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## Broadband ultrafast photoprotection by oxybenzone across the UVB and UVC spectral regions†

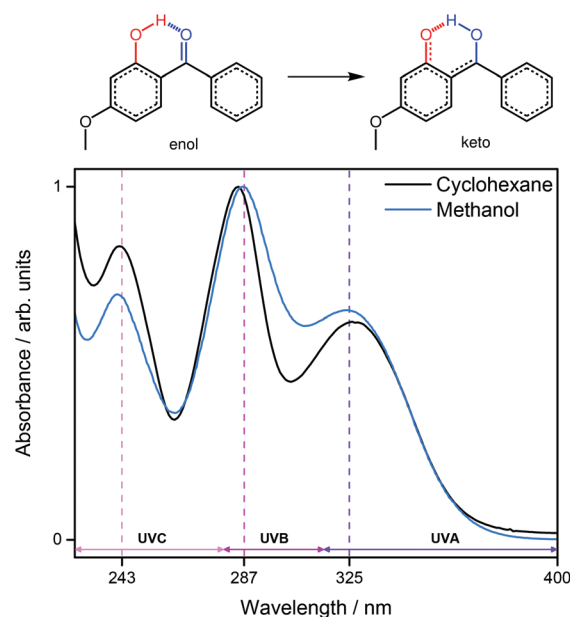
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Recent studies have shed light on the energy dissipation mechanism of oxybenzone, a common ingredient in commercial sunscreens. After UVA photoexcitation, the dissipation mechanism may be understood in terms of an initial ultrafast excited state enol  $\rightarrow$  keto tautomerisation, followed by nonadiabatic transfer to the ground electronic state and subsequent collisional relaxation to the starting enol tautomer. We expand on these studies using femtosecond transient electronic absorption spectroscopy to understand the non-radiative relaxation pathways of oxybenzone in cyclohexane and in methanol after UVB and UVC excitation. We find that the relaxation pathway may be understood in the same way as when exciting in the UVA region, concluding that oxybenzone displays proficient broadband non-radiative photoprotection, and thus photophysically justifying its inclusion in sunscreen mixtures.

Ultraviolet (UV) radiation is a vital requirement for the production of vitamin D. However, its benefit remains in a delicate balance given that overexposure to UV is known to be a dominant cause of cancers such as malignant melanomas.<sup>1–3</sup> To protect against overexposure, the human body has developed an extensive network of naturally synthesised UV light absorbing biopolymers, the concentrations of which can fluctuate through the regulation of melanogenesis.<sup>4</sup> Thus, in instances of high UV exposure, melanogenesis is up-regulated and the result is tanning of the skin. This natural photoprotective mechanism is a delayed response to overexposure from which DNA damage could have already occurred. This has led to the development of sunscreens that utilise a combination of UV chromophores to act preemptively to reduce the risk of overexposure and thus the probability of skin photodamage.<sup>5</sup>

Particularly in the current age, shifts in perceptions on tanning and increased travel and tourism have meant the correct use of sunscreens has never been more important.<sup>6,7</sup>

Oxybenzone (OB), shown in Fig. 1, has been the subject of numerous studies due to its use as a UV chromophore in many commercial sunscreens.<sup>5,10–13</sup> It displays intense, broad absorption bands in each of the UVA (400–315 nm), UVB (315–280 nm) and UVC (280–100 nm) regions (Fig. 1) making it suitably flexible to absorb incident (broadband UV) solar radiation, and, in the case of UVC, artificial UV sources.<sup>14–16</sup> Importantly, OB has also been shown to remain photostable after several hours of irradiation,<sup>17</sup> although some controversy



**Fig. 1** Enol-OB and keto-OB tautomers. Studies have suggested that enol-OB undergoes an ultrafast tautomerisation into a keto tautomer after UVA irradiation.<sup>8,9</sup> The UV-visible spectrum for enol-OB in solution in cyclohexane (black) and methanol (blue) displays three absorption maxima at ca. 325 nm, 287 nm and 243 nm, each highlighted by a vertical dashed line.

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†Electronic supplementary information (ESI) available: Global fits for spectral ranges  $355 \leq \lambda \leq 415$  and  $415 < \lambda \leq 650$ . Support plane analysis on extracted lifetimes. Solvent response transient absorption spectra. Single transient fits for 366 nm and 380 nm. Cross correlation measurements are given. See DOI: 10.1039/C5PP00217F



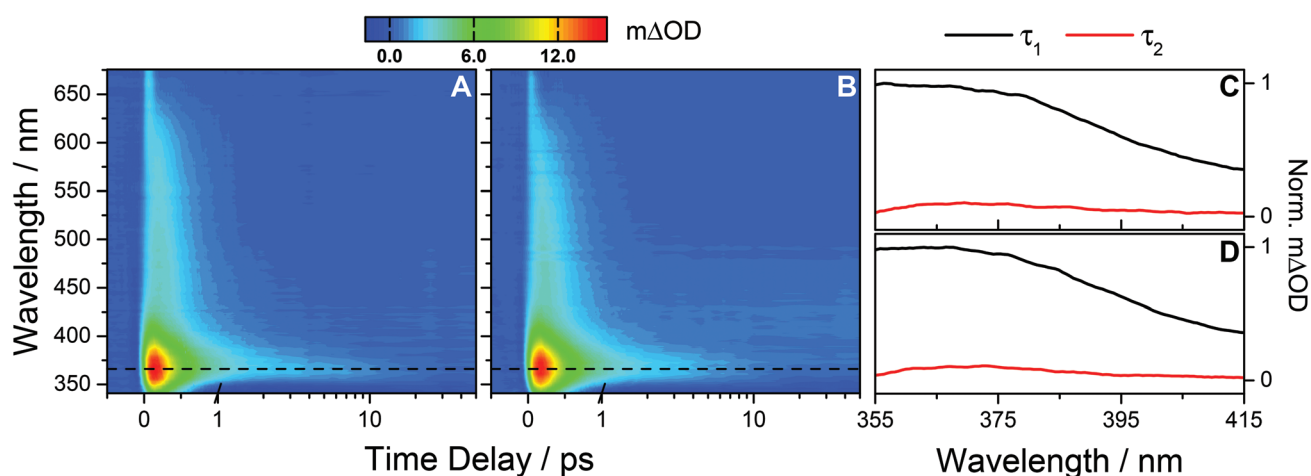
remains surrounding its suitability as a sunscreen constituent with respect to adverse dermatological effects and endocrine disruption.<sup>13,18–20</sup> Recent *ab initio* electronic structure calculations on OB suggest that ultrafast dynamics are responsible for its efficacy as a sunscreen.<sup>13</sup> Specifically, these studies identify internal conversion (IC) *via* a barrierless electron-driven excited state hydrogen atom transfer (ESHT) as a plausible energy deposition mechanism along the enol  $R_{O-H}$  reaction coordinate, transforming to the keto conformer with increasing O–H bond distance. (Fig. 1, red).<sup>13</sup> Furthermore, theoretical studies of similar species, that contain hydrogen donor (OH)–acceptor (CO) sites in close proximity, indicate similar ultrafast relaxation mechanisms may be operative.<sup>21,22</sup> Support for the operation of such ESHT processes is provided by recent ultrafast transient absorption studies on OB following excitation in the UVA region<sup>8,9</sup> and by earlier work which identified ESHT as a relaxation process in related molecules like 2-(2'-hydroxyphenyl)benzothiazole (HBT) and 2-(2'-hydroxy-5'-methylphenyl)benzotriazole (TINUVIN).<sup>23–25</sup> The energy dissipation mechanism described following UVA excitation of OB involves successive ultrafast steps beginning with IC from the initially excited  $^1\pi\pi^*$  ( $S_2$ ) state to the lower lying  $^1n\pi^*$  ( $S_1$ ) state, followed by ultrafast enol  $\rightarrow$  keto tautomerisation (ESHT) on the  $S_1$  potential energy surface (PES). A slower C–C twist then facilitates nonadiabatic coupling to form highly vibrationally excited ground state ( $S_0$ ) molecules, relaxing predominantly to the enol-OB conformer through a combination of ground state hydrogen transfer (GSHT) and vibrational energy transfer (VET). Other recent studies have suggested some probability for excited state OB molecules to undergo intersystem crossing (ISC) to long-lived triplet states<sup>12</sup> as well as homolytic O–H bond fission yielding phenoxy radicals.<sup>9</sup> An interesting question that remains following these previous studies is whether OB dissipates electronic energy with similar efficiency

when photoexcited at shorter wavelengths – a property that should be required of any organic sunscreen constituent.

In this communication we utilise femtosecond pump–probe transient electronic (UV-visible) absorption spectroscopy (TEAS) to probe the ultrafast energy dissipation of OB following photoexcitation on absorption maxima within the UVB and UVC regions *ca.* 287 nm (4.32 eV) and *ca.* 243 nm (5.10 eV) respectively (Fig. 1). The observed dynamics are, again, highly suggestive of ultrafast excited state enol  $\rightarrow$  keto tautomerisation, followed by IC and VET to reform the ground state enol tautomer, as documented in the case of UVA photoexcitation.<sup>8,9</sup>

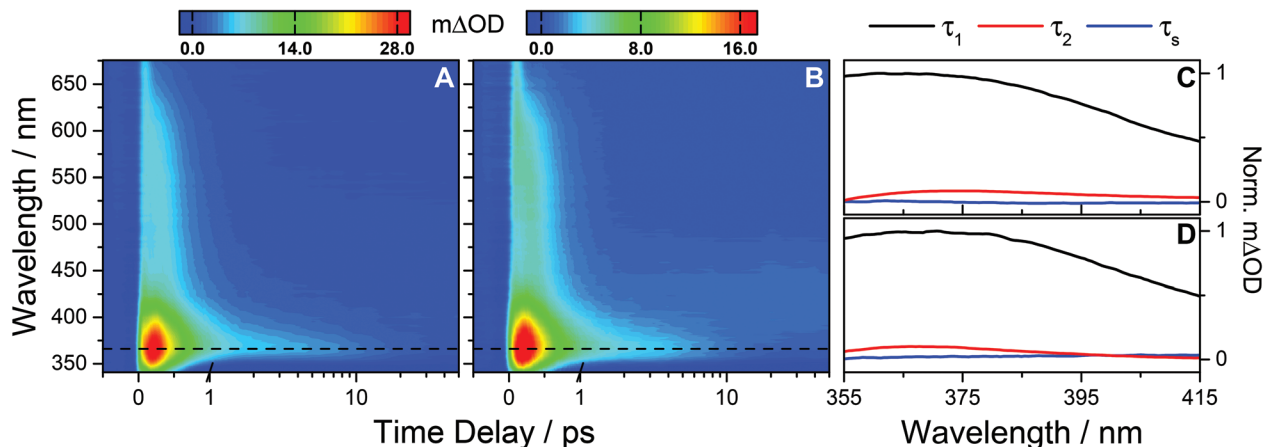
For all TEAS measurements, 10 mM solutions of OB (98%, Sigma-Aldrich), in either cyclohexane (>99%, VWR) or methanol ( $\geq 99.6\%$ , Sigma-Aldrich), were recirculated and delivered using a flow-through cell (Harrick Scientific), which is equipped with two  $\text{CaF}_2$  windows and a 100  $\mu\text{m}$  thick PTFE spacer. The sample is photoexcited using 287 nm or 243 nm,  $\sim 50$  fs pump pulses with fluences of 1–2  $\text{mJ cm}^{-2}$ . Probe pulses are derived from a broadband white light continuum (340 to 675 nm) and set to a time delay of up to 1.7 ns relative to the pump pulse. Probe pulse polarisation is held at the magic angle ( $54.7^\circ$ ) relative to the pump polarisation. Further experimental details are reported in ref. 8 and 26. All transient absorption spectra (TAS) are chirp corrected using the KOALA package<sup>27</sup> and reported lifetimes are determined using global fitting with uncertainties reported to a 95% confidence interval ( $2\sigma$ ) using support plane analysis, see ESI† for details.<sup>8,28–31</sup> All ‘static’ UV-visible spectroscopic measurements were taken using a Cary 50 UV-visible spectrophotometer with a 1 cm path length quartz cuvette, and  $\sim \mu\text{M}$  OB-cyclohexane and OB-methanol solutions.

Using a photoexcitation wavelength of 287 nm, the TAS of OB-cyclohexane and OB-methanol for a range of pump–probe time delays are shown in Fig. 2A and B, respectively. We con-



**Fig. 2** (A) Raw TAS following 287 nm photoexcitation of OB-cyclohexane displaying an absorption peak at 366 nm (dashed line) and a broad absorption extending out to  $\sim 650$  nm. The TAS for negative and early ( $\leq 1$  ps) time delays are shown on a linear scale, and longer time delays ( $> 1$  ps) are displayed on a logarithmic scale. The colour map indicates the change in optical density ( $\Delta\text{OD}$ ). Similar observations are seen for OB-methanol (B). (C) and (D) The corresponding decay associated spectra for OB-cyclohexane and OB-methanol respectively from the bi-exponential global fit across the spectral range 355–415 nm returning a femtosecond lifetime  $\tau_1$  and a picosecond lifetime  $\tau_2$ , as summarised in Table 1.





**Fig. 3** (A) Raw TAS following 243 nm photoexcitation of OB-cyclohexane displaying an absorption peak at 366 nm (dashed line) and a broad absorption extending out to  $\sim 650$  nm. The TAS for negative and early ( $\leq 1$  ps) time delays are shown on a linear scale, and longer time delays ( $>1$  ps) are displayed on a logarithmic scale. Similar observations are seen for OB-methanol (B). (C) and (D) The corresponding decay associated spectra for OB-cyclohexane and OB-methanol respectively from the bi-exponential global fit across the spectral range 355–415 nm returning a femtosecond lifetime  $\tau_1$  and a picosecond lifetime  $\tau_2$  as summarised in Table 1, with a solvent dependant lifetime,  $\tau_s$ , which is  $\sim 400$  ps for cyclohexane and  $\sim 3$  ns for methanol. We note the small amplitude of  $\tau_s$  reflecting the minor contribution of this component to the DAS.

sider OB-cyclohexane first. At early pump-probe time delays ( $<500$  fs) the TAS are dominated by two features. There is an intense absorption peak centred on  $\sim 366$  nm and a broad absorption extending out to  $\sim 650$  nm. Based on the analysis of similar absorption features observed after UVA excitation, and guided by *ab initio* calculations, this broad absorption is assigned to the OB excited state absorption (ESA),<sup>8,9,13</sup> likely to originate from a transition to a dense manifold of high-lying electronic states *i.e.*  $S_n \leftarrow S_1$ .<sup>32,33</sup> By 2 ps, the majority of the broad ESA feature has decayed away, leaving the 366 nm absorption peak. A negative feature between 340 and 350 nm is also present which we assign, through comparison with the static UV-visible of OB (Fig. 1), to a ground state bleach (GSB). By 50 ps, the 366 nm feature has almost decayed to the baseline, without full recovery of the GSB up to the maximum available pump-probe time delay of 1.7 ns. Similar dynamics are found for OB-methanol with a small observed blue-shift in the features discussed thus far.

Quantitative insight into the dynamical processes observed in the TAS is obtained by employing a global fitting procedure.<sup>28,29</sup> To recover the dynamics observed in the TAS, two exponential fitting functions are required over the spectral range of  $355 \leq \lambda \leq 415$  nm. This short wavelength limit is required to avoid the GSB, and includes only the 366 nm absorption peak. Furthermore, for all UVB measurements, a broad feature around time zero, unresolvable within our instrument response ( $\sim 100$  fs, see ESI†) and which shifts by  $\sim 50$  fs to longer time delays for UVC measurements (*vide infra*), is excluded by the omission of very early delay times ( $<250$  fs). For OB-cyclohexane, the transient signal is fitted using two lifetimes,  $\tau_1 = 391 \pm 12$  fs and  $\tau_2 = 8.6 \pm 2.7$  ps, with corresponding decay associated spectra (DAS) as shown in Fig. 2C. For OB-methanol, the corresponding lifetimes are  $\tau_1 =$

$382 \pm 8$  fs and  $\tau_2 = 6.0 \pm 1.0$  ps, with corresponding DAS as shown in Fig. 2D. The dynamics of the ESA feature for  $\lambda > 415$  nm is characterised by a single exponential decay function with a lifetime which closely resembles the  $\tau_1$  value for OB in the corresponding solvent (see ESI†).

Analogous experiments were performed for photoexcitation at 243 nm. TAS of OB-cyclohexane and OB-methanol are shown in Fig. 3A and B respectively. We observe similar spectral features as described in the above UVB case. Again, quantitative insight into the prevailing dynamical processes is obtained through global fitting. In the case of the UVC measurements, time delays  $<300$  fs were excluded in order to avoid the broad feature at very early time delays (*vide infra*). To recover the dynamics of the TAS over the spectral range of  $355 \leq \lambda \leq 415$  nm, three exponential functions are required; one of which is a long-lived solvent response with a lifetime  $\sim 400$  ps for cyclohexane and  $\sim 3$  ns for methanol, predetermined from fitted solvent-only TAS (see ESI†). The lifetimes of the other two processes are determined analogously to the UVB case. These are determined to be  $\tau_1 = 392 \pm 10$  fs and  $\tau_2 = 11.0 \pm 4.4$  ps for OB-cyclohexane and  $\tau_1 = 371 \pm 9$  fs and  $\tau_2 = 7.8 \pm 1.8$  ps for OB-methanol. The corresponding DAS are shown in Fig. 3C and D, respectively. The ESA feature for  $\lambda > 415$  nm is described by two exponential functions; a solvent response as given above, and another with a lifetime closely mirroring the  $\tau_1$  described above.

For all pump excitation wavelengths, TEAS measurements identify three dynamical features (excluding any solvent responses) which closely match those previously seen for OB photoexcited at 325 nm.<sup>8</sup> Two of these processes, a short lived process  $\tau_1$  and a longer lived process  $\tau_2$ , are extracted from the TAS by the global fitting of the 366 nm absorption maxima. The third feature is that previously described; a broad feature



which persists over time delays  $100 < \Delta t < 200$  fs, and which appears to shift  $\sim 50$  fs to longer time delays with increasing excitation energy (see Fig. 4A). In the case of 243 nm photoexcitation, there is a suspicion of the emergence of a 'double hump' structure in this broad feature, that is not evident in the corresponding data obtained when exciting in the UVA or UVB regions.<sup>8</sup> Furthermore, for all excitation wavelengths, we note that the GSB has not fully recovered, seen from the non-zero baseline out to the maximum delay time of our experiment (1.7 ns), see Fig. 4B.

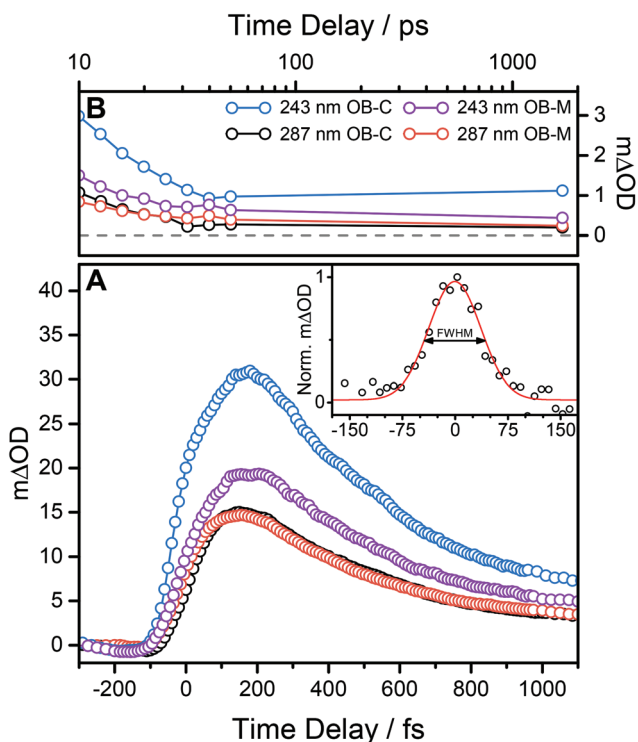
We can begin to rationalise these features drawing on *ab initio* calculations performed in previous studies.<sup>8,13</sup> Firstly, given the similarity of the TAS measured in both cyclohexane and methanol, we conclude that the intramolecular hydrogen bond in OB is preserved in both solvents, in both the ground<sup>13</sup> and excited states – as depicted in Fig. 1. Further support for this conclusion is provided by previous studies of HBT<sup>34</sup> and 4-*tert*-butylcatechol<sup>35</sup> for example. Both of these latter studies

show markedly different dynamics when photoexcited in polar and non-polar solvents, consistent with loss of the intramolecular hydrogen bond in suitably polar solvents. The signal broadening for higher energy excitations observed in Fig. 4, may be understood as IC from a photoexcited  $S_3$  ( $2^1\pi\pi^*$ ) state to  $S_1$  in the case of UVB excitation,<sup>13</sup> or, in the case of UVC excitation, from a higher lying  $S_n$  state ( $n \geq 4$ ) to  $S_1$ ; followed by ballistic ESHT on the  $S_1$  PES. We note that ESHT may occur before IC or indeed concomitantly. This mechanism fits well with the observed shift of the very early time feature (see Fig. 4A) given that IC from higher electronic states will take longer to relax to the  $S_1$  enol-OB tautomer before ESHT occurs. In such cases, one might speculate the appearance of a double hump, as these processes begin to emerge from our experimental temporal resolution of  $\sim 100$  fs.

After this ESHT, which drives enol  $\rightarrow$  keto tautomerisation, a slower rotation about the C–C bond may occur which facilitates IC back to the  $S_0$  state through a  $S_1/S_0$  conical intersection (nonadiabatic transfer); similar observations have been reported previously in the case of HBT.<sup>34</sup> We associate this process with the lifetime  $\tau_1$  which is reassuringly similar to those reported for comparable systems undergoing enol  $\rightarrow$  keto tautomerisation, which too require a twisted geometry rearrangement to facilitate IC.<sup>34,36–38</sup>

Thus we suggest that, following IC and formation of vibrationally hot  $S_0$  keto molecules, GSHT allows for keto  $\rightarrow$  enol tautomerisation. The enol tautomer subsequently cools *via* VET with the surrounding solvent, which we therefore attribute to the lifetime,  $\tau_2$ .<sup>8</sup> The absorption feature centred at 366 nm reflects this last step, whereby vibrationally hot molecules in the  $S_0$  state absorb the white light probe. The apparent sharpness of this feature is attributed to the spectral overlap of the (negative-going) GSB with the ESA. This model supports the determined globally fitted lifetimes, as summarised in Table 1. There is constancy in the  $\tau_1$  process when OB is excited by either a UVA, UVB or UVC pump pulse. As  $\tau_1$  represents the rotation around the C–C bond, this lack of change suggests that rotational motion about the C–C bond is unaffected by the increase in vibrational energy in the  $S_1$  molecules.

One might anticipate that, following  $S_1 \rightarrow S_0$  nonadiabatic transfer, the extra vibrational energy in the vibrationally hot  $S_0$  keto tautomer would lead to an increase in VET lifetime ( $\tau_2$ ). However, within a  $2\sigma$  uncertainty, our measurements of  $\tau_2$



**Fig. 4** (A) 366 nm transient displaying early time dynamics of OB-cyclohexane (OB-C) and OB-methanol (OB-M) after 287 nm and 243 nm excitation. For OB excitation at 243 nm we observe a shift of the 366 nm plateau relative to the same feature following 287 nm excitation. A representative normalised 366 nm cross correlation signal is shown in the inset of panel (A) for a methanol-only scan after 243 nm photoexcitation (black circles). The full width at half maximum is determined to be  $87 \pm 5$  fs from a Gaussian fit (red line). (B) Another observed feature for all TAS, that is also clearly shown for these transients, is the non-zero baseline out to the maximum available pump–probe time delay of 1.7 ns indicating incomplete GSB recovery, where the grey dashed line indicates zero. The transients shown are taken by integrating a 5 nm 'slice' about 366 nm.

**Table 1** Summary of the lifetimes of observed, dynamical processes of oxybenzone when photoexcited by 325 nm, 287 nm and 243 nm radiation. 325 nm values are taken from previous work.<sup>8</sup>

$\lambda$ /nm	Cyclohexane		Methanol	
	$\tau_1$ /fs	$\tau_2$ /ps	$\tau_1$ /fs	$\tau_2$ /ps
325	$375 \pm 13$	$7.8 \pm 2.8$	$368 \pm 13$	$4.9 \pm 1.9$
287	$391 \pm 12$	$8.6 \pm 2.7$	$382 \pm 8$	$6.0 \pm 1.0$
243	$392 \pm 10$	$11.0 \pm 4.4$	$371 \pm 9$	$7.8 \pm 1.8$



remain constant over different excitation energies, *cf.*  $7.8 \pm 2.8$  ps to  $8.6 \pm 2.7$  ps to  $11.0 \pm 4.4$  ps after UVA, UVB and UVC excitation respectively for OB-cyclohexane, indicating an insensitivity to this extra vibrational energy. Furthermore, we note that the  $\tau_2$  lifetime is solvent dependent in that the extracted lifetimes for OB-methanol are consistently shorter than the OB-cyclohexane counterparts. This implies higher VET efficiencies for OB-methanol than OB-cyclohexane, consistent with a more strongly (polar) interacting (hydrogen bonding) solvent.<sup>39</sup> Reassuringly, the lifetime  $\tau_2$  returned from fitting a single transient at 380 nm is consistently shorter than the 366 nm counterpart. This is unsurprising given that the higher rate of VET is likely to be associated with the more vibrationally excited  $S_0$  molecules, which absorb at longer wavelengths (*i.e.* red-shifted from 366 nm). Further details pertaining to this can be found in the ESI.†

The aforementioned non-zero baseline, that shows incomplete GSB recovery, implies that some of the hot  $S_0$  keto molecules do not reform the enol-OB tautomer, and instead form some long-lived (>1.7 ns) photoproduct. Given previous condensed phase observations, triplet state formation is one possible assignment,<sup>12</sup> although recent *ab initio* electronic structure calculations of likely triplet photoproducts coupled with transient vibrational absorption studies suggest such triplet states are negligible.<sup>8</sup> Other studies suggest that the formation of phenoxy radicals by O–H bond fission may explain incomplete GSB recovery,<sup>9</sup> although we note that neither previous similar studies<sup>8</sup> nor the current work detected any absorption attributable to the phenoxy radical within our signal-to-noise. Guided by recent *ab initio* calculations and experimental studies of similar systems,<sup>8,36–38,40</sup> we suggest the photoproduct to be a *trans*-keto tautomer formed from the extended rotation of the C–C bond of the vibrationally hot keto tautomer on the  $S_0$  PES. Ultimately, time-resolved infrared studies may hold the answers to the incomplete recovery of the GSB, but are beyond the scope of the present work. Finally we note that, irrespective of the initial photoexcitation energy (UVA, UVB or UVC) the observed early time dynamics reflect motion on the  $S_1$  PES, which is another manifestation of Kasha's rule.<sup>41</sup>

To conclude, we have determined a sequence of ultrafast relaxation steps by which OB molecules photoexcited in the UVB and UVC region reform the ground state enol-OB molecules with high efficiency. The non-radiative decay involves initial enol  $\rightarrow$  keto tautomerisation (*i.e.* ESHT) followed by a nonadiabatic transfer to the ground state, a reverse keto  $\rightarrow$  enol tautomerisation (*i.e.* GSHT) and subsequent vibrational cooling. This study extends previous work on OB relaxation dynamics after UVA excitation to reveal photostability with a similar relaxation mechanism across a broad region in the UV-visible spectrum. This indicates that OB is highly suitable as a UV absorber in sunscreens given its propensity to dissipate energy non-radiatively after photoexcitation in both UVA and UVB regions. The present results also suggest that OB is a suitable absorber in the UVC region, which could be important for industrial applications. We hope this work will serve as a

stimulus for further *ab initio* calculations and time-resolved infrared experiments to fully characterise the energy dissipation mechanism OB displays.

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