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# Spiro donor–acceptor TADF emitters: naked TADF free from inhomogeneity caused by donor acceptor bridge bond disorder. Fast rISC and invariant photophysics in solid state hosts†

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We have studied the thermally activated delayed fluorescence (TADF) properties of the spiro-bridged donor–acceptor molecule, 10-phenyl-10*H*,10'*H*-spiro[acridine-9,9-anthracen]-10'-one, (ACRSA) in guest–host films and used it as a probe to explore the details of host effects on the TADF mechanism in typical OLED host materials. Linked by the rigid spiro C–C bond rather than a flexible C–N bond, we observe no inhomogeneous effects arising from distributions of donor–acceptor bridge dihedral angles. ACRSA displays no time dependent 'apparent' red shift of the prompt or delayed charge transfer (CT) emission. Moreover, using a range of different hosts, we show that the ground state dipole moment (dielectric value) of the host has very little effect on the ACRSA CT energy, *i.e.* there is no so-called 'solid state solvatochromism'. This leads to weak stabilisation of the CT state in all hosts, but has a very small singlet triplet gap and very fast and efficient monoexponential rISC rates in films (reaching nearly  $10^7$  s<sup>-1</sup> in zeonex host). We observe no power law decaying DF tail because there is no dispersion of the dihedral angle between donor and acceptor units. The previously much lower reported rISC rates in ACRSA are instead reattributed to intermolecular excimer states. The intermolecular species give rise to additional slow TADF contributions and broaden the overall CT emission band at 10% ACRSA loading and in neat films. Harnessing the rapid and homogenous rISC displayed by isolated ACRSA molecules may unlock higher efficiencies and – crucially – extended operational lifetimes in future TADF OLEDs.

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## Introduction

We recently revealed the complex solution state molecular photophysics<sup>1</sup> of the deceptively simple thermally activated delayed fluorescence (TADF) molecule 10-phenyl-10*H*,10'*H*-spiro[acridine-9,9'-anthracen]-10'-one (ACRSA), see Scheme S1 (ESI†).<sup>2–4</sup> This is an important TADF molecule for two reasons. First, the acridine donor (D) and anthracenone acceptor (A) units of the ACRSA molecule are bridged by a spiro carbon, not the usual C–N bond found in most D–A TADF molecules. This connecting group, imparts the required orthogonality between the D and A,<sup>5</sup> but is also considered to be significantly stronger than bridging C–N bonds. This is one of the suspected main causes of the short operational lifetimes achievable by TADF emitters in devices. Moving away from this architecture to a spiro system for example could greatly enhance emitter lifetime. Second, the spiro bridge is

significantly more rigid than C–N bonds. Consequently, there should be far less intrinsic inhomogeneity arising from distortions about this dihedral bond between D and A, which in C–N linked materials results in a distribution of CT state energies. This in turn results in a distribution of singlet–triplet energy gaps, and dispersion in reverse intersystem crossing (rISC) rates ranging up to 4 orders of magnitude in many cases.<sup>6</sup> This conformational disorder, originating from the rotation around the C–N bond has also been investigated in other studies.<sup>7–10</sup> Furthermore, a molecular design which restricts this torsional motion was reported as an approach to avoid this disorder effect.<sup>11–13</sup> However, by using a spiro TADF emitter, the rigidity enforced by the tetrahedral spiro carbon atom gives potential immunity to external conformational distortion of the molecule caused by host interactions. By removing such sources of heterogeneity, the molecular photophysics of ACRSA can be used as an ideal molecule to probe how external factors affect the TADF and rISC mechanisms, including local environment, packing and intermolecular interactions. This should prevent degradation arising through long triplet residency times.

The TADF method for harvesting triplet states has attracted great interest for OLEDs providing a means of harvesting triplet

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excitons without the need for scarce and expensive heavy metals.<sup>14,15</sup> To achieve TADF, rISC must be efficient therefore TADF molecules must have a small energy gap between the lowest energy excited singlet ( $S_1$ ) and triplet ( $T_1$ ) states, *i.e.* small electron exchange energy.<sup>15</sup> One way to achieve this is *via* charge transfer excited states, as found in donor-acceptor (D-A) molecules, that have effective electronic decoupling of ground and excited orbitals, *e.g.* the D and A units are oriented orthogonally, minimising electron exchange resulting in a very small singlet-triplet energy gap of <50 meV.<sup>16</sup> However, when the electron exchange energy is so small, the singlet and triplet CT (charge transfer) orbitals become degenerate and transition between them forbidden because there is no way to change the electron orbital angular momentum during the spin flip transition.<sup>17</sup> To facilitate TADF a third (triplet) excited state, very close in energy to the  $^3CT$  state is required which has a different orbital character which can mix non-adiabatically with  $^3CT$  to mediate a spin flip through second order vibronic coupled spin orbit coupling and couple the triplet back to the singlet manifold *i.e.* rISC.<sup>18,19</sup> As we previously demonstrated experimentally and Lyskov and Marian theoretically,<sup>20</sup> this mechanism is achieved in ACRSA.

ACRSA was first shown to have both efficient solution state photoluminescence (PL) efficiency of 81%, with a high device EQE of 16.5% by Nasu and Adachi *et al.*<sup>2</sup> However, there are rather few reports of other spiro bridged D A TADF emitters,<sup>21-26</sup> as compared to D-A and D-A-D type emitters exploiting the N-C bridging bond motif. Lyskov and Marian have presented a very detailed quantum chemistry analysis of ACRSA highlighting the complex interactions between charge transfer and local states, the role of mixing of high lying excitonic transitions to overcome the forbidden nature of radiative decay from these states, and the potential for rapid nanosecond rISC mediated by  $n-\pi^*$  to  $\pi-\pi^*$  ISC transitions, see Scheme S1 (ESI<sup>†</sup>).<sup>20</sup> Following on from this we have made an in-depth study on the molecular photophysics of the ACRSA molecule in solution to experimentally verify many of these theoretical predictions as well as showing unique photophysics. Throughout this new report all comparison to solution state properties of ACRSA refer to our previous paper,<sup>1</sup> and the terminology used in describing the energy levels of ACRSA is given Scheme S1 (ESI<sup>†</sup>), along with a briefly summary of the results from our solution state measurements.

In this new work, we study the molecular photophysics of ACRSA in a range of solid host materials to elucidate the perturbations arising from solid-state interactions. ACRSA's unique spiro derived properties give unprecedented new insight into this realm to answer several major questions of host interactions that form an important part of the overall response of the system. Importantly, we believe these new results offer new understanding on the current short devices working lifetimes suffered by TADF emitters.

## Results

Excitation dependent steady state emission spectra of 1% wt ACRSA in zeonex, UGH, DPEPO or neat film and the chemical

structures of the hosts are shown in Fig. S1 and S2 (ESI<sup>†</sup>). In zeonex, the spectra consist of a poorly structured emission, peak energy 2.8 eV (460 nm) and a higher energy knee at 3.05 eV (410 nm). Comparing to previous solution spectra,<sup>1</sup> we see that zeonex gives an environment very similar to MCH solution where the  $^1CT$  and  $^1LE$  states are highly mixed giving strongly overlapped emission. Whereas, in DPEPO and UGH films the  $^1CT$  band dominates, peak at 2.4 eV (520 nm), with the appearance of a blue shoulder at *ca.* 450 nm which is highly dependent on excitation wavelength. We note that in solid-state the  $^1CT$  emission is much less red shifted compared to that observed in previously results in toluene solution (a very low polarity solvent). In solution, it is clear that the ACRSA  $^1CT$  state causes a large change in dipole moment. As a result of excitation (CT formation) and subsequent strongly red shift, even in low polarity solvents such as toluene, due to rapid re-organisation of the solvent molecules in the solvent shell around the ACRSA molecule which relaxes the Coulomb energy of the excited state.<sup>27</sup> Whereas the red shift in all solid films is only around 30 meV, which is far less than toluene (280 meV). As previously discussed by us<sup>28,29</sup> and shown by Northey *et al.*<sup>30</sup> a solid host cannot rearrange around the excited guest molecule in the way a solvent molecule can (after excitation) and thus there is little relaxation of the Coulomb energy of the CT state in solid state, *i.e.* no analogous solid state solvatochromism. The host can organise around the emitter molecule in the ground state if it has a large ground state dipole moment, but no further rearrangement can occur in the excited state. As a consequence, there can be no directly analogous 'solid state solvatochromic' effect. Lyskov *et al.*<sup>†</sup> calculate that the GS dipole moment of ACRSA is modest, 5.46 D, and so only a small ground state effect should be expected. This is very clear from these ACRSA results.

Little emission contribution from the  $^1LE$  excited state is observed from measurements made in vacuum (Fig. S3, ESI<sup>†</sup>), which was also found in solution where degassing results in a very large  $^1CT$  DF emission contribution that totally dominated the emission. Film measurements made in air as a function of excitation wavelength do show consistent photophysics with degassed solution measurements, which implies poor oxygen diffusion through the dense films (Fig. S2, ESI<sup>†</sup>). With 330 nm excitation, we observe emission around 350 nm from the  $^1B_1$  exciton state (acridine donor unit fluorescence<sup>31</sup>), which was very strong in MCH solution for example. Excitation at 350 nm shows well-resolved features on the blue edge of the CT band, which is consistent with the observed  $^1LE$  emission in solution. Excitation at 380 nm yields effectively the main CT band (550 nm) and a blue shoulder at 450 nm. From these observations, it is clear that all films behave very similarly to each other and to MCH solution. However, in the films, the overlap between the  $^1LE$  and  $^1CT$  emission is very large leading to poor spectral resolution. Excitation at 350 nm gives the highest intensity of  $^1LE$  emission confirming that the states are present. Moreover, we know from solution (and film results below) that the  $^1LE$  state is quenched rapidly by ISC. In film, this leads to an unstructured emission band from the  $^1LE$  that is very consistent with emission from a hot, vibrationally unrelaxed



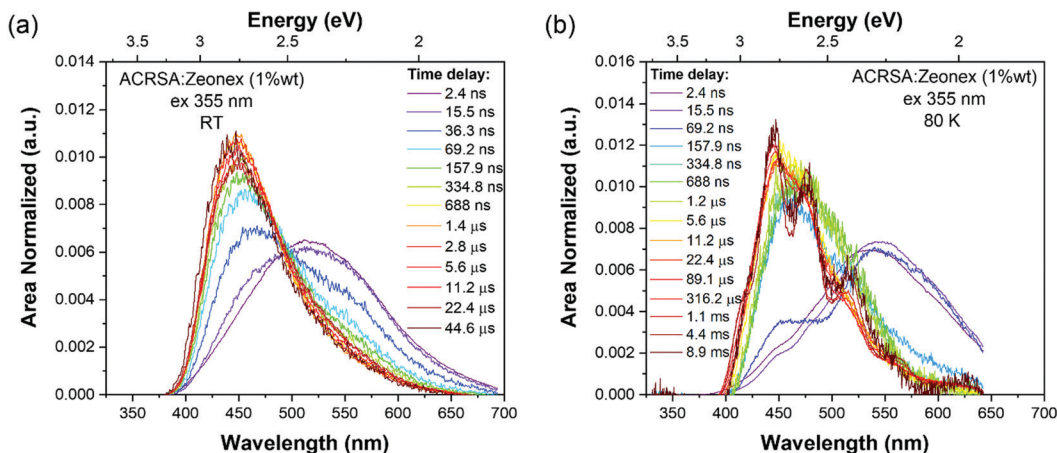


Fig. 1 Time evolution of the emission spectra (area normalised). Laser excitation at 355 nm excitation wavelength, for ACRSA in zeonex (1% loading) film measured at; (a) 300 K, (b) at 80 K. Normalised data shown in Fig. S23 (ESI†).

state as previously postulated by Greene *et al.*<sup>32</sup> in rather similar acceptor molecules. Thus, it appears to look more Gaussian band shaped than the highly structure band observed in MCH where vibrational cooling to the fluid environment is far more efficient. Such hot emission from local states in the film seems to be the main reason why vibrationally structured emission is lost compared to solution emission bands in CT materials.

As in MCH solution, zeonex films (1% wt ACRSA loading) presented an instantaneous and very broad (half width *ca.* 125 nm) emission band, centred at 520 nm (2.38 eV) (Fig. 1). This band decays with lifetime of 6.2 ns (Fig. S4, ESI†). We believe that this metastable state arises from a highly conformationally distorted structure, similar to that theoretically described by Fan *et al.*<sup>33</sup> We aim to give a more detailed report about this in a further study. The central energy position is similar to the CT band observed in toluene solution. As this broad band decays, we observe prompt and DF emission centred around 448 nm, having multi-component lifetime of 93.7 ns, 1.2  $\mu$ s and 8.3  $\mu$ s. This complex, long lived emission may signify two competing rISC channels from two closely spaced (in energy) excited states, however the dominant channel is very fast, approaching  $1 \times 10^7$  s<sup>-1</sup>, Table 1. This we believe

strongly reflects the mixed  $n\pi^*/\pi\pi^*$  (<sup>1</sup>LE/<sup>1</sup>CT) character of the <sup>1</sup>CT state and very small energy gap to local triplet states in zeonex. Spectrally it is difficult to resolve these states given the complexity of the overall emission decay, but if we compare to MCH solution we can deduce that the 93.7 ns and 1.21  $\mu$ s are prompt and DF CT emission, respectively. We believe that zeonex acts very much like the solution state, because the asymmetrically branched polyolefin structure of zeonex which prevents crystallisation and creates a large free volume within the polymer network that the guest molecules occupy with little or no hindrance to conformation motions.

Zeonex films measured at 80 K are more complex. From the kinetic trace (Fig. 2), four distinct decay processes are seen with well-resolved decay times. The initial transient, very broad 520 nm band is not so clearly observed, pointing to this species being strongly thermally activated. A more complex structure is observed at the first few nanoseconds suggesting possible multiple spectral components. This suggests multiple transient emitting states potentially from different metastable molecular configurations.<sup>33</sup> We do observe a CT like emission, peak *ca.* 550 nm decays much more slowly (*ca.* 250 ns) than at 300 K (Fig. S5, ESI†). After 100 ns, dual emission is observed with a shoulder at 470–500 nm observed, that decays with lifetime of

Table 1 Lifetime and amplitude results from data fitting of ACRSA in Zeonex (1%), UGH, DPEPO (1% and 10%) and neat film measured at 300 K; with 355 nm excitation wavelength

ACRSA:HOST	Zeonex 1%	DPEPO 1%	UGH 1%	DPEPO 10%	UGH 10%	NEAT
$\tau_{LE}$ (ns)	6.2	7.7	4.8	6.3	5.4	4.4
$\tau_{PF/1}$ (ns)/ $A_{PF/1}$ (%)	93.7	196	287	74.8/69	69.2/62	105/87
$\tau_{PF/2}$ (ns)/ $A_{PF/2}$ (%)	—	—	—	393/31	400/38	759/13
$\tau_{DF/1}$ ( $\mu$ s)/ $A_{DF/1}$ (%)	1.2/72	—	—	—	—	—
$\tau_{DF/2}$ ( $\mu$ s)/ $A_{DF/2}$ (%)	8.3/28	7.3	5.4	8.7	7.0	6.6
$\tau_{AV/PF}$ (ns)	93.7	196	287	299	328	437
$\tau_{AV/DF}$ ( $\mu$ s)	6.3	7.3	5.4	8.7	7.0	6.6
$k_{rad}$ (s <sup>-1</sup> ) $\times 10^6$	1.02	0.70	0.43	1.45	1.15	3.28
$k_{ISC}$ (s <sup>-1</sup> ) $\times 10^6$	9.66	4.41	3.06	11.9	13.3	6.23
$k_{HISC}$ (s <sup>-1</sup> ) $\times 10^6$	8.67	1.00	1.50	1.06	1.81	0.44
<sup>1</sup> CT (nm)/(cm <sup>-1</sup> )	394/3147	419/2959	416/2980	417/2973	417/2973	424/2924
<sup>3</sup> LE (nm)/(cm <sup>-1</sup> )	398/3115	403/3076	405/3061	433/2863	426/2910	448/2767
$ \Delta E_{ST} $ (meV)	32	117	81	110	63	157



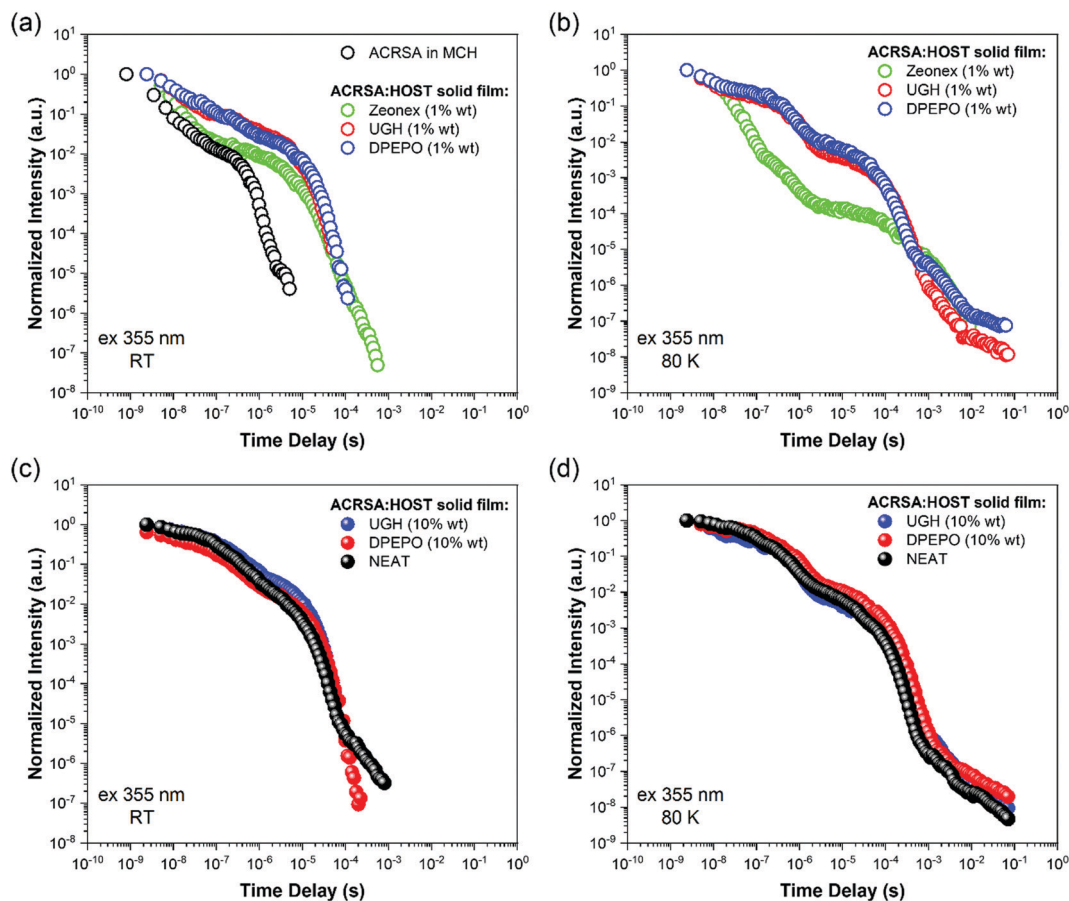


Fig. 2 Temperature dependent emission decay kinetics for ACRSA films. Time resolved emission decay for ACRSA in different host matrices ((a) and (b) at 1% loading, (c) and (d) at 10% loading) and in neat film, recorded at 300 K, left, and 80 K right. Decay trace of ACRSA in 20  $\mu\text{M}$  MCH solution is shown in (a) for reference. Efficient thermally activated delayed fluorescence is always observed even at 80 K indicative of small energy splitting and strong vibrational coupling. Laser excitation at 355 nm.

458 ns. Also, from 100 ns on, a further emission band at 450 nm (peak) is observed. After 25  $\mu\text{s}$ , a highly structure emission band dominates, clearly phosphorescence with lifetime of 1 ms from a local excited triplet state, having a highly developed vibronic progression. This emission is at lower energy than the singlet  $^1\text{CT}$  (*ca.* 3.1 eV) having onset energy of 2.97 eV. The energy spacing of the vibronic replicas is 0.19 eV ( $1530\text{ cm}^{-1}$ ), corresponding to the energy of a C=O stretch. This identifies the emission as phosphorescence from the anthracenone  $^3\text{LE } \pi-\pi^*$  triplet localise around the C=O bond.<sup>20</sup> This phosphorescence is consistent with the ‘early’ ms phosphorescence observed in UGH and DPEPO films at 80 K (Fig. S6 and S7, ESI<sup>†</sup>). We also observe a blue shoulder to this emission, onset 3.12 eV. This shoulder does not fit with the 0.19 eV vibronic progression of the main band, and is not truncated by self-absorption. We believe this to be a second underlying higher energy phosphorescence,<sup>18,34</sup> potentially from the  $^3\pi\pi^*$  state of the acridine donor triplet.<sup>31</sup> At 300 K in zeonex, we see that this higher energy, unstructured phosphorescence is observed from 20  $\mu\text{s}$  onwards, indicating that thermal equilibrium between these two triplet populations has changed.<sup>34</sup> For further comparison, a full data set measured with 337 nm excitation, which

excites the  $^1\text{B}_1$  exciton state as opposed to 355 nm excitation that directly excites the  $^1\text{LE}$  and  $^1\text{CT}$  states, giving excess energy to the molecules (Fig. S8 and Table S1, ESI<sup>†</sup>). A clear result of this excess energy given to the molecule is seen in zeonex films where 337 nm excitation yields higher intensity DF from the  $^1\text{CT}$  band (510 nm) because of the increased triplet production which 337 nm excitation produces.

Making measurements at very low ACRSA loadings of 0.01% clearly show that the 550 nm (2.27 eV) band is not present at low ACRSA loadings, (Fig. S9, ESI<sup>†</sup>) and so we ascribe it to an intermolecular species, as will be seen is readily observed in small molecule hosts in the next section. At 0.01% ACRSA loading in zeonex we find characteristic decay lifetimes of 4 ns and 90 ns prompt emission components and a DF component of 1.3  $\mu\text{s}$ . Because of limited signal to noise ratio at these very low ACRSA concentrations we can observe long lived emission at 430–440 nm but cannot obtain an accurate lifetime.

1% wt ACRSA in UGH and DPEPO films were also investigated with 355 nm and 337 nm excitation. The time resolved emission spectra, Fig. 3, show that the CT emission undergoes minimal time evolution with 337 nm excitation and a very small



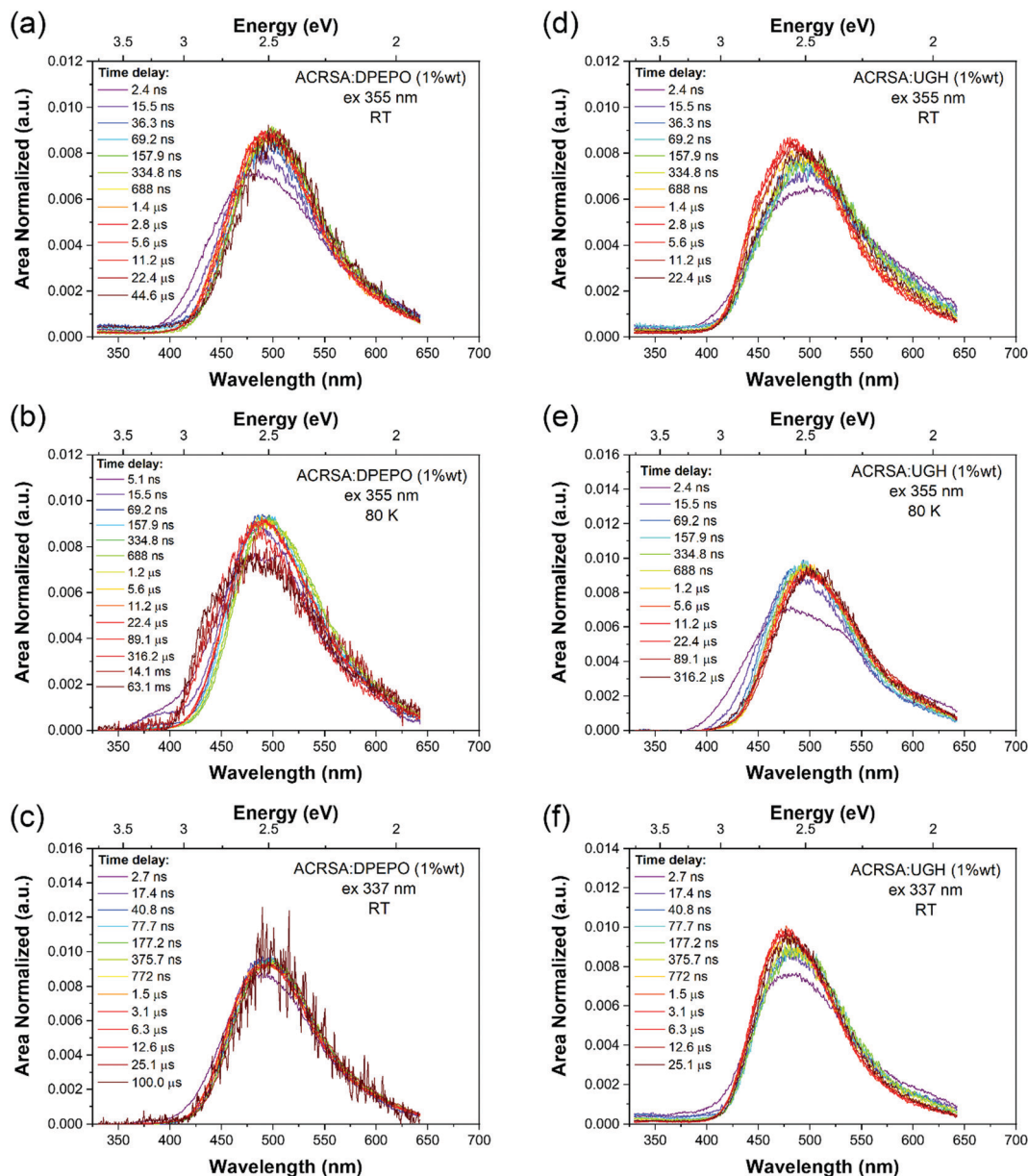


Fig. 3 Time evolution of the emission spectra (area normalised). Time resolved emission spectra from ACRSA at 1% loading in UGH and DPEPO host matrices, measured at room temperature and at 80 K, with 355 nm and 337 nm excitation. Normalised data shown in Fig. S24 (ESI†).

degree with 355 nm excitation. As we predicted for the rigid spiro-TADF emitter, the CT energy has constant onset at 2.98 eV (UGH) and 2.95 eV (DPEPO), peak at ca 2.5 eV, Table 1 and Fig. S10 and S11 (ESI†). Exciting at 337 nm into the  $1^1B_1$  exciton state, both UGH and DPEPO films behave in a rather simple fashion (Fig. 3). The spectra remain relatively constant, showing only a minimal temporal shift and no change in band shape. Moreover, photoluminescence decay in both films can be fitted with a simple single exponential expression for prompt and DF. In UGH, we find 30 meV inhomogeneous broadening (from the red shift of emission over 20  $\mu$ s) and in DPEPO 35 meV. In the first 10 ns, a small blue edge contribution is seen which we ascribe to  $1^1LE$  emission and a small  $1^1B_1$  exciton contribution at 350–400 nm. As in MCH solution,

a weak red tail which decays very rapidly within the first 10 ns is also observed in UGH and DPEPO films. With 355 nm excitation (into the  $1^1LE/1^1CT$  mixed transitions), during the first 100 ns, the emission from UGH films is rather broad, centred at ca 500 nm with a pronounced (but small contribution) red tail in the 550–650 nm region. The major emitting species, especially DF, is a single band, centred at ca. 484 nm, which has some structure, onset 420 nm (2.95 eV). Whereas in DPEPO the earliest emission, is more to the blue, onset 400 nm (3.10 eV), consistent with emission from the  $1^1LE$  state. However, the major contribution to emission from 100 ns onwards is centred at 490 nm. Overall, we see that the band in UGH has more spectral weight around 450–475 nm whereas in DPEPO the majority spectral weight is from 475 nm to 510 nm. At 80 K,



the ACRSA UGH and DPEPO films also show well resolved phosphorescence with lifetime of 0.6 ms, Fig. S6 and S7 (ESI<sup>†</sup>), having the same spectral shape and onset as observed in zeonex, confirming this to be the anthracenone <sup>3</sup>LE  $\pi$ - $\pi^*$  triplet, the lowest monomolecular triplet state of ACRSA independent of host.

Excitation dependent steady state emission spectra of 10% wt ACRSA in UGH, DPEPO and neat film were also studied, Fig. S12 (ESI<sup>†</sup>). Emission is found to be much less excitation dependent. In all cases the <sup>1</sup>CT band dominates, peak at 2.4 eV (520 nm), with the appearance of a blue shoulder at *ca.* 450 nm, which is more dependent on excitation wavelength. We note that the <sup>1</sup>CT emission of ACRSA in UGH, DPEPO and neat film is also much less red shifted compared to that observed in toluene (a very low polarity solvent).

At 300 K, in the DPEPO, UGH (10% wt loading) and neat film (Fig. 4), there is little fast <sup>1</sup>LE emission observable (similar to degassed solution samples). Moreover, a pair of prompt <sup>1</sup>CT decays are observed, lifetimes of 69.2 ns onset at 3 eV

and 400 ns onset 2.78 eV, Fig. 4a–c and Fig. S13–S15 (ESI<sup>†</sup>). This is very different behaviour to that found in the 1% wt loading films. The shorter prompt lifetime species (peak *ca.* 2.6 eV) is consistent with the solution state <sup>1</sup>CT transition and 1% ACRSA zeonex films. Whereas the long lifetime species that grows in very slowly (peak at 2.4 eV) comes from a new excited state species found only at 10% ACRSA loading, which presented a redshifted and broader band as compared to the 1% wt loading films. An isoemissive point is observed in the area normalised time resolve spectra at *ca.* 505 nm. This implies that there are only two distinct emissive species in the films. Further, we also can clearly see that the blue edge of the emission decays with a concomitant growth in the red edge intensity, Fig. S13–S15 (ESI<sup>†</sup>). This indicates that the blue species converts to the red species in competition with radiative decay.<sup>35</sup>

We take this new behaviour in the highly loaded films to indicate formation of some weakly interacting excimeric like state between nearest neighbours. Such a weakly interacting

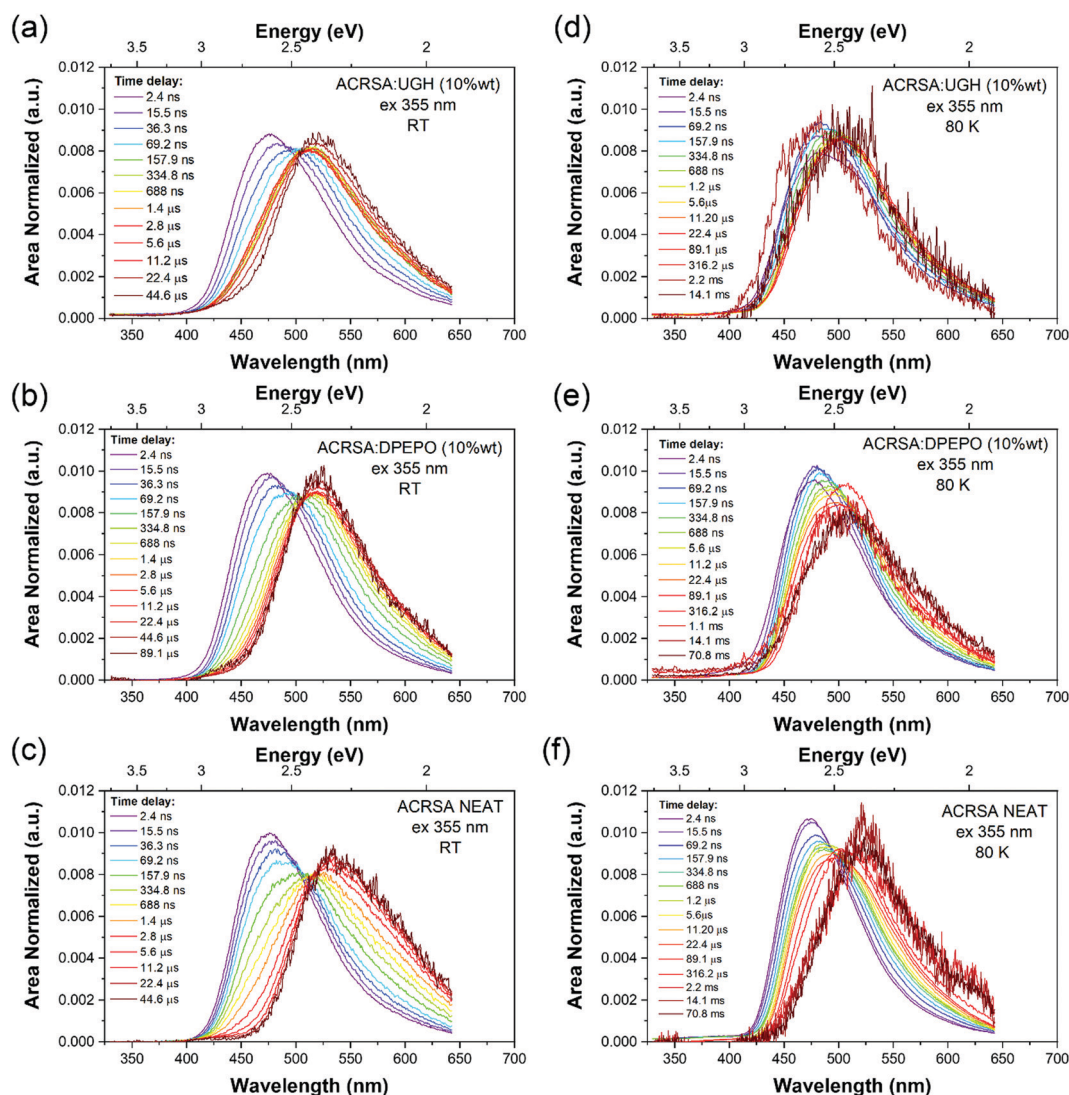


Fig. 4 Time evolution of the emission spectra (area normalised). Laser excitation at 355 nm excitation wavelength, for ACRSA in UGH, DPEPO (10% loading) and neat film measured at; (a–c) 300 K, (d–f) at 80 K. Normalised data shown in Fig. S25 (ESI<sup>†</sup>).



state would be in keeping with intermolecular interactions between two spiro molecules. The very long ‘prompt’ lifetime of this species would also fit with a weakly allowed transition decoupled from its ground state *via* delocalisation across two molecules. The evolution of the time resolved spectra are totally incompatible with any molecular confirmation change (especially given that we are observing a rigid spiro molecule) or from dispersive rISC rates as found in non-spiro D–A TADF molecules.<sup>28</sup> DF in DPEPO, UGH (10% wt loading) and neat film show the same trends and comes from the low energy CT feature, with a single lifetime component of 6–9  $\mu\text{s}$ . UGH films have a slightly faster rISC rate than DPEPO films, but in all cases are  $k_{\text{rISC}}$  is high, around  $1 \times 10^6 \text{ s}^{-1}$ .

80 K measurements, Fig. 4d–f, we again find very similar kinetics and decay times in prompt and delayed emission in 10% wt loading and neat films, Table S2 (ESI<sup>†</sup>). Two emission bands are observed, at *ca.* 470 nm (peak) early times, and at 500–510 nm (peak) later times, with an isoemissive point seen (Fig. S16–S18, ESI<sup>†</sup>). However, the full spectral evolution takes *ca.* 2 orders of magnitude longer to achieve at 80 K, indicating that this mechanism is thermally activated. Also, given the redder species is at higher energy compared to at 300 K, we take this as evidence of thermally activated excimer formation. The UGH films show more detailed spectral information (Fig. S16, ESI<sup>†</sup>). The early time emission, onset *ca.* 420 nm (peak 475 nm) shows some structure indicative of a CT state with a high degree of local character and has strong resemblance to the main emission band observed in zeonex films. We do not observe fast <sup>1</sup>LE state emission in solid films because the prompt and delayed CT emission intensity is so much greater.

At very late times, emission has a different band shape and it is centered at 520 nm. From the decay kinetics this is phosphorescence. As can be seen in neat film spectra recorded at 70 ms, Fig. 4f, this distinct new band at the longest times, onset 2.73 eV is ascribed to the phosphorescence of the excimeric species. In UGH there is also evidence of dual phosphorescence from the monomeric species, with a blue component, onset 405 nm (3.06 eV) which decays with lifetime of 0.7 ms to leave a very long lived phosphorescence at 426 nm onset (2.91 eV) that has a lifetime  $> 70 \text{ ms}$ ,<sup>18,34</sup> from the excimeric species. The very long lived monomeric phosphorescence should be from the acridine <sup>3</sup>LE state as it was seen in zeonex films at room temperature. Looking at the absolute intensity of DF as a function of temperature, at 80 K it is about an order of magnitude less than at 300 K, indicative of an efficient thermally activated rISC mechanism. The majority of triplet energies quoted here are taken from the 80 K data, see Fig. S5–S7, S16–S18 and Table S2 (ESI<sup>†</sup>).

Comparing the kinetic decay traces for both 10% wt and 1% wt loaded films, and scaled for film ACRSA absorbance (Fig. S19, ESI<sup>†</sup>), we see that essentially the kinetics are identical. However, 10% loaded films show quenching of the prompt emission and relatively less DF which implies that in 10% films a fraction of the initial excited states are quenched.

In all cases, independent of ACRSA loading, host and even neat film, we find a linear excitation power dependence for the DF

indication that in all cases the DF arises from a monomolecular process, fully in line with rISC, Fig. S20 and S21 (ESI<sup>†</sup>).

## Discussion

We find that ACRSA is a very efficient TADF material with very high rISC rates. Comparing 1% to 10% loaded ACRSA films, has highlighted new photophysical properties for spiro-TADF materials, as well as showing the weak perturbations from host matrices. But most importantly we clearly see that there is no so-called “solid state solvatochromism”. At 1% loading films, the time resolved emission spectra show little or no red shift with time and the energy of the CT state, which differs by only 30 meV in DPEPO compared to UGH. Moreover, the CT emission band is at the same energy as found in MCH solution and in zeonex films, indicating clearly how the static dielectric properties of the host (typically confused as host polarity) has minimal effect on the CT state. This definitively shows that for D–A TADF materials which show large apparent time dependent red shifts of their CT emission, is behaviour that clearly comes from inhomogeneity in the D–A dihedral angle. This different D–A dihedral angle of CT state (with different energies) have different lifetimes, which gives rise to dispersive ISC, rISC and radiative decay rates, leading to an apparent time dependent red shift of emission. 1% ACRSA UGH films behave almost identically to ACRSA dissolved in MCH, both with 355 nm and 337 nm excitation, although we never observe a very fast and highly structured emission from the local <sup>1</sup>LE state that we observed in MCH solutions, which may indicate that in solid state this emission comes from a vibrationally unrelaxed states at early times.<sup>32</sup> We do though observe weak structure on the emission band ascribed to high LE character of the state. In DPEPO films (1% wt ACRSA loading), we also observe very similar behaviour as UGH films. The main CT emission is seen to be only 30 meV lower in energy than in UGH host, but has more spectral weight in the red which indicates marginally more CT character. However, with 355 nm excitation where the <sup>1</sup>LE and <sup>1</sup>CT states are excited directly, we see a fast blue component ascribed to <sup>1</sup>LE emission. With 337 nm excitation, the <sup>1</sup>B<sub>1</sub> exciton state is excited and this blue component is much weaker, again as we previously observed in solution, indicating that the <sup>1</sup>B<sub>1</sub> exciton state couples primarily to the <sup>1</sup>CT state as well as decaying by fast intersystem crossing.

At 10% loading films we observe rather different photophysics, but again the different hosts, including neat film, show little difference. Two well-resolved bands in the time resolved (area normalised) spectra are observed, with clear isoemissive points. The redder state decays more slowly with long DF lifetimes, whereas the bluer band decays much faster with fast DF, and is the same species as we observe in 1% films (and MCH solution). Clearly, these two bands are not the result of host perturbation through static dielectric/polarizability effects. More importantly, neat film shows near identical behaviour, with the blue band seen at exactly the same energy as in UGH and DPEPO. We also observe that at 80 K there is very little



effect on the photophysics apart from the evolution of the spectra and the lifetime of the red state which increases by two orders of magnitude. From this we conclude that the blue emission seen in 1% wt loading films, 10% wt loading films and MCH solution must be a monomolecular state that has mixed LE/CT character, which is totally unaffected by host environment. As we have previously shown, the degree of charge separation and the amount of local wavefunction character dictates the photophysical character of the excited state in such D–A molecules,<sup>36,37</sup> we therefore propose that this blue emission band is a mixed  $^1n\pi^*$  (acceptor)  $^1\pi\pi^*$  CT excited state having a high degree of charge transfer character. This type of highly mixed LE/CT excited state is as predicted in the DFT/MCRI calculations by Lyskov and Marian.<sup>20</sup> This state has a high radiative decay rate because of the strong coupling to the ground state (imparted by the high local character) and fast rISC indicating a very small S–T gap. Thus, this vibronically coupled TADF system is very efficient and fast, as seen in the kinetic fit data for UGH films, where >90% of the DF decays with a lifetime less than 5  $\mu$ s. Gibson and Penfold calculated that the arrangement of  $^3LE$  above  $^3CT$  but close to  $^1CT$  would be the optimum state configuration for fast rISC,<sup>19</sup> which we believe is the case for ACRSA.

In highly loaded (10% ACRSA) and neat films, a second red CT band is also observed. Obviously, at higher loading we must have a proportion of ACRSA molecules that interact with nearest neighbours, as well as ‘isolated’ molecules. Such intermolecular interactions must be very similar in all hosts and neat film as the resulting red emission band is observed at very similar energies in all hosts. Most telling is that we observe dual emission from monomeric and this red species simultaneously at high loading, but only from the monomeric species at low loading, again clearly showing that the red shifted species is not a result of ‘solid state solvatochromism’ (*oxymoron*) but is an excimer state.<sup>38</sup> Further, the FWHM of the ACRSA emission band in 1% wt loading film (100 nm) (Fig. S3, ESI<sup>†</sup>) is much narrower to that of 10% wt loading films (150 nm) (Fig. S12, ESI<sup>†</sup>) and the FWHM of the steady state emission is also 150 nm. Thus, we can see that the steady state emission spectrum is a weighted average of these two contributions at 10% loading, causing the steady state emission band to red shifted slightly in different hosts. This is in agreement with our previous observations of intermolecular interactions causing emission red shifts in film.<sup>12,39</sup> We see from the X-ray structure for ACRSA, given in Fig. S22, that even in this spiro compound, there is face to face overlap of adjacent acceptor groups. We do not believe this to be enough to sustain strongly coupled dimer states, but instead introduces an excimeric type perturbative effect between two acceptors, effectively increasing the delocalisation of the acceptor (a weakly coupled excimer). This increases the effective accepting strength leading to a red shifted intramolecular CT state. In this way the intermolecular interaction mimics polarity but is not affected in anyway by different hosts. It also explains why this red CT is well defined with a single lifetime and energy in all hosts. We note that previously Méhes *et al.*<sup>22</sup> studying a slightly different spiro TADF system also concluded that a bimolecular species was forming in film, but

postulated that it was an exciplex between adjacent spiro molecules. We believe that our model of the interacting acceptors, which affect the intramolecular ICT state, has a far greater probability of forming in spiro compounds compared to full exciplex formation. In Fig. 6, we label this state as ‘CT<sub>excimer</sub> state’ simply as a way to differentiate it from the isolated molecular  $^1CT$  state. To enable rISC from these CT states there must be coupling to an energetically close triplet state. Since in the 10% loaded films, we observe a long-lived emitting species (onset at 2.73 eV) which is not phosphorescence from the  $^3LE$  state at 2.9 eV. We assume that this excimeric species has a low energy triplet state giving rise to the 2.73 eV phosphorescence (again noted as an excimer triplet in Fig. 6), yielding a  $\Delta E_{ST}$  of *ca.* 150 meV and thus the rISC rate will be slower for this low energy CT states as we observe. Cho *et al.*<sup>40</sup> have shown theoretically that dimer states of TADF molecules can have appreciable rISC rates and not greatly reduce overall performance. Thus, we believe that this is the first definitive evidence for TADF from a dimer/excimer like state.

In zeonex, the presence of intermolecular states described above becomes very clear. The photophysics is very similar to that observed in MCH solution including to observation of a highly red shifted transient emitting species. This we ascribe to the large free volume in this polymer host enabling the phenyl ring bond to the acridine nitrogen to potentially achieve a non-90° dihedral angle in the excited state weakening the donor strength leading to a blue shift of the CT state (higher LE character). This is in line with the theoretical calculations of Fan *et al.* in ACRSA.<sup>33</sup> Whereas, in tightly packed small molecule hosts this is more rigidly fixed in the ground state configuration.

The time resolved emission in zeonex is though dominated by two main, closely spaced (in energy) emitting states. The  $^1LE$  state at *ca.* 430 nm peak (3.14 eV onset) and the mixed (vibrationally coupled)  $n\pi^*/\pi\pi^*$   $^1CT$  state at 460 nm.<sup>20</sup> The  $^1LE$  state has a fast radiative decay rate but is rapidly quenched by allowed SOC to the  $\pi\pi^*$   $^3LE_A$  triplet state, as in solution.<sup>1</sup> From the observed local phosphorescence, which is highly structured and very well differentiated spectroscopically from the  $^1CT$  state, it is clear that the energy gap between the CT states and the mediating local triplet state is very small, calculated at 32 meV, giving rise to the measured very fast rISC rate approaching  $1 \times 10^7$  s<sup>-1</sup> (1.2  $\mu$ s decay component) and thus highly efficient TADF. We also observe a slow DF decay (8  $\mu$ s decay component) having the spectrum of the  $^1LE$  state. Given the strong allowed coupling between the  $n\pi^*$   $^3LE_A$  and  $\pi\pi^*$   $^3LE_A$  triplet states and allowed SOC between  $\pi\pi^*$   $^3LE_A$  and  $n\pi^*$   $^1LE$  singlet state,<sup>1,20</sup> we attribute this slow DF contribution to be from reverse upper triplet state ISC (USrISC) with thermally activated reverse IC driving the initial  $n\pi^*$   $^3LE_A$  and  $\pi\pi^*$   $^3LE_A$  step, see Fig. 6. Thus, at long times we only see  $^1LE$  delayed emission because all the CT states have been depopulated *via* fast rISC (TADF). Even at RT, we observe the spectra to blue shift, which we attribute to emergence of phosphorescence from the local  $^3LE$  states. At 80 K, this slow USrISC becomes slower because of reduced thermal activation (of the RIC) whereas the CT DF from rISC is still fast indicative of the very





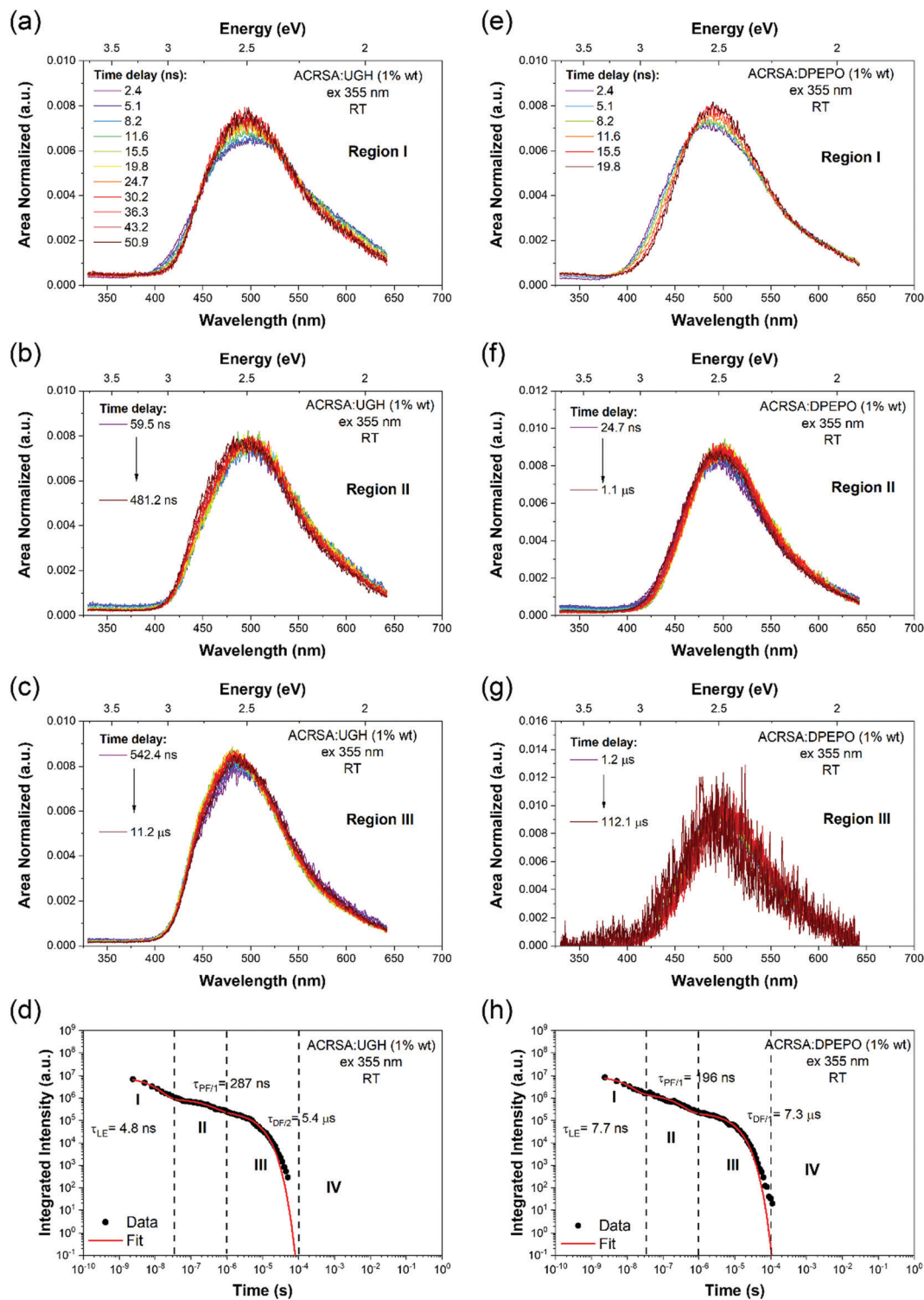


Fig. 5 Comparison of long time DF decay kinetics for ACRSA in DPEPO host showing the single exponential decay of the DF with no associated power law decay component. (a)–(c) Area normalised time resolved spectra of ACRSA in UGH (10% wt loading), shown in the three main regimes indicated in (d). (d) Fitting of kinetic decay results at room temperature. (e)–(g) Area normalised time resolved spectra of ACRSA in UGH (1% wt loading), (h) fitting of kinetic decay results at room temperature. Laser excitation at 355 nm.

small energy gaps and strong vibrational coupling between the CT and local triplet states. We observe very strong local phosphorescence because local triplet states are depopulated

through the USrISC mechanism. We do not observe USrISC in solution at RT because the triplets are more rapidly quenched by solvent collisions.



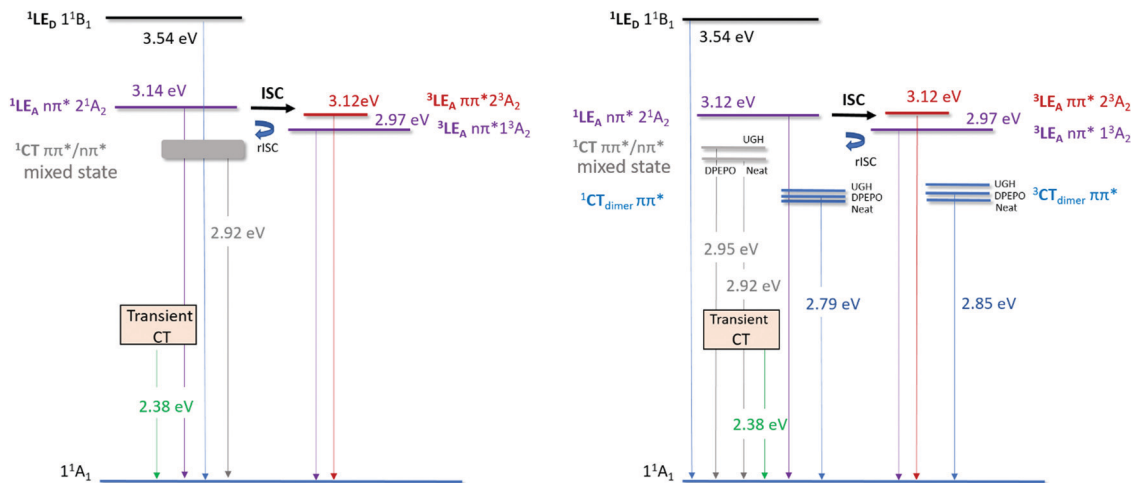


Fig. 6 Excited state energy diagram for ACRSA in different host matrices; left at 1% ACRSA loading in UGH, DPEPO and Zeonex; right, 10% wt ACRSA loading in UGH and DPEPO and neat film. All quoted energies are taken from the spectral on-set of the associated emission bands at 300 K. Triplet energies measured at 80 K and/or long delay time spectra. We follow nomenclature used by Lyskov and Marian.<sup>20</sup>  $1^1B_1$  is the local excited (LE) singlet state associated with the donor unit.  $2^1A_2$  is the LE singlet state associated with the acceptor unit (transition with  $n\pi^*$  character).  $1^1CT$  is the singlet state with charge transfer character.  $2^3A_2$  and  $1^3A_2$  is the LE triplet state associated with the acceptor unit, these transitions have  $\pi\pi^*$  and  $n\pi^*$  character, respectively.

With 337 nm excitation both the  $1^1LE$  and  $1^1CT$  band are observed with an isoemissive point is observed, whereas with 355 nm only the blue band, onset 3.14 eV, ascribed to  $1^1LE$  local emission, is observed which is preferentially excited by 355 nm excitation. Again, possible vibrationally hot emission results in the loss of vibrational structure.<sup>32</sup> The red  $1^1CT$  shoulder with 337 nm excitation arises because we excite *via* the  $1^1B_1$  exciton state which populates both the  $1^1LE$  and  $1^1CT$  states preferentially *via* IC.<sup>1</sup> The red band matches that seen in all other hosts and MCH and is ascribed to the  $1^1CT$  state which has a long prompt lifetime of 60–90 ns, but as the energy gap to the coupling local triplet states is very small, it has fast rISC which decays with a half-life of 1.2  $\mu$ s giving a rISC rate of  $10^7$   $s^{-1}$ .

At 1% ACRSA loading in zeonex we also see a weak 550 nm emission contribution. Measurements at very low ACRSA loadings (0.01%) show this band not to be present and we conclude that this is from an intermolecular species, as we observe in UGH and DPEPO. We assume that the solubility of ACRSA in zeonex is rather poor in this case.

Thus, the major effect of the solid state host on TADF molecules is to allow stable intermolecular interactions to occur as the guest concentration increases. In ACRSA, we observe this as a long lived low energy CT emission band in 10% wt loading and neat films. The steady state emission spectra then represents a true envelop of both CT bands giving greatly increased spectral width compared to the individual CT emission bands observed in the time resolved data. Potentially a very large emission FWHM of a TADF emitter indicates a high contribution from such intermolecular states<sup>39</sup> fully in line with our previous work on the role of dimer/excimer states in TADF emitters. We find, irrespective of host, or ACRSA loading that the DF decays mono-exponentially (as always seen in solution) and no long time, non-exponential decay time tail is observed in the decay kinetics,

which is usually observed in D–A TADF molecules with a C–N bridge between D and A, Fig. 5. Recently, Serevičius *et al.* have demonstrated monoexponential decay can be observed in a specifically designed C–N linked D–A molecule.<sup>9</sup> From our clear observation, we can also unambiguously ascribe these long lifetime non-exponential DF tails as arising from conformationally distorted (about the D–A dihedral bond) molecules. This gives a broad dispersion of dihedral angles and thus a range of electron exchange energies (*i.e.*  $\Delta E_{ST}$  gaps) leading to rISC rates ranging anywhere between *ca.*  $10^6$   $s^{-1}$  and  $10^2$   $s^{-1}$ . This then yields the typical power law DF tail. In ACRSA at 1% wt loading films where there is no spectral shift at all, this is abundantly clear. This then will have a major consequence on device lifetime because these slow DF states are much more prone to annihilation with charge carriers leading to degradation. We believe that in ACRSA, lies the template for long lifetime TADF molecules. The immunity to conformational distortion caused by host packing prevents slow rISC rate sites greatly reducing possible degradation from triplet polaron annihilation.

All of our photophysical observations for ACRSA in solid state are summarised in the state energy diagrams shown in Fig. 6.

## Conclusions

We have studied the spiro-TADF emitter, ACRSA in different host matrices at 1% wt and 10% wt loading and compared the results to those from neat films. Previously using solution measurements, we identify all of the emissive singlet and triplet states of ACRSA and how they are perturbed by solvent polarity. In the solid state at 1% wt loading, we observe a single CT emission, centred around 485 nm which is both independent of



the host matrix and shows no spectral relaxation over more than 20  $\mu\text{s}$ . Clearly this shows that with the rigid spiro bridge between D and A unit there is little inhomogeneity in the dihedral angle between D and A and therefore we see no inhomogeneity in CT state energy or rISC rate, and no power law tail in the DF decay kinetics. As we proposed, this is very different to D–A TADF systems with a C–N bond bridge between D and A which is soft and gives rise to large inhomogeneous rISC rates and CT energies in solid state. This causes a large degree of the unwanted spectral broadening in these D–A TADF systems. Moreover, it also shows unambiguously that there is no extrinsic perturbation from the host due to dielectric constant, *i.e.* so called “solid state solvatochromism”. At 1% ACRSA loading we see virtually no difference between DPEPO, UGH and neat films. Further, we see that the pure  $^1\text{CT}$  emission in these films is far less red shifted than it is observed in toluene solution. In zeonex, the energy gap between the CT states and the vibronic coupling mediating local triplet state is very small, 32 meV, giving rise to the measured very fast rISC rate approaching  $1 \times 10^7 \text{ s}^{-1}$  and thus highly efficient TADF. We also observe DF from the local singlet state which we propose arises through upper triplet state rISC with thermally activated reverse IC from the strong allowed coupling between the  $n\pi^* \text{ } ^3\text{LE}_A$  and  $\pi\pi^* \text{ } ^3\text{LE}_A$  triplet states and allowed SOC between  $\pi\pi^* \text{ } ^3\text{LE}_A$  and  $n\pi^* \text{ } ^1\text{LE}$  singlet state.

In small molecule hosts at low ACRSA loadings, where the higher dielectric constant of the host materials helps to stabilise the CT state, we observe very efficiently along with mono-exponential DF decay times indicating a homogeneous system. However, surprisingly for a spiro material, at 10% wt loading and in neat film we clearly see strong effects due to intermolecular interaction between neighbouring ACRSA molecules. We believe that these are weakly coupled, not a fully excimeric state, more a co-facial overlap of neighbouring acceptor units which give a more delocalised and thus stronger acceptor. This then gives rise to a lower energy CT state that also contributes to a slow DF because of the large gap between the  $^3\text{CT}$  and coupling  $^3\text{LE}$  state which mediates rISC. This state also gives rise to a low energy phosphorescence and the associated triplet state (possibly a  $^3\text{LE}$  state of the delocalised acridine pair) might be able to mediate rISC independently as well. At low temperatures, in zeonex, we observed a highly structured phosphorescence with 0.19 eV vibronic replicas pointing to strong coupling to the acceptor C=O stretching mode, identifying this lowest energy local triplet state as an acceptor  $^3n\pi^*$  state. We also see a long lived, structureless phosphorescence at higher energy which is assigned to the acceptor  $^3\pi\pi^*$  state. The thermal equilibrium between these two triplet populations is clearly temperature dependent.

ACRSA defines a new blue print for highly efficiency TADF emitters,  $k_{\text{rISC}} > 1 \times 10^7 \text{ s}^{-1}$ , as it is also a very robust molecule because it removes the weak C–N bridging bond between D and A. This rigidity has great importance, because it is immune to distortions of the D A dihedral bond (driven through host packing) it does not have a long tail of slowly decay DF states. The long residency times of triplet excitation in these distorted

states will be far more prone to triplet polaron annihilation preventing a major cause of emitter degradation. This should yield far longer device operating lifetimes. Moreover, because of excimer formation at high loadings, as originally used in devices, we believe EQEs of 16% reflect quenching by excimers and by optimisation of the emitter concentration true state of the art device performance can be achieved. We are actively pursuing this currently.

## Author contributions

L. Franca and A. Danos made the presented optical and time resolved spectroscopy measurements. A. Monkman devised the research, undertook the data analysis and supervised the work. A. Monkman wrote the manuscript with L. Franca and A. Danos.

## Conflicts of interest

The authors declare no competing financial interests.

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