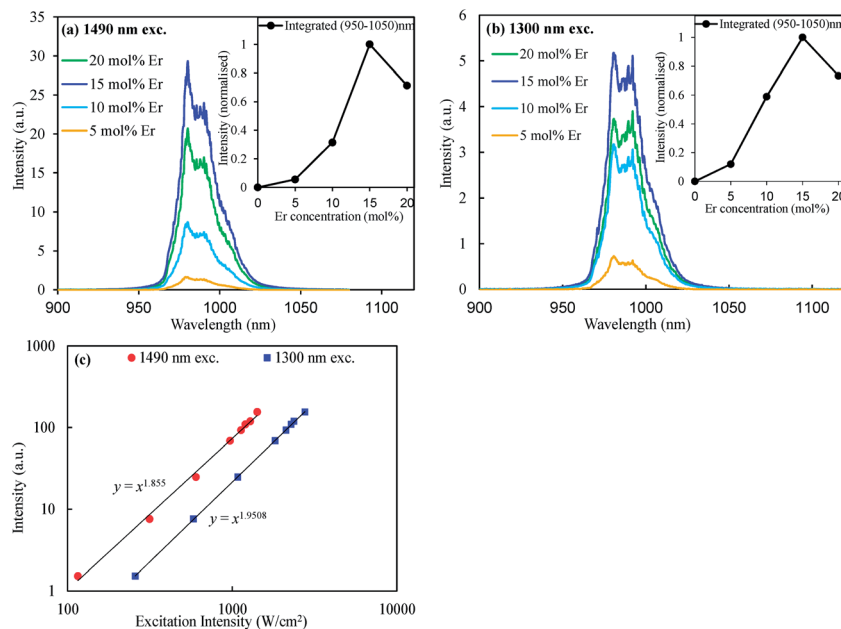


Fig. 3 Upconversion spectra of the CaZrO₃:0.2 mol% Ni codoped with different mol% of Er excited at (a) 1490 nm, and (b) 1300 nm. (c) Power-dependent UC intensity of the CaZrO₃:0.2 mol% Ni²⁺, 15 mol% Er³⁺ under 1490 and 1300 nm excitations. Inset in the (a) and (b) are the dependence of the UC intensities integrated over 950–1050 nm with the Er³⁺ ion concentration.

Ni concentrations, the upconversion intensity decreased rapidly. The reason of the decrease would be due to remarkable non-radiative relaxation through concentration and defect-related quenching as oxygen vacancies may increase at higher Ni concentrations.⁴¹ The energy transfer efficiency is expected to be independent of the Ni²⁺ concentrations because the Ni²⁺ concentration is very low compared to the Er³⁺ concentrations. Thus increased Ni²⁺ absorption enhanced the overall energy transfer and hence intensified the Er³⁺ upconversion. However, effects of high absorption and energy transfer efficiency were overcome by nonradiative loss at a relatively higher Ni²⁺

concentration as mentioned above and reduced the overall upconversion intensity. Inset in the Fig. 5 shows the effect of Nb⁵⁺ concentration on the upconverted emission integrated over 950–1050 nm. By introducing proper amount of Nb⁵⁺ ions (twice the amount of Ni²⁺) the upconverted emission intensity was almost doubled. Since, Ni²⁺ ions were introduced at the Zr⁴⁺ sites, charge imbalance reduced the solubility of Ni²⁺ ions into the CaZrO₃ host (no Ni-phase was detected due to low doping

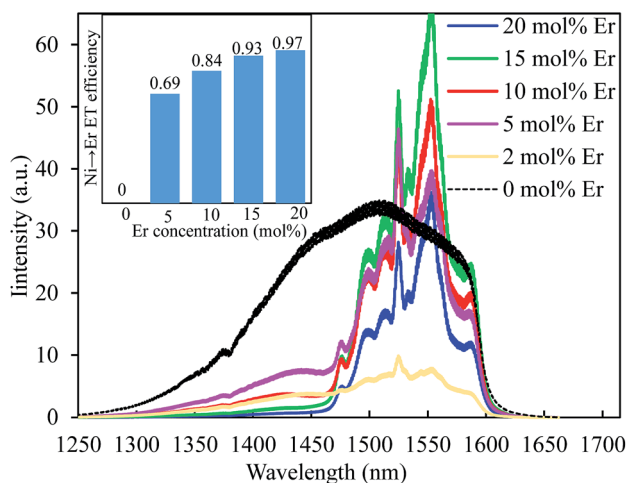


Fig. 4 Stokes emission spectra of the Ni (Nb) only doped (dotted line) and different mol% of Er (Li) codoped CaZrO₃ samples excited at 1180 nm.

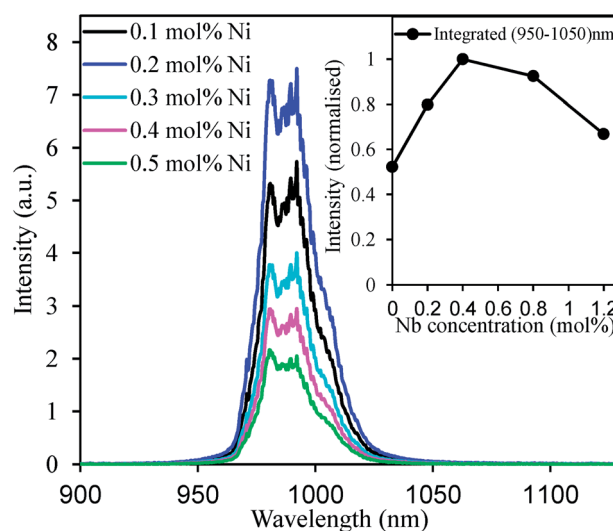


Fig. 5 Upconversion spectra of the CaZrO₃:15 mol% Er codoped with different mol% of Ni excited at 1300 nm. Inset is the effect of Nb⁵⁺ doping concentrations on the upconverted emission intensity (integrated over 950–1050 nm) of CaZrO₃:15 mol% Er³⁺, 0.2 mol% Ni²⁺ samples.



concentration) or stabilised the Ni^{3+} ions to reduce the stress of charge imbalance. Substitution of two Nb^{5+} ions at Zr^{4+} sites for each Ni^{2+} addition balanced the total charge that might have increased the Ni-solubility as well as stabilize the Ni^{2+} . As a result, improved upconverted emission intensity was detected.

To evaluate the $\text{Ni} \rightarrow \text{Er}$ energy transfer rate quantitatively, time-resolved decay profiles were recorded. Typically, the Ni^{2+} emission at 1400 nm was recorded for the Ni only doped (no Er) and Ni, Er-codoped (0.2 mol% Ni, 10 mol% Er) samples under the 1180 nm pulsed excitation. The results are presented in Fig. 6. The time-dependent Ni^{2+} emission intensity of the Ni only doped (no Er) sample $I_{\text{Ni}(\text{no Er})}(t)$ was well fitted using a double exponential eqn (1) mentioned below:³⁹

$$I_{\text{Ni}(\text{no Er})}(t) = a_{\text{Ni}} \exp[-t/\tau_{\text{Ni}}^{(1)}] + (1 - a_{\text{Ni}}) \exp[-t/\tau_{\text{Ni}}^{(2)}] \quad (1)$$

The Ni^{2+} emission lifetimes τ_{Ni} are determined by

$$\tau_{\text{Ni}} = (a_{\text{Ni}}\tau_{\text{Ni}}^{(1)^2} + (1 - a_{\text{Ni}})\tau_{\text{Ni}}^{(2)^2}) / (a_{\text{Ni}}\tau_{\text{Ni}}^{(1)} + (1 - a_{\text{Ni}})\tau_{\text{Ni}}^{(2)}) \quad (2)$$

From the fitting data, τ_{Ni} was found to be ~ 0.6 ms for 0.2 mol% Ni (Nb) only doped CaZrO_3 , which is comparable with the previously reported values for other oxide hosts.^{47,48}

When the Er^{3+} acceptors were introduced, the Ni^{2+} emission intensity decreased more rapidly, because of the energy transfer from the Ni^{2+} donors to the Er^{3+} acceptors. The effect of the energy transfer on the time evolution of the emission intensity of the Ni^{2+} donors is expressed as $\exp[-\gamma_{\text{Ni} \rightarrow \text{Er}} \sqrt{t}]$ when energy migration among the donors is negligibly weak,^{38,39} whereas it is described by an exponential function $\exp[-w_{\text{Ni} \rightarrow \text{Er}} t]$ for significant migration.^{49–51} In the present case, the acquired data are fitted by the former better than the latter, *i.e.*,

$$I_{\text{Ni}(\text{Er co-doped})}(t) = \exp[-\gamma_{\text{Ni} \rightarrow \text{Er}} \sqrt{t}] I_{\text{Ni}(\text{no Er})}(t) \quad (3)$$

This is consistent with the facts that the Ni concentration is as low as 0.2 mol% and the energy transfer is completed within sub-milliseconds. Using eqn (1) and (3), the energy transfer efficiency $\eta_{\text{Ni} \rightarrow \text{Er}}$ is derived as follows,

$$\eta_{\text{Ni} \rightarrow \text{Er}} = 1 - \int_0^\infty dt I_{\text{Ni}(\text{Er co-doped})}(t) / \int_0^\infty dt I_{\text{Ni}(\text{no Er})}(t) \quad (4)$$

The efficiency obtained from the time-resolved data was about 86%, which is very close to the result obtained from the steady-state Ni-emission intensities as mentioned earlier.

Fig. 7 shows the effect of monovalent ions (Li^+ , Na^+ , K^+) substitution on the upconversion intensity of the Er, Ni codoped CaZrO_3 upconverters. Addition of equivalent amount of alkali ions to that of doped Er^{3+} ions remarkably increased the upconversion intensity. Because the Er^{3+} ions (0.106 nm) were substituted for the Ca^{2+} (0.112 nm) sites, there was imbalance of the total charge. As a result, stress was produced that limits the solubility of the Er^{3+} ions in CaZrO_3 host, which was confirmed by the existence of segregated Er_2O_3 impurity phase in the XRD patterns as stated previously (Fig. S2†). When the equivalent amount of monovalent ions such as Li^+ , Na^+ , K^+ etc. were added, the charge neutrality was maintained, and consequently the Er^{3+} impurity phase was remarkably reduced (see Fig. S2†). There were no significant differences in the absorption peak positions of the Er^{3+} and Ni^{2+} but remarkable enhancement of the peak intensities around the Er^{3+} and Ni^{2+} absorption bands was observed (see Fig. S6†). The enhanced absorption of Er^{3+} is obviously due to increased numbers of the Er^{3+} ions inside the CaZrO_3 lattice, resulting in more intense upconversion. However, the upconversion intensity was found in the order of $\text{Li}^+ \gg \text{K}^+ > \text{Na}^+ > \text{no alkali ions}$, and could not be explained by charge compensation only. The f-f electronic transitions of rare earth ions are partially forbidden. As a result, they exhibit low

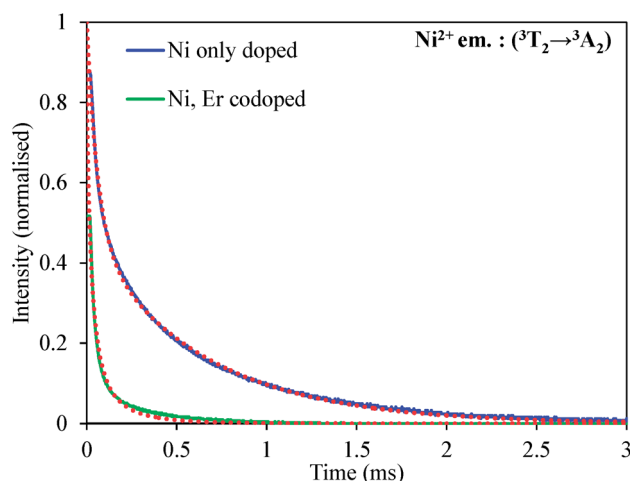


Fig. 6 Decay profiles of the Ni^{2+} emission at 1400 nm for the Ni (Nb) only doped (no Er) and Ni, Er (Nb, Li) codoped CaZrO_3 samples under the 1180 nm pulsed excitation. Red dotted lines are the results of fitting.

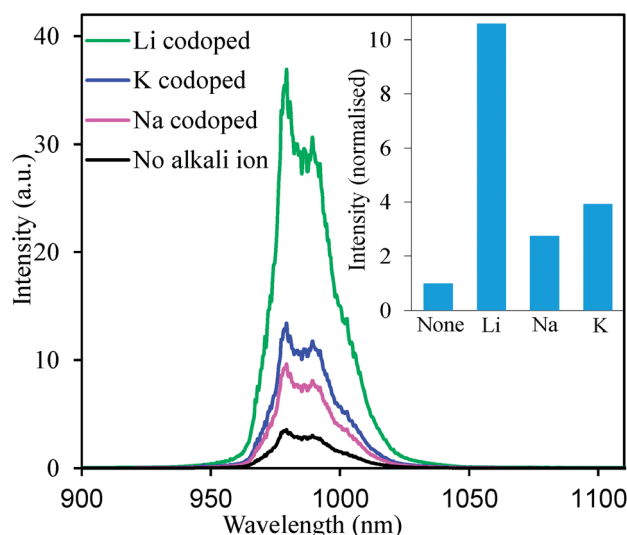


Fig. 7 Upconversion spectra of $\text{CaZrO}_3:\text{Er}^{3+}, \text{Ni}^{2+}$ codoped with alkali ions excited at 1180 nm. Inset is the variation of integrated upconversion intensity over 950–1050 nm with Li^+ , Na^+ , K^+ ions substitutions.



transition probabilities in presence of inversion symmetry.⁴⁴ It has been reported that substitution of larger or smaller ions into host cation sites remarkably distorts the crystal geometry from the ideal symmetry and hence changes the crystal field splitting.⁴⁴ As a result, the f-f transition probabilities increase and absorption originating from the Er^{3+} pronounce. The similar trend was observed in the present case too, where the Er^{3+} absorption bands were pronounced more by the substitution of the smaller Li^+ ions (0.092 nm) at the Ca^{2+} sites (0.112 nm) than the similar sized Na^+ ions (0.118 nm). Substitution with K^+ ions (0.151 nm) is also expected to intensify the upconversion as with Li^+ ions, however, K^+ ions are much larger than Ca^{2+} ions and it is difficult to occupy the Ca^{2+} sites. Thus, some of the added K^+ ions (hence the same number of the Er^{3+} ions) did not exist inside the CaZrO_3 crystals and remained at the surfaces. Nonradiative transitions from surface ions (here Er^{3+} ions) will be pronounced due to rapid migration of excited energy to the surface defects.⁵² Thus, lower Er^{3+} concentration inside the CaZrO_3 lattice and existence of considerable amount of Er^{3+} ions at the surfaces made the K^+ -substituted samples less efficient compared to the Li^+ -substituted samples. Furthermore, the K^+ -substituted sample exhibited lower crystallinity as seen by weaker XRD peak intensity depicted in Fig. S2.† It is well known that luminescence is crystallinity dependent; higher the crystallinity the better is the luminescence performance.

Fig. 8 shows the absorption, excitation and upconverted emission spectra of the CaZrO_3 :15 mol% $\text{Er}(\text{Li})$, 0.2 mol% $\text{Ni}(\text{Nb})$ sample. The excitation spectrum was obtained from the upconverted emission intensity normalized by the square of the excitation photon flux (due to two-photon process), namely, the emission intensity at a constant excitation photon flux. The upconverter developed here provides broadband sensitivity ranging from 1060 nm to longer than 1600 nm as seen from the excitation spectrum. It further suggests that quite strong upconversion was achieved at the Er-absorption band while

comparatively weak at the Ni-absorption band, which is in accordance to the absorbance at the corresponding wavelengths. We have compared the upconversion intensities of such a broadband sensitive upconverter developed here with the previously reported similar upconverter (see ESI S7†) and concluded that the upconversion emission of the CaZrO_3 : Er , Ni -based upconverter is about 1.7 times stronger. Further, the new upconverter developed here can efficiently absorb 1350–1450 nm light even though the solar irradiation at these wavelengths is weak.

The photon flux in the solar spectrum integrated over the Er^{3+} absorption band (1450–1600 nm) is about $2.0 \times 10^{20} \text{ m}^{-2} \text{ s}^{-1}$, which is equivalent to $\sim 1.93 \text{ mA cm}^{-2}$ current density gain for c-Si solar cells assuming perfect upconversion.⁴⁵ Introduction of the Ni^{2+} sensitizers increases the absorbed photon flux by $6.0 \times 10^{20} \text{ m}^{-2} \text{ s}^{-1}$ (1060–1450 nm), leading to an additional gain of $\sim 5.94 \text{ mA cm}^{-2}$ that is more than 3 times the gain originating from the Er^{3+} absorption alone. The optimum current density gain of present silicon solar cell is about 40 mA cm^{-2} .⁵³ Assuming perfect absorption and upconversion, the newly developed CaZrO_3 : Ni^{2+} , Er^{3+} upconverter can contribute $\sim 7.3 \text{ mA cm}^{-2}$ current density gain of present c-Si solar cells that corresponds to nearly $\sim 18\%$ increase in gain. However, to realize such efficient broadband-sensitive upconverters, and hence higher c-Si solar cell efficiency, further improved Ni^{2+} and Er^{3+} absorption is essential and it can be achieved, for example, by coupling with plasmons at the desired wavelengths and such experiments will be conducted in near future.

5. Conclusions

To improve conversion efficiency of c-Si solar cells, we have developed a new class of broadband-sensitive upconverter, CaZrO_3 : Ni^{2+} , Er^{3+} that absorbs photons in a broad wavelength range *viz.* 1060–1600 nm and emits photons at 980 nm to be efficiently absorbed by c-Si solar cells. We used 6-coordinated Ni^{2+} located at the octahedron centers (Zr^{4+} -sites) and 12-coordinated Er^{3+} located at the Ca^{2+} sites, such that efficient energy transfer from the Ni^{2+} to the Er^{3+} was achieved. Further, optimization of the Ni^{2+} and Er^{3+} concentrations and addition of charge compensators to reduce the stress caused by the Er^{3+} and Ni^{2+} doping significantly improved the upconversion efficiency. Assuming ideal upconversion, the newly developed CaZrO_3 : Ni^{2+} , Er^{3+} upconverter can contribute $\sim 7.3 \text{ mA cm}^{-2}$ current density gain of present c-Si solar cells ($\sim 40 \text{ mA cm}^{-2}$) which corresponds to nearly $\sim 18\%$ increase in gain. Such broadband-sensitive absorption of sub-bandgap photons above the absorption edge of c-Si solar cells and upconverting them to the most sensitive range (980 nm) of the cell offer a clear route towards surpassing the limiting conversion efficiency of single-junction solar cells, and further enable new applications in sub-bandgap detector sensitization and many more.

Acknowledgements

This work was partially supported by Advanced Low Carbon Technology Research and Development Program (ALCA), Japan Science and Technology Agency.

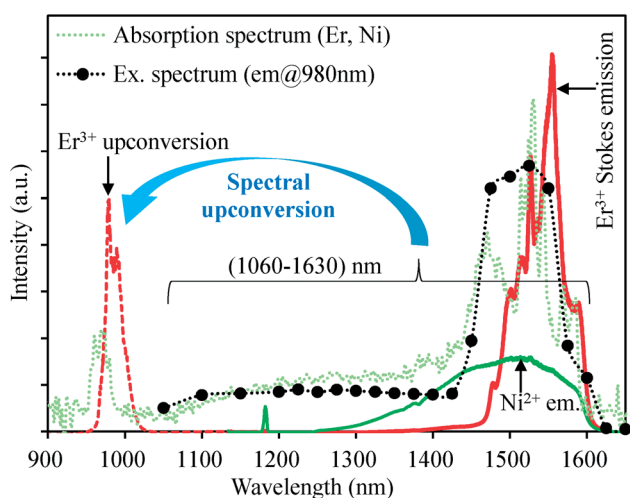


Fig. 8 Illustration of broadband-sensitive upconversion in the CaZrO_3 : Ni^{2+} , Er^{3+} upconverter along with the Ni and Er Stokes emissions excited at 1180 nm.



References

- W. Shockley and H. J. Queisser, *J. Appl. Phys.*, 1961, **32**, 510–519.
- T. Markvart, *Phys. Status Solidi A*, 2008, **205**, 2752–2756.
- V. Badescu and A. M. Badescu, *Renewable Energy*, 2009, **34**, 1538–1544.
- J. Wild, A. Meijerink, J. K. Rath, W. G. J. H. M. Sark and R. E. I. Schropp, *Energy Environ. Sci.*, 2011, **4**, 4835.
- X. Huang, S. Han, W. Huang and X. Liu, *Chem. Soc. Rev.*, 2013, **42**, 173–201.
- H. N. Luitel, R. Chand and T. Watari, *Displays*, 2016, **42**, 1–8.
- H. N. Luitel, K. Ikeue, R. Okuda, R. Chand, Y. Mitsunori, T. Torikai and T. Watari, *Opt. Mater.*, 2014, **36**, 591–595.
- J. C. Goldschmidt and S. Fischer, *Adv. Opt. Mater.*, 2015, **3**, 510–535.
- A. Shalav, B. S. Richards, T. Trupke, K. W. Krämer and H. U. Güdel, *Appl. Phys. Lett.*, 2005, **86**, 10–13.
- S. Fischer, R. Martín-Rodríguez, B. Fröhlich, K. W. Krämer, A. Meijerink and J. C. Goldschmidt, *J. Lumin.*, 2014, **153**, 281–287.
- G. E. Arnaoutakis, J. Marques-Hueso, A. Ivaturi, S. Fischer, J. C. Goldschmidt, K. W. Krämer and B. S. Richards, *Sol. Energy Mater. Sol. Cells*, 2015, **140**, 217–223.
- J. L. Wu, F. C. Chen, S. H. Chang, K. S. Tan and H. Y. Tuan, *Org. Electron.*, 2012, **13**, 2104–2108.
- G. B. Shan and G. P. Demopoulos, *Adv. Mater.*, 2010, **22**, 4373–4377.
- Y. Chen, J. Zhou, Y. Jiao, W. He, H. Wang, X. Hao, J. Lu and S. Yang, *J. Lumin.*, 2013, **134**, 504–507.
- J. de Wild, T. F. Duindam, J. K. Rath, A. Meijerink, W. G. J. H. M. van Sark and R. E. I. Schropp, *IEEE Journal of Photovoltaics*, 2013, **3**, 17–21.
- N. C. Dyck and G. P. Demopoulos, *RSC Adv.*, 2014, **4**, 52694–52701.
- J. A. Briggs, A. C. Atre and J. A. Dionne, *J. Appl. Phys.*, 2013, **113**, 124509.
- C. M. Johnson, S. Woo and G. J. Conibeer, *IEEE Journal of Photovoltaics*, 2014, **4**, 799–806.
- W. Zou, C. Visser, J. A. Maduro, M. S. Pshenichnikov and J. C. Hummelen, *Nat. Photonics*, 2012, **6**, 560–564.
- S. Ye, E. H. Song, E. Ma, S. J. Zhang, J. Wang, X. Y. Chen, Q. Y. Zhang and J. R. Qiu, *Opt. Mater. Express*, 2014, **4**, 638.
- A. Nattestad, Y. Cheng, R. Macqueen, T. Schulze, F. Thompson, A. Mozer, B. Fu, T. Khoury, M. Crossley, K. Lips, G. Wallace and T. Schmidt, *J. Phys. Chem. Lett.*, 2013, **4**, 2073–2078.
- T. F. Schulze and T. W. Schmidt, *Energy Environ. Sci.*, 2014, **8**, 103–125.
- T. F. Schulze and T. W. Schmidt, *Energy & Environmental Science*, 2015, **8**, 103–125.
- C. Strümpel, M. McCann, C. Del Canizo, I. Tobias and P. Fath, *Proc. 20th European Photovoltaic Solar Energy Conference*, Barcelona, Spain, 2005, pp. 43–43.
- T. Suzuki, G. S. Murugan and Y. Ohishi, *J. Lumin.*, 2005, **113**, 265–270.
- N. Vasileva, P. A. Gerus, V. Sokolov and V. G. Plotnichenko, *J. Phys. D: Appl. Phys.*, 2012, **45**, 485301.
- M. Alfredsson, F. Corà, D. P. Dobson, J. Davy, J. P. Brodholt, S. C. Parker and G. D. Price, *Surf. Sci.*, 2007, **601**, 4793–4800.
- E. P. Dubrovina, V. A. Sandulenko, M. I. Demchuk, N. V. Kuleshov and V. P. Mikhailov, *Chem. Phys. Lett.*, 1990, **170**, 473–477.
- M. Saiful Islam and R. Andrew Davies, *J. Mater. Chem.*, 2004, **14**, 86–93.
- J. Koetke, G. Huber and K. Petermann, *J. Lumin.*, 1991, **48–49**, 564–568.
- R. Shannon, *Acta Crystallogr., Sect. A: Cryst. Phys., Diffraction, Theor. Gen. Crystallogr.*, 1976, **32**, 751–767.
- R. Zhang, H. Lin, D. Chen, Y. Yu and Y. Wang, *J. Alloys Compd.*, 2013, **552**, 398–404.
- J. F. Donegan, F. J. Bergin, T. J. Glynn, G. F. Imbusch and J. P. Remeika, *J. Lumin.*, 1986, **35**, 57–63.
- F. Auzel, *Chem. Rev.*, 2004, **104**, 139–173.
- T. Y. Eeu, X. G. Pang, T. Q. Leow, I. Zuhairi and R. Hussin, *Adv. Mater. Res.*, 2014, **895**, 265–268.
- S. Kuck, *Appl. Phys. B: Lasers Opt.*, 2001, **72**, 515–562.
- J. F. Suyver, A. Aebischer, D. Biner, P. Gerner, J. Grimm, S. Heer, K. W. Krämer, C. Reinhard and H. U. Güdel, *Opt. Mater.*, 2005, **27**, 1111–1130.
- T. Förster, *Ann. Phys.*, 1948, **437**, 55–75.
- J. R. Lakowicz, *Principles of Fluorescence Spectroscopy*, Third edn, Springer, New York, 2006.
- R. Martín-Rodríguez, S. Fischer, A. Ivaturi, B. Fröhlich, K. W. Krämer, J. C. Goldschmidt, B. S. Richards and A. Meijerink, *Chem. Mater.*, 2013, **25**, 1912–1921.
- Y. Takeda, S. Mizuno, H. N. Luitel and T. Tani, *Appl. Phys. Lett.*, 2016, **108**, 04390.
- J. Fu, Q. Zhang, Y. Li and H. Wang, *J. Alloys Compd.*, 2009, **485**, 418–421.
- R. Grimes, *Atomistic Simulation*, <http://abulafia.mt.ic.ac.uk/shannon/ptable.php>.
- Q. Huang, H. Yu, E. Ma, X. Zhang, W. Cao, C. Yang and J. Yu, *Inorg. Chem.*, 2015, **54**, 2643–2651.
- Solar spectrum AM1.5, <http://rredc.nrel.gov/solar/spectra/am1.5/>.
- Materials Project, <https://www.materialsproject.org>.
- B. Wu, J. Qiu, E. Wu and H. Zeng, *Opt. Mater.*, 2013, **35**, 983.
- G. Dong, M. Liang, H. Qin, G. Chai, X. Zhang, Z. Ma, M. Peng and J. Qiu, *Phys. Chem. Chem. Phys.*, 2012, **14**, 13594.
- J. A. Caird, A. J. Ramponi and P. R. Staver, *J. Opt. Soc. Am. B*, 1991, **7**, 1391–1403.
- A. I. Burshtein, *Soviet Physics-JETP*, 1972, **35**, 882.
- C. Parent, C. Lurin, G. Le Flem and P. Hagenmuller, *J. Lumin.*, 1986, **36**, 49–55.
- S. Som, A. K. Kunti, V. V. Kumar, S. Dutta, M. Chowdhury, S. K. Sharma, J. J. Terblans and H. C. Swart, *J. Appl. Phys.*, 2014, **115**, 193101.
- M. A. Green, K. Emery, Y. Hishikawa, W. Warta and E. D. Dunlop, *Prog. Photovoltaics*, 2016, **24**, 3–11.

