

Cite this: *Nanoscale*, 2022, **14**, 17254

## Recharging upconversion: revealing rubrene's replacement

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One of the major limitations of solid-state perovskite-sensitized photon upconversion to date is that the only annihilator successfully paired with the perovskite sensitizer has been rubrene, raising the question of whether this approach of triplet sensitization is universal or limited in scope. Additionally, the inherent energetic mismatch between the perovskite bandgap and the rubrene triplet energy has restricted the apparent anti-Stokes shift achievable in the upconversion process. To increase the apparent anti-Stokes shift for upconversion processes, anthracene derivatives are of particular interest due to their higher triplet energies. Here, we demonstrate successful sensitization of the triplet state of 1-chloro-9,10-bis(phenylethynyl)anthracene using the established formamidineium methylammonium lead triiodide perovskite  $\text{FA}_{0.85}\text{MA}_{0.15}\text{PbI}_3$ , resulting in upconverted emission at 550 nm under 780 nm excitation. We draw a direct comparison to rubrene to unravel the underlying differences in the upconversion processes.

Received 26th September 2022,

Accepted 8th November 2022

DOI: 10.1039/d2nr05309h

rsc.li/nanoscale

### Introduction

Perovskite-sensitized triplet fusion upconversion (UC) via triplet-triplet annihilation (TTA) first emerged in 2019.<sup>1,2</sup> Since its inception, focus has been placed primarily on understanding the mechanism of triplet generation and role of the

fabrication conditions on the device performance.<sup>3–11</sup> High absorption cross sections<sup>12</sup> and long charge carrier diffusion lengths in the perovskite sensitizer<sup>13–15</sup> in addition to the underlying triplet sensitization mechanism based on free charge carrier injection to the bound triplet state of the annihilator<sup>1–3</sup> promise a path towards the realization of efficient solid-state UC devices.<sup>16</sup> However, to date, the only triplet annihilator that has been successfully paired with bulk perovskites in the solid state and reported in bulk lead halide perovskite-sensitized UC has been rubrene, allowing for near-infrared-to-yellow UC.<sup>17,18</sup>

Rubrene has long been the 'gold standard' of solid-state near-infrared-to-visible UC due to its known high UC efficiencies in solution as well as solid state configurations.<sup>19–23</sup> However, the inherent ~0.4 eV energy loss caused by the mismatched energy alignment between the perovskite bandgap (1.55 eV) and triplet energy ( $T_1 = 1.14$  eV)<sup>24,25</sup> negates the benefits of using a sensitizer that does not rely on intersystem crossing to generate the prerequisite triplet states. The resulting small apparent anti-Stokes shift from 800 nm (1.55 eV) to 605 nm (2.05 eV) in the case of rubrene doped with ~1% dibenzotetraphenylperiflanthene (DBP) is far from optimized. In the ideal case, the TTA-UC process should be able to double the sensitizer bandgap energy due to minimal losses when generating the triplet state  $E_{\text{bandgap}} = E(T_1)$  and achieve an apparent anti-Stokes shift equal to the triplet energy  $E(T_1)$  resulting in an emitted singlet energy  $E(S_1) = 2E(T_1)$ . Hence, even when allowing for small energy losses in the TTA processes, achieving upconverted emission around 450 nm (2.8 eV) should be targeted when using the established forma-

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midinium methylammonium lead triiodide (FA<sub>0.85</sub>MA<sub>0.15</sub>PbI<sub>3</sub>, FAMA) perovskite with a bandgap of  $E_{\text{bandgap}} = 1.55$  eV as the triplet sensitizer.

In this desired triplet energy regime, anthracene derivatives are of general interest<sup>26–30</sup> as these allow an increase in triplet energy and reduced energy losses during hole extraction due to generally deeper highest occupied molecular orbital (HOMO) energy levels.<sup>31,32</sup> While 9,10-diphenylanthracene (DPA) is the common workhorse in solution-phase visible-to-blue UC,<sup>33–38</sup> it is incompatible with the FAMA sensitizer used here due to its high triplet energy level ( $T_1 = 1.77$  eV) and strong excimer formation reduces the achievable UC yield and apparent anti-Stokes shift in solid state.<sup>39</sup> However, the green dye 1-chloro-9,10-bis(phenylethynyl)anthracene (1-CBPEA), which has been commercially used in glowsticks, has been reported as an efficient triplet annihilator with peak emission at 490 nm (2.53 eV) in solution.<sup>40–43</sup> Similarly, Castellano and co-workers have successfully demonstrated TTA-UC using 2-chloro-9,10-bis(phenylethynyl)anthracene (2-CBPEA) which has similar optical properties as 1-CBPEA and a FAMA-compatible upper bound of the triplet energy level set at  $T_1 < 1.61$  eV.<sup>44,45</sup> A previous study has given a triplet energy for 1-CBPEA of  $T_1 = 1.20$  eV in solution,<sup>40</sup> which appears to be at odds with TTA-UC occurring successfully as the requirement of  $E(S_1) \leq E(2T_1)$  is not fulfilled.<sup>40–43</sup> While slightly endothermic UC is in principle possible, we suggest a marginally higher native triplet energy level in the range of  $T_1 = 1.3 - 1.2$  eV in solution, estimated at roughly half of the singlet energy  $S_1$  (*vide infra*). However, additional effects on the singlet and triplet energy surfaces caused by intermolecular coupling and different rotational conformations upon condensation into the solid state cannot be ruled out; this effect has previously been observed for the parent molecule 9,10-bis(phenylethynyl)anthracene (BPEA), where different molecular arrangements and rotational conformations influence the singlet and triplet energy surfaces and therefore the rates of both TTA and singlet fission (SF).<sup>46–49</sup>

In this contribution, we investigate the triplet sensitization at the FAMA/1-CBPEA interface and the resulting UC process comparing the performance to the established FAMA/rubrene/1%DBP (FAMA/Rub) device. We find that FAMA perovskites can act as triplet sensitizers for 1-CBPEA resulting in upconverted emission at 550 nm (2.4 eV) highlighting the fact that triplet generation at the perovskite/organic semiconductor (OSC) interface is a general phenomenon and not limited to rubrene. However, despite the larger apparent anti-Stokes shift of the FAMA/1-CBPEA UC device, the overall performance is slightly diminished and the threshold intensity for efficient UC is increased. These effects can in large part be traced back to the significantly reduced triplet lifetime of 1-CBPEA in comparison to rubrene and a reduced OSC quantum yield. However, further optimization of the energetic driving force for charge extraction by compositional<sup>8</sup> and interfacial tuning<sup>4</sup> or addition of a dopant dye to improve the OSC quantum yield<sup>21</sup> bear promise in further advancing this next-generation perovskite-sensitized UC.

## Results and discussion

The general optical characterization and comparison of 1-CBPEA and rubrene doped with ~1% DBP (RubDBP) OSC thin films are depicted in Fig. 1. Fig. 1a shows the absorption spectra of 1-CBPEA and RubDBP. In the latter case, the expected vibronic progression of rubrene is visible, as well as the vibronic feature corresponding to DBP at 600 nm. For 1-CBPEA, in addition to the expected vibronic progression, we find a broad red shifted feature at 520 nm (2.4 eV), which we attribute to an aggregation-induced redshift due to intermolecular coupling in the solid state.<sup>47</sup> As reported previously, the emission profile of RubDBP (Fig. 1b) is governed by the dopant dye DBP (605 nm), indicating efficient Förster resonance energy transfer (FRET) between rubrene and DBP. The 1-CBPEA emission does not show the vibronic progression observed in solution, rather is dominated by the red-shifted aggregate emission peaking at 550 nm. A comparison of the photoluminescence (PL) intensity normalized by the optical density of the OSC at the excitation wavelength (405 nm) indi-

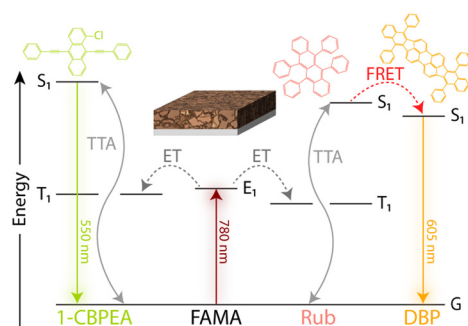


**Fig. 1** Comparison of the RubDBP and 1-CBPEA OSC thin films. (a) UV-visible absorbance spectra of RubDBP and 1-CBPEA, where a redshifted aggregation-induced absorption feature can be seen at 520 nm. (b) Emission spectra of the OSC films normalized to their optical density at the excitation wavelength of 405 nm. 1-CBPEA emits at 550 nm and RubDBP is dominated by DBP emission at 605 nm. Absorption spectra at selected delay times for RubDBP (c) and 1-CBPEA (d) thin films. The relevant spectral features are identified for both OSC films. Rubrene exhibits the characteristic  $S_1$  excited state absorption (ESA) at 445 nm and  $T_1$  ESA features at 485 nm and 515 nm and a bleach corresponding to DBP is observed at 605 nm. Two bleach features are found for 1-CBPEA, which we attribute to the ground state bleach (GSB) at 440 nm and 530 nm, respectively. The  $S_1$  ESA is found at 470 nm and 500 nm,  $T_1$  ESA features emerge at 475 nm and 510 nm. Excess pump laser scatter is denoted by the gray box.

cates that the PL quantum yield of the RubDBP film is  $\sim 1.5$  times that of 1-CBPEA.

Transient absorption (TA) spectroscopy gives further insight into the spectral signatures of the singlet and triplet states under direct excitation. Rubrene is known to undergo both SF and TTA, due to its triplet energy  $T_1$  at half of the singlet energy  $S_1$ . Therefore, direct excitation of the RubDBP OSC thin film at 400 nm yields the singlet signature dominated by the  $S_1$  excited state absorption (ESA) at 445 nm which rapidly evolves to reveal the  $T_1 \rightarrow T_n$  ESA features characteristic of rubrene triplet state at 485 nm and 515 nm.<sup>24,25</sup> Furthermore, a DBP-related bleach feature can be observed at 605 nm which we attribute to stimulated emission under direct excitation or to the DBP ground state bleach (GSB). To the best of our knowledge, SF has not been previously reported in 1-CBPEA; however, has been shown in the parent molecule BPEA.<sup>47</sup> Due to the similarity of the TA spectra observed here for 1-CBPEA and previous reports for BPEA, we suggest that SF also occurs in the 1-CBPEA OSC thin films under direct excitation at 400 nm.<sup>47</sup> In particular, we find bleach features at 440 nm and 530 nm which we attribute to the GSB or stimulated emission of 1-CBPEA. Overlapping ESA features corresponding to the  $S_1$  state are found at 470 nm and 500 nm. Additional redshifted ESA features rapidly emerge at 475 nm and 510 nm, which can be attributed to the spin-allowed  $T_1 \rightarrow T_n$  optical transitions of triplet states generated by SF on an ultrafast timescale.<sup>44,47</sup>

To investigate the charge extraction at the perovskite/OSC interface and resulting TTA-UC process, bilayer devices consisting of a  $\sim 30$  nm thin FAMA sensitizer film and either RubDBP or 1-CBPEA as the annihilator were fabricated. The anticipated energy flow resulting in UC in each device is highlighted in Fig. 2. Upon photoexcitation of the perovskite at 780 nm, free charge carriers are generated which transfer to the OSC and recombine to populate the triplet state of either 1-CBPEA or RubDBP. Following triplet generation, two triplets can combine *via* TTA to yield the desired higher energy singlet



**Fig. 2** Energy flow schematic for UC process in the bilayer devices with the corresponding chemical structures for the OSC molecules. Incident 780 nm photons are absorbed by the perovskite, where charge transfer (ET) populates the triplet state  $T_1$  of 1-CBPEA (left) or Rub (right). Interaction of two triplet states results in TTA, yielding the higher energy  $S_1$  state. For Rub, an additional energy transfer (FRET) occurs to the DBP  $S_1$  energy level. Resulting UC emission wavelengths are 550 nm for 1-CBPEA and 605 nm for RubDBP.

state. After creation of the singlet state, the photon can be emitted directly from the  $S_1$  state (1-CBPEA) or the higher energy state can be harvested by the DBP dopant (RubDBP) with subsequent emission from the  $S_1$  state of DBP.

The absorption spectra of the bilayer devices and a FAMA control film are shown in Fig. 3a. The absorption onset of the perovskite is unchanged at 800 nm, and the OSC-related vibronic progression near 500 nm is observed indicating successful fabrication of the bilayers. Further investigation by time-correlated single-photon counting yields information into the charge carrier extraction at the FAMA/OSC interface where we monitor the perovskite PL decay dynamics ( $>800$  nm) under 780 nm excitation (Fig. 3b). In comparison to the FAMA control, both the RubDBP and 1-CBPEA show quenched PL decay dynamics, indicating successful charge extraction. Anthracene derivatives generally have deeper highest occupied molecular orbital (HOMO) levels in comparison to their tetraene analogues.<sup>31,32,50</sup> The HOMO of the 1-CBPEA parent molecule BPEA has been reported at  $-5.49$  eV,<sup>51</sup> deeper with respect to the HOMO level of rubrene at  $-5.4$  eV.<sup>50</sup> Electron withdrawing groups such as the chlorine in 1-CBPEA, further deepen the HOMO level and simultaneously reduce the bandgap.<sup>52-54</sup> As a result, we expect the driving force for initial hole extraction from the perovskite valence band to the OSC HOMO to be inherently reduced in 1-CBPEA in contrast to RubDBP when utilizing the same FAMA perovskite sensitizer composition. Hence, the increased amount of quenching and more rapid PL decay for RubDBP over 1-CBPEA can be attributed to a higher driving force for charge extraction for RubDBP.

Thus far, we have shown that charge extraction occurs at the perovskite/OSC interface. However, the observation of charge extraction at the perovskite/OSC interface is not sufficient for triplet generation since single charge transfer would similarly reduce the PL lifetime of the perovskite. To investigate whether the triplet of 1-CBPEA is indeed populated following the hole extraction process at the perovskite/OSC interface, we turn to TA spectroscopy to probe the ESA corresponding to the  $T_1 \rightarrow T_n$  transition.<sup>9</sup> Fig. 4a shows the TA



**Fig. 3** (a) Absorption spectra for the FAMA perovskite thin film (black) and bilayer devices (RubDBP: pink, 1-CBPEA: green). (b) Perovskite PL decays for the FAMA and bilayer devices under 780 nm excitation at an excitation density of  $44.2$  mW  $\text{cm}^{-2}$ . A 780 nm notch filter and an 800 nm long pass filter were used to remove laser scatter and isolate the perovskite emission.



**Fig. 4** Absorbance spectra at selected delay times for (a) FAMA, (b) RubDBP, and (c) 1-CBPEA in the spectral range of 440–650 nm showcasing the higher energy perovskite photobleach (PB2) at 495 nm and photoinduced absorption (PIA) at 550 nm. The onset of the ground state bleach PB1 can be seen at 650 nm. Residual  $\text{PbI}_2$  signals at 510 nm and 520 nm are present in the FAMA only spectra due to surface treatment with acetonitrile to increase interfacial charge extraction. In (b), both the Rub  $T_1$  and polaron signatures are present at 520 nm and 550 nm signifying successful triplet sensitization. The  $T_1 \rightarrow T_n$  transitions at 460 nm and 510 nm for 1-CBPEA are highlighted in (c). All TA measurements were taken under 700 nm pump. (d) UC emission for the bilayer films with photos of the corresponding bilayer films under 780 nm excitation.

spectra for the underlying FAMA sensitizer at selected delay times under 700 nm excitation in the relevant spectral region from 440–650 nm for the triplet-related ESA. The two characteristic bleach features of the perovskite are denoted as PB1 and PB2 respectively, while the perovskite photoinduced absorption feature centred at 550 nm is labelled as PIA. An additional overlapping bleach and photoinduced absorption attributed to  $\text{PbI}_2$  is found at 510 and 520 nm, respectively. This excess  $\text{PbI}_2$  is caused by a post-fabrication acetonitrile treatment of the perovskite surface. We have previously shown that this solvent treatment increases triplet generation in rubrene due to n-type doping of the perovskite, which increases the hole density at the surface due to interfacial band bending.<sup>55</sup> Once interfaced with the OSCs, additional spectral features corresponding to the  $T_1 \rightarrow T_n$  ESA transitions previously shown in Fig. 1 for both OSCs emerge and can be detected on a sub-nanosecond timescale for both RubDBP (Fig. 4b) and 1-CBPEA (Fig. 4c). Hence, we confirm that the charge extraction observed by PL spectroscopy indeed yields triplet excitons in the OSC layers.

To validate that the triplet generation in the OSC layer results in upconverted emission from the OSC singlet state, we probe the emission of the bilayer devices under 780 nm excitation (Fig. 4d). The RubDBP UC device yields the expected strong emission with a peak emission at 605 nm (Fig. 4d,

pink).<sup>2</sup> The 1-CBPEA UC device yields an upconverted emission spectrum peaking at 550 nm, which increases the apparent anti-Stokes shift achievable by perovskite-sensitized UC by 0.2 eV (Fig. 4d, green) relative to the FAMA/RubDBP UC device. The insets show photographs of the corresponding devices under 780 nm excitation where the strong upconverted emission is visible for both UC devices. A stronger intensity of upconverted emission is found for RubDBP than for 1-CBPEA which can be traced back to two factors: (i) a higher triplet generation in rubrene than in 1-CBPEA as indicated by the stronger quenching of the perovskite PL in the case of RubDBP. (ii) A higher native OSC quantum yield: the observed upconverted emission is directly dependent on the underlying OSC PL quantum yield which dictates the fraction of singlet states generated in the TTA-UC process which undergo radiative recombination.<sup>2,21</sup>

Lastly, we investigate the properties of the upconverted PL. TTA-UC exhibits a characteristic power dependence: at low incident powers, the upconverted PL intensity increases quadratically. In this regime, the triplets decay primarily through other decay pathways. However, above the threshold intensity  $I_{th}$  for efficient UC, the upconverted PL intensity increases linearly with incident power.<sup>35</sup> Here, TTA is the predominant triplet decay pathway and the upconversion process becomes efficient. Increasing the excitation power further results in a saturation regime, where the UC process becomes less efficient. To include TTA-UC devices into photovoltaics, an  $I_{th}$  significantly lower than the integrated solar spectrum at the wavelength region of interest is required, on the order of  $I_{th} < 10 \text{ mW cm}^{-2}$ .

Mathematically, the power threshold  $I_{th}$  can be expressed by the following equation as described by Monguzzi *et al.*,<sup>35</sup>

$$I_{th} = \frac{(k_T)^2}{\gamma_{TT}\alpha(E)\phi_T}$$

Here,  $k_T$  is the annihilator triplet decay rate,  $\gamma_{TT}$  is the second order rate constant characterizing the TTA process for a given annihilator,  $\alpha(E)$  is the absorption coefficient of the perovskite, and  $\phi_T$  represents the efficiency of the generation of the OSC triplet state.<sup>35,56,57</sup>

For both annihilators investigated here, we find the characteristic turnover from a slope  $\alpha = 2\beta$  to  $\alpha = \beta$ , where  $\beta$  is the inherent power dependency of the underlying perovskite PL ( $\beta = 1.3$ ). We extract a threshold intensity  $I_{th} = 18 \text{ mW cm}^{-2}$  for RubDBP and  $I_{th} = 195 \text{ mW cm}^{-2}$  for 1-CBPEA. As a note, in this study, the underlying perovskite sensitizer films are kept significantly thinner than the 100 nm optimal thickness we have previously determined<sup>1</sup> to allow for sufficiently low optical density for TA studies.<sup>9</sup> Upon increasing the perovskite thickness, we have previously shown that the  $I_{th}$  is further reduced due to the increase of  $\alpha(E)$ .<sup>1</sup> Hence, increasing the sensitizer film thickness will push the  $I_{th}$  of the 1-CBPEA UC device towards viable intensity thresholds for device applications.

A back-of-the-envelope calculation allows for further insight into the UC process for the two annihilators and the cause of

the discrepancy in the  $I_{th}$ . The optical density  $\alpha(E)$  at 780 nm for both devices is similar and therefore, is not a determining factor. We estimate a slightly lower triplet population yield  $\phi_T$  for 1-CBPEA in comparison to RubDBP based on the magnitude of perovskite PL quenching, which in part accounts for a higher threshold intensity  $I_{th}$  of 1-CBPEA. However, the effects of the annihilator properties on the efficiency threshold cannot be overlooked and  $k_T$  or  $\gamma_{TT}$  must be the underlying cause of the differences in the  $I_{th}$ .

To determine the underlying cause of the difference in the  $I_{th}$ , we investigate the dynamics of the UC PL (Fig. 5b). The UC PL dynamics are generally governed by two separate processes: (i) the initial rise time which gives the characteristic time of singlet generation. This time scale is a convolution of the triplet generation, TTA-UC and  $S_1$  decay. (ii) The long-lived decay, which is rate-limited by the triplet decay. In both cases, we find a rapid rise of the UC PL. RubDBP exhibits two rise times  $\tau_{rise,1} = 50$  ns and  $\tau_{rise,2} = 400$  ns, due to rapid TTA-UC occurring close to the interface and diffusion-mediated TTA-UC far from the interface.<sup>6</sup> 1-CBPEA on the other hand shows a single characteristic rise of  $\tau_{rise} = 28$  ns, indicating rapid generation of the emissive singlet state. However, the dynamics of the UC PL shown in Fig. 5b also clearly highlight a major difference between the annihilators. Within the 6  $\mu$ s time window shown, the UC PL of RubDBP has barely started decaying, while the UC PL of 1-CBPEA has already fully decayed. Estimating the triplet lifetime in the solid-state devices based on a tail fit results in a long-lived decay for RubDBP ( $\tau_{RubDBP} > 10$   $\mu$ s) and a much more rapid decay for 1-CBPEA ( $\tau_{1-CBPEA} = 1$   $\mu$ s). Therefore, based on the UC PL

dynamics alone ( $k_{T,1-CBPEA} = 10^6$  s<sup>-1</sup>, while  $k_{T,RubDBP} < 10^5$  s<sup>-1</sup>), we would expect the  $I_{th}$  of RubDBP to be a minimum of two orders of magnitude lower than that of 1-CBPEA.

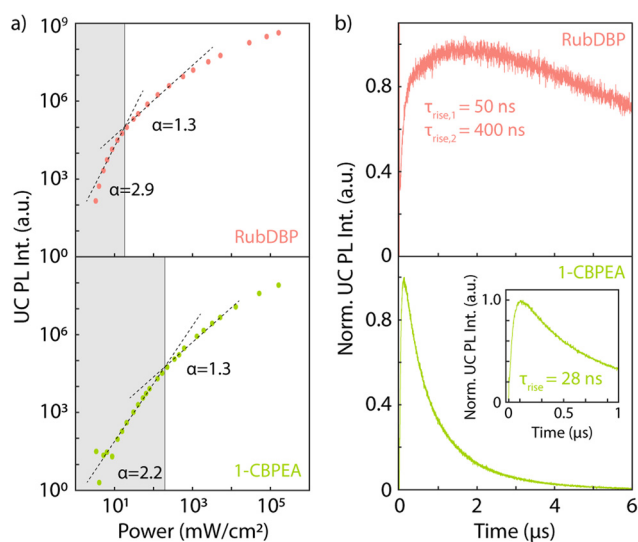
To summarize thus far, the trends for  $k_T$  and  $\phi_T$  predict a significantly larger difference in the  $I_{th}$  than we observe experimentally. The last key factor in the calculation of the  $I_{th}$  is the second order rate constant  $\gamma_{TT}$ , which is unique to each individual annihilator. According to Monguzzi *et al.*,<sup>35</sup>  $\gamma_{TT}$  is dependent on the triplet exciton diffusion length, the probability of TTA generating a singlet state, as well as the exciton interaction distance. To counterbalance the decrease in the  $I_{th}$  caused by the significantly longer-lived triplets in RubDBP and higher yield of triplets  $\phi_T$ , 1-CBPEA must have a much larger  $\gamma_{TT}$  to account for the relative differences in the  $I_{th}$ . Hence, the triplet diffusion rate or probability of TTA-UC occurring in 1-CBPEA must be higher than in the current state-of-the-art solid-state annihilator rubrene. This further solidifies the promise of this new annihilator 1-CBPEA, despite the currently superior performance of RubDBP.

Circling back to the previous discussion of the triplet energy  $T_1$  of 1-CBPEA, due to the observation of both TTA-UC and SF, we acknowledge that the previously reported triplet energy of  $T_1 \approx 1.2$  eV is indeed accurate in the solid-state devices investigated here.<sup>40</sup> This amounts to half of the singlet energy  $S_1$ :  $E(T_1) = 0.5E(S_1)$  – the fundamental requirement for both SF and TTA-UC to occur in the same material. However, this result begs the question whether the TTA-UC process here occurs through the true singlet state  $S_1$  of 1-CBPEA ( $\sim 2.6$  eV in solution) and which then relaxes to the aggregate-induced red-shifted singlet state at (2.4 eV) or, whether the lower singlet energy of the aggregated state in solid state enables TTA-UC in this material in the first place. As a result, we emphasize that the translation of solution-phase annihilator properties to the solid state is not straight-forward, as the effect of intermolecular coupling cannot be ignored. In fact, we suggest that the properties of possible annihilators must first be investigated in their aggregated form, prior to making an educated decision on whether they are viable candidates for TTA-UC.

## Conclusions

We have demonstrated successful perovskite-sensitized solid-state TTA-UC with a novel annihilator 1-CBPEA in the solid state, enabling near-infrared-to-green UC. While we acknowledge that the direct comparison between the RubDBP and 1-CBPEA UC devices does highlight that RubDBP at present still outperforms its replacement 1-CBPEA in terms of efficiency and intensity threshold, the 0.2 eV increase in apparent anti-Stokes shift cannot be ignored. Further device improvements by compositional tuning of the perovskite to better match the promise increased triplet quantum yields to boost the UC efficiency.

Additionally, improvements to the OCS film itself may yield increased TTA-UC efficiencies. Wasielewski and co-workers have shown that for the parent molecule BPEA, the fabrication



**Fig. 5** (a) Power dependent upconverted emission for the RubDBP (top, pink) and 1-CBPEA (bottom, green) bilayer devices. The intersect of the dashed lines indicating the different TTA-UC regimes yields the  $I_{th}$  (vertical grey lines). Calculated intensity threshold  $I_{th}$  values for RubDBP and 1-CBPEA films are 18.2 mW cm<sup>-2</sup> and 195 mW cm<sup>-2</sup>, respectively. (b) UC PL dynamics for the RubDBP (top) and 1-CBPEA (bottom) films (repetition rate: 50 kHz, power density: 120 mW cm<sup>-2</sup>). A magnification of the early time rise for 1-CBPEA is shown in the inset.

method directly influences the crystal arrangement, and hence the OSC PL quantum yield due to a change in rates of SF.<sup>47</sup> Work by Moth-Poulsen and co-workers highlight the role of the molecular conformation on the singlet and triplet energy surfaces,<sup>46</sup> which dictate the rates of SF and TTA. Together, these results emphasize that a clear fundamental understanding of the molecular conformation at the nanoscale and the effects of the resulting intermolecular interactions are required to further advance this technology. Since 1-CBPEA is also capable of both SF and TTA-UC, the balance between these two processes must be shifted towards TTA. Hence, tailoring the fabrication conditions to favour TTA over SF will be necessary.

In conclusion, we have demonstrated that triplet generation at the perovskite/OSC interface by charge injection is a universal process and is not inherently limited to the perovskite/rubrene interface. Moving forward, focus should be placed on developing new sensitizer/annihilator pairs to expand the library of viable partners.

## Experimental

### Device synthesis

Glass substrates were cleaned *via* sonication for 15 min in each of the respective solutions: 2% Hellmanex, deionized water, and acetone. Following sonication, the substrates were cleaned by UV-ozone (Ossila) treatment for 15 min. Precursor solutions of PbI<sub>2</sub> (1.2 M, TCI), MAI (1.2 M, Dyenamo), and FAI (1.2 M, Dyenamo) were prepared in anhydrous DMF : DMSO (9 : 1, v/v, Sigma-Aldrich) in a 1 : 1.09 ratio. The precursor solution was then diluted to 0.24 M (~30 nm) and spin-coated at 1000 rpm for 10 s and 5000 rpm for 30 s. Anhydrous chlorobenzene (Sigma-Aldrich) was used as the antisolvent. The films were then annealed at 120 °C for 30 min prior to solvent treatment with acetonitrile (Sigma-Aldrich).

Rubrene (Rub 99.99%, Sigma-Aldrich), dibenzotetraphenylperiflanthene (DBP, 98%, Sigma-Aldrich), and 1-chloro-9,10-bis(phenylethynyl)anthracene (1-CBPEA, 99%, Sigma-Aldrich) were used without further purification. A 10 mg mL<sup>-1</sup> solution of Rub in anhydrous toluene (Sigma-Aldrich) was prepared and doped with ~1% DBP, similarly a 10 mg mL<sup>-1</sup> solution of 1-CBPEA in anhydrous toluene (Sigma-Aldrich) was prepared. All annihilator solutions were spin-coated onto the perovskite substrates at 6000 rpm for 20 s then annealed at 100 °C for 1 min. Films were encapsulated with a cover slip using a two-part epoxy (Devcon) under an inert nitrogen atmosphere (<0.5 ppm O<sub>2</sub>) prior to removal from the glovebox.

### Steady-state absorption spectroscopy

A Thermo Scientific Evolution 220 Spectrophotometer was used to collect the steady-state absorption spectra.

### Steady-state emission spectroscopy

Direct excitation emission spectra were collected with a 405 nm continuous wave (CW) laser (LDH-D-C-405, PicoQuant) at a power density of 92 W cm<sup>-2</sup> using a 425 nm long-pass

filter (Chroma Tech) to remove excess laser scattering. Upconverted emission spectra were taken with a 780 nm CW laser (LDH-D-C-780, PicoQuant) at a power density of 150 W cm<sup>-2</sup> using a 700 nm short-pass laser (ThorLabs) to remove laser scatter. An Ocean Insight emission spectrometer (HR2000+ES) was used to collect all spectra.

### Time-resolved emission spectroscopy

Perovskite PL decays were measured through time-correlated single photon counting (TCSPC) with a 780 nm picosecond pulsed laser (LDH-D-C-780, PicoQuant) with a 125 kHz repetition frequency at a power density of 44.2 mW cm<sup>-2</sup>. An 800 nm long-pass filter (ThorLabs) and 780 nm notch (ThorLabs) were used to remove laser scatter as well as isolate the perovskite emission. OSC PL decays were taken under 405 nm picosecond pulsed excitation (LDH-D-C-405, PicoQuant) at a repetition frequency of 1 MHz and power density of 4.97 W cm<sup>-2</sup>. A 425 nm long pass filter (ThorLabs) was used to remove laser scattering. Upconverted PL decays were measured under a 780 nm picosecond pulsed laser (LDH-D-C-780, PicoQuant) at a repetition frequency of 50 kHz at a power density of 120 mW cm<sup>-2</sup>. A combination of a 650 nm short pass (ThorLabs), 700 nm short pass (ThorLabs), and 780 nm notch (ThorLabs) were used to isolate the upconverted PL emission and remove scattering. A MultiHarp 150 event timer (PicoQuant) connected to a single-photon avalanche photodiode (Micro Photon Devices) was used to collect photon arrival times for all measurements. Laser powers were measured with a silicon power meter (PM100-D, ThorLabs), and spot sizes were determined using the razor blade method (90 : 10).

### PL power dependencies

All power dependent measurements were taken using a 780 nm CW laser (LDH-D-C-780, PicoQuant). Laser powers were measured using a silicon power meter (PM100-D, ThorLabs) and arriving photons were counted for 25 s by a MultiHarp 150 event timer (PicoQuant) with a single-photon photodiode (Micro Photon Devices) connected. For the UC PL power dependencies, both a 780 nm notch filter (ThorLabs) and 600/40 nm (centre/width) band pass filter (ThorLabs) were used to isolate the UC signal and remove laser scattering. Laser powers were attenuated with neutral density filters (ThorLabs).

### Transient absorption spectroscopy

A HELIOS Fire transient absorption spectrometer (Ultrafast Systems) was used for all transient absorption measurements. Femtosecond laser pulses were generated by an Astrella-V-F-1K amplifier where the Vitara-S Coherent Ti:Sapphire laser used was amplified using a 1 kHz Coherent Revolution-50 pump laser. All resulting laser pulses were 5 mJ with a full width half max of 100 fs at 800 nm. Pump and probe beams were directed through an optical parametric amplifier (OperaA Solo, Coherent) and delay stage, respectively. The visible probe (400 nm to 780 nm) was generated *via* a sapphire crystal for

bilayer and perovskite only measurements while a CaF<sub>2</sub> crystal generated the ultraviolet probe for the OSC-only measurements (320 nm to 650 nm). Excess laser scattering was minimized through a dual chopper system, and neutral density filters were used to attenuate pump power. For all measurements, 3 spectra were collected with a 2 seconds integration time for each delay position with an exponential point collection method starting at 0.01 ps, resulting in a total of 150 points. Perovskite only measurements were taken under 700 nm pump at 4.7 mW, and bilayer films were taken under 700 nm pump with a power of 6.0 mW. A 400 nm pump was used for the OSC measurements at a power of 0.6 mW for both. The TA maps were processed through the Surface Xplorer software package from Ultrafast Systems in addition to MATLAB.

## Author contributions

L. N. conceptualized the study. C. M. S. and L. N. fabricated the samples. C. M. S. performed the experiments. C. M. S. and L. N. analysed the data. L. N. wrote the manuscript with contributions by C. M. S.

## Conflicts of interest

FSU has filed a provisional application for a US patent based on this technology that names L. N. as an inventor.

## Acknowledgements

The authors acknowledge funding by Florida State University. This project used resources provided by the Materials Characterization Laboratory (FSU075000MAC) at the FSU Department of Chemistry and Biochemistry. Ultrafast transient absorption measurements were performed on a spectrometer supported by the National Science Foundation under Grant No. CHE-1919633.

## References

- 1 S. Wieghold, A. S. Bieber, Z. A. VanOrman, L. Daley, M. Leger, J.-P. Correa-Baena and L. Nienhaus, *Matter*, 2019, **1**, 705–719.
- 2 L. Nienhaus, J.-P. Correa-Baena, S. Wieghold, M. Einzinger, T.-A. Lin, K. E. Shulenberger, N. D. Klein, M. Wu, V. Bulović, T. Buonassisi, M. A. Baldo and M. G. Bawendi, *ACS Energy Lett.*, 2019, **4**, 888–895.
- 3 S. Wieghold and L. Nienhaus, *J. Phys. Chem. Lett.*, 2020, **11**, 601–607.
- 4 L. Wang, J. J. Yoo, T.-A. Lin, C. F. Perkinson, Y. Lu, M. A. Baldo and M. G. Bawendi, *Adv. Mater.*, 2021, **33**, 2100854.
- 5 K. Prashanthan, B. Naydenov, K. Lips, E. Unger and R. W. MacQueen, *J. Chem. Phys.*, 2020, **153**, 164711.
- 6 S. Wieghold, A. S. Bieber, Z. A. VanOrman and L. Nienhaus, *J. Phys. Chem. Lett.*, 2019, **10**, 3806–3811.
- 7 S. Wieghold, A. S. Bieber, J. Lackner, K. Nienhaus, G. U. Nienhaus and L. Nienhaus, *ChemPhotoChem*, 2020, **4**, 704–712.
- 8 A. S. Bieber, Z. A. VanOrman, H. K. Drozdick, R. Weiss, S. Wieghold and L. Nienhaus, *J. Chem. Phys.*, 2021, **155**, 234706.
- 9 C. R. Conti, A. S. Bieber, Z. A. VanOrman, G. Moller, S. Wieghold, R. D. Schaller, G. F. Strouse and L. Nienhaus, *ACS Energy Lett.*, 2022, **7**, 617–623.
- 10 A. S. Bieber, Z. A. VanOrman, S. Wieghold and L. Nienhaus, *J. Chem. Phys.*, 2020, **153**, 084703.
- 11 Z. A. VanOrman, J. Lackner, S. Wieghold, K. Nienhaus, G. U. Nienhaus and L. Nienhaus, *Appl. Phys. Lett.*, 2021, **118**, 203903.
- 12 M. M. Lee, J. Teuscher, T. Miyasaka, T. N. Murakami and H. J. Snaith, *Science*, 2012, **338**, 643–647.
- 13 S. D. Stranks, G. E. Eperon, G. Grancini, C. Menelaou, M. J. P. Alcocer, T. Leijtens, L. M. Herz, A. Petrozza and H. J. Snaith, *Science*, 2013, **342**, 341–344.
- 14 G. Xing, N. Mathews, S. Sun, S. S. Lim, Y. M. Lam, M. Grätzel, S. Mhaisalkar and T. C. Sum, *Science*, 2013, **342**, 344–347.
- 15 W. Tress, N. Marinova, O. Inganäs, M. K. Nazeeruddin, S. M. Zakeeruddin and M. Graetzel, *Adv. Energy Mater.*, 2015, **5**, 1400812.
- 16 J. Alves, J. Feng, L. Nienhaus and T. W. Schmidt, *J. Mater. Chem. C*, 2022, **10**, 7783–7798.
- 17 Z. A. VanOrman and L. Nienhaus, *ACS Energy Lett.*, 2021, **6**, 3686–3694.
- 18 S. Wieghold, Z. A. VanOrman and L. Nienhaus, *Adv. Opt. Mater.*, 2021, **9**, 2001470.
- 19 M. Mahboub, Z. Huang and M. L. Tang, *Nano Lett.*, 2016, **16**, 7169–7175.
- 20 L. Nienhaus, M. Wu, N. Geva, J. J. Shepherd, M. W. B. Wilson, V. Bulović, T. Van Voorhis, M. A. Baldo and M. G. Bawendi, *ACS Nano*, 2017, **11**, 7848–7857.
- 21 M. Wu, D. N. Congreve, M. W. B. Wilson, J. Jean, N. Geva, M. Welborn, T. Van Voorhis, V. Bulović, M. G. Bawendi and M. A. Baldo, *Nat. Photonics*, 2016, **10**, 31–34.
- 22 M. Wu, J. Jean, V. Bulović and M. A. Baldo, *Appl. Phys. Lett.*, 2017, **110**, 211101.
- 23 N. Geva, L. Nienhaus, M. Wu, V. Bulović, M. A. Baldo, T. Van Voorhis and M. G. Bawendi, *J. Phys. Chem. Lett.*, 2019, **10**, 3147–3152.
- 24 D. G. Bossanyi, Y. Sasaki, S. Wang, D. Chekulaev, N. Kimizuka, N. Yanai and J. Clark, *J. Mater. Chem. C*, 2022, **10**, 4684–4696.
- 25 L. Ma, K. Zhang, C. Kloc, H. Sun, M. E. Michel-Beyerle and G. G. Gurzadyan, *Phys. Chem. Chem. Phys.*, 2012, **14**, 8307–8312.
- 26 A. Monguzzi, R. Tubino and F. Meinardi, *J. Phys. Chem. A*, 2009, **113**, 1171–1174.
- 27 K. Wang, S. Huang, P. Ding, Z. Liang, S. Chen, L. Li, C. Ye and X. Wang, *RSC Adv.*, 2021, **11**, 17755–17759.

- 28 J.-H. Kim, F. Deng, F. N. Castellano and J.-H. Kim, *ACS Photonics*, 2014, **1**, 382–388.
- 29 P. B. Merkel and J. P. Dinnocenzo, *J. Lumin.*, 2009, **129**, 303–306.
- 30 M. Häring, R. Pérez-Ruiz, A. J. von Wangelin and D. D. Díaz, *Chem. Commun.*, 2015, **51**, 16848–16851.
- 31 M. Y. Vorona, N. J. Yutronkie, O. A. Melville, A. J. Daszczyński, J. S. Ovens, J. L. Brusso and B. H. Lessard, *Materials*, 2020, **13**, 1961.
- 32 Y.-Y. Lyu, J. Kwak, O. Kwon, S.-H. Lee, D. Kim, C. Lee and K. Char, *Adv. Mater.*, 2008, **20**, 2720–2729.
- 33 Z. A. VanOrman, A. S. Bieber, S. Wiegold and L. Nienhaus, *Chem. Mater.*, 2020, **32**, 4734–4742.
- 34 C. J. Imperiale, P. B. Green, M. Hasham and M. W. B. Wilson, *Chem. Sci.*, 2021, **12**, 14111–14120.
- 35 A. Monguzzi, J. Mezyk, F. Scotognella, R. Tubino and F. Meinardi, *Phys. Rev. B*, 2008, **78**, 195112.
- 36 T. Miyashita, P. Jaimes, T. Lian, M. L. Tang and Z. Xu, *J. Phys. Chem. Lett.*, 2022, **13**, 3002–3007.
- 37 E. M. Rigsby, T. Miyashita, P. Jaimes, D. A. Fishman and M. L. Tang, *J. Chem. Phys.*, 2020, **153**, 114702.
- 38 Z. Huang and M. L. Tang, *J. Phys. Chem. Lett.*, 2018, **9**, 6198–6206.
- 39 A. Nandi, B. Manna and R. Ghosh, *Phys. Chem. Chem. Phys.*, 2019, **21**, 11193–11202.
- 40 W. Wu, H. Guo, W. Wu, S. Ji and J. Zhao, *J. Org. Chem.*, 2011, **76**, 7056–7064.
- 41 S. Guo, L. Xu, K. Xu, J. Zhao, B. Küçüköz, A. Karatay, H. G. Yaglioglu, M. Hayvali and A. Elmali, *Chem. Sci.*, 2015, **6**, 3724–3737.
- 42 F. Zhong and J. Zhao, *Dyes Pigm.*, 2017, **136**, 909–918.
- 43 S. Guo, J. Sun, L. Ma, W. You, P. Yang and J. Zhao, *Dyes Pigm.*, 2013, **96**, 449–458.
- 44 C. Mongin, S. Garakyaraghi, N. Razgoniaeva, M. Zamkov and F. N. Castellano, *Science*, 2016, **351**, 369–372.
- 45 T. N. Singh-Rachford and F. N. Castellano, *Inorg. Chem.*, 2009, **48**, 2541–2548.
- 46 V. Gray, A. Dreos, P. Erhart, B. Albinsson, K. Moth-Poulsen and M. Abrahamsson, *Phys. Chem. Chem. Phys.*, 2017, **19**, 10931–10939.
- 47 Y. J. Bae, G. Kang, C. D. Malliakas, J. N. Nelson, J. Zhou, R. M. Young, Y.-L. Wu, R. P. Van Duyne, G. C. Schatz and M. R. Wasielewski, *J. Am. Chem. Soc.*, 2018, **140**, 15140–15144.
- 48 B. Manna, A. Nandi and R. Ghosh, *J. Phys. Chem. C*, 2018, **122**, 21047–21055.
- 49 Y. J. Bae, J. A. Christensen, G. Kang, J. Zhou, R. M. Young, Y.-L. Wu, R. P. Van Duyne, G. C. Schatz and M. R. Wasielewski, *J. Chem. Phys.*, 2019, **151**, 044501.
- 50 G. Ji, G. Zheng, B. Zhao, F. Song, X. Zhang, K. Shen, Y. Yang, Y. Xiong, X. Gao, L. Cao and D.-C. Qi, *Phys. Chem. Chem. Phys.*, 2017, **19**, 6546–6553.
- 51 B. vander Zee, S. Paulus, R.-Q. Png, P. K. H. Ho, L.-L. Chua, G.-J. A. H. Wetzelaer and P. W. M. Blom, *Adv. Electron. Mater.*, 2020, **6**, 2000367.
- 52 J. F. Tannaci, M. Noji, J. McBee and T. D. Tilley, *J. Org. Chem.*, 2007, **72**, 5567–5573.
- 53 S. Liu, L. Zheng, M. Chen, Y. Sun, P. Wang, S. Li, H. Wu, X. Zhang and W. Hu, *J. Mater. Chem. C*, 2021, **9**, 4217–4222.
- 54 M. Aydemir, G. Haykır, A. Battal, V. Jankus, S. K. Sugunan, F. B. Dias, H. Al-Attar, F. Türksoy, M. Tavaslı and A. P. Monkman, *Org. Electron.*, 2016, **30**, 149–157.
- 55 C. M. Sullivan, A. S. Bieber, H. K. Drozdick, G. Moller, J. E. Kuszynski, Z. A. VanOrman, S. Wiegold, G. F. Strouse and L. Nienhaus, *Adv. Opt. Mater.*, 2022, **n/a**, 2201921.
- 56 A. Monguzzi, R. Tubino, S. Hoseinkhani, M. Campione and F. Meinardi, *Phys. Chem. Chem. Phys.*, 2012, **14**, 4322–4332.
- 57 A. Ronchi, P. Brazzo, M. Sassi, L. Beverina, J. Pedrini, F. Meinardi and A. Monguzzi, *Phys. Chem. Chem. Phys.*, 2019, **21**, 12353–12359.