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# Pressure-driven phase transformations on $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$ huntite carbonate†

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Magnesium and calcium carbonate minerals are significant reservoirs of Earth's carbon and understanding their behavior under different conditions is crucial for elucidating the mechanisms of deep carbon storage. Huntite,  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$ , is one of the two stable calcium magnesium carbonate phases, together with dolomite. The distinctive cation coordination environment of Ca atoms compared to calcite-type and dolomite structures makes huntite a comparatively less dense phase. Here we examine the behavior of a polycrystalline natural huntite sample under room-temperature compression up to 38 GPa. Synchrotron X-ray diffraction and Raman spectroscopy experiments were carried out in a diamond-anvil cell using He as a highly hydrostatic pressure transmitting medium. XRD results suggest that the initial  $R\bar{3}2$  huntite structure persists up to 21 GPa. The Raman experiment agrees with this result but also suggests the appearance of structural defects from 10 GPa on. Birch–Murnaghan equation of state parameters were fit to the pressure–volume huntite data resulting in zero-pressure volume  $V_0$  of 611.7(2) Å<sup>3</sup>, a bulk modulus  $B_0$  of 99.5(11) GPa and a pressure derivative of the bulk modulus of  $B'_0 = 3.51(11)$ . At 21 GPa, huntite transforms to another trigonal phase ( $R\bar{3}$ ), designated here as huntite II. This phase persists up to at least 38 GPa, the maximum pressure reached in this study. The major structural differences between huntite and the huntite-II phase involve the tilting of the  $[\text{CO}_3]$  units with respect to the basal plane and a rotation, which cause a progressive change in the coordination number of the Ca atoms, from 6 to 9. DFT calculations complement the experimental data, providing new insights into the structural response to high-pressure conditions of this magnesium–calcium double carbonate mineral.

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## Introduction

The Earth's deep carbon cycle plays a critical role in regulating the planet's climate, atmospheric composition, and overall geochemical processes.<sup>1–4</sup> Central to this cycle are carbonate minerals, which act as significant reservoirs of carbon. Among these, magnesium (Mg) and calcium (Ca) carbonates are particularly important due to their abundance and stability in various geological settings. These minerals not only sequester carbon dioxide ( $\text{CO}_2$ ) from the atmosphere but also facilitate its

long-term storage within the Earth's crust and mantle. Understanding the behavior of Mg and Ca carbonates under different conditions is crucial for elucidating the mechanisms of deep carbon storage and transport.

Magnesite ( $\text{MgCO}_3$ , space group (SG)  $R\bar{3}c$ ), calcite ( $\text{CaCO}_3$ , SG.  $R\bar{3}c$ ) and dolomite ( $\text{MgCa}(\text{CO}_3)_2$ , SG.  $R\bar{3}$ ) are trigonal stable phases at room conditions, where the Mg and Ca atoms are octahedrally coordinated by O atoms. Trigonal structures are also adopted by other carbonates with small divalent cations (Fe, Ni, Zn, Mn, Co or Cd).<sup>5–7</sup> Aragonite, another  $\text{CaCO}_3$  polymorph, is an orthorhombic phase (SG.  $Pmcn$ ) where the Ca atoms have 9 O neighbor atoms and this structure-type is adopted by carbonates with larger divalent cations (Sr, Ba or Pb).<sup>8,9</sup> The  $\text{Ca}^{2+}$  cation size delimits the border of the (trigonal) calcite- to (orthorhombic) aragonite-type structures. The chemical stability of carbonates of divalent cations within the trigonal and orthorhombic groups therefore varies with cation radius, which is also directly related with its solubility behavior.<sup>10</sup> Ca carbonates are at the center of the trigonal–orthorhombic structural discontinuity and their crystal chemistry, in particular, the local environment around the Ca atoms, is not yet well understood.

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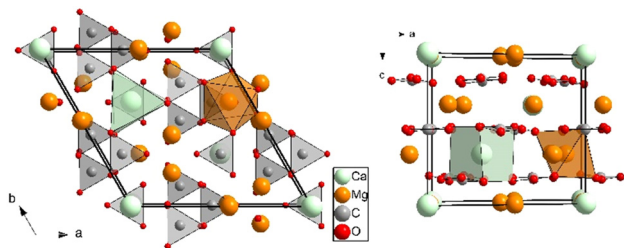
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**Fig. 1** Projections along the *c* (left) and *b* axes (right) of the crystalline *R*32  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$  huntite structure at room conditions. Thick black lines indicate the unit-cell. Light green, orange, light gray, and red spheres represent Ca, Mg, C, and O atoms, respectively. A  $[\text{CaO}_6]$  trigonal prism and a  $[\text{MgO}_6]$  octahedron are depicted to illustrate the different coordination sphere around the cation atoms.  $[\text{CO}_3]$  trigonal planar units are also represented.

Huntite  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$  mineral is produced in low-temperature surface or near surface environments by precipitation from magnesium-rich aqueous solutions or as a result of the interaction of such solutions with other carbonates.<sup>11–13</sup> Its composition lies halfway between  $\text{MgCO}_3$  magnesite and  $\text{MgCa}(\text{CO}_3)_2$  dolomite, and it is also a trigonal phase (SG. *R*32) but its structure consists of  $[\text{CO}_3]$  carbonate units roughly perpendicular to the *c* axis, 3/4 of these carbonate groups having the same orientation (see Fig. 1). This means that, whereas the Mg atoms are octahedrally coordinated by oxygen atoms, the Ca atoms are at the center of trigonal prisms. This particular oxygen arrangement around the Ca atoms produces a relatively low-density structure, with a density of  $2.875 \text{ g cm}^{-3}$  at ambient conditions (similar to that of dolomite and smaller than the density average of dolomite and magnesite,  $\sim 2.94 \text{ g cm}^{-3}$ ).

Magnesium and calcium carbonates have been widely studied at high-pressure (HP) and high-temperature (HT) conditions. The structural behavior of the Mg and Ca end-members of the  $\text{MgO-CaO-CO}_2$  carbonate system under compression and heating differ significantly.  $\text{MgCO}_3$  magnesite is stable throughout a wide pressure and temperature range, with just one HP-HT phase transition reported above 115 GPa and 2200 K.<sup>14,15</sup>  $\text{CaCO}_3$ , on the other hand, undergoes numerous high pressure phase transitions below 30 GPa.<sup>16,17</sup> Dolomite carbonate,  $\text{MgCa}(\text{CO}_3)_2$ , an intermediate composition between magnesite and calcite, has two room-temperature high-pressure polymorphs above 17 GPa and 35 GPa<sup>18,19</sup> and high P,T phase at 46 GPa and 1800 K.<sup>19</sup> The composition of  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$  huntite lies halfway between  $\text{MgCO}_3$  magnesite and  $\text{MgCa}(\text{CO}_3)_2$  dolomite and it is reported to be structurally stable up to 12 GPa at room temperature.<sup>20</sup> When heated below 6 GPa, huntite decomposes into  $\text{CaCO}_3$ ,  $\text{MgO}$  and  $\text{CO}_2$  at approximately 420 °C independently of pressure.<sup>20</sup>

Huntite's unique composition and structure provide valuable insights into the stability and behavior of mixed carbonates under various environmental conditions. In particular, the study of huntite upon compression will enhance our understanding of the coordination chemistry of the Ca atoms in carbonates and the stability fields of Mg and Ca carbonates. This work studies the structural and vibrational properties of

huntite up to 30 and 38 GPa, respectively, extending more than 3 times the previously reported pressure range. Synchrotron powder X-ray diffraction (XRD) and Raman spectroscopy measurements using quasi hydrostatic pressure medium, reveal a phase transition and DFT calculations confirm the stability of the dense polymorph.

## Experimental details

Naturally occurring huntite samples from the Tea Tree Gully in the South Australia State (Australia) were kindly provided by the Yale Peabody Museum (Specimen YPM MIN 035091). The chemical composition of the polycrystalline huntite sample was characterized by means of energy dispersive X-ray spectroscopy on a Philips XL30 scanning electron microscope. Only traces ( $<0.2$  atom %) of Sr and Na were found besides the Mg, Ca, C, and O atoms present in the  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$  huntite composition. XRD measurements carried out at room conditions confirmed that the sample has the previously reported structure.

Room temperature HP XRD experiments up to 30 GPa were carried out in a membrane-driven diamond-anvil cell (DAC) equipped with diamonds of 350  $\mu\text{m}$  of culet diameter. The huntite sample was placed together with a ruby chip at the center of a 150  $\mu\text{m}$  diameter 40  $\mu\text{m}$  thick pressure chamber drilled in a preindented rhenium gasket. High-purity helium (He) gas was also loaded in the DAC at room temperature using the Sanchez Technologies gas loading apparatus. Helium is fluid up to 12 GPa at room temperature<sup>21,22</sup> and acted as a quasi-hydrostatic medium in the pressure range of the study.<sup>23</sup> Pressure was determined using the ruby fluorescence method.<sup>24</sup> *In situ* angle-dispersive powder X-ray diffraction measurements were carried out at MSPD beamline at ALBA-CELLS synchrotron using an X-ray wavelength of 0.4246 Å.<sup>25</sup> This beamline provides an X-ray beam focused down to  $\sim 20 \times 20 \mu\text{m}^2$ , and the diffracted signal was collected with a Rayonix CCD detector. The detector parameters were calibrated with  $\text{LaB}_6$  powder standard, and integration to conventional  $2\theta$ -intensity data was carried out with the Dioptas software.<sup>26</sup> The indexing and refinement of the powder patterns were performed using the Unitcell,<sup>27</sup> Powdercell<sup>28</sup> and Fullprof<sup>29</sup> program packages.

Raman spectra were measured in 2–3 GPa steps upon compression and decompression covering a range between ambient pressure and 38 GPa. A frequency-doubled 532.14 nm Nd:YAG Oxxius laser (LCX-532S) was focused on the sample with a spot size of 6  $\mu\text{m}$ . Spectra were collected in backscattering geometry, using a grating spectrometer (Acton, SP-2356) equipped with a Pixis 256E charge coupled device (CCD) detector and a Mitutoyo microscope objective. The spectral resolution of the spectrometer is  $3 \text{ cm}^{-1}$ .<sup>30,31</sup> The laser power was set to 80 mW and spectra were collected for 300 s in a frequency window of 100–1300  $\text{cm}^{-1}$ , using a grating of 1800 grooves per mm.

## Computational details

Density-functional theory (DFT) calculations were carried out in the huntite structure using Quantum ESPRESSO (QE), version 6.5.<sup>32</sup> The projector augmented wave (PAW) method<sup>33</sup> was used in combination with the B86bPBE-XDM functional<sup>34–38</sup>



implemented in QE. PAW datasets were obtained from the pslibrary<sup>39</sup> for all atoms, with number of valence electrons equal to 10 (Ca), 4 (C), 10 (Mg), and 6 (O). The kinetic energy cutoff for the plane-wave expansions of the wavefunction and the electron density were 100 Ry and 1000 Ry, respectively, and a uniform  $k$ -point grid of  $2 \times 2 \times 2$  was used for the huntite structure and doubled dolomite structures. The experimental zero-pressure geometry of huntite was relaxed at zero pressure and 50 GPa. All geometry relaxations were carried out with tight convergence parameters:  $10^{-5}$  Ry in the energy and  $10^{-4}$  Ry bohr<sup>-1</sup> in the forces. The equilibrium volumes resulting from these initial relaxations were used to set up a uniformly spaced volume grid with 41 points spanning the corresponding pressure range. Constant-volume geometry relaxations were carried out at each of the 41 points.

To probe the dynamical instability of the calculated structure, phonon frequency calculations were carried out with the density-functional perturbation theory method.<sup>40</sup> Uniform reciprocal-space  $q$ -point grids with size  $2 \times 2 \times 2$  were used. As described below, our calculations predict that the  $R32$  structure becomes dynamically unstable under pressure resulting in a broken-symmetry  $R3$  structure, in agreement with experiment. For this reason, we calculated both the  $R32$  structure, dynamically unstable at high pressure, and the  $R3$  structure, which coincides with  $R32$  at low pressure and is stable at high pressure.

The equation of state and the transition pressures were calculated with the gibbs2 code,<sup>41,42</sup> using the data generated by Quantum ESPRESSO. To compare to the experimental data, Raman spectra were simulated at selected pressures in both the  $R32$  and  $R3$  phases. The QE equilibrium structures at a given pressure was relaxed using the crystal17 program,<sup>43</sup> employing the pob-DZVP-rev2 double-zeta basis sets.<sup>44</sup> Constant-volume geometry relaxations were carried out using the B3LYP-D3<sup>45–47</sup> functional with the same  $k$ -point grid as in the QE calculation. Raman intensities were then obtained using the coupled-perturbed Kohn–Sham method.<sup>48,49</sup>

## Results and discussion

### XRD measurements

The structural evolution of huntite,  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$ , at high pressure and room temperature has been studied by means of *in situ* powder synchrotron X-ray diffraction measurements. The experiments were performed in hydrostatic conditions using He as pressure transmitting medium. Fig. 2 illustrates a selection of patterns at different pressures, which shows that there are no obvious changes in the XRD profiles except for the shift of the peaks to higher  $2\theta$  angles as a result of unit cell compression. High-pressure synchrotron powder XRD patterns present some texturing effects due to uneven crystal sizes (see the bottom of Fig. 2). Despite the fact that, in our case, this can cause small inaccuracies in the relative intensities of the diffraction maxima, we were able to perform full structural Rietveld refinements. The lattice parameters and the atomic coordinates at the lowest pressure (1.05 GPa) are consistent with those of the previously

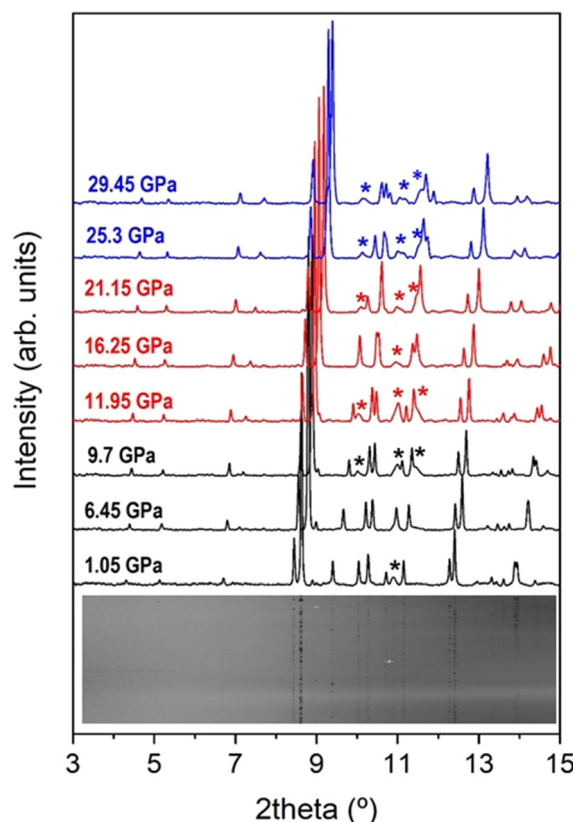


Fig. 2 Selected powder XRD patterns of huntite up to 29.5 GPa ( $\lambda = 0.4246$  Å). All the patterns can be indexed with a trigonal unitcell and no significant peak intensity changes are observed in the studied pressure range. The diffraction signals corresponding to the rhenium of the gasket are marked with asterisks. The different colors denote the pressure ranges where indications of phase transitions were found either by discontinuities in lattice parameters dimensions or changes in the Raman spectra. The unrolled (or cake) image of the raw data from the CCD detector at 1.05 GPa and room temperature is given to illustrate the quality of our data.

reported huntite structure<sup>12,13,20</sup> (see Fig. S1 and Table S1 of ESI†). Raman spectroscopic results at room conditions, as we will discuss later, and DFT calculations also confirm that the low-pressure phase is the  $R32$  huntite structure.<sup>12,13,20</sup>

A short description of huntite needs to be done in order to understand its atomic arrangement and the phase transformations that could occur at high pressures. The lattice parameters of huntite are related to those of calcite (or magnesite) by  $a_{\text{hunt}} \sim 2 \times a_{\text{calc}}$ ;  $c_{\text{hunt}} \sim c_{\text{calc}}/2$ . The relationship to calcite arises from the ordering of the cations on the expanded cation layer, with a stoichiometric Mg/Ca ratio of 3/1.

Two crystallographically distinct carbonate groups are present in the structure: one parallel to the  $ab$  basal plane of the cell and three symmetry-related  $[\text{CO}_3]$  slightly canted to the basal plane ( $6.4^\circ$  tilt), the orientation of the former rotated  $30^\circ$  with respect to the others (see Fig. 1). In this atomic arrangement, Mg atoms are octahedrally coordinated by O atoms and Ca atoms are hexa-coordinated adopting a trigonal prism configuration, unlike the calcite, magnesite and dolomite structures, where all the cations are forming  $[\text{MO}_6]$  octahedral units.



All the high-pressure XRD patterns can be indexed and fitted using a hexagonal unit cell. The experimental lattice parameters and unit cell volumes at different pressures are collected in Table 1 and their evolution as a function of pressure is represented in Fig. 3 and 4, respectively. As can be seen, the lattice parameters vary smoothly with increasing pressure up to 21 GPa, in reasonable agreement with theory (also depicted in Fig. 3 and 4), which supports the absence of first-order phase transitions in this pressure range.

The calculated volume is overestimated by 1.6% with respect to the experimental value at room conditions, as a consequence of the 0.7% overestimation of the *a* and *b* axis lengths. The *a/c* ratio is also affected by this small difference, being its 0.6% larger in the simulations but presenting a similar slope upon compression. The pressure evolution of the lattice parameters (Fig. 3) shows that the *c* axis is the most compressible. A linearization of the 3rd-order Birch–Murnaghan equation of state (through the substitution of the volume for the cube of a lattice parameter) was used to analyze the axial compression,<sup>50</sup> providing evidence of the strong anisotropy in this compound. Thus, the bulk moduli values obtained from data up to 21 GPa are  $B_a = 151(3)$  GPa ( $B'_{0a} = 6.0(3)$ ) and  $B_c = 56.8(5)$  GPa ( $B'_{0c} = 2.46(4)$ ), for the *a* and *c* axes, respectively, which indicate that the *a* axis is prominently more uncompressible and that the huntite bulk compressibility is mainly dominated by that of the *c* axis. The *a/c* ratio increases with pressure according to the expression  $a/c = 1.2163(2) + 0.00405(5) \cdot P - 3.07(2) \cdot P^2$  (see Fig. 5). This response to external pressure arises from the fact that the relatively incompressible  $[\text{CO}_3]$  carbonate units are

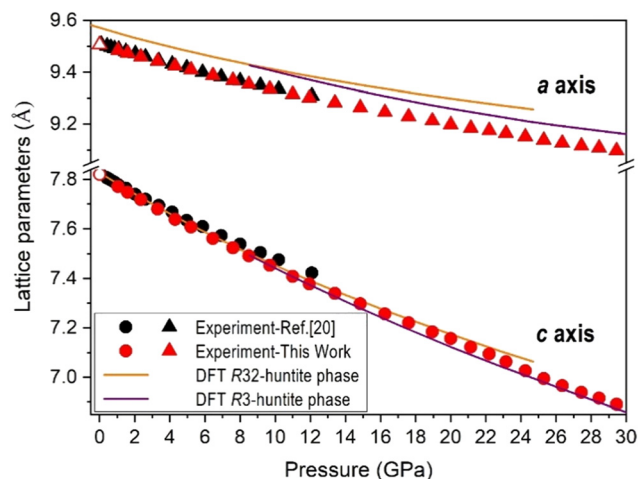


Fig. 3 Evolution of the lattice parameters of  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$  huntite with pressure. Triangles and circles represent the dimensions of the *a* and *c* axes of the trigonal structure. Red and black symbols correspond to results of the present study and previous results from ref. 20, respectively. Solid and empty symbols correspond to upstroke and downstroke data, respectively. The evolution of the DFT calculated lattice parameters of both the initial R32 and the predicted R3 phases as a function of pressure are represented as orange and purple lines, respectively.

arranged roughly parallel to the *ab* plane and, therefore, the compressibility of the *c* axis can be directly attributable to that of  $[\text{CaO}_6]$  trigonal prisms and  $[\text{MgO}_6]$  octahedra. The unit cell volumes as a function of pressure can be explained with a third-order Birch–Murnaghan equation of state (BM-EoS)<sup>51</sup> with a zero-pressure volume  $V_0 = 611.7(2) \text{ Å}^3$ , a bulk modulus of  $B_0 = 99.5(11)$  GPa, and a bulk modulus first-pressure derivative of  $B'_0 = 3.51(11)$ . These experimental results show a slightly larger compressibility than that reported using a less-hydrostatic pressure-transmitting medium (a mixture 4 : 1

Table 1 Experimental unit cell parameters and volumes of huntite and huntite-II phases at different pressures

Pressure (GPa)	Space group	<i>a</i> axis (Å)	<i>c</i> axis (Å)	<i>a/c</i> ratio	Volume (Å <sup>3</sup> )
1.05	R32	9.4834(6)	7.7712(8)	1.2203(2)	605.27(10)
1.6	R32	9.4731(6)	7.7484(9)	1.2226(2)	602.19(11)
2.35	R32	9.4590(7)	7.7185(11)	1.2255(3)	598.077(13)
3.3	R32	9.4418(8)	7.6802(9)	1.2294(2)	592.95(12)
4.3	R32	9.4232(7)	7.6382(11)	1.2337(3)	587.38(13)
5.2	R32	9.4092(4)	7.6070(7)	1.2369(2)	583.25(8)
6.45	R32	9.3869(4)	7.5609(6)	1.2415(2)	576.97(7)
7.6	R32	9.3686(4)	7.5236(7)	1.2452(2)	571.89(8)
8.5	R32	9.3527(5)	7.4909(7)	1.2485(2)	567.47(8)
9.7	R32	9.3349(5)	7.4527(7)	1.2526(2)	562.43(8)
11	R32	9.3141(6)	7.4083(7)	1.2573(2)	556.59(9)
11.95	R32	9.2995(8)	7.3776(12)	1.2605(3)	552.55(14)
13.4	R32	9.2818(16)	7.3391(19)	1.2647(5)	547.6(2)
14.85	R32	9.2634(7)	7.2974(6)	1.2694(2)	542.30(9)
16.25	R32	9.2450(10)	7.2558(9)	1.2742(3)	537.07(12)
17.6	R32	9.2284(9)	7.2201(8)	1.2782(3)	532.512(11)
18.95	R32	9.2123(15)	7.184(2)	1.2824(6)	528.0(3)
20.0	R32	9.1976(8)	7.1561(7)	1.2853(2)	524.28(10)
21.15	R3	9.1841(15)	7.1217(18)	1.2896(5)	520.2(2)
22.2	R3	9.1754(16)	7.0936(18)	1.2935(6)	517.2(2)
23.15	R3	9.1643(6)	7.0624(6)	1.2976(2)	513.67(8)
24.25	R3	9.1507(6)	7.0255(8)	1.3025(2)	509.47(9)
25.3	R3	9.1380(8)	6.9945(9)	1.3065(3)	505.82(11)
26.35	R3	9.1268(12)	6.9655(13)	1.3103(4)	502.49(16)
27.45	R3	9.1157(10)	6.9381(10)	1.3139(3)	499.29(13)
28.45	R3	9.1077(8)	6.9151(8)	1.3171(3)	496.76(10)
29.45	R3	9.0971(11)	6.8906(10)	1.3202(4)	493.85(13)

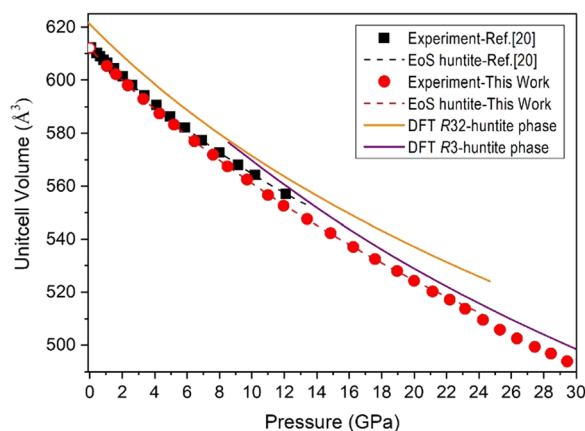


Fig. 4 Unit-cell volume vs. pressure for  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$ . Black squares and red circles represent experimental data from a previous study<sup>20</sup> and the present study, respectively, and the dashed lines the fitted equation of state. Solid and empty symbols correspond to upstroke and downstroke data, respectively. The evolution of the DFT calculated unitcell volumes of both the initial R32 and the predicted R3 phases as a function of pressure are represented as orange and purple lines, respectively.



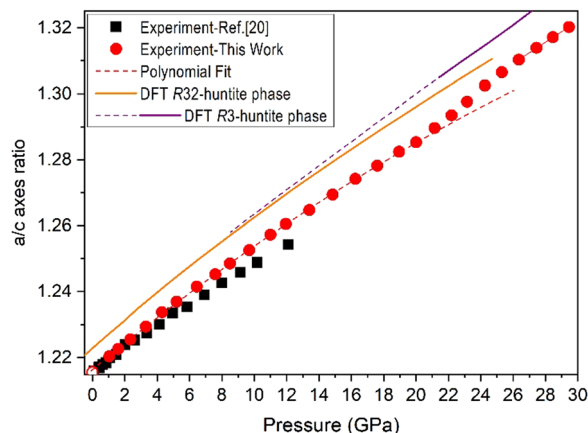


Fig. 5 Evolution of the  $a/c$  ratio of  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$  huntite with pressure. Black squares and red circles represent experimental data from a previous study<sup>20</sup> and the present study, respectively. The dashed line is a 2nd-order polynomial fitting to the low-pressure phase. The evolution of the DFT calculated ratios of both the initial  $R32$  and the predicted  $R3$  phases as a function of pressure are represented as orange and purple lines, respectively.

methanol:ethanol) up to 12 GPa<sup>20</sup> and compare well with those obtained from our DFT calculations below 10 GPa:  $V_0 = 621.046 \text{ \AA}^3$ ,  $B_0 = 99.55 \text{ GPa}$ , and  $B'_0 = 4.48$ . The bulk modulus of  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$  huntite is slightly larger than that of  $\text{MgCa}(\text{CO}_3)_2$  dolomite ( $B_0 = 94 \text{ GPa}$ )<sup>52</sup> and it is only slightly smaller to that of  $\text{MgCO}_3$  magnesite ( $B_0 = 103\text{--}108 \text{ GPa}$ ).<sup>53</sup> Huntite's intermediate bulk modulus results from its mixed cation structure, where one fourth of calcium atoms and three-fourth magnesium atoms alternate within the crystal lattice. The trigonal prism coordination of the Ca atoms does not seem to have a significant effect on compressibility despite its higher polyhedral volume.

At approximately 21 GPa, a discontinuity in the evolution of the hexagonal lattice parameters starts to occur. Between this pressure and 25 GPa, the  $a/c$  ratio increases more pronouncedly as a consequence of a larger decrease of the  $c$ -axis, which suggests the existence of a progressive second-order phase transition between hexagonal crystal structures. Above 25 GPa, that sudden increase in slope is moderated and the  $a/c$  ratio increases smoothly according to the following linear expression:  $a/c = 1.2259(5) + 0.00321(17) \cdot P$ . The intensities of the XRD peaks of the HP patterns show small changes upon compression, which points to small atomic displacements that could reduce the symmetry of the unit cell lattice. DFT calculations, as we will see later in detail, predict that the  $R32$  structure is dynamically unstable above 8 GPa, and suggest a lower symmetry  $R3$  at high pressures. Using the atomic coordinates given by the simulations for the HP  $R3$  structure as a reference model, we performed a Rietveld refinement at 25.3 GPa, above the observed transition (see Fig. S2 and Tables S2, S3 in ESI†). The results show the atoms slightly rearrange as a consequence of the disappearance of the 2-fold axes, which entails both a rotation and a larger tilting of the  $[\text{CO}_3]$  groups causing a distortion of the cation-centered polyhedra. In fact, Ca atoms can be considered as 9-fold coordinated by O atoms in the HP  $R3$  phase ( $3 \times 2.18 \text{ \AA}$ ,  $3 \times 2.37 \text{ \AA}$  and  $3 \times 2.68 \text{ \AA}$  at 25.3 GPa), whereas in the low-pressure  $R32$  phase Ca

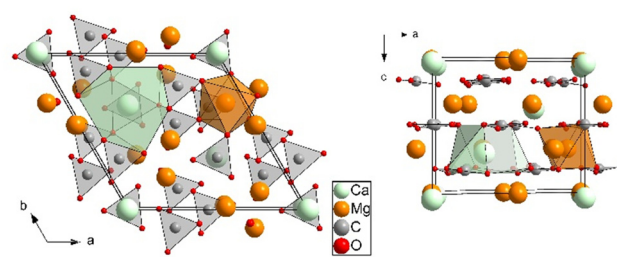


Fig. 6 Projections along the  $c$  (left) and  $b$  axes (right) of the crystalline high-pressure  $R3 \text{ Mg}_3\text{Ca}(\text{CO}_3)_4$  huntite structure. Thick black lines indicate the unit-cell. Light green, orange, light gray, and red spheres represent Ca, Mg, C, and O atoms, respectively. A  $[\text{CaO}_9]$  polyhedron and a  $[\text{MgO}_6]$  octahedron are depicted to illustrate, in particular, the increase in the coordination sphere around the Ca atoms.  $[\text{CO}_3]$  trigonal planar units are also represented and are significantly more tilted than in the initial huntite phase.

are hexacoordinated forming trigonal prisms ( $6 \times 2.37 \text{ \AA}$  at 1.05 GPa). The new Ca-centered polyhedra are depicted in Fig. 6 (to be compared to Fig. 1). The phase transition is reversible upon decompression, as inferred by the unitcell volume and  $a/c$  ratio.

### Raman spectroscopy measurements

We performed Raman spectroscopy measurements of huntite up to 38 GPa also using He as hydrostatic pressure transmitting medium. In general, the spectra of carbonates can be divided into two regions.<sup>54,55</sup> Those bands at frequencies above  $600 \text{ cm}^{-1}$  are mainly due to the internal motions of the molecular carbonate ion. Those below  $600 \text{ cm}^{-1}$  are generic lattice modes. Group theoretical analyses (point group  $D_3$ ) predict 26 Raman active modes  $\Gamma_R = 7A_1 + 19E$ .<sup>56,57</sup>

Raman scattering spectra at some representative pressures are shown in Fig. 7 (top) and (bottom). The huntite sample from Tea Tree Gully, Australia, gave spectra with well-defined bands. A complete band assignment was made by Scheetz and White<sup>57</sup> by carefully comparing IR and Raman measurements with the predictions of the factor-group calculation. Assignments can be summarized as follows:

(i) There are two crystallographically distinct carbonate ions in the huntite structure. A unique  $[\text{CO}_3]$  group which lies in the basal plane of the trigonal cell and has point symmetry  $D_3$ . The remaining three  $[\text{CO}_3]$  groups are arranged around the three-fold axis and are slightly tilted towards the basal plane. All three are crystallographically equivalent with site symmetry  $C_2$ . The orientation of these two types of carbonate groups is different, one rotated  $30^\circ$  with respect to the other. Only one strong symmetric C–O stretching band appears at  $1120 \text{ cm}^{-1}$  in the Raman spectra at room and low pressures (below 3.9 GPa) due to the fact that the bands that arise from the two different crystallographic carbonate units and from combinations of the three symmetry-related equivalent  $[\text{CO}_3]$  groups have similar frequencies.

Above 3.9 GPa, this band splits into two well-defined Raman bands due to the different pressure-induced vibrational frequency shifts of the C–O stretching bands caused by the change in its atomic environment. The evolution of the Raman band frequencies with increasing pressure is depicted in Fig. 8. In this



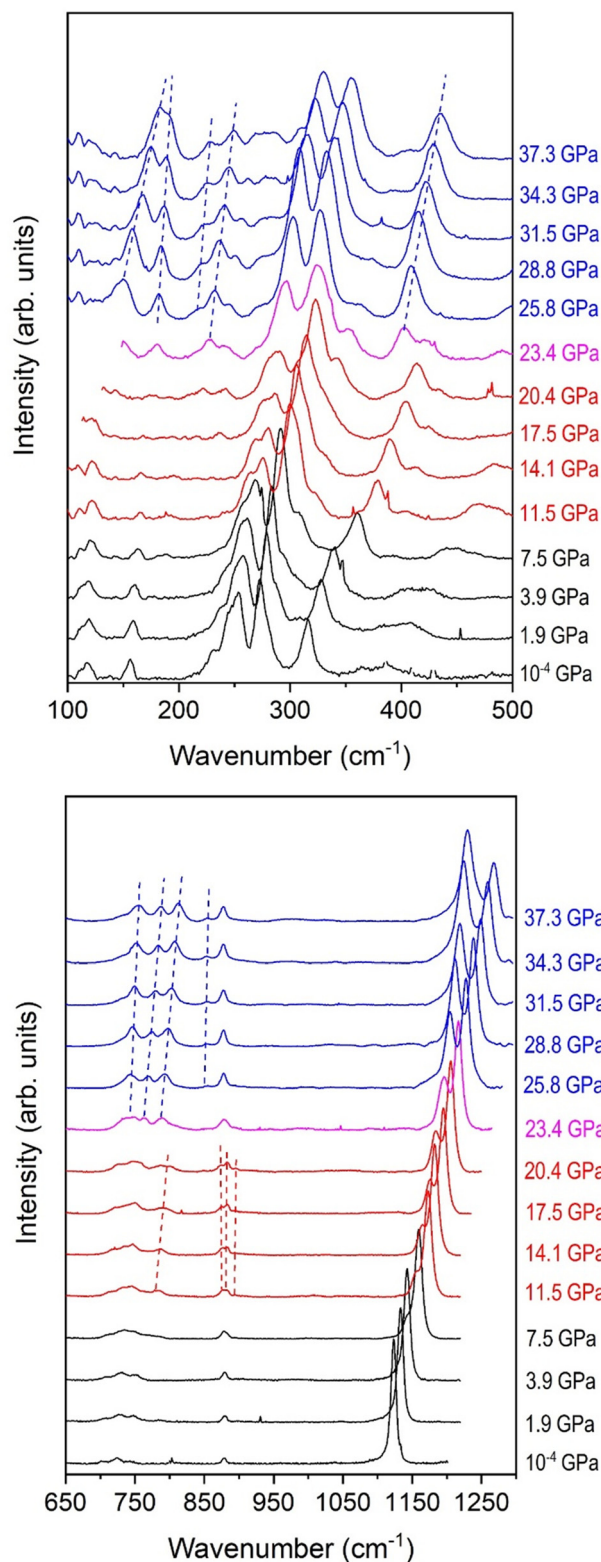


Fig. 7 Room-temperature Raman spectra of  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$  at selected pressures on upstroke in the frequency regions 100–500  $\text{cm}^{-1}$  (top) and 650–1300  $\text{cm}^{-1}$  (bottom). The black and blue spectra correspond to the initial *R32* huntite and the high-pressure *R3* phase, respectively. Red spectra correspond to the pressure range where most of the bands can be assigned to the initial *R32* huntite phase but small peaks and a band splitting in the  $[\text{CO}_3]$  bending frequency region appear. Dashed lines show the appearance and evolution of new Raman bands.

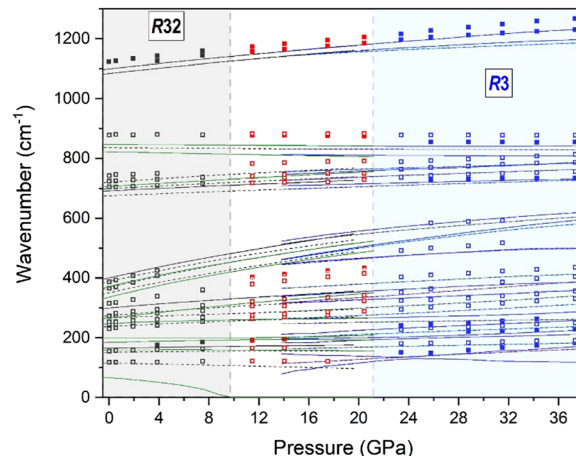


Fig. 8 Experimental (symbols) and calculated (lines) pressure dependence of the Raman-active modes of  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$ . Black and blue symbols represent bands of the initial *R32* huntite and the high-pressure *R3* phase, respectively. Red symbols correspond to the pressure range where most of the bands can be assigned to the initial *R32* huntite phase but small peaks and a band splitting in the  $[\text{CO}_3]$  bending frequency region appear.

figure it can be seen that the frequency difference between the two C–O stretching bands progressively and slowly increases upon compression. On the higher frequency region, at around 1460  $\text{cm}^{-1}$ , an antisymmetric C–O stretching mode appears. This band was observed at room conditions but, unfortunately, in static high pressure experiments the first-order Raman band of diamond anvils prevents access to the 1300–1500  $\text{cm}^{-1}$  frequency range. The bands between 705 and 880  $\text{cm}^{-1}$  correspond to  $[\text{CO}_3]$  bending modes. At 11.5 GPa, the Raman mode of  $\sim 880 \text{ cm}^{-1}$  splits into two components and a new Raman band appear at 782  $\text{cm}^{-1}$ . Above 23 GPa, where the axial ratio discontinuity occurs, more changes occur in the vibrational spectra, appearing a new Raman band at 854  $\text{cm}^{-1}$  and a triplet with well-defined bands between 725 and 830  $\text{cm}^{-1}$ .

(ii) A detailed classification of the lattice modes, those below 420  $\text{cm}^{-1}$ , was done by Scheetz and White based on factor-group analysis.<sup>57</sup> In Fig. 7 and 8 it can be seen the general blueshift of the bands of huntite with increasing pressure, which do not present significant broadening as expected by the crystallinity of the solid and the hydrostaticity provided by the pressure transmitting medium. Above 23 GPa, there are changes in the relative intensities of the modes in this low frequency region and the appearance of new modes like the one at 403  $\text{cm}^{-1}$ . These observations could be explained by the continuous change towards a higher coordination number of the Ca atoms induced by pressure, as revealed in XRD experiments and suggested in DFT calculations (see below).

The calculated Raman spectra are in reasonable agreement with those experimentally measured at low and high pressures, as shown in Fig. 9. The initial huntite spectrum was observed on pressure release, confirming the reversibility of the phase transitions.

### Phase transitions

The behavior of the *c/a* ratio above 21 GPa, as characterized by XRD, as well as the changes observed in the Raman spectra,



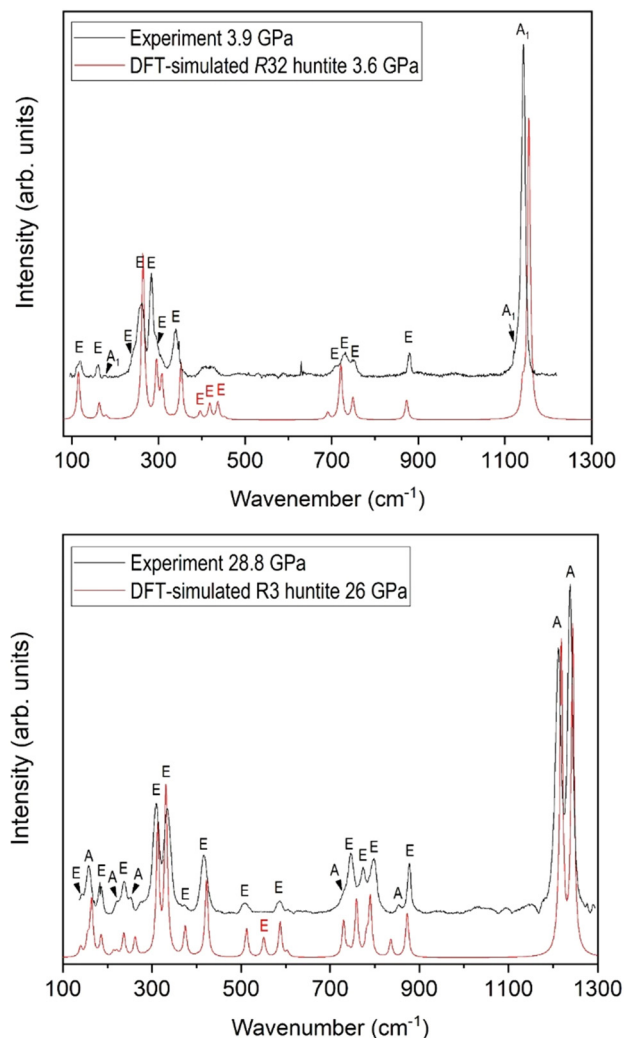


Fig. 9 Comparisons between calculated (red lines) and experimental (black lines) Raman spectra of the low-pressure  $R32$  (top) and the high-pressure  $R3$  (bottom) phases of  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$  at 3.9 and 28.8 GPa. Calculated Raman spectra were used for the mode assignment of Raman frequencies in the experiment.

suggest the existence of a second order phase transition that takes place between 21 and 25 GPa. Although this is the main structural change observed, there are other subtleties which is worth commenting on.

DFT calculations reveal that the initial  $R32$  structure of  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$  is dynamically unstable above 8 GPa, the pressure where the first additional features start to appear in the Raman spectra. The dynamical instability of the system at this pressure is revealed by the existence of imaginary phonon frequencies at Gamma. Subsequent geometry relaxation of the structure at the same volume leads to the broken-symmetry  $R3$  structure, which has been verified to be dynamically stable by performing further DFPT phonon frequency calculations. The high-pressure  $R3$  structure is energetically more favorable above this pressure, as indicate the E–V and H–P graphs represented in Fig. 10.

The question arises as to whether the changes observed in the Raman spectra at 10 GPa correspond to an additional phase

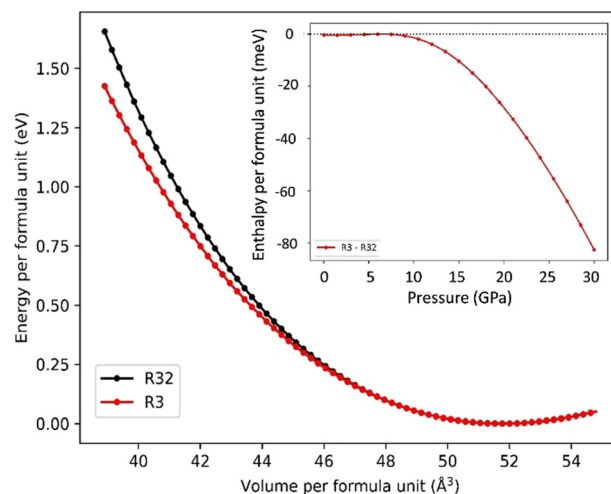


Fig. 10 DFT energy as a function of the volume per  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$  formula unit for the initial  $R32$  phase and the HP  $R3$  phase. Inset: Enthalpy difference as a function of pressure, showing the stability with respect to  $R32$   $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$  huntite.

transition or whether they are related to precursor effects to the transition to the  $R3$  structure at 21 GPa. We have explored two possibilities.

Firstly, we envisaged a smooth transition from 10 GPa onwards, which would produce gradual changes on the Raman spectra as the distortion of the new phase increases. The twelve  $A_2$  modes of the huntite  $R32$  phase are IR active but Raman forbidden. Space group correlations between the 32 and 3 point groups show that the  $A_2$  modes give rise to  $A$  modes in the  $R3$  structure, which are both Raman and IR active. The Raman tensor of the new modes would increase in parallel to the distortion associated with the  $R3$  phase, and for some modes it would not result in relevant intensities up to 21 GPa. We have plotted the calculated pressure dependence of the IR active modes in Fig. 8 as green lines. At first sight, at 10 GPa there are some of the new modes which are close enough to the calculated values to be assigned to IR modes, such as the one observed at approx.  $865\text{ cm}^{-1}$ . However, this line of reasoning cannot explain why there are some modes, such as the one appearing at around  $400\text{ cm}^{-1}$  at 12 GPa, which disappears from 22 GPa on.

Another possibility is to consider a progressive appearance of structural defects starting at 10 GPa. The structural defects would relax the momentum conservation rule and allow Raman dispersion from phonons with wavevectors away from the Brillouin zone center. The most relevant vibrations involve wavevectors at edge points and flat bands, whose associated density of states is large. A similar reasoning was used in ref. 58 to explain Raman features appearing prior to the phase transition in the  $\text{CuAlO}_2$  delafossite. A relevant section of the huntite phonon dispersion curves is shown in Fig. 11. States with a large density of states are highlighted with red dashed lines. There are three relevant contributions to the Raman scattering that could explain the progressive transformation of the Raman peak at  $890\text{ cm}^{-1}$  into a triplet from 10 GPa on. Interestingly,





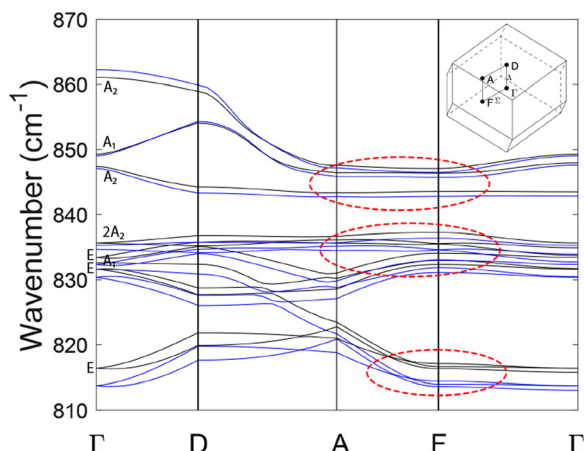


Fig. 11 Dispersion curves along a path in the Brillouin zone, including  $\Gamma(0\ 0\ 0)$ ,  $D(0\ 0\ 1/3)$ ,  $A(-1/2\ 1\ 1/3)$  and  $F(-1/2\ 1\ 0)$  points. Coordinates are given in fractional coordinates of the conventional (hexagonal) unit cell. Black and blue lines correspond to calculations at 7.6 and 11.1 GPa, respectively. Dashed red lines highlight states with a high density of states which may contribute to Raman dispersion.

the low energy side of the triplet is predicted to be soft ( $-0.8\text{ cm}^{-1}\text{ GPa}^{-1}$ ), as it is actually observed in the experiment ( $-0.3\text{ cm}^{-1}\text{ GPa}^{-1}$ ). Since the precursor defects do not significantly affect long-range order, this mechanism would be compatible with the XRD results, which show no evidence of the phase transition up to 21 GPa.

## Conclusions

We report an experimental and theoretical work on the behavior of  $\text{Mg}_3\text{Ca}(\text{CO}_3)_4$  huntite under high pressure in which we have studied the stability of the initial  $R32$  phase from a mechanical and dynamical perspective. Our experimental XRD and Raman spectroscopy results confirm that there is a  $R32$  to  $R3$  phase transition between 21 and 25 GPa, which entails small atomic displacements allowed by the symmetry break and a change in the coordination environment of the Ca atom. This atom goes from being hexacoordinated by O atoms and adopting a trigonal prism configuration in the initial huntite phase to an irregular 9-fold coordination in the high-pressure phase. DFT calculations also predict the dynamical instability of the initial  $R32$  structure at high pressures and the stability of the  $R3$  phase. The Raman study also reveals a pressure range (11.5–20.5 GPa) where significant changes in the Raman spectrum are observed, although most of the Raman peaks can still be assigned to the initial  $R32$  huntite phase. The analysis of the phonon dispersion curves suggests that these changes could be related to structural defects that are precursors to the phase transition. DFT calculations reveal that the broken-symmetry  $R3$  phase is dynamically stable in the 8–30 GPa pressure range. After decompression, the recovered sample is the initial huntite phase, showing the reversibility of the structural transition.

## Author contributions

Conceptualization, D. S. P., investigation, D. S. P., R. Ch. J., B. D. B. N., G. B., J. P. P., L. B. A. O. R., and C. P., formal analysis, R. Ch. J., D. S. P., and A. O. R., writing – original draft, D. S. P., writing – review & editing, all the authors.

## Data availability

Crystallographic data for huntite-II has been deposited at the CCDC database under number 2393118. The data supporting this article have been included as part of the ESI.†

## Conflicts of interest

There are no conflicts to declare.

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