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On the multiphoton ionisation photoelectron spectra of phenol†

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The phenol molecule is a prototype for non-adiabatic dynamics and the excited-state photochemistry of biomolecules. In this article, we report a joint theoretical and experimental investigation on the resonance enhanced multiphoton ionisation photoelectron (REMPI) spectra of the two lowest ionisation bands of phenol. The focus is on the theoretical interpretation of the measured spectra using quantum dynamics simulations. These were performed by numerically solving the time-dependent Schrödinger equation using the multi-layer variant of the multiconfiguration time-dependent Hartree algorithm together with a vibronic coupling Hamiltonian model. The ionising laser pulse is modelled explicitly within the ionisation continuum model to simulate experimental femtosecond 1+1 REMPI photoelectron spectra. These measured spectra are sensitive to very short lived electronically excited states, providing a rigorous benchmark for our theoretical methods. The match between experiment and theory allows for an interpretation of the features of the spectra at different wavelengths and shows that there are features due to both 'direct' and 'indirect' ionisation, resulting from non-resonant and resonant excitation by the pump pulse.

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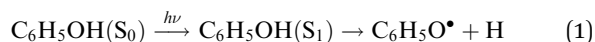
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1 Introduction

Phenol (C₆H₅OH) is an important molecular motif in many biologically relevant molecules, for example the amino acid tyrosine, where it plays a key role in the catalysis of the water-splitting reaction in photosystem II,¹ and the green fluorescent and photoactive yellow proteins chromophores.^{2–4} As a result, the photophysics and photochemistry of phenol have attracted considerable attention as prototypes for understanding the photochemistry of important biomolecules.^{5–25}

Ultraviolet (UV) photoexcitation of phenol results in cleavage of the O–H bond to produce the phenoxyl radical (C₆H₅O•) and a H atom,



The H-atom detachment dynamics of phenol in the gas-phase involves the participation of the three lowest lying singlet electronic states: the S₀ ground state of ¹ππ character, the optically bright S₁ state of ¹ππ* character, and the optically dark S₂ state of ¹πσ* character.^{5,6,8} The S₁ state is bound,

whereas the S₂ state is dissociative along the O–H bond stretch, which results in conical intersections (CIs) or crossings between the S₂/S₁ and S₂/S₀ states (see Fig. 1). UV-excitation to the bright S₁ state thereby results in a transfer of population to the dark S₂ state and ultimately back to the S₀ ground state, leading to a rapid quenching of fluorescence.⁶

CIs and non-radiative decay mechanisms are ubiquitous in the photochemistry of polyatomic molecules.²⁶ Interestingly, both the S₂/S₁ and S₂/S₀ CIs involve the same coupling coordinates: the O–H stretch (r_{O–H}) and CCOH dihedral angle (θ).⁸ Electronic structure calculations have revealed that the S₂ (¹πσ*) state corresponds to a Rydberg-type 3s orbital in the Franck–Condon (FC) region with significant antibonding σ* character with respect to the O–H bond that eventually opens the pathway for hydrogen abstraction.^{5,6} With UV-excitation below the S₂/S₁ CI (for excitation wavelengths 275 nm > λ > 248 nm), nonadiabatic tunneling-facilitated O–H bond fission persists, accounting for a much slower timescale for photodissociation.^{16,17,20} The excited state dynamics of phenol thus represent a classic example of nonadiabatic transitions between adiabatic electronic states mediated by CIs, in which the relative position of the ¹πσ* state with respect to ¹ππ* state governs the dynamics, as also found in the photodissociation of other heteroatomic molecules.^{6,11,12}

Although numerous studies have focused on the photodissociation of phenol following photoexcitation to the bright S₁ and dark S₂ states,^{5–20,24,25} relatively little attention has been

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Fig. 1 One-dimensional cuts of the (diabatic) potential energy surfaces of phenol from the vibronic coupling model along the O–H stretching mode ($r_{\text{O-H}}$) for the S_0 ($^1\pi\pi$) ground state, and first two singlet excited states, S_1 ($^1\pi\pi^*$) and S_2 ($^1\pi\sigma^*$). Potential energy cuts of the first two doublet states D_0 and D_1 of the phenol cation are also shown. The S_2/S_1 and S_2/S_0 states cross each other at 2.3 bohr and 3.95 bohr, respectively. 1+1 photon excitation from S_0 to D_0 and D_1 states, via the bright S_1 state is illustrated by vertical arrows. The inset depicts a schematic representation of the optimized geometry of phenol in the ground state.

paid to the photoionisation dynamics. Experimentally, the most direct way of probing ionisation is to use photoelectron spectroscopy (PES), where measurement of the electron kinetic energy (eKE) distribution provides a direct measure of the binding energy of the molecular orbital from which the electron is removed, as well as information about the vibrational energy distribution, both prior to and after photoionisation. In this regard, multiphoton ionisation²⁷ has emerged as a powerful tool. In particular, resonance-enhanced multiphoton ionisation (REMPI) PES provides valuable insight into the electronically excited states that are involved in the UV photochemistry and photophysics^{27,28} and can reveal excited-state dynamics of short-lived states.²⁹

In a conventional REMPI-PES experiment, a molecule is ionised by sequential absorption of two (or more) photons. In 1+1 REMPI, the first photon promotes the molecule to specific vibrational levels of an excited electronic state and a second photon ionises the molecule (as illustrated by arrows in Fig. 1). The spectra are often reported as a function of electron binding energy (eBE) defined as the difference between the total photon energy and the measured eKE. Analyzing the eBE distribution of photoelectrons can shed light on the electronic structure, electronic character and relaxation dynamics of the resonance-enhancing state; *e.g.*, the $S_2 \rightarrow S_1$ internal conversion is reported to be the dominant relaxation pathway in the two-photon ionisation of phenol at ~ 207 nm, while increasing the

duration of the ionising pulse led to increases in photoelectron yield at high electron binding energy.²⁹

Here, we present calculated photoelectron spectra determined using numerical quantum wavepacket dynamics simulations to solve the time-dependent Schrödinger equation with an explicit treatment of the light-field. The results are compared with measured photoelectron spectra obtained following 1+1 REMPI with UV femtosecond laser pulses at six different wavelengths, most of which were reported in earlier work,²² spanning the two UV absorption bands.

Since vibronic (electron-vibrational) coupling plays a crucial role in the electronic spectra of molecules,^{30,31} we model the ionisation spectra by taking into account the vibronic coupling effects which in turn can help understand the experimental spectrum. To do this, the theoretical simulations are based on constructing a vibronic coupling Hamiltonian³¹ and carefully investigating the associated nonadiabatic dynamics using the multi-layer variant of the multi-configurational time-dependent Hartree (ML-MCTDH) algorithm.^{32,33} The model includes five electronic states: S_0 , S_1 , S_2 , D_0 and D_1 . An ionisation continuum model is invoked in conjunction with photoelectron cross-section calculations to model the photoelectron spectra in the presence of a laser pulse. From the model Hamiltonian, eighteen vibrations are found to be required for the treatment of the short-time dynamics. These include the O–H bond stretch and C–C–O–H torsion that contain CIs in the singlet manifold along with the modes that provide couplings to at least first-order. Of these 18 modes, from an analysis of the coupling strengths (both inter- and intra-state) it was found that the key dynamics can be treated with only 10 vibrations.

The motivation for this work was: (i) to shed light on the participation of intermediate states in the two-photon excitation process, (ii) to quantify the power of our theoretical model in reproducing the energy-resolved photoelectron spectra, and (iii) to provide a preliminary understanding prior to measuring and computing future time-resolved photoelectron spectra of phenol.

The remainder of the paper is organized as follows. The vibronic coupling Hamiltonian model and the quantum dynamical simulations are described in Section 2.1. The experimental details are presented in Section 2.2. The 1+1 REMPI spectra obtained experimentally and theoretically are then presented and discussed in Section 3. Finally, we summarize and conclude with future prospects in Section 4.

2 Methodology

2.1 Theoretical framework

2.1.1 Model Hamiltonian. A simple and efficient way to simulate the non-adiabatic dynamics associated with a system comprising several electronic states is by constructing a vibronic coupling Hamiltonian.^{31,34} This is done by assuming a diabatic electronic basis, with the diabatic potentials and couplings between states expressed by (low-order) Taylor series in mass- and frequency-scaled normal mode coordinates Q ,



around a particular point. For photoexcitation processes this is chosen as the equilibrium geometry of the ground state Q_0 . If the Taylor series expansion for the diabatic potentials is truncated at the second order while the couplings are truncated at first order, it is termed a linear vibronic coupling Hamiltonian model. The “global gauge” of the diabatic basis is chosen by setting it to be identical to the adiabatic basis at the Franck–Condon (FC) point. This transforms the singular non-adiabatic operator in the adiabatic representation to local potential-type operator in the chosen diabatic representation.^{35,36}

The molecular Hamiltonian $\mathbf{H}_{\text{mol}}$ for a set of N electronic states is an $N \times N$ matrix which can be written as a sum of terms

$$\mathbf{H}_{\text{mol}} = \mathbf{H}^{(0)} + \mathbf{W}^{(0)} + \mathbf{W}^{(1)} + \mathbf{W}^{(2)} \quad (2)$$

where the zeroth-order Hamiltonian $\mathbf{H}^{(0)}$ is the ground-state Hamiltonian represented by an (on-diagonal) nuclear kinetic energy operator $\mathbf{T}_N$ and potential energy $\mathbf{V}_0$, taken as a harmonic approximation to the ground-state potential:

$$\mathbf{H}^{(0)} = \mathbf{T}_N + \mathbf{V}_0 = \sum_k \frac{\omega_k}{2} \left(-\frac{\partial^2}{\partial Q_k^2} + Q_k^2 \right) \quad (3)$$

where ω_k is the frequency of the normal mode Q_k and the summation runs over all the vibrational modes k . The ground-state frequencies were calculated at the MP2/aug-cc-pVDZ level of theory and are listed in the ESI† in Table S1.

The set of diabatic potential matrices $\mathbf{W}$ describe the changes in the excited-state surfaces with respect to the ground state by including the effects of electronic excitation and vibronic coupling. The model contains $N = 5$ states. These are the ground and lowest two excited singlet states, along with the lowest two cation states. To this end, the zeroth-order matrix $\mathbf{W}^{(0)}$ is a diagonal matrix containing the vertical excitation and ionisation energies E_k at Q_0

$$W^{(0)} = E_k. \quad (4)$$

Since we choose the diabatic and adiabatic basis to coincide at the FC geometry, there is no off-diagonal counterpart of E_k . The values for these energies were taken from different sources and ultimately adjusted so that the calculated absorption and photoelectron spectra match experiment. Table 1 lists the vertical excitation energy values for the different states. These compare well with other studies.¹⁶

The first-order matrix elements are expressed as

$$W_{ii}^{(1)} = \sum_k \kappa_k^{(i)} Q_k \quad (5)$$

$$W_{ij}^{(1)} = \sum_k \lambda_k^{(ij)} Q_k; \quad i \neq j \quad (6)$$

where the linear intrastate coupling constants $\kappa_k^{(i)}$ are related to the gradients of the adiabatic potentials with respect to the nuclear coordinates at the FC point, and the linear interstate coupling constants $\lambda_k^{(ij)}$ provide the non-adiabatic (vibronic) couplings between the electronic states that are close in energy.

Table 1 Vertical excitation energies (VEEs) and ionisation energies (VIEs) $E^{(0)}$ for the 5 states included in the vibronic coupling model. The column labelled model are the final values

		Vertical excitation (eV)				Osc. strength
		CASSCF ^a	CCSD ^b	Expt.	Model	CCSD ^b (a.u.)
1	$S_0(^1A_1)$	0.0	0.0	0.0	0.0	—
2	$S_1(^1B_2\pi\pi^*)$	4.82	4.93	4.51 ^c	4.859	0.0205
3	$S_2(^1B_1\pi\sigma^*)$	5.94	5.75	5.12 ^d	5.379	0.0002
4	$D_0(^2B_1)$	7.986	8.333	8.508 ^e	8.608	—
5	$D_1(^2A_2)$	8.670	9.164	9.280 ^f	9.380	—

^a SA(3)-CAS(8,8)/6-31+G* for the singlet and SA(2)-CAS(7,8)/6-31+G* for the doublet states. ^b EOM-CCSD/6-31+G* for the singlet and IP-EOM-CCSD/6-31+G* for the doublet states. ^c Bist, Brand and Williams, *J. Mol. Spectrosc.*, 1966, **21**, 76. ^d Bist, Brand and Williams, *J. Mol. Spectrosc.*, 1967, **24**, 413. ^e Lipert and Colson *JCP*, 1990, **92**, 3240. ^f Palmer *et al.*, *J. Mol. Struct.*, 1979, **52**, 293.

The second-order matrix elements are given by

$$W_{ii}^{(2)} = \sum_k \frac{1}{2} \gamma_{kk}^{(i)} Q_k^2 + \sum_{k < m} \gamma_{km}^{(i)} Q_k Q_m \quad (7)$$

where the bilinear (quadratic) intrastate coupling constants γ are related to the second derivatives of the adiabatic potentials with respect to nuclear coordinates at the FC point. The constants $\gamma_{kk}^{(i)}$ account for the frequency changes upon the electronic transition, while $\gamma_{km}^{(i)}$ are responsible for the so-called Duschinsky rotation of the normal modes in the excited state. The bilinear (quadratic) interstate coupling constants $\mu_{kk}^{(ij)}$ and $\mu_{km}^{(ij)}$ that provide higher order coupling between electronic states are usually of minor importance and therefore suppressed here for simplicity.

Consideration of molecular symmetry simplifies the model by imposing restrictions on the modes that will appear in the summations. For non-vanishing interstate coupling constants $\lambda_k^{(ij)}$, the product of symmetries of the two states (Γ_i and Γ_j) and the vibrational mode (Γ_k) must contain the totally symmetric irreducible representation of the point group of the molecule (Γ_A), i.e.,

$$\lambda_k^{(ij)} \neq 0, \quad \text{if } \Gamma_i \otimes \Gamma_k \otimes \Gamma_j \supset \Gamma_A. \quad (8)$$

Similar conditions holds for intrastate coupling constants

$$\gamma_{km}^{(i)} \neq 0, \quad \text{if } \Gamma_k \otimes \Gamma_m \otimes \Gamma_i \supset \Gamma_A \quad (9)$$

and only totally symmetric vibrational modes can couple to the electronic motion to first order

$$\kappa_k^{(j)} \neq 0, \quad \text{if } \Gamma_k \supset \Gamma_A \quad (10)$$

The linear intrastate $\kappa_k^{(i)}$ and interstate $\lambda_k^{(ij)}$ coupling constants, along with the bilinear interstate coupling constants $\gamma_{km}^{(ij)}$ are reported in the ESI†

For vibrational modes exhibiting significant anharmonicity, a harmonic approximation to the diabatic potentials can be a poor description in which case they should be replaced by anharmonic potentials such as Morse or quartic functions.^{21,37,38} An anharmonic description to the intersecting singlet potential surfaces (S_0 , S_1 , and S_2) was constructed by replacing vibration ν_{33} with the O–H bond stretch, r_{OH} and the vibration ν_2 with the



C–C–O–H bond torsion angle, θ . The diabatic potentials for these modes were then taken from the work of Lan *et al.*⁸ For S_0 and S_1 this was a Morse potential for r_{OH} , while for S_2 an avoided-crossing potential was used to provide the barrier. These were combined with cosine series for θ . For the cationic states (D_0 and D_1), which are states 4 and 5 in the model, the following diabatic potentials were constructed

$$V_4(r, \theta) = v_{40}(r) + J_1 - J_2 \cos(2\theta) - J_3 \cos \theta \quad (11)$$

where

$$v_{40}(r) = D_e^4 [1 - \exp(-a_4(r - r_4))]^2 + a_{40}, \quad (12)$$

and

$$V_5(r, \theta) = v_{50}(r) + K_1 - K_2 \cos(2\theta) - K_3 \cos \theta \quad (13)$$

where

$$v_{50}(r) = D_e^5 [1 - \exp(-a_5(r - r_5))]^2 + a_{50}. \quad (14)$$

The values of the parameters are reported in the ESI.[†] The cut through the model surfaces along the O–H bond is shown in Fig. 1. The crossings in the singlet manifold are clear and the cuts compare well to previous high level calculations.¹⁶

The remaining coupling parameters entering the model Hamiltonian are obtained by fitting the *ab initio* energy points along each normal mode to the adiabatic form of the diabatic electronic Hamiltonian through a least-squares fitting procedure. This is done by using the VCHAM program³⁹ of the Quantics package.^{40,41} Ten modes were found to have significant coupling (the mode numbers are given in Fig. 3). The *ab initio* energies were calculated at the CASSCF level of theory, employing an active space of either 7 or 8 electrons in 8 orbitals, *i.e.*, CAS(7,8) or CAS(8,8) with a 6-31+G* basis set. The active space comprises the oxygen lone pair, $1\pi\pi$, $1\pi\pi^*$, and $1\pi\sigma^*$ orbitals as shown in Fig. 2. The electronic structure calculations were carried out using the Gaussian 09 program.⁴² The model for the cation states is similar to that previously published in a study of the photoelectron spectrum of phenol.²¹ These parameters then underwent some minor adjustments to make the calculated absorption and photoelectron spectra match experiments. The final values are listed in the ESI.[†]

2.1.2 Quantum dynamical simulations. Quantum dynamical simulations were performed by numerically solving the time-dependent Schrödinger equation using the multi-layer (ML) variant^{32,33,43} of the multiconfiguration time-dependent Hartree (MCTDH) algorithm.^{44,45} The ML-MCTDH approach

enables wavepacket dynamics to be performed for many degrees of freedom.^{32,46,47} In the original MCTDH approach, the multiconfigurational ansatz for the nuclear wavefunction $\Psi(Q, t)$ for a system with f degrees of freedom is expressed as a Hartree product of time-dependent basis functions or single-particle functions (SPFs) as

$$\Psi(Q_1, \dots, Q_f, t) = \sum_{j_1=1}^{n_1} \dots \sum_{j_f=1}^{n_f} A_{j_1 \dots j_f}(t) \prod_{\kappa=1}^f \phi_{j_\kappa}^{(\kappa)}(Q_\kappa, t) \quad (15)$$

where $Q_1, \dots, Q_f$ are the nuclear coordinates for the vibrational modes, $A_{j_1 \dots j_f}$ denote the time-dependent expansion coefficients and $\{\phi_{j_\kappa}^{(\kappa)}\}$ denote the SPFs for each degree of freedom. The SPFs are further expressed as linear combinations of a time-independent primitive basis $\chi_{i\kappa}^{(\kappa)}$ that depend on the particle coordinate Q_κ

$$\phi_{j_\kappa}^{(\kappa)}(Q_\kappa, t) = \sum_{i_\kappa=1}^{N_\kappa} c_{i_\kappa j_\kappa}^{(\kappa)}(t) \chi_{i_\kappa}^{(\kappa)}(Q_\kappa). \quad (16)$$

In the ML-MCTDH scheme, the wavefunction ansatz corresponds to a hierarchical expansion of SPFs in a recursive way to form an ML-tree structure where each layer l acts as a set of SPFs for the layer above ($l-1$) as

$$\begin{aligned} &\phi_m^{l-1; \kappa_1 \dots \kappa_{l-1}}(Q_{\kappa_{l-1}}^{l-1; \kappa_1 \dots \kappa_{l-2}}, t) \\ &= \sum_{j_1=1}^{n_1} \dots \sum_{j_{p_{\kappa_l}}=1}^{n_{\kappa_l}} A_{m; j_1 \dots j_{p_{\kappa_l}}}^{l; \kappa_1 \dots \kappa_{l-1}}(t) \prod_{\kappa_l=1}^{p_{\kappa_l}} \phi_{j_{\kappa_l}}^{(l; \kappa_1 \dots \kappa_{l-1})}(Q_{\kappa_l}^{l; \kappa_1 \dots \kappa_{l-1}}, t) \end{aligned} \quad (17)$$

where $\kappa_1, \dots, \kappa_{l-1}$ denote the indices of the logical degrees of freedom which are a combination of several vibrational modes starting from each node on the top layer down to a particular primary coordinate. The hierarchical expansion is finally truncated by a direct product expansion of the SPFs of the last layer in terms of time-independent primitive basis functions (similar to eqn (16)).

ML expansion of the wavepacket can be intuitively visualized using ML-tree diagrams.³³ Fig. 3 depicts the ML-tree structure used in the present work, in which a circle represents a node that stands for a set of A-coefficients and a square represents a set of time-independent primitive basis functions. The number in the circles denotes the node number in the layered structure, while within nodes, the number next to the link lines represent the numbers of SPFs used on the node. The number on the lines between a circle and a square represent the number of primitive functions or grid points used for that particular degree of freedom. Primitive basis functions are usually chosen as discrete variable representation (DVR) functions. Herein, we adopted Hermite DVRs for the vibrational modes. The depth of the tree is four layers where the first layer separates the electronic (el) and vibrational (r, theta, and v_x) degrees of freedom. The electronic continuum (Elcont) is added to model the kinetic energy of the ejected electron. The electronic degree of freedom represents the manifold of electronic states; the wavefunction is expanded in the set of electronic states where each electronic state-dependent component is expanded in the

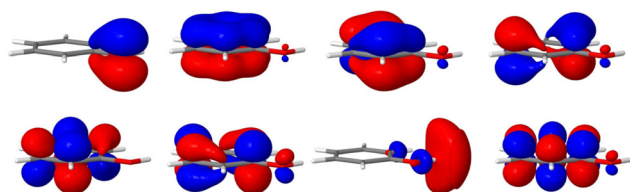


Fig. 2 Molecular orbitals of phenol used for the CASSCF calculations. The active space employed has 7 electrons in 8 orbitals comprising the oxygen lone pair, $1\pi\pi$, $1\pi\pi^*$ and $1\pi\sigma^*$ orbitals.



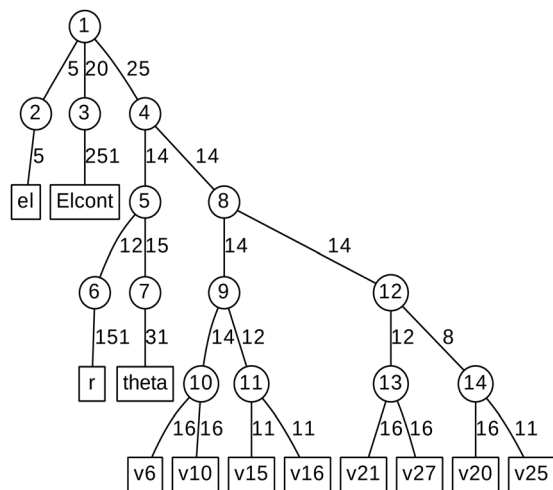


Fig. 3 Diagrammatic representation of the ML-MCTDH wavefunction. The circle represents a node on each layer and the square represents primitive bases or grids. 'el' indicates electronic degrees of freedom, 'Elcont' indicates the electronic continuum, 'r' indicates O–H bond stretch, 'theta' indicates C–C–O–H bond torsion, and 'v_x' indicates the vibrational normal modes (see text for details).

ML-MCTDH form. The particular choice of ML-tree strongly influences the numerical effort.

The ML-MCTDH equations of motion (EOMs) for the SPFs and expansion coefficients are derived variationally using the Dirac–Frenkel variational principle and applying the recursive algorithm of Manthe.^{33,48} This leads to an optimal description of the evolving wavepacket for a given choice of basis and layering scheme. The EOMs are integrated following variable mean field technique with Adams–Bashforth–Moulton predictor–corrector integrator of order 6 and accuracy 10^{-5} . Standard convergence tests were carried out to assess the precision of the integrator and the number of SPFs and primitive basis functions used. This is usually done by monitoring the population of the natural orbitals – the largest population of the least occupied natural orbital was noted to be $\sim 10^{-3}$. The ground state wavepacket is obtained by relaxation method⁴⁹ following imaginary time propagation of an initial wavepacket.^{45,50} All ML-MCTDH simulations were performed using the Quantics suite of programs.⁴⁰

Within Fermi's golden-rule limit,⁵¹ assuming a direct transfer of the initial wavepacket to the final state, a spectrum can be directly obtained as the Fourier transform of the autocorrelation function $C(t)$ as

$$I(\omega) \propto \omega \int_{-\infty}^{\infty} dt C(t) e^{i\omega t} \quad (18)$$

where

$$C(t) = \langle \Psi(0) | \Psi(t) \rangle = \langle \Psi(t/2) | \Psi(t/2) \rangle. \quad (19)$$

This is the Condon approximation, where the nuclei are assumed to remain stationary during the fast electronic transition. To reduce spurious structures (Gibbs phenomenon) arising due to a finite time propagation T in the Fourier transform,

the autocorrelation function is multiplied by a weight function $\cos(\pi t/2T)$ which ensures it vanishes as $t \rightarrow T$.^{45,52} Experimental line broadening effects are obtained by further multiplying the autocorrelation function with a damping function $\exp(-t/\tau)$, where the damping time τ needs to be tuned. Within the (ML-)MCTDH framework, the photoelectron spectrum can also be simulated by explicitly including the pump and probe laser pulses. This is based on the pioneering work of Seel and Domcke⁵³ where a continuum of free-electron states (Elcont in Fig. 3) is included explicitly to represent the kinetic energy of the ionised electron. The electronic model space now comprises besides the manifold of electronic states, an ionisation continuum associated with each of the cationic states.

$$\Psi(\mathbf{Q}, E, t) = S_0 + S_1 + S_2 + \int_0^{\infty} dE (D_0(E) + D_1(E)) \quad (20)$$

where the energy E of the continuum states is treated as an additional degree of freedom, which is discretized by a DVR grid with SPFs on the top of the grid. The SPFs are represented by sine-DVR functions which effectively discretises the continuum. The photoelectron spectra is then calculated as the population of the continuum states in the long time limit as⁵⁴

$$I(E) \propto |\langle \Psi(t \rightarrow \infty) | D_0(E) \rangle|^2 + |\langle \Psi(t \rightarrow \infty) | D_1(E) \rangle|^2. \quad (21)$$

Herein, we choose 251 DVR points spanning the energy range 0.0 to 4.0 eV. The idea of invoking an electronic continua to model the outgoing electron also finds fruitful application in the simulation of ultrafast charge migration phenomenon.⁵⁵

The light field is added to the molecular Hamiltonian to get the full Hamiltonian for the system

$$\mathbf{H} = \mathbf{H}_{\text{mol}} + \mathbf{H}_{\text{L}}(t) \quad (22)$$

where the light-matter Hamiltonian is

$$\begin{aligned} \mathbf{H}_{\text{L}}(t) = & \left(\sum_{f=2,3} \mu_{f1} + \sum_{f=4,5} (I_{f2} + I_{f3}) \right) \\ & \times \sum_{p=1,2} s_p G_p(t - t_p) \cos(\omega_p(t - t_p)) \end{aligned} \quad (23)$$

where μ_{21} , μ_{31} are the transition dipole moments from the ground state to S_1 and S_2 , respectively, I_{42} , I_{52} the ionisation propensities from S_1 to the cation states D_0 and D_1 , and I_{43} , I_{53} the ionisation propensities from S_2 . The two pulses are defined by a maximum pulse amplitude, s_p , frequency, ω_p , and a Gaussian envelope

$$G_p(t - t_p) = \frac{1}{\tau_p} \sqrt{\frac{4 \ln 2}{\pi}} \exp\left(-\frac{4 \ln 2}{\tau_p^2} (t - t_p)^2\right) \quad (24)$$

The maximum amplitude of each pulse is thus at t_p and the time delay between them is $t_2 - t_1$.

In these simulations, both pulses are centred at $t_1 = t_2 = 0$. The width of the pulses is given by the FWHM, τ_p . Here, both pulses use $\tau_p = 50$ fs. It should be noted that this is shorter than the experimental pulse widths, which were 200 fs. This was chosen to keep the simulations short, and means that the bandwidth is slightly broader than the experiment, but this



should be insignificant for the overall dynamics seen. The strength parameters for both pulses were set to 0.01 a.u., *i.e.* a weak pulse. The strength of the interaction is then modulated by the values of the transition dipoles and ionisation propensities. The transition dipole from S_0 to S_1 is the dominant factor for the excitation. μ_{21} was taken to be 1.0. The transition dipole from the ground state to S_2 is a factor of 10 smaller than that of the S_1 state, as reflected in the oscillator strengths listed in Table 1 which depend on the square of the transition dipole. For this reason a value for $\mu_{31} = 0.1$ was taken. Due to the symmetry of the states, the transition dipoles are in different directions, with the S_1 transition dipole in the x, y plane and the S_2 along the z -axis. The model thus assumes unpolarised light.

The ionisation propensity from the S_0 and S_1 states to the D_0 and D_1 cationic states can be obtained by computing the photoionisation cross-section.²² This is done by using ezDyson code^{56,57} (version 5.0) where the photoionisation matrix elements $\langle \phi_{I\alpha}^D | \mu | \psi_k \rangle$ (for each transition at the FC point) are evaluated numerically on a grid and averaged isotropically over all molecular orientations. ψ_k is the initial state wavefunction, μ is the molecular dipole operator and $\phi_{I\alpha}^D$ is the Dyson orbital. The Dyson orbitals are one-electron functions defined as the overlap between the neutral and cationic states as⁵⁸

$$\phi_{I\alpha}^D = \sqrt{N} \int \Phi_I^{(N)*}(\mathbf{r}_1, \dots, \mathbf{r}_N) \Phi_\alpha^{(N-1)}(\mathbf{r}_1, \dots, \mathbf{r}_{N-1}) d\mathbf{r}_1 \dots d\mathbf{r}_{N-1}. \quad (25)$$

These calculations were performed using the equation-of-motion ionisation-potential method combined with the coupled-cluster theory with single and double excitations (EOM-IP-CCSD)⁵⁹ and aug-cc-pVDZ basis set⁶⁰ using Q-Chem 5.4 program.⁶¹ The results are shown in Fig. 4. The ionisation cross-sections are seen to be fairly independent of photon energy across the range of interest.

For accurate simulations, ionisation propensities should be used in simulations as a propensity surface to take into account the geometry dependence of the dipole moment and Dyson

orbitals. However, for simplicity, the propensities used here were taken as constants based on the values in Fig. 4 calculated at the Franck–Condon point. This is equivalent to the Condon approximation used for the simulation of excitation spectra and should be a good approximation in the diabatic picture used for the Hamiltonian in this work as the orbitals should change slowly with changing nuclear geometry. For the Hamiltonian, the S_1 and S_2 ionisation propensities were taken to be equal and given a value of 1.0, except for the S_2 – D_1 ionisation which was given a value of 0.0.

2.2 Experimental details

Our molecular beam velocity-map imaging (VMI) photoelectron spectrometer has been described in detail elsewhere.⁶² Briefly, a molecular beam of phenol was created by expanding 1.8 bar of helium carrier gas through phenol heated to 70 °C in the sample tube of a pulsed Even-Lavie valve, operating at a repetition rate of 500 Hz. The molecular beam was collimated by a 1 mm skimmer as it couples into the differentially pumped VMI chamber where it is intersected with UV femtosecond laser pulses in the range 275–200 nm. Wavelengths in the range 275–235.5 nm were generated by sum frequency mixing the output of an optical parametric amplifier (Coherent Opera-F) with the fundamental of a Ti:sapphire regenerative amplifier (Coherent Legend) to produce tunable visible radiation, which was subsequently frequency doubled. The full-width at half-maximum (FWHM) sech^2 pulse durations have been measured to be in the range of around 175 fs at the relevant wavelengths through autocorrelation measurements using the nonresonant ionisation of butadiene at 235.5 nm. The 200 nm light was generated using a Coherent 4th harmonic generation system and previous measurements have suggested the sech^2 pulse duration is on the order of 200 fs FWHM. Pulse energies are $< 2 \mu\text{J}$ per pulse.⁶³ The photon flux was attenuated to keep the photoelectron count-rates below 5 photoelectrons per pulse to avoid detector saturation and multiphoton processes. The known binding energy of phenol⁶⁴ was used to confirm the absence of space-charge effects. Photoelectron images were recorded for 600 000 laser shots. Background images (without phenol) were also recorded for 600 000 laser shots and subtracted from the photoelectron images. Photoelectron spectra were recovered from the background-subtracted data using the pBSEX image inversion algorithm,⁶⁵ and the energy scale was calibrated by recording the 2+1 resonance-enhanced multiphoton ionisation (REMPI) spectrum of Xe at 249.6 nm for spectra recorded using wavelengths in the range 275–235.5 nm. The resolution was $\Delta E/E \approx 3\%$. For the 200 nm spectrum, the 3-photon ionisation of xenon at 200 nm was used to calibrate the detector, and the resolution was $\Delta E/E \approx 6\%$.

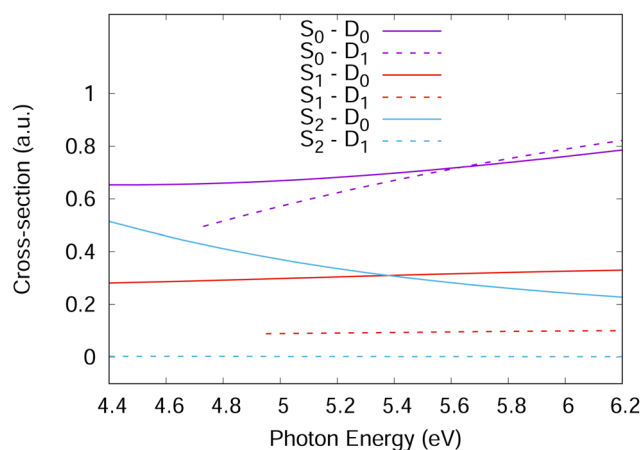


Fig. 4 One-photon photo-ionisation cross-sections of phenol as a function of incident photon energy corresponding to ionisation from S_1 and S_2 to D_0 and D_1 . For ionisation from S_0 , the cross-section for the two-photon process is shown.

3 Results and discussion

Since the intermediate electronic state plays a crucial role in the 1+1 REMPI spectra, we first calculate the UV absorption spectra to check that the bound excited states are well described by the





Fig. 5 The gas-phase UV-vis absorption spectrum of phenol plotted against a calculated spectrum corresponding to $S_1(^1B_2) \leftarrow S_0$ excitation (upper panel). An energy scale (in eV) is added corresponding to the wavelength (in nm) as a guide to the eye. The photoelectron spectra of phenol corresponding to $D_0 \leftarrow S_0$ and $D_1 \leftarrow S_0$ excitations (lower panel). The experimental photoelectron spectrum is a single photon He(I) spectrum adapted from ref. 66. Here we assume a vertical excitation in the theoretical simulations and damping times of 150 fs, 27 fs and 35 fs, respectively are considered.

model. Fig. 5(a) shows the gas-phase experimental absorption spectrum and a calculated absorption spectrum corresponding to $S_1 \leftarrow S_0$ excitation. In the simulation, a Franck-Condon excitation of the initial wavepacket to the S_1 state was assumed and a damping time of 150 fs applied to the resultant auto-correlation function. The $S_2(\pi\sigma^*)$ state is dark, with a much smaller oscillator strength than S_1 and does not provide significant intensity. The S_1 band has a strong 0_0^0 transition at 275.1 nm (4.5 eV) followed by another relatively strong absorption at 268.2 nm (experiment) or 267.3 nm (theory). The band origin and the relative spacing of the simulated spectrum matches well with the experimental spectrum, which suggest that the vibronic coupling Hamiltonian model can capture the excited-state dynamics involving the intermediate electronic state.

Before investigating the REMPI-PES, we simulated the photoelectron spectra assuming a vertical excitation of the initial wavepacket to each of the ionic states D_0 and D_1 . The resulting spectrum is given as a sum of individual spectrum from separate wavepacket propagations on the cationic states. This gives the vibrational fine structure associated with the photoelectron spectrum and checks the accuracy of the vibronic coupling model. Fig. 5(b) shows the resulting photoelectron spectra of these two lowest energy bands – the band between 8.2 and 9.0 eV corresponds to the $D_0(^2B_1) \leftarrow S_0(\tilde{X}^1A_1)$ transition, and the band between 9.2 and 10.0 eV corresponds to the $D_1(^2A_2) \leftarrow S_0(\tilde{X}^1A_1)$ transition. It is to be noted that both the spectral bands have been shifted to account for the zero point energy of the neutral

molecule, and damping times of 27 fs and 35 fs were chosen for the 2B_1 and 2A_2 bands, respectively. The spectrum indicates the 2B_1 band to be well-structured with progressions from several vibrational modes, while the 2A_2 band is mostly broad and lacks structure. The experimental photoelectron spectrum of phenol at 21 eV of Debies and Rabalais⁶⁶ is superimposed as this experiment used a single-photon ionisation scheme and the spectrum will therefore only contain signal from direct photoionisation from the ground electronic state. A good match between experiment and theory is noted with respect to the band origins 0_0^0 located at 8.37 eV and 9.3 eV and the relative spacing between the peaks.

With the benchmarked model Hamiltonian for the valence and cation states, we then simulated the one-colour two-photon REMPI spectra following excitation with UV femtosecond laser pulses. The laser pulse is now explicitly included in the simulations using the light-matter Hamiltonian of eqn (23). The state populations from the simulations with different laser pulse frequencies are shown in Fig. 6. At 275 nm, only the S_1 state is



Fig. 6 The time evolution of the S_1 and S_2 state populations following excitation from the S_0 state for different wavelengths. The electric field corresponding to a 235.5 nm pulse is also plotted as a guide to the eye. The depletion of the S_0 state population (scale on the right-hand axis) for the 235.5 nm pulse indicates a shoulder region (ca. 30 fs) similar to S_1 . The S_2 population for the 235.5 nm pulse is multiplied by a factor of 10 for better visualization.



populated, and the population follows the envelope of the laser pulse. This is typical of non-resonant excitation. Previous experiments have found a long-lived state at this excitation energy, but in our model the lowest state in S_1 lies just above 275 nm in energy. In contrast, for pulses with wavelengths 265.5 nm to 249.6 nm the population of S_1 rises and remains fairly constant after the pulse finishes, consistent with the very long decay lifetimes (around 1 nanosecond) observed experimentally in this wavelength range and previously assigned to tunneling under the S_2/S_1 CI.^{17,67} This is resonant excitation. No population of S_2 is seen in any of these simulations as the energy is too low for direct excitation, and also too low for the molecule to access the S_2/S_1 CI. For this reason the only contributions in the photoelectron spectra arise from ionisation from S_1 , going to either D_0 or D_1 .

The spectra obtained from the simulations from the populations of the continuum states using eqn (21), are plotted as a function of two-photon electron binding energy (eBE), *i.e.*

$$\text{eBE} = 2h\nu - \text{eKE} \quad (26)$$

where ν is the frequency of the incident pulses and eKE the kinetic energy of the emitted photoelectrons. Fig. 7(a) shows the experimental (sky blue) and simulated (black) 1+1 REMPI-PES spectra using UV pulses in the 275–200 nm range. Six different wavelengths were chosen to cover the key region of energy as 275 nm (~ 4.5 eV) corresponds to the $S_1 \leftarrow S_0$ vertical excitation energy and 249.6 nm (~ 4.96 eV) corresponds to the S_2/S_1 CI. Although the vertical excitation energy to the S_2

($\pi\sigma^*$) state lies around 5.64 eV, $S_2 \leftarrow S_0$, this transition possesses almost no oscillator strength indicating a negligibly small absorption coefficient and therefore it is optically dark. However, the S_2 state may still become involved in the resonance enhanced ionisation due to S_2/S_1 curve crossing and vibronic coupling. The first ionisation threshold corresponding to $D_0 \leftarrow S_0$ ionisation lies at 8.37 eV, and the second ionisation threshold corresponding to $D_1 \leftarrow S_0$ ionisation lies at 9.3 eV.

The 1+1 REMPI-PES at 275 nm (~ 4.5 eV) comprises a two-peak structure which corresponds to $D_0 \leftarrow S_0$ excitation *via* the intermediate excited state. But for shorter wavelengths, the 2-photon excitation energy becomes sufficient to include a contribution from the D_1 state. The calculated spectra in Fig. 7 are also broken down into contributions from the final cation states. Starting from 265.5 nm, spectral signatures from both $D_0 \leftarrow S_0$ (in purple lines) and $D_1 \leftarrow S_0$ (in green lines) are observed.

In general the shape and widths of the calculated spectra qualitatively match those of the experimental signals. The calculated spectra are, however, slightly red-shifted by approximately 0.1 eV and the calculated spectra have more structure. The lack of quantitative agreement is due to the simplicity of the model used in the calculations. For efficiency, the simulations only included 10 vibrational modes, which leads to a more structured spectra due to the missing intramolecular vibrational relaxation. The model is able to reproduce the vertical absorption and photoelectron spectra quite well, but the energies and vibrational frequencies do not match exactly the real values resulting in discrepancies in the relative intensities of



Fig. 7 1+1 REMPI-PES of phenol at six different wavelengths. The experimental spectra are depicted in sky blue lines. The simulated spectra (in black lines) are computed as a sum of the individual contributions arising from $D_0 \leftarrow S_0$ (in purple lines) and $D_1 \leftarrow S_0$ (in green lines), respectively. (a) 2-photon electron binding energy. The vertical line represents the $D_0 \leftarrow S_0$ ionisation threshold. (b) 1-photon electron binding energy. The vertical line represents the $D_0 \leftarrow S_1$ ionisation threshold.



spectral peaks that may become significant in the more sensitive calculations including a pulse. The use of the Condon approximation for the excitation may also be significant in the latter case where the excitation process must be described accurately. It is also possible that the simple pulse used in the calculations does not adequately match the experimental pulse.

The calculated spectrum at 200 nm is clearly not a good match. This is because the S_3 state was not incorporated in our vibronic coupling Hamiltonian for simplicity, and for excitation wavelengths shorter than 200 nm it becomes relevant since the vertical excitation energy to bound S_3 ($2\pi\pi^*$) state is around 203 nm (~ 6.1 eV). The missing spectral features at higher energies in the simulated spectrum compared to the experimental spectrum reflect the omission of this state.

Despite the lack of quantitative agreement, the main features of the calculated spectra allow an analysis of the experimental spectra, and in particular allow the assignment of features due to resonant and non-resonant ionisation. Returning to the spectra in Fig. 7, from 275 nm to 249.6 nm, the photoelectron spectrum peak moves to higher energy with higher pulse energy. In the experimental spectrum, a shoulder grows in at a fairly constant energy at the low end of the spectrum, and this is seen as a peak around 8.6 eV that becomes more intense with increasing frequency in the calculated spectra. This peak and the main peak are due to ionisation to D_0 , while the ionisation to D_1 is much lower in intensity, growing in at the high energy side with increasing pulse frequency.

At 235.5 nm, there are clearly two peaks that are both due to ionisation to D_0 . One is centred around 8.6 eV with the higher energy peak at 9.2 eV. Looking at the state populations in Fig. 6, at this frequency there is now some direct, non-resonant, excitation of the S_2 state, but it is very small and does not contribute significantly to the photoelectron spectrum. The low energy peak, as well as the low energy shoulders at lower photon energies, are thus the signals from non-resonant ionisation of S_1 . The larger peak that shifts with photon energy is due to the resonant ionisation from S_0 via S_1 .

The resonant ionisation peak shifts due to the vibrational energy carried over from the neutral excitation to the ion. This energy is $E_{\text{vib}} = h\nu - S_n$, where S_n is the ground vibrational state energy for the neutral state being excited. Thus the electron kinetic energy is $eKE = 2h\nu - (D_n + E_{\text{vib}})$, where D_n is the ground vibrational state energy for the cation state being ionised into. This leads to the one-photon binding energy

$$D_n - S_n = h\nu - eKE \quad (27)$$

which is plotted in Fig. 7(b). This peak is seen to be fairly constant in this representation, and provides an estimate of the S_1 adiabatic ionisation energy of 3.9 eV.

As mentioned, at energies below 248 nm the S_2/S_1 conical intersection is not accessible and no dissociation is seen on the time-scale of the simulations (200 fs). At 235.5 nm, some population enters the $S_2(\pi\sigma^*)$ state and dissociation is observed with this small amount of nuclear density associated with S_2

moving out along the O–H coordinate. At 200 nm, Fig. 6 shows that the S_2 state becomes populated after the S_1 state due to internal conversion through the intersection. The S_1 population in fact decays through the conical intersection entirely into S_2 , which then decays as the O–H bond breaks.

4 Conclusions

We have performed a joint theoretical and experimental study to probe the excited-state dynamics of phenol from 1+1 REMPI spectra of the first two ionisation bands, carrying out quantum wavepacket calculations to model and interpret the experimental spectra.

Building on our earlier work modelling the photoelectron spectrum of phenol,²¹ we constructed a new vibronic coupling Hamiltonian in normal mode coordinates for the lowest two valence states of phenol. The electronic structure calculations were carried out using the CASSCF method and the model parameters entering the Hamiltonian were determined by least-square fitting of the *ab initio* points. The absorption and photoelectron spectrum calculated using the two model Hamiltonians and quantum dynamics simulations show an excellent agreement with experimental spectra while considering a vertical excitation in the FC region.

To model the REMPI spectra, the Hamiltonians for the valence and ionised states of phenol were combined and laser pulses with different frequencies explicitly included in the simulations. The outgoing electron is modelled by the ionisation continuum model which populates electron kinetic energy states and allows the excitation energy dependent photoelectron spectra to be calculated. The calculated spectra are in good agreement with the experimental spectra, except at 200 nm where the model needs to be further improved by including the S_3 state, which dominates the experimental spectrum at this wavelength.

From an analysis of the state populations, it is clear that the photoelectron signal is entirely due to ionisation from the $S_1(\pi\pi^*)$ state. Ionisation takes place to both D_0 and D_1 cation states. The lower D_0 band has regions that can be assigned to high energy 'direct', non-resonant, ionisation that gives rise to a peak with constant electron binding energy and 'indirect' ionisation due to ionising S_1 after resonant excitation that gives to a peak which shifts with excitation energy.²⁸

Femtosecond time-resolved photoelectron spectroscopy (TRPES) is a powerful tool to track the evolution of electronic structure following photoexcitation and it is particularly advantageous when multiple excited states are involved in the dynamics.^{28,68–71} Here, we used femtosecond 1+1 REMPI at a series of wavelengths to see the photon energy dependence of the photoelectron signal and thus probe the short time dynamics. Adding in time delays between the excitation and ionisation pulses would gain further information by following the relaxation dynamics as a function of time.

Finally, although understanding the electronic structure and relaxation dynamics of isolated molecular chromophores in the



gas-phase lies at the heart of photochemical reactions, the environment can also tune the dynamics. To that end, studies have also looked into the role of an aqueous environment and highlighted the possibility of formation of solvated electrons on ultrafast timescales.²² Currently, liquid-jet photoelectron spectroscopy experiments are becoming feasible that can reveal the effect of complex environments.⁷² The present work strengthens our understanding on the UV photoresponse of phenol in the gas-phase and will serve as a benchmark to study the effect of complex environments in which the chromophores are usually embedded. This work will also serve as a test-bed to conduct TRPES studies on phenol.

Conflicts of interest

There are no conflicts to declare.

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References

- 1 B. A. Barry, *Photochem. Photobiol.*, 1993, **57**, 179–188.
- 2 J. J. van Thor, *Chem. Soc. Rev.*, 2009, **38**, 2935–2950.
- 3 A. Acharya, A. M. Bogdanov, B. L. Grigorenko, K. B. Bravaya, A. V. Nemukhin, K. A. Lukyanov and A. I. Krylov, *Chem. Rev.*, 2017, **117**, 758–795.
- 4 A. Henley and H. H. Fielding, *Int. Rev. Phys. Chem.*, 2019, **38**, 1–34.
- 5 A. L. Sobolewski and W. Domcke, *J. Phys. Chem. A*, 2001, **105**, 9275–9283.
- 6 A. L. Sobolewski, W. Domcke, C. Dedonder-Lardeux and C. Jouvet, *Phys. Chem. Chem. Phys.*, 2002, **4**, 1093–1100.
- 7 C.-M. Tseng, Y. T. Lee and C.-K. Ni, *J. Chem. Phys.*, 2004, **121**, 2459–2461.
- 8 Z. Lan, W. Domcke, V. Vallet, A. L. Sobolewski and S. Mahapatra, *J. Chem. Phys.*, 2005, **122**, 224315.
- 9 M. Abe, Y. Ohtsuki, Y. Fujimura, Z. Lan and W. Domcke, *J. Chem. Phys.*, 2006, **124**, 224316.
- 10 M. G. D. Nix, A. L. Devine, B. Cronin, R. N. Dixon and M. N. R. Ashfold, *J. Chem. Phys.*, 2006, **125**, 133318.
- 11 M. N. R. Ashfold, B. Cronin, A. L. Devine, R. N. Dixon and M. G. D. Nix, *Science*, 2006, **312**, 1637–1640.
- 12 M. N. R. Ashfold, A. L. Devine, R. N. Dixon, G. A. King, M. G. D. Nix and T. A. A. Oliver, *Proc. Natl. Acad. Sci. U. S. A.*, 2008, **105**, 12701–12706.
- 13 M. G. D. Nix, A. L. Devine, R. N. Dixon and M. N. R. Ashfold, *Chem. Phys. Lett.*, 2008, **463**, 305–308.
- 14 M. L. Hause, Y. H. Yoon, A. S. Case and F. F. Crim, *J. Chem. Phys.*, 2008, **128**, 104307.
- 15 A. Iqbal, M. S. Y. Cheung, M. G. D. Nix and V. G. Stavros, *J. Phys. Chem. A*, 2009, **113**, 8157–8163.
- 16 R. N. Dixon, T. A. A. Oliver and M. N. R. Ashfold, *J. Chem. Phys.*, 2011, **134**, 194303.
- 17 G. M. Roberts, A. S. Chatterley, J. D. Young and V. G. Stavros, *J. Phys. Chem. Lett.*, 2012, **3**, 348–352.
- 18 K. R. Yang, X. Xu, J. Zheng and D. G. Truhlar, *Chem. Sci.*, 2014, **5**, 4661–4680.
- 19 X. Zhu and D. R. Yarkony, *J. Chem. Phys.*, 2016, **144**, 024105.
- 20 C. Xie, J. Ma, X. Zhu, D. R. Yarkony, D. Xie and H. Guo, *J. Am. Chem. Soc.*, 2016, **138**, 7828–7831.
- 21 M. P. Taylor and G. A. Worth, *Chem. Phys.*, 2018, **515**, 719–727.
- 22 J. W. Riley, B. Wang, J. L. Woodhouse, M. Assmann, G. A. Worth and H. H. Fielding, *J. Phys. Chem. Lett.*, 2018, **9**, 678–682.
- 23 A. Henley, J. W. Riley, B. Wang and H. H. Fielding, *Faraday Discuss.*, 2020, **221**, 202–218.
- 24 G. Christopoulou, T. Tran and G. A. Worth, *Phys. Chem. Chem. Phys.*, 2021, **23**, 23684–23695.
- 25 N. Giri and S. Mahapatra, *J. Chem. Phys.*, 2022, **156**, 094305.
- 26 S. Matsika and P. Krause, *Annu. Rev. Phys. Chem.*, 2011, **62**, 621–643.
- 27 M. N. R. Ashfold and J. D. Howe, *Annu. Rev. Phys. Chem.*, 1994, **45**, 57–82.
- 28 H. H. Fielding and G. A. Worth, *Chem. Soc. Rev.*, 2018, **47**, 309–321.
- 29 C. P. Schick, S. D. Carpenter and P. M. Weber, *J. Phys. Chem. A*, 1999, **103**, 10470–10476.
- 30 L. S. Cederbaum, W. Domcke, H. Köppel and W. Von Niessen, *Chem. Phys.*, 1977, **26**, 169–177.
- 31 H. Köppel, W. Domcke and L. S. Cederbaum, *Adv. Chem. Phys.*, 1984, **57**, 59–246.
- 32 H. Wang and M. Thoss, *J. Chem. Phys.*, 2003, **119**, 1289–1299.
- 33 U. Manthe, *J. Chem. Phys.*, 2008, **128**, 164116.
- 34 G. A. Worth, H. D. Meyer, H. Köppel, L. S. Cederbaum and I. Burghardt, *Int. Rev. Phys. Chem.*, 2008, **27**, 569–606.
- 35 M. Baer, *Chem. Phys.*, 2000, **259**, 123–147.
- 36 G. A. Worth and L. S. Cederbaum, *Annu. Rev. Phys. Chem.*, 2004, **55**, 127–158.
- 37 A. Lehr, S. Gómez, M. A. Parkes and G. A. Worth, *Phys. Chem. Chem. Phys.*, 2020, **22**, 25272–25283.
- 38 S. Faraji, M. Vazdar, V. Sivarajana Reddy, M. Eckert-Maksic, H. Lischka and H. Köppel, *J. Chem. Phys.*, 2011, **135**, 154310.
- 39 C. Cattarius, A. Markmann and G. A. Worth, *The VCHAM program*, 2007, <https://www.pci.uni-heidelberg.de/tc/ustr/mctdh/>.
- 40 G. A. Worth, K. Giri, G. W. Richings, I. Burghardt, M. H. Beck, A. Jäckle and H.-D. Meyer, *The Quantics Package, Version 1.2*, University of Birmingham, Birmingham, UK, 2016.



- 41 G. A. Worth, *Comput. Phys. Commun.*, 2020, **248**, 107040.
- 42 M. J. Frisch, G. W. Trucks, H. B. Schlegel, G. E. Scuseria, M. A. Robb, J. R. Cheeseman, G. Scalmani, V. Barone, G. A. Petersson, H. Nakatsuji, *et al.*, *Gaussian 09, Revision A.02*, Gaussian, Inc., 2009.
- 43 O. Vendrell and H.-D. Meyer, *J. Chem. Phys.*, 2011, **134**, 044135.
- 44 H.-D. Meyer, U. Manthe and L. S. Cederbaum, *Chem. Phys. Lett.*, 1990, **165**, 73–78.
- 45 M. H. Beck, A. Jäckle, G. A. Worth and H.-D. Meyer, *Phys. Rep.*, 2000, **324**, 1–105.
- 46 T. Westermann, R. Brodbeck, A. B. Rozhenko, W. Schoeller and U. Manthe, *J. Chem. Phys.*, 2011, **135**, 184102.
- 47 J. Schulze, M. F. Shibl, M. J. Al-Marri and O. Kühn, *J. Chem. Phys.*, 2016, **144**, 185101.
- 48 U. Manthe, *J. Chem. Phys.*, 2009, **130**, 054109.
- 49 R. Kosloff and H. Tal-Ezer, *Chem. Phys. Lett.*, 1986, **127**, 223–230.
- 50 H.-D. Meyer and G. A. Worth, *Theor. Chem. Acc.*, 2003, **109**, 251–267.
- 51 R. Schinke, *Photodissociation Dynamics*, Cambridge University Press, Cambridge, 1991.
- 52 A. Raab, G. A. Worth, H.-D. Meyer and L. S. Cederbaum, *J. Chem. Phys.*, 1999, **110**, 936–946.
- 53 M. Seel and W. Domcke, *J. Chem. Phys.*, 1991, **95**, 7806–7822.
- 54 G. A. Worth, R. E. Carley and H. H. Fielding, *Chem. Phys.*, 2007, **338**, 220–227.
- 55 D. Dey, A. I. Kuleff and G. A. Worth, *Phys. Rev. Lett.*, 2022, **129**, 173203.
- 56 S. Gozem and A. I. Krylov, *ezDyson*, <https://iopenshell.usc.edu/downloads/ezdyson>.
- 57 S. Gozem, A. O. Gunina, T. Ichino, D. L. Osborn, J. F. Stanton and A. I. Krylov, *J. Phys. Chem. Lett.*, 2015, **6**, 4532–4540.
- 58 G. Wu, S. P. Neville, O. Schalk, T. Sekikawa, M. N. R. Ashfold, G. A. Worth and A. Stolow, *J. Chem. Phys.*, 2015, **142**, 074302.
- 59 A. I. Krylov, *Ann. Rev. Phys. Chem.*, 2008, **59**, 433–462.
- 60 R. A. Kendall, T. H. Dunning Jr. and R. J. Harrison, *J. Chem. Phys.*, 1992, **96**, 6796–6806.
- 61 Y. Shao, Z. Gan, E. Epifanovsky, A. T. Gilbert, M. Wormit, J. Kussmann, A. W. Lange, A. Behn, J. Deng and X. Feng, *et al.*, *Mol. Phys.*, 2015, **113**, 184–215.
- 62 A. R. McKay, M. E. Sanz, C. R. S. Mooney, R. S. Minns, E. M. Gill and H. H. Fielding, *Rev. Sci. Instrum.*, 2010, **81**, 123101.
- 63 O. M. Kirkby, M. A. Parkes, S. P. Neville, G. A. Worth and H. H. Fielding, *Chem. Phys. Lett.*, 2017, **683**, 179–185.
- 64 K. Fuke, H. Yoshiuchi, K. Kaya, Y. Achiba, K. Sato and K. Kimura, *Chem. Phys. Lett.*, 1984, **108**, 179–184.
- 65 G. A. Garcia, L. Nahon and I. Powis, *Rev. Sci. Instrum.*, 2004, **75**, 4989–4996.
- 66 T. Debies and J. Rabalais, *J. Electron Spectrosc. Relat. Phenom.*, 1972, **1**, 355–370.
- 67 R. A. Livingstone, J. O. F. Thompson, M. Iljina, R. J. Donaldson, B. J. Sussman, M. J. Paterson and D. Townsend, *J. Chem. Phys.*, 2012, **137**, 184304.
- 68 D. M. Neumark, *Annu. Rev. Phys. Chem.*, 2001, **52**, 255–277.
- 69 A. Stolow and J. G. Underwood, *Adv. Chem. Phys.*, 2008, **139**, 497–583.
- 70 P. Chakraborty, Y. Liu, S. McClung, T. Weinacht and S. Matsika, *J. Phys. Chem. Lett.*, 2021, **12**, 5099–5104.
- 71 M. S. Schuurman and V. Blanchet, *Phys. Chem. Chem. Phys.*, 2022, **24**, 20012–20024.
- 72 W. G. Fortune, M. S. Scholz and H. H. Fielding, *Acc. Chem. Res.*, 2022, **55**, 3631–3640.

