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1	Analysis of Mass Dependent and Mass Independent Selenium Isotope Variability
2	in Black Shales
3	
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15	Abstract
16	The measurement of selenium isotope ratios is of increasing interest for understanding
17	redox conditions in present and past surface environments. Se has six stable isotopes,
18	and is therefore well suited for isotope analysis by double spiking. However due to
19	relatively large interferences on every isotope, and complex chemical purification
20	methods that frequently do not generate 100% yields, rigorously determining the
21	accuracy of measurements is critical. Here we present analyses of USGS shale
22	standards (SCo-1 and SGR-1b), as representatives of material which might be of
23	interest to Se isotope studies. We have made analyses using two separate double
24	spikes ( <sup>74</sup> Se- <sup>78</sup> Se and <sup>78</sup> Se- <sup>82</sup> Se), and compare them to previously published results. In
25	addition, we present models of the effects of uncorrected interferences on double

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26	spike inversions. This leads us to propose $\delta^{82/76}Se$ (parts per thousand deviation of
27	$^{82}\text{Se}/^{76}\text{Se}$ from NIST SRM-3149) values of -0.22 $\pm$ 0.15 for SCo-1, and +0.25 $\pm$ 0.17
28	for SGR-1b. Further, we present a new method of measuring Se isotopes by
29	desolvation nebulisation. Se sensitivity is enhanced by a factor of 100-200 times by
30	doping solutions with pure Mg, leading to almost a factor of two less material
31	required compared to the more standard hydride generation. Interferences are
32	different compared to our standard hydride generation protocol, but analyses of
33	double spiked NIST-3149 shows that this method can generate accurate isotope ratios.
34	Finally, mass independent fractionation (MIF) of sulphur isotopes has generated
35	considerable interest for constraining the early oxygenation of the atmosphere. Given
36	the chemical similarities between S and Se, Archean shales with S MIF might be
37	expected to exhibit Se MIF. However, within our analytical uncertainty of $\pm 0.4$ -0.5
38	(parts per 10,000), there is no resolvable Se MIF in these samples, indicating different
39	atmospheric cycling of Se and S.
40	
41	Keywords
42	Ocean redox; Se isotopes; black shales; mass dependent fractionation; mass
43	independent fractionation; hydride generation; desolvation
44	

## 45 **1.0 Introduction**

Selenium and its stable isotopes are increasingly generating interest in the geochemical community for their potential to trace present and past surface water redox processes<sup>1-7</sup>. As for S, microbial reduction of Se is thought to be the dominant cause of natural Se isotope fractionation<sup>3, 8</sup>. The chemistries of S and Se have notable similarities but, significantly, the redox transitions between the highest naturally

#### Journal of Analytical Atomic Spectrometry

occurring oxidation state of selenium, Se(VI), and its common reduced forms, Se
(IV), Se(0) and Se(-II), occur at higher pE than for the dominant sulphur redox states,
S(VI) and S(-II)<sup>3, 6, 9-12</sup>. Selenium isotope measurements may therefore be suited to
trace changes in ocean redox at relatively oxidised conditions, after the great
oxidation event.

Of the existing means of analysis, gas-source spectrometry requires prohibitively large quantities of Se<sup>6</sup>. Negative thermal ionisation mass spectrometry has proven to be useful<sup>2</sup>, but it has only been since the development of multi-collector plasma ionisation mass spectrometers that interest in Se isotope measurements for geochemical purposes has begun to grow<sup>13</sup>. Even so, Se ionises poorly in a plasma and Se is generally introduced as a Se-hydride gas, thereby avoiding the sensitivity loss due to a hydrous carrier matrix<sup>13, 14</sup>. However, hydride generation has its own inherent problems, as elements that directly interfere with Se isotopes (e.g. Ge, As), as well as those that may cause matrix effects (e.g. Sb), also readily form hydrides, and therefore pre-analysis chemical purification must take place, as well as interference corrections. Coupled with interferences stemming from the plasma carrier gas (ArAr and Kr), these factors make Se isotope measurements challenging.

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Purification chemistry for Se from common geological matrices has been described by a number of authors, using the now fairly well established thiol cotton fibre method<sup>7, 9, 13, 15-17</sup>. However, given the difficulty in obtaining 100% yields from this chemistry, the potential exists for isotopic fractionation of the sample to occur during sample preparation. This concern can be circumvented by using a double-spiked approach, which corrects for any mass-dependent fractionation during sample processing if the samples and spike are equilibrated prior to chemical purification<sup>1, 3, 7</sup>, <sup>14, 16, 18-21</sup>. Double spiked analyses are also robust against possible differential 

#### Journal of Analytical Atomic Spectrometry

instrumental fractionation of samples with residual matrix compared to clean solution standards, which is a major consideration in sample-standard bracketing methods. Selenium has six stable isotopes, apparently providing ample choice for selecting the four isotopes required for double spiked analysis. The potential drawback of this method, however, is that at these four isotopes must be fully corrected for any interferences, compared to only two required in sample-standard bracketing. The benefits of double spiking outlined above should outweigh those of sample-standard bracketing in measuring mass dependent isotope variability, but interferences must be resolved or accurately corrected. Here we examine a number of these issues, with reference to our own work and to also to different studies in the literature. Since Se is strongly associated with organic matter<sup>8, 22, 23</sup>, shale is the most 

commonly used rock type for Se isotope analyses of ancient past environments. Therefore well-characterised shale standards are required in order for Se isotope studies to be cross-calibrated. However, published values for the common USGS shale standard SGR-1 differ by up to 0.74% on the <sup>82</sup>Se/<sup>76</sup>Se ratio<sup>7, 13, 17, 19, 24, 25</sup>. Here we examine the Se isotope composition of two USGS shales (SGR-1 and SCo-1) with multiple, independently constrained, double spike inversions, with the aim of providing useful isotopic reference compositions. We also examine some Archean rocks for the potential of Se mass independent fractionation (MIF), given the presence of sulphur MIF in these rocks.

#### **2.0 Double spike**

98 Selenium has six stable isotopes (given that  $^{82}$ Se has a half life of  $\sim 10^{20}$  years 99 and can be considered stable):  $^{74}$ Se,  $^{76}$ Se,  $^{77}$ Se,  $^{78}$ Se,  $^{80}$ Se and  $^{82}$ Se. Their natural 100 abundances are detailed in Table 1. Two separate double spikes were trialled in this

#### Journal of Analytical Atomic Spectrometry

study. Both were made up from enriched single isotope standards from Cambridge Isotope Laboratories, Inc., with certified enrichment factors between 92% for <sup>82</sup>Se, 97% for <sup>78</sup>Se and >99% for <sup>74</sup>Se. The enriched material was dissolved individually in concentrated HNO<sub>3</sub> to ensure no reductive loss of Se, diluted, and then mixed gravimetrically, according to the compositions suggested by the Double Spike Toolbox<sup>26</sup>. The double spikes were calibrated by standard addition, by mixing NIST-3149 and the double spike in various proportions, which were then used to constrain the pure spike composition. This composition was also verified by measuring a pure double spike by sample-standard bracketing with NIST-3149.

Initially we used a  $^{74}$ Se- $^{78}$ Se spike (calibrated ratios:  $^{74}$ Se/ $^{78}$ Se =0.856,  $^{74}$ Se/ $^{77}$ Se =90.445,  $^{78}$ Se/ $^{77}$ Se =105.645), with  $^{77}$ Se and  $^{82}$ Se as the other isotopes for the inversion. This approach has the disadvantage that it has a significant Ge interference on <sup>74</sup>Se, which can be up to 70% of the signal. We also used a  $^{78}$ Se- $^{82}$ Se spike (calibrated ratios:  ${}^{78}$ Se/ ${}^{82}$ Se =0.185,  ${}^{78}$ Se/ ${}^{76}$ Se =110.847,  ${}^{82}$ Se/ ${}^{76}$ Se =598.23), with an inversion scheme that additionally uses <sup>76</sup>Se and <sup>77</sup>Se measurements. The latter procedure allows analyses at lower Se intensities, because the interfered  $^{74}$ Se is not required. However, it does use <sup>76</sup>Se, which has three significant interferences: <sup>40</sup>Ar<sup>36</sup>Ar, <sup>76</sup>Ge and <sup>75</sup>AsH, albeit the sum of these contributions is typically smaller ( $\sim$ 3–10%) relative to the mass 76 intensity than <sup>74</sup>Ge to mass 74. The <sup>78</sup>Se-<sup>82</sup>Se spike was used most extensively, although results from both spikes are presented below to provide multiple constraints on the compositions of the shale standards.

Journal of Analytical Atomic Spectrometry Accepted Manuscript

All isotope ratios are presented as  $\delta^{82/76}$ Se, which is the ‰ deviation of the <sup>82</sup>Se/<sup>76</sup>Se ratio of the sample from that of NIST SRM-3149 (Lot No. 992106). This reference material is a NIST 10.042 ± 0.051 mg/g single element standard solution, which is not isotopically certified, although its absolute composition has been

reported<sup>27</sup>, and has become the isotopic reference standard of choice for Se<sup>1, 5, 7, 17, 19</sup>, <sup>28</sup>. Our standard is from the same lot as for some other studies<sup>5, 17, 28</sup>. The <sup>74</sup>Se-<sup>78</sup>Se double spike inversion does not use <sup>76</sup>Se, so these ratios are converted to  $\delta^{82/76}$ Se notation from  $\delta^{82/77}$ Se using the exponential mass fractionation law.

130 Concentrations of samples were determined by isotope dilution as part of the 131 double spike inversion, with an inaccuracy of  $\pm 4\%$  (likely dominated by weighing 132 inaccuracy), as determined by repeat analyses of the USGS shales described below.

#### **3.0 Experimental**

## *3.1 Hydride-generation MDF chemistry*

The chemical purification of Se for both isotope ratio analysis and isotope dilution has been well described in a number of studies. In general we follow the methods of Rouxel et al. (2002) and Elwaer and Hintelmann (2008). Briefly, this involves dissolution of the powdered shale samples in repeated steps of HF-HNO<sub>3</sub> and HNO<sub>3</sub>-H<sub>2</sub>O<sub>2</sub>, without exceeding 80°C in order to avoid loss of Se by volatilisation. The double spike was normally introduced to the sample at the end of dissolution (during the final HNO<sub>3</sub> step to allow spike-sample equilibration), but before chemistry, so as to be able to respike the same dissolved material if initial sample/spike mixtures were sub-optimal. Our experiments show that identical isotope ratios and concentrations result from adding the spike at various stages during dissolution. Because shales can contain a significant amount of organic matter, we also experimented with dissolution of rock powders at high temperature and pressure, using an Anton-Paar HPA-S ashing system (2 hours at 220°C and 100 bar, in concentrated HF and HNO<sub>3</sub>).

150	Following dissolution, the samples were reduced in 4 M HCl for 1 hour at
151	80°C, and then passed three times through 0.2 g of Thioglycollic Cotton Fibre
152	$(TCF)^{13}$ . Se was then desorbed from the TCF by placing the TCF in a centrifuge tube
153	with 1 ml concentrated $HNO_3$ and then heating the tube in a boiling water bath for 20
154	minutes. The oxidised TCF was then diluted with 3 ml MQ $H_2O$ and centrifuged.
155	Multiple extractions ensure higher yields, but also dilute the sample more, because
156	completely drying the sample down can cause the formation of insoluble thiol salts. It
157	is, however, possible to extract Se several times (typically three times), and dry the
158	resulting solution down to a pre-determined volume, in order to increase the Se
159	concentration in the solution (by decreasing the solution volume). Following this,
160	concentrated HCl was added to make the sample up to 1.6M HCl, and heated at 80°C
161	for one hour, to reduce the Se(VI) to Se(IV), ready for hydride generation <sup>14, 16, 17</sup> .
162	Yields of 100% can be obtained for experiments with pure Se, but comparisons of Se
163	beam intensity on the mass spectrometer with concentrations derived from the double
164	spike inversion clearly show that additional matrix can lead to lower yields. The
165	yields from purifying shale samples are $\sim 80-100\%^1$ , and can vary randomly with
166	individual samples. This would likely not be observable in this, or indeed in other,
167	studies unless some form of isotopic spiking were involved. Ashing samples leads to
168	lower yields (50-60%), possibly because the high temperatures cause Se loss, as
169	described elsewhere <sup>13, 29</sup> . However, this effect does not cause resolvable differences in
170	the isotope ratio, due to addition of double spike prior to ashing. The total procedural
171	blank from this method is $\sim$ 0.4 pg Se, which is too low to determine a typical blank
172	isotope ratio.
173	The TCE chemistry efficiently removes major elements (e.g. nost-chemistry

Journal of Analytical Atomic Spectrometry Accepted Manuscript

173The TCF chemistry efficiently removes major elements (e.g. post-chemistry174Al/Se of SCo-1  $\sim 10^{-5}$ , determined on a Thermo Finnigan Element 2 sector-field ICP-

#### Journal of Analytical Atomic Spectrometry

Journal of Analytical Atomic Spectrometry Accepted Manuscr

MS, by calibrating against a set of synthetic multi-element standards), as well as
matrix elements that could compete for hydride generation (S/Se of SCo-1 is reduced
from $\sim$ 700 to 0.002 by chemical purification, a similar final ratio to NIST 3149).
Complete elemental scans of purified SCo-1 have showed that Sb has relatively high
post-chemistry abundance (Sb/Se reduced from 2.8 to 0.01). Such matrix effects
should not affect double spiked measurements, but may influence sample-standard
bracketing (SSB), because Sb forms hydrides relatively easily <sup>13</sup> , and therefore its
possible influence should be monitored in SSB studies. Abundances of elements with
nuclides that isobarically interfere with the Se isotope spectrum are also reduced by
the chemistry: As/Se in SCo-1 is reduced from 13.5 to 0.01, and Ge/Se from 1.7 to 10 <sup>-</sup>
<sup>4</sup> . Nevertheless these abundances are sufficient to cause significant perturbation of the
Se isotope ratios if not corrected and this effect can be greater in other shale samples
as both As/Se and Ge/Se can be several orders of magnitude higher than in SCo-1 and
SGR-1.
3.2 Analyses
All analyses were performed on a Thermo Finnigan multicollector inductively
coupled plasma mass spectrometer, with a hydride-generation introduction system
(HG-MC-ICP-MS) in the laboratories of the Bristol Isotope Group, University of
Bristol. In order to reduce levels of oxide and hydride interferences, standard "H"
skimmer cones were used, rather than the more sensitive "X" cones. Since various
NiO isotopologues produced from commonly used Ni cones can interfere on Se
isotopes (e.g. ${}^{58}\text{Ni}{}^{16}\text{O} \sim 0.03$ pA, typically $\sim 0.2-1\%$ of the ${}^{74}\text{Se}$ beam), all cones used
were aluminium, which reduced the intensity of $^{58}\mathrm{Ni}^{16}\mathrm{O}$ by an order of magnitude.

199 This also has the advantage of a  $\sim$ 15% sensitivity increase, likely because of the 200 higher working temperatures of Al cones.

Samples were introduced using a Cetac HGX 200 hydride generator. This system features a frosted glass rod that allows enhanced mixing between the sample and the hydride forming reagent (NaBH<sub>4</sub>) and therefore a high efficiency of Se-hydride generation. However, this system also has the disadvantage that the frosted glass rod causes long wash-out times - on the order of tens of minutes to obtain natural ratios after spiked samples. In order to avoid using large volumes of rinse acid,  $\sim$ 75% of this rinse was performed while aspirating air, leaving the continuously pumped NaBH<sub>4</sub> to clean the glass rod.

At the beginning of an analytical session, the samples were taken up into 1.6 M HCl, ready for introduction into the hydride generator. For analysis, samples were mixed with 1% NaBH<sub>4</sub> dissolved in 0.014 M NaOH in the hydride generator. All solutions were pumped through the hydride generator, using a multi-channel peristaltic pump. Using an uptake rate of  $\sim$ 300 µl/min, a 200 ng/ml Se solution gave an intensity of  $\sim 30$  pA on <sup>76</sup>Se. Including uptake time, the analyses requires 2 ml of sample and ideally a total of 400 ng of Se. Up to 3–4 times less Se can be analysed, although uncertainty begins to increase at the lower end of this spectrum, as the role of interferences starts to become critical (see Section 4.1). The amount of Se required is generally similar to, or slightly higher, than reported by other recent studies<sup>7, 17, 19</sup>, and considerably more than reported for a collision-cell instrument<sup>13</sup>. Interestingly, this is the first Se isotope study to employ a Neptune, rather than Nu or Isoprobe MC-ICP-MS, potentially hinting that the relative sensitivity for interferences (such as argides, as detailed in Section 4.1) is somewhat higher on a Neptune. All measurements were bracketed with on-peak background measurements, and doublespiked NIST-3149 was analysed every 3–4 measurements. Each individual sample
and blank measurement consisted of 50 ratios (210 s integration time).

 **4.0 Interference corrections** 

A series of interference corrections were performed on measured samples. The justification and evaluation of these corrections are detailed below but, briefly, analyses were background corrected, to remove contributions from the plasma gas and carrier matrix, followed by correction for Ge, SeH and AsH interferences. Most of these corrections are standard for Se isotope analysis<sup>6, 7, 13, 14, 17, 18, 30</sup>, although we explore potential pitfalls in these corrections to a greater degree than previous studies. The formulae used for the novel approach of simultaneous correction of these interferences, together with the double spike inversion, are given in the appendix.

All nine faraday cups of the Neptune were used for the measurements, which allowed analysis of all Se isotopes and requisite Se-free isobaric interferences without peak jumping. The dominant beams collected in each cup, including the main interferences, are detailed in Table 1.

## *4.1 ArAr corrections (background correction)*

Argon dimers from the plasma gas interfere on several Se isotopes. Most notably, typical  ${}^{40}$ Ar ${}^{40}$ Ar intensities of ~40 pA mean that its influence on  ${}^{80}$ Se cannot be accurately corrected, but this is not the case for the minor Ar dimers. The most straightforward method for correcting for Ar dimer interference is by subtracting the background measured in clean acid ("on-peak background"). It is alternatively possible to correct for ArAr by measuring the  ${}^{80}$ Se-corrected  ${}^{40}$ Ar ${}^{40}$ Ar beam, and using this to calculate the other ArAr beams<sup>5, 17</sup>. However, this method would require

#### Journal of Analytical Atomic Spectrometry

an appropriate functional relationship between  ${}^{40}\text{Ar}{}^{40}\text{Ar}$  and the other dimers; an assumption of an exponential mass dependent relationship as used for elemental species might be the most obvious. Although the exponential form is widely used for inter-relating atomic species its effectiveness for dimers is not well documented and need not necessarily be similarly appropriate. Further complications arise from the need to correct for  ${}^{80}\text{Kr}$  or  ${}^{79}\text{BrH}$  interferences on mass 80.

In contrast, an 'on-peak' background correction has the advantage that all machine-generated backgrounds are corrected (i.e. also those from Kr and ArArH), but this approach requires long wash-out times (up to 45-60 min) in order to return to a normal background after analysis of a spiked solution. An assumption also needs to be made that the ArAr beams remain identical when aspirating Se-bearing and Se-free acidic solutions, although this is a common approach in MC-ICP-MS. However, we have tested this assumption, by connecting the hydride generator to an Element 2 ICP-MS run at high resolution (M/ $\Delta$ M ~ 10,000, 5-95% peak edge width), with which it is possible to resolve Se and ArAr peaks (albeit with a loss of almost an order of magnitude of sensitivity). As shown in Figure 1, the ArAr peaks remain at the same intensity whether there is a Se peak present or not, justifying the use of a background correction. The <sup>40</sup>Ar<sup>37</sup>Cl, which is entirely dominated by the HCl matrix, interference on mass 77 is similarly corrected. The potential effect of uncorrected ArAr interferences is further explored in Section 4.4.

Journal of Analytical Atomic Spectrometry Accepted Manuscript

Nevertheless, a minimum Se intensity threshold is necessary, below which the background intensities of even the minor dimers, particularly  ${}^{36}$ Ar ${}^{40}$ Ar (typically ~0.3 pA), cannot be corrected for with sufficient accuracy. A 10% background contribution to the measured mass 76 intensity, using our inversion including  ${}^{76}$ Se, causes a ~2.5‰  $\delta^{82/76}$ Se inaccuracy in measurements on NIST SRM-3149 (Fig. 2), whereas for a

background contribution of ~5% the measurements are within error of the reference value. In contrast, the  ${}^{38}$ Ar ${}^{40}$ Ar background does not exceed ~0.1% of the 78 beam in either double spike system and therefore does not compromise the analyses.

*4.2 Ge corrections* 

Germanium-73 was monitored to correct for Ge interferences on <sup>74</sup>Se and <sup>76</sup>Se. Since fractionation of Se and Ge are likely subtly different, this correction will likely only work well at low Ge/Se relative abundances. It is therefore critical that tests were undertaken using Se standards doped with variable amounts of Ge to establish the limit of Ge/Se at which the correction is robust. The results of these tests, shown in Figure 3, demonstrate that a Se mass bias can be used to correct Ge interferences accurately up to a Ge/Se intensity ratio of  $\sim 0.1$  in the purified samples. The analysed Se fractions of the shale standards reported here have Ge/Se < 0.05 (mostly <0.03). Such correction problems are likely duplicated in any method that uses the Se mass bias to correct Ge interferences, and studies analysing samples with higher post-chemistry Ge/Se should monitor two Ge isotopes for accurate mass bias generation.

## *4.3 Hydride corrections (AsH, SeH, GeH, ArArH)*

The use of a hydride generator results in large amounts of free hydrogen gas entering the plasma, which may subsequently bind to other elements introducing hydride molecular ions into the mass spectrometer. The method we employ to correct for hydride ions is to determine the degree of SeH formation by monitoring the background-corrected <sup>82</sup>Se and mass 83, assumed to consist of only <sup>82</sup>SeH. The background correction ("on-peak background") on mass 82 corrects for <sup>82</sup>Kr and <sup>81</sup>BrH, assuming Br dominantly stems from the carrier matrix (HCl), based on scans

#### Journal of Analytical Atomic Spectrometry

via an Element 2 ICP-MS, and on mass 83 corrects for  ${}^{83}$ Kr interference. Depending on the tuning of the hydride generator (e.g. concentrations of NaBH<sub>4</sub>) and of the mass spectrometer, the resulting SeH/Se is ~ 10<sup>-3</sup> to 10<sup>-4</sup>. Typically, tuning was optimised to minimise hydride formation. The correction of  ${}^{76}$ SeH on  ${}^{77}$ Se, and of  ${}^{77}$ SeH on  ${}^{78}$ Se is straightforward (see Appendix). The hydride correction is mass bias corrected for the difference in mass between  ${}^{82}$ Se/ ${}^{82}$ SeH and the other hydride-forming isotopes.

Arsenic hydride is a significant interference on  $^{76}$ Se. For example, an uncorrected AsH interference at an As/Se = 0.2 will decrease the final  $\delta^{8276}$ Se by 1 to 5‰, depending on the sample introduction and mass spectrometer tuning. Since As is monoisotopic we assume, out of necessity, that the As/AsH ratio is identical to <sup>82</sup>Se/<sup>82</sup>SeH. This relationship has been assessed using tests with the NIST-3149 standard doped with As, with ratios up to As/Se  $\sim 1.5$  (As/<sup>76</sup>Se  $\sim 16$ ; Fig. 4). Post chemistry, samples rarely exceed As/Se  $\sim 0.3$  but occasionally shales with very high original As/Se (>100) may exceed this. Such samples were discarded, as the influence of the interference is uncalibrated at such ratios, although it is possible that the correction will still be accurate.

Journal of Analytical Atomic Spectrometry Accepted Manuscript

Doping experiments have also shown that the GeH/Ge ratio is identical within uncertainty to that of SeH/Se (separate analyses of GeH formation ( $^{74}$ Ge/ $^{74}$ GeH) from a pure Ge solution and SeH formation from a pure Se solution performed within a few minutes of each other, with the same tuning parameters). However, given the low abundance of Ge in purified samples (typically ~0.1 pA on mass 73), no GeH correction was performed.

A further interfering hydride on the Se isotopic spectrum is ArArH. The degree of argide dimer hydride formation is different from that of SeH formation. We determined this from measurements of <sup>40</sup>Ar<sup>40</sup>ArH/<sup>40</sup>Ar<sup>40</sup>Ar at 'medium resolution'

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 $(M/\Delta M \sim 6000, 5-95\%$  peak edge width), which resolves the hydride from <sup>81</sup>Br. We 324 325 found ArArH/ArAr  $\sim 0.01$ , 1–2 orders of magnitude greater than SeH/Se, which 326 makes direct correction difficult because, given the number of faraday cups on a 327 Neptune, a peak jump would have to be performed to measure the degree of argide 328 dimer hydride formation at mass 81. Our observations on ArArH formation, however, 329 have a bearing on studies that perform argide corrections based on the beam at mass 80 (e.g.<sup>17, 21</sup>), rather than a background correction, unless, for example, this is 330 experimentally calibrated<sup>17</sup>. Fortunately, as demonstrated above, the levels of argide 331 332 dimer, and therefore also the levels of ArArH, are identical within measurement 333 uncertainty in blank acid and sample and, therefore, ArArH is also corrected for in the 334 background subtraction. The formulae for the simultaneous correction of all the above 335 interferences as part of the double spike inversion are given in the appendix.

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## 337 *4.4 Double spike interference models*

338 Although individual interferences can be corrected in the tests described 339 above, it is nevertheless necessary to determine whether natural samples have had all 340 interferences corrected. Generally this is hard to accomplish, especially with an 341 isotopic system with few published studies or standards. However, the effectiveness 342 of our correction procedure can be checked by spiking a natural sample with variable 343 amounts of double spike, to generate different spike/sample ratios (fspike - defined as 344 the mixture M in Rudge et al., 2009). Clearly, consistent results should be obtained regardless of  $f_{\text{spike}}$ , if the interferences are correctly accounted for. These results can 345 346 also be compared with a similar test conducted on a variably spiked pure Se standard. 347 Since double spike inversions respond non-linearly to interferences, and because 348 changing the sample/spike ratio effectively varies the ratio of potential interferences

#### Journal of Analytical Atomic Spectrometry

to Se (at very high  $f_{spike}$  only two Se isotopes are present, whereas at very low  $f_{spike}$  Se is present in virtually natural abundances), a large  $f_{spike}$  range provides the opportunity to test the accuracy of the isotopic data.

The results of variably spiking NIST-3149 and the USGS shales SGR-1 and SCo-1 are shown in Figure 5. This exercise not only shows the optimal sample/spike ratio ("cocktail"<sup>26</sup>), but also shows that isotope ratios resulting from variable spiking of the NIST standard and variable spiking of shales coincide. This demonstrates that all interferences, both instrumental- and sample-generated, are fully corrected for. In order to demonstrate how sensitive spiked samples are to uncorrected interferences. we also show forward models for the <sup>78</sup>Se-<sup>82</sup>Se double spike inversion with uncorrected interferences on various isotopes. Clearly the isotope with the greatest potential for uncorrected interferences is <sup>76</sup>Se (in an inversion using <sup>76</sup>Se, <sup>77</sup>Se, <sup>78</sup>Se and <sup>82</sup>Se), where explicit <sup>76</sup>Ge and <sup>75</sup>AsH corrections are required on top of a significant back-ground correction (for  ${}^{40}Ar^{36}Ar$ ). The variability of  $\delta^{82/76}$ Se against f<sub>spike</sub> (Fig. 6) in models with incompletely corrected interference is strikingly different from the invariant results of rock standards run at different f<sub>spike</sub>, demonstrating that these interferences are correctly accounted for in the samples. This model indicates that even an uncorrected 1‰ (0.1%) interference on <sup>76</sup>Se can push the final  $\delta^{82/76}$ Se by several permil from the true ratio, which demonstrates the necessity for careful interference correction. Modelling shows that the effect of the more minor argon dimer isotopologues is insignificant compared to that of <sup>40</sup>Ar<sup>36</sup>Ar. 

Journal of Analytical Atomic Spectrometry Accepted Manuscript

Figure 6 also shows the model results for an uncorrected interference at mass which affects the  $\delta^{82/76}$ Se to a lesser extent, but still demonstrates that any existing interferences are fully corrected in our data reduction.

#### **5.0** Analytical Results and Discussion

# 375 5.1 Shale $\delta^{82/76}$ Se and analytical uncertainty

The various tests described above give confidence in the accuracy of our analyses. Two USGS shale standards (SCo-1 and SGR-1b) were repeatedly analysed, including multiple dissolutions and chemical purification. Importantly, both were analysed using both double spikes described above, as a further test on sensitivity to inaccuracy of different aspects of our data reduction. The results are shown in Figure 7.

The internal reproducibility of an analysis is strongly dependent on the sample/background ratio (Fig. 2). Since the beam from a hydride generator is inherently unstable, due to the instability in reagent flow generated by the peristaltic pump, a low-intensity beam will be more strongly affected by background noise. This instability results in periodic beam fluctuations that will cause a degradation in the internal reproducibility which can be larger than the external reproducibility. Thus, for a <5% background contribution on <sup>76</sup>Se, the internal reproducibility is  $\pm (0.03 -$ 0.1)% on  $\delta^{82/76}$ Se, whereas for a 10% background contribution the reproducibility is  $\pm 1\%$  (Fig. 2). Samples were not run with a background contribution >5%. External uncertainty, based on repeated analyses (including dissolution and chemistry) of both USGS shales is  $\pm 0.17\%$  (2sd) on  $\delta^{82/76}$ Se.

The  $\delta^{82/76}$ Se values of both SCo-1 ( $\delta^{82/76}$ Se = -0.22 ± 0.15, 2sd, n=18) and SGR-1 (0.25 ± 0.17, 2sd, n=16) from this study generally compare well to those published by Rouxel et al. (2002), Schilling et al. (2011) and Layton-Matthews et al. (2006, 2013). The results from SGR-1 from this study and those of Stücken et al. (2013) are within analytical uncertainty of each other, but the population means of this study and that of Mitchell et al. (2012) are distinct from each other at the 95%

confidence level (student t-test, using data from Mitchell 2013, pers. comm.). These studies present a mix of double spike<sup>7, 19</sup> and sample-standard bracketing analyses<sup>13, 17</sup>, <sup>24, 25</sup>. We note that the study of Rouxel et al. (2002) used a collision cell instrument that is no longer manufactured, which likely had a very different interference spectrum to correct. In addition, the different dissolution methods trialled in this study ("standard" HF-HNO<sub>3</sub> dissolution vs. high temperature ashing) resulted in identical isotope ratios within analytical uncertainty. The values for both shales using these different methods are reported in Table 2. Given that we have determined these values with two separate double spike inversion systems, we recommend these USGS shales as useful and easily accessible rock standards for assessing accuracy, or at least inter-study reproducibility, for future Se isotope analyses. We can, as yet, offer no explanation for the differences in published values of SGR-1, especially compared to those reported by Mitchell et al., 2012. Further work is needed to determine whether SGR-1 is isotopically heterogeneous, or whether some studies have unresolved analytical issues.

Journal of Analytical Atomic Spectrometry Accepted Manuscript

- 415 5.2 Desolvation analysis of Se isotopes

Generally the sensitivity of an MC-ICP-MS for Se is ~50 times better when using a hydride generator than a desolvation nebuliser such as a Cetac Aridus. However, we have discovered that doping a Se (or sulphur) solution with magnesium (in this case, the Mg isotope standard DSM-3) enhances the sensitivity when using "wet" (Elemental Scientific Inc. Apex) or "dry" (Cetac Aridus) plasma by up to a factor of 100-200 (Fig. 8a). The greatest Se sensitivity occurs at Mg/Se mass ratios of 10–20. When normalised to uptake rate ( $\sim$ 50 µl/min), the sensitivity of Mg-doped Se using an Aridus is 1.6-1.9 times greater than that of a hydride generator. Here, Se was

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424	introduced as selenate, in 2% HNO <sub>3</sub> . Yield enhancement of Se when using ICP-MS
425	has been reported when doping the solution with $arbon^{31}$ or in H <sub>2</sub> SO <sub>4</sub> - or HCl-based
426	matrices <sup>32</sup> . However, these enhancements are only a factor of 2-3 (two orders of
427	magnitude less than this study), and is thought to be due to the transfer of electrons to
428	the C ions, because Se has a lower ionisation potential than $C^{31}$ . Clearly this effect
429	cannot be causing the sensitivity enhancement by Mg, because Mg has a lower
430	ionisation potential than Se. Recently, sulphate sensitivity has been enhanced by a
431	factor of ~20, by doping with Na <sup>33</sup> . That study was also not able to explain these
432	effects, and proposed that it might be related to the desolvating membrane. Our
433	experiments with a membrane-less Apex, as well as a desolvating Aridus, now
434	demonstrate this not to be the case. To our knowledge, sensitivity enhancement by
435	Mg doping has not been reported to date. The enhancement effect also means that Se
436	concentration analyses in any matrix containing Mg may be highly inaccurate unless
437	either isotope dilution is used, or concentration analyses are performed by hydride
438	generation ICP-MS.

The sensitivity enhancement effect of Mg could however be used to analyse Se isotope ratios in low concentration samples. Advantages of using desolvation introduction systems over hydride generation include faster washout times, more stable beams, and an introduction system that is more widely used. An obvious drawback is that in dry plasma mode, argide dimer production is higher. However, this can be overcome by tuning gas flows, torch positions and RF generator power until the background on <sup>76</sup>Se is sufficiently low (<5%) for corrections to be accurate.

Double-spike was used to analyse the Mg-doped Se solutions. Differences in mass bias style (i.e. deviations from exponential mass bias) due to Mg addition to pure Se solutions are not resolvable in double spike inversions. As shown in Figure

#### Journal of Analytical Atomic Spectrometry

8b, mixtures of the NIST standard with amounts of spike yield consistent results
whether run by hydride generator or with Mg-addition and desolvation, which implies
that, whilst the relative importance of different interferences using the two techniques
might change, our procedure corrects for these accurately.

The intensity of the main interferences (Ge and As) are little affected by Mg doping, and desolvation nebulisation is not an efficient means of generating Ge beams. Thus a solution with Ge/Se = 0.2, when doped with Mg, has a measured Ge/Se= 0.05. In contrast, the sensitivity of As when using desolvation nebulisation is not as poor as for Ge, and the yield of Mg-doped SeH is different from that of AsH in the same solution, so that the SeH yield cannot be used to correct for AsH. This means that a double spike scheme that avoids <sup>76</sup>Se is advisable, for example the alternative inversion scheme detailed in this study, or in other studies<sup>19</sup>. Nevertheless, this method provides the potential for Se isotope analysis by laboratories with no hydride generation facilities.

Journal of Analytical Atomic Spectrometry Accepted Manuscript

## 464 5.3 Mass Independent Fractionation of Selenium

Sulphur mass independent fractionation (MIF) has been reported in samples from the Archean, and is generally thought to reflect photolysis of S-bearing tropospheric gases by ultraviolet radiation and the production of elemental sulphur particles in the atmosphere<sup>34-36</sup>. Thus, the widespread occurrence of significant non-zero  $\Delta^{33}S (\Delta^{33}S = ({}^{33}S/{}^{32}S)_{sample}/({}^{33}S/{}^{32}S)_{V-CDT}-[({}^{34}S/{}^{32}S)_{sample}/({}^{34}S/{}^{32}S)_{V-CDT}]^{0.515})$  in rocks deposited prior to  $\sim 2.4 \text{ Ga}^{34}$  suggests that atmospheric O<sub>2</sub> was  $< 10^{-5}$  of present atmospheric levels<sup>37, 38</sup>, and that reducing gases such as methane or hydrogen were considerably more abundant than today<sup>39</sup>. Given the chemical similarities between S and Se, partly through the volcanic cycle<sup>6</sup>, and the ability of both to form compounds

that are gases, it is possible that Se isotopes could demonstrate MIF in a mannersimilar to S isotopes.

To test the hypothesis that Se MIF may be preserved in Archean shales, we have performed Se isotope analyses on five Neoarchean shales from a well-preserved drill core (GKF01) from the ~2.65-2.5 Ga Ghaap Group, South Africa<sup>40</sup>, all of which record large  $\Delta^{33}$ S fractionations (-1.7 to +6.9‰; Table 3<sup>41</sup>). Our analyses were identical to those described above for hydride generation mass dependent fractionation, except that samples were not double-spiked, and new plasma interface cones were employed in the mass spectrometer that had not been used for any double spiked analyses. While no memory effect from spike tracer has been observed on the cones in this study, changing cones countered any such possibility. In addition, the hydride generator was thoroughly cleaned prior to analyses, and all sample introduction tubing was replaced. Finally, sample analyses were bracketed by analyses of NIST-3149, in order to be able to correct for any non-exponential law instrumental mass fractionation. Due to the potential mass-dependent fractionation effects of the chemistry, MDF results are not reported from these samples.

There are no MIF Se data reported in any previous studies. Those studies that have used SSB to determine mass-dependent fractionation (MDF)<sup>4, 13, 17, 24, 25</sup> neither report sufficient statistical detail nor, with one exception<sup>13</sup>, the results of repeat analyses, making it impossible to retrospectively calculate the uncertainty of the MIF. In principal, SSB-determined MDF ratios may be used for reporting MIF, but care needs to be taken to propagate the uncertainties, and their correlations, to correctly determine the uncertainty in the MIF. In general, the correlations between the isotope ratios are such that explicit MIF measurements using internal normalisation are significantly more precise than MDF measurements. The study by Rouxel et al.

#### Journal of Analytical Atomic Spectrometry

499 (2002), wherein isotope measurements on repeat analyses of a MERCK Se standard 500 are reported, allows us one point of comparison of the uncertainty to our own MIF 501 data. We have calculated the MIF on each analysis of the MERCK standard 502 (excluding two that did not measure <sup>77</sup>Se, see table 4 in <sup>13</sup>) using the exponential 503 fractionation law to normalise to <sup>82</sup>Se/<sup>78</sup>Se yielding a reproducibility of ±0.20‰ (2 $\sigma$ , 504 n=13) for MIF  $\delta^{82/77}$ Se compared to our MIF reproducibility of ±0.03–0.04‰ as 505 detailed below.

Given the low relative intensities of interferences on <sup>77</sup>Se, <sup>78</sup>Se and <sup>82</sup>Se, the <sup>82</sup>Se/<sup>78</sup>Se ratio was used to internally normalise the reported <sup>82</sup>Se/<sup>77</sup>Se ratio, using the exponential mass bias law<sup>42</sup>. We have also examined the <sup>82</sup>Se/<sup>78</sup>Se normalised <sup>76</sup>Se/<sup>77</sup>Se ratio. We present our internally normalised results in epsilon form (deviation in parts per 10,000, to bring units in-line with other isotopic MIF results (e.g., <sup>43</sup>):

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$$\varepsilon^{x/77} Se_{\frac{82}{78}} = \left( \begin{pmatrix} xSe / {}^{77}Se_{\frac{82}{78}} \end{pmatrix}_{spl} - 1 \\ \hline \begin{pmatrix} xSe / {}^{77}Se_{\frac{82}{78}} \end{pmatrix}_{std} - 1 \end{pmatrix} \times 10^4$$
 (1)

513 where the normalising ratio is included as a subscript (i.e.,  ${}^{82}\text{Se}/{}^{78}\text{Se})^{43}$ , and x refers to 514  ${}^{76}\text{Se}$  or  ${}^{82}\text{Se}$ , *spl* is the sample, and *std* the standard. The interferences on  ${}^{76}\text{Se}$  (from 515 Ge, As and ArAr) result in variable  $\varepsilon^{76/77}Se_{\frac{82}{78}}$  Thus for 500 ng/g NIST-3149,

516 
$$\varepsilon^{76/77} Se_{\frac{82}{78}}$$
 varies by ±1.2 , and for SGR-1b (n=2, chemistry =2) it varies by ±1.7

517 This suggests that these interferences are also a large factor in the uncertainty on the 518 double spike inversions described above. However the  $\varepsilon^{82/77}Se_{\frac{82}{78}}$  ratio reproduces 519 much better; the analytical uncertainty for NIST-3149 is ±0.31 , and for SGR-1b is ±

520 0.40 (Table 3). This uncertainty is therefore likely to be the external reproducibility 521 of our analyses, assuming that the Eocene SGR-1b does not exhibit any Se MIF.

522 The Archean shales exhibit a mean  $\varepsilon^{82/77} Se_{\frac{82}{78}}$  of -0.39 ± 0.27 (n=5). All

523 shales (including SGR-1b) exhibit slightly negative  $\varepsilon^{82/77} Se_{\frac{82}{78}}$  values relative to NIST-

3149, which may suggest that either NIST-3149 itself is not typical of common terrestrial Se. This has been observed for several other highly purified standards analysed to high precision for  $MIF^{43-45}$ . This effect is potentially as a function of the purification by non-exponential processes, and/or that during purification, a species other than the element is being fractionated (e.g. in this case selenate- or selenite-oxide rather than pure Se). In any case, the critical observation is that the Archean shales are within the range of SGR-1b, and therefore do not exhibit any resolvable mass independent fractionation, or correlation between  $\varepsilon^{82/77}Se_{\frac{82}{70}}$  and  $\Delta^{33}S$  (Fig. 9). 

This may be because Se gas-phase reactions are not photocatalysed in a mass independent manner under anoxic atmospheric conditions, but could also be due to homogenisation of an atmospheric Se MIF signal prior to preservation in the sediment<sup>46</sup>. In addition, oxidised Se gases tend to be less volatile than similar sulphurbased compounds.

#### **6.0 Conclusions**

539 This study presents data using two different double spikes and variably-spiked 540 samples to accurately constrain selenium isotope ratios in two different USGS shale 541 standards. Our measurements suggest that the  $\delta^{82/76}$ Se ( ${}^{82}$ Se/ ${}^{76}$ Se deviation from 542 NIST-3149) values of SCo-1 = -0.22 ± 0.15‰ (n=18), and of SGR-1b = +0.25 ±

#### Journal of Analytical Atomic Spectrometry

543	0.17‰ (n=16). Measurements obtained by other studies using a mix of double spike
544	or sample-standard bracketing, generally agree well with our robustly defined values.
545	We also present a new method for measuring Se isotopes (for isotope dilution
546	or isotope ratios) using desolvation or "moist" plasma. Doping Se solutions with Mg
547	leads to an increase in Se sensitivity by a factor of 100-200, almost 2 times greater
548	than the uptake-rate normalised sensitivity of a hydride generator.
549	Finally, we have assessed Archean shales that demonstrate sulphur mass
550	independent isotope fractionation (MIF) for Se MIF. The <sup>82</sup> Se/ <sup>77</sup> Se ratios, normalised
551	to <sup>82</sup> Se/ <sup>78</sup> Se, show no significant deviations from analyses of the Eocene standard
552	SGR-1b (-0.39 $\pm$ 0.27 , parts per 10,000), and no correlation with existing $\Delta^{33}S$
553	measurements in Archean samples. Thus, Se does not show resolvable MIF
554	behaviour in samples that are likely candidates to record a strong variability in MIF
555	signature.
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557	Acknowledgments
558	The authors would like to thank Thomas Johnson and Kristin Mitchell for discussions
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561 562 563 564 565 566 567 568 569 569 570	<ol> <li>S. K. Clark and T. M. Johnson, <i>Journal of Environmental Quality</i>, 2010, <b>39</b>, 2200-2210. DOI: <u>10.2134/jeq2009.0380</u>.</li> <li>T. M. Johnson, <i>Chemical Geology</i>, 2004, <b>204</b>, 201-214. DOI: 10.1016/j.chemgeo.2003.11.015.</li> <li>T. M. Johnson, M. J. Herbel, T. D. Bullen and P. T. Zawislanski, <i>Geochimica et Cosmochimica Acta</i>, 1999, <b>63</b>, 2775-2783. DOI: 10.1016/s0016-7037(99)00279-3.</li> <li>O. Rouxel, Y. Fouquet and J. N. Ludden, <i>Geochimica et Cosmochimica Acta</i>, 2004, <b>68</b>, 2295-2311. DOI: 10.1016/j.gca.2003.11.029.</li> <li>H. J. Wen and J. Carignan, <i>Geochimica et Cosmochimica Acta</i>, 2011, <b>75</b>,</li> </ol>
571	1411-1427. DOI: <u>10.1016/j.gca.2010.12.021</u> .

Journal of Analytical Atomic Spectrometry Accepted Manuscri

- Isotopes, ed. C. M. Johnson, B. L. Beard and F. Albarede, Geochemical Society.
- 2004, pp. 289-317.

5	75 7. K. Mitchell, P. R. D. Mason, P. Van Cappellen, T. M. Johnson, B. C. Gill, J.
5	D. Owens, J. Diaz, E. D. Ingall, G. J. Reichart and T. W. Lyons, <i>Geochimica et</i>
3	Cosmochimica Acta, 2012, <b>89</b> , 302-317. DOI: 10.1016/j.gca.2012.03.038.
5	8. A. Fernandez-Martinez and L. Charlet, <i>Reviews in Environmental Science and</i>
5	79 Biotechnology, 2009, <b>8</b> , 81-110.
5	9. A. S. Ellis, T. M. Johnson, M. J. Herbel and T. D. Bullen, <i>Chemical Geology</i> ,
5	2003, <b>195</b> , 119-129. DOI: 10.1016/s0009-2541(02)00391-1.
5	182 10. M. J. Herbel, J. S. Blum, R. S. Oremland and S. E. Borglin, <i>Geomicrobiology</i>
3	<i>Journal</i> , 2003, <b>20</b> , 587-602. DOI: 10.1080//13851163.
5	11. M. J. Herbel, I. M. Jonnson, K. S. Oremland and I. D. Bullen, <i>Geochimica et Cosmochimica Acta</i> 2000 64 3701 3709 DOI: 10.1016/s0016.7037(00)00456.7
5	12 T M Johnson and T D Bullen <i>Geochimica et Cosmochimica Acta</i> 2003 <b>67</b>
5	<sup>1</sup> 87 413-419
5	188 13 O Rouxel I Ludden I Carignan L Marin and Y Fouquet Geochimica et
5	Cosmochimica Acta 2002 66 3191-3199
5	190 14 N Flwaer and H Hintelmann Analytical and Rioanalytical Chemistry 2007
5	<b>389</b> 1889-1899 DOI: 10 1007/s00216-007-1537-z
5	<sup>192</sup> 15 D Savard L P Bedard and S -J Barnes <i>Talanta</i> 2006 <b>70</b> 466-571
5	10. N Elwaer and H Hintelmann <i>Journal of Analytical Atomic Spectrometry</i>
5	2008 <b>23</b> 733-743 DOI: 10.1039/b801673a
5	<sup>195</sup> 17 E E Stücken J Foriel B K Nelson R Buick and D C Catling <i>Journal of</i>
5	Analytical Atomic Spectrometry 2013 <b>28</b> 1734–1749 DOI: 10.1039/c3ia50186h
5	197 18. N. Elwaer and H. Hintelmann. <i>Journal of Analytical Atomic Spectrometry</i> .
5	2008 <b>23</b> 1392-1396 DOI: 10.1039/b808645c
5	199 19. K. Schilling, T. M. Johnson and W. Wilcke. Soil Science Society of America
6	Journal, 2011, <b>75</b> , 1354-1364, DOI: 10.2136/sssaj2010.0377.
6	20. J. M. Zhu, T. M. Johnson, S. K. Clark and X. K. Zhu, Chinese Journal of
6	02 Analytical Chemistry, 2008, <b>36</b> , 1385-1390.
6	21. J. M. Zhu, T. M. Johnson, S. K. Clark, X. K. Zhu and X. L. Wang,
6	Geochimica et Cosmochimica Acta, 2014, <b>126</b> , 228–249.
6	105 22. H. Fan, H. Wen, R. Hu and H. Zhao, <i>Geochimica et Cosmochimica Acta</i> ,
6	2011, <b>75</b> , 7725-7740. DOI: 10.1016/j.gca.2011.09.027.
6	T. R. Kulp and L. M. Pratt, <i>Geochimica et Cosmochimica Acta</i> , 2004, <b>68</b> ,
6	3687-3701. DOI: 10.1016/j.gca.2004.03.008.
6	D. Layton-Matthews, M. I. Leybourne, J. M. Peter and S. D. Scott, <i>Journal of</i>
6	Analytical Atomic Spectrometry, 2006, <b>21</b> , 41-49. DOI: 10.1039/b501704a.
6	D. Layton-Matthews, M. I. Leybourne, J. M. Peter, S. D. Scott, B. Cousens
6	and B. M. Eglington, <i>Geochimica et Cosmochimica Acta</i> , 2013, <b>117</b> , 313–331.
6	13 26. J. F. Rudge, B. C. Reynolds and B. Bourdon, <i>Chemical Geology</i> , 2009, <b>265</b> ,
6	420-431. DOI: 10.1016/j.chemgeo.2009.05.010.
6	15 27. J. Wang, T. X. Ren, H. Lu, T. Zhou and M. T. Zhao, <i>International Journal of</i>
6	Mass Spectrometry, 2011, <b>308</b> , 65-70. DOI: 10.1016/j.ijms.2011.07.023.
6	28. J. Carignan and H. J. Wen, <i>Chemical Geology</i> , 2007, <b>242</b> , 347-350. DOI:
6	118 10.1016/j.chemgeo.2007.03.020.
6	29. Y. K. Chau and J. P. Riley, <i>Analytica Chimica Acta</i> , 1964, <b>33</b> , 36-49.
6	30. N. Elwaer and H. Hintelmann, <i>Talanta</i> , 2008, <b>75</b> , 205-214. DOI:
6	10.1016/j.talanta.2007.10.046.
6	E. H. Larsen and S. Sturup, <i>Journal of Analytical Atomic Spectrometry</i> , 1994,
6	<b>9</b> , 1099-1105. DOI: 10.1039/ja9940901099.

 $\begin{array}{c} 10 \\ 11 \\ 12 \\ 13 \\ 14 \\ 15 \\ 16 \\ 17 \\ 18 \\ 19 \\ 20 \\ 21 \\ 22 \\ 23 \\ 24 \\ 25 \end{array}$ 

1		
2		
3	624	32. KS. Park, ST. Kim, YM. Kim, Y. Kim and W. Lee, <i>Bulletin of the Korean</i>
4	625	<i>Chemical Society</i> , 2002, <b>23</b> , 1389-1393.
5	626	33. G. Paris, A. L. Sessions, A. V. Subhas and J. F. Adkins, <i>Chemical Geology</i> ,
6	627	2013, <b>45</b> , 50–61.
7	628	34. J. Farguhar, H. M. Bao and M. Thiemens, <i>Science</i> , 2000, <b>289</b> , 756-758, DOI:
8	629	10.1126/science.289.5480.756.
9	630	35 I Farguhar I Savarino S Airieau and M H Thiemens <i>Journal of</i>
10	631	Geophysical Research-Planets 2001 106 32829-32839 DOI:
11	632	10 1029/2000je001437
12	633	36 I Farguhar N.P. Wu, D.F. Canfield and H. Oduro, Economic Geology 2010
14	634	<b>105</b> 500 533
15	625	27 A A Dayloy and L E Kasting Astrobiology 2002 2 27 41 DOL
16	626	57. A. A. Faviov and J. F. Kasning, <i>Astrobiology</i> , 2002, 2, 27-41. DOI.
17	(27	10.1069/155110/02/55021521.
18	637	38. M. H. Themens, in Annual Review of Earth and Planetary Sciences. 2006,
19	638	vol. 34, pp. 217-262.
20	639	39. K. Zahnle, M. Claire and D. Catling, <i>Geobiology</i> , 2006, 4, 271-283. DOI:
21	640	10.1111/j.1472-4669.2006.00085.x.
22	641	40. S. Schroder, J. P. Lacassie and N. J. Beukes, <i>South African Journal of</i>
23	642	<i>Geology</i> , 2006, <b>109</b> , 23–54.
24	643	41. A. L. Zerkle, M. Claire, S. D. Domagal-Goldman, J. Farquhar and S. W.
25	644	Poulton, Nature Geoscience, 2012, 5, 359-363. DOI: 10.1038/ngeo1425.
26	645	42. W. Russell, D. A. Papanastassiou and T. Tombrello, <i>Geochimica et</i>
27	646	<i>Cosmochimica Acta</i> , 1978, <b>42</b> , 1075–1090.
28	647	43 R C J Steele T Elliott C D Coath and M Regelous <i>Geochimica et</i>
29	648	Cosmochimica Acta 2011 75 7906-7925 DOI: 10.1016/j.gca.2011.08.030
30	649	44 R W Carlson M Boyet and M Horan Science 2007 <b>316</b> 1175–1178
32	650	45 F Movnier A Agranier D C Hezel and A Bouvier Farth and Planetary
32	651	Science Letters 2010 300 359_366
34	652	A6 I Eargubar and D A Wing Earth and Planatamy Saionan Lattons 2002 212
35	652	40. J. Farquiar and D. A. wing, Earth and Finnetary Science Letters, 2005, 215, $1.12$ , DOI: 10.1016/s0012.821 $w(02)00206.6$
36	033	1-15. DOI: 10.1010/S0012-821X(03)00290-0.
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656 Figure 1

## Journal of Analytical Atomic Spectrometry



666 Figure 2

667	Se isotope ratios determined with decreasing Se intensity, and therefore increasing
668	background contribution of <sup>40</sup> Ar <sup>36</sup> Ar. This demonstrates the minimum signal
669	to noise ratio for accurate data when using <sup>76</sup> Se for the double spike inversion.
670	The error bars represent the 2 s.e. of the analyses.
671	

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#### 687 Figure 3

Ge doping experiments on NIST-3149. In the insert the solid line represents 0‰, and
the dashed lines the external uncertainty (2 s.d.).

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Arsenic doping experiments on NIST-3149. The data represent experiments at several
 different mass spectrometer tuning settings, and therefore different levels of hydride

formation. The error bars represent the 2 s.e. internal uncertainty of the analyses.



733 Figure 5

734	Results of variably spiking standards and samples. The approach is detailed in the
735	text. Both SGR-1 and SCo-1 data have been differenced to their "true" $\delta^{82/76}$ Se
736	value, as detailed in Table 2, in order to directly compare these data to NIST
737	and model data. This approach shows that both NIST and shale samples
738	behave identically to each other, demonstrating that all interferences are
739	corrected. The NIST data also indicate the ideal sample-spike mix, where the
740	uncertainty is smallest ( $f_{spike} \sim 0.55 - 0.75$ ).
741	







742 743 Figure 6

Modelled results for variable spiking (see Fig. 5) if interferences are left uncorrected.
For example, the black lines show results for a 2‰ uncorrected interference
on <sup>76</sup>Se (A) and <sup>82</sup>Se (B), respectively. The data from variably spiked NIST
(Fig. 5) are shown for comparison.



749 Figure 7

	$\mathcal{O}^{\circ}$
750	Results from this study's analyses of both USGS shales (closed symbols), compared
751	to published values (open symbols). (1): Schilling et al., 2011; (2): Stüeken et
752	al., 2013; (3): Mitchell et al., 2012; (4) Rouxel et al., 2002; (5) Layton-
753	Matthews et al., 2006, 2013. The error bars in all cases represent the 2 s.d.
754	analytical reproducibility. The small open diamonds represent this study's
755	individual analyses.
756	-



**"spike**758 Figure 8. a) Mg doping of Se solutions when using an Aridus introduction system.
759 Mg/Se is given as a mass ratio. b) Results of variably spiking Mg-doped
760 NIST-3149, also using an Aridus. The error bars represent the 2s.e. internal
761 reproducibility.
762



 $\begin{array}{rcl} & & & & \Delta^{33}S \\ \hline 764 & Figure 9. Mass independent Se isotopes (in parts per 10,000) plotted against mass \\ \hline 765 & & independent S isotopes. The error bars on the individual data points represent \\ \hline 766 & & 2 s.e internal reproducibility. The 2 s.d. external reproducibility can be \\ \hline 767 & & observed from the repeated measurements of SGR-1b. \end{array}$ 



 $\begin{array}{c} 10 \\ 11 \\ 12 \\ 13 \\ 14 \\ 15 \\ 16 \\ 17 \\ 18 \\ 19 \\ 20 \\ 21 \\ 22 \\ 23 \\ 24 \\ 25 \end{array}$ 

# Journal of Analytical Atomic Spectrometry

	<u> </u>	ups L4	L3	L2	L1	С	H1	H2	H3	H4
		<sup>73</sup> Ge	<sup>74</sup> Se	<sup>75</sup> As	<sup>76</sup> Se	<sup>77</sup> Se	<sup>78</sup> Se	<sup>80</sup> Se	<sup>82</sup> Se	<sup>82</sup> Se
	natı	ural	0.89%		9.36%	7.63%	23.78%	6 <u>49.61</u> %	8.73%	
D	ouble spil	ke 1	46%				54%		0.40/	
D	ouble spil	ke 2					16%		84%	
Fralic	interferen	ated	<sup>74</sup> Ge		<sup>76</sup> Ga					
Ехрис	illy correc	.1eu	Ue		<sup>75</sup> AsH					
					11011	<sup>76</sup> SeH	<sup>77</sup> SeH			
Backgro	und correc	cted	<sup>36</sup> Ar <sup>38</sup> A	r	<sup>40</sup> Ar <sup>36</sup> Ar		<sup>40</sup> Ar <sup>38</sup> A	r <sup>40</sup> Ar <sup>40</sup> Ar		
0					<sup>38</sup> Ar <sup>38</sup> Ar					
							<sup>78</sup> Kr	<sup>80</sup> Kr	<sup>82</sup> Kr	<sup>83</sup> K
						<sup>40</sup> Ar <sup>37</sup> Cl				
			~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~					<sup>79</sup> BrH	<sup>81</sup> BrH	
Table I.	. Details	of the cu	ip confi	guration	and dot	ible sp	ikes us	ed, as we	en as th	e ma
1	interfere	ences on e	each ma	SS.						
		dissolution	n n	$\delta^{82/76}$ Se	2 sd	[Se]	µg/g			
SCo-1		standard	14	-0.23	0.16					
74~ 79~		ashing	3	-0.20	0.13					
/4Se-/8Se	spike	standard	1	-0.09*						
		mean	18	-0.22	0.15	0.	79			
SGR-1b		standard	12	0.26	0.15					
		ashing	2	0.19	0.06					
74~ 70	amil	standard	2	0.21*	0.09					
′ <sup>4</sup> Se- <sup>/8</sup> Se	spike									
′ <sup>4</sup> Se- <sup>/8</sup> Se	e spike	mean	16	0.25	0.17	3.	10			
<sup>/*</sup> Se- <sup>/*</sup> Se Table 2.	. Results	mean s of the tw	16 vo USG	0.25 S shale s	0.17 standard	<u>3.</u> s analy	<u>10</u> /sed in	this stud	v. * ind	licate
<sup>(*</sup> Se- <sup>78</sup> Se	Results	mean s of the two	<u>16</u> νο USG ο δ <sup>82/76</sup> 9	0.25 S shale s	$\frac{0.17}{\text{standard}}$	3. s analy Unles	<u>10</u> /sed in	this stud	y. * ind ults we	licate
<sup>/*</sup> Se- <sup>/8</sup> Se Table 2.	Results	$\frac{\text{mean}}{\text{s of the two}}$	$\frac{16}{\text{vo USG}}$ o $\delta^{82/76}$	0.25 S shale s Se from	$\frac{0.17}{\text{standard}}$	<u>3.</u> s analy . Unles	<u>10</u> /sed in ss indic	this stud ated, res	y. * ind ults we	licate re
$\frac{7^{4}\text{Se}-7^{8}\text{Se}}{\text{Table 2.}}$	. Results ratios co obtained	mean s of the two ponverted t l using a <sup>7</sup>	$\frac{16}{\text{vo USG}} \\ \text{o } \delta^{82/76} \\ \text{Se}^{78} \\ \text{Se}^{-82} \\ \text{Se}^$	0.25 S shale s Se from se spike.	$\frac{0.17}{\text{standard}}$ $\delta^{82/77}$ Se	3. s analy . Unles	<u>10</u> /sed in ss indic	this stud ated, res	y. * ind ults we	licate re
<sup>7*</sup> Se- <sup>78</sup> Se Table 2.	. Results ratios co obtained	mean s of the two nverted to l using a <sup>7</sup>	$\frac{16}{\text{vo USG}}$ o $\delta^{82/76}$ Se- <sup>82</sup> S	0.25 S shale s Se from le spike.	0.17 standard $\delta^{82/77}$ Se	<u>3.</u> s analy . Unles	10 /sed in ss indic	this stud ated, res	y. * ind ults we	licate re
<sup>7*</sup> Se- <sup>78</sup> Se Table 2.	Results . Results ratios co obtained	mean s of the two onverted t l using a 7	$\frac{16}{\text{vo USG}} \\ \text{o } \delta^{82/76} \\ \text{Se}^{78} \\ \text{Se}^{-82} \\ \text{Se}^$	0.25 S shale s Se from e spike.	$rac{0.17}{ ext{standard}}$	3. s analy . Unles	<u>10</u> /sed in ss indic	this stud ated, res	y. * ind ults we	licate re
<sup>7*</sup> Se- <sup>78</sup> Se Table 2.	Results ratios co obtained	mean s of the two onverted to l using a <sup>7</sup>	$\frac{16}{\text{vo USG}}$ o $\delta^{82/76}$ Se- <sup>82</sup> S	$\frac{0.25}{\text{S shale s}}$ Se from the spike.	0.17 standard $\delta^{82/77}$ Se	3. s analy . Unles	$\frac{10}{10}$ /sed in ss indic	this stud ated, res	y. * ind ults we	licate re
Table 2.	Results ratios co obtained	mean s of the two onverted to l using a <sup>7</sup>	$\frac{16}{\text{vo USG}}$ o $\delta^{82/76}$ Se- <sup>82</sup> S	$\frac{0.25}{\text{S shale s}}$ Se from the spike. $\varepsilon^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/77}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7}Se^{76/7$	$\frac{0.17}{\text{standard}}$	$\frac{3.}{\text{s analy}}$ . Unles	$\frac{10}{\sqrt{\text{sed in}}}$ s indic	this stud ated, res	y. * ind ults we	licate re
Table 2.	Results ratios co obtained	mean s of the two onverted t l using a <sup>7</sup> ation	$\frac{16}{\text{vo USG}}$ o $\delta^{82/76}$ Se <sup>78</sup> Se <sup>-82</sup> S	$\frac{0.25}{\text{S shale s}}$ S shale so from the spike. $\varepsilon^{76/77}Se$	$\frac{0.17}{\text{standard}}$	$\frac{3.}{\text{s analy}}$ . Unles	$\frac{10}{\text{ysed in}}$	this stud ated, res	y. * ind ults we	licate re
<sup>7*</sup> Se- <sup>78</sup> Se Table 2.	Results ratios co obtained	mean s of the two onverted to l using a 7	$     \frac{16}{\text{vo USG}} \\         o \delta^{82/76} \\         Se^{-82} \\         Se^{-82} \\         \Delta^{33} \\         (‱)     $	$\frac{0.25}{\text{S shale s}}$ Se from the spike. $\varepsilon^{76/77}Se$ (%)	$\frac{0.17}{\text{standard}}$ $\delta^{82/77}\text{Se}$ $\frac{2^{2}82}{78} - 2 \text{ se}$	$\frac{3.}{\text{s analy}}$ . Unles	$\frac{10}{\sqrt{\text{sed in}}}$	this stud rated, res 2  se	y. * ind ults we	licate re
<sup>7*</sup> Se- <sup>78</sup> Se Table 2. 1 Sample SGR-1b	Results ratios co obtained	mean s of the two onverted t l using a <sup>7</sup> ation	$\frac{16}{\text{vo USG}}$ o $\delta^{82/76}$ Se- $^{82}$ S $^{78}$ Se- $^{82}$ S $^{33}$ S	$\frac{0.25}{\text{S shale s}}$ Se from be spike. $\varepsilon^{76/77}Se$ $(\%)$ $-1.654$	$\frac{0.17}{\text{standard}}$ $\delta^{82/77}\text{Se}$ $\frac{2^{82}}{78} = 2 \text{ se}$ $\frac{2^{82}}{78} = 0.410$ $3 = 0.52^{7}$	$\frac{3}{5} \text{ s analy}$ Unless $\frac{\varepsilon^{82/7}}{\varepsilon^{82/7}}$	$\frac{10}{\text{/sed in}}$ s indice $7Se_{\frac{82}{78}}$ $\frac{7}{8}e_{\frac{82}{78}}$	this stud ated, res 2  se 0.276 0.393	y. * ind ults we	licate re
<sup>7*</sup> Se- <sup>78</sup> Se Table 2. 1 Sample SGR-1b SGR-1b	Results ratios co obtained	mean s of the two onverted to l using a <sup>7</sup> ation	$     \frac{16}{\text{vo USG}} \\         o \delta^{82/76} \\         Se^{-82} \\         Se^{-82} \\         \Delta^{33} \\         (‰)     $	$\frac{0.25}{\text{S shale s}}$ Se from the spike. $\varepsilon^{76/77}Se$ $\frac{(\%\infty)}{-1.65^{2}}$	$\begin{array}{r} 0.17 \\ \text{standard} \\ \delta^{82/77} \text{Se} \\ \end{array}$	$\frac{3.}{\text{s analy}}$ . Unles	$\frac{10}{7\text{ sed in}}$	this stud ated, res 2 se 0.276 0.393	y. * ind ults we	licate
<sup>7*</sup> Se- <sup>78</sup> Se Table 2. 1 Sample SGR-1b SGR-1b SGR-1b	Results ratios co obtained	mean s of the two onverted t l using a 7	$\frac{16}{\text{vo USG}}$ o $\delta^{82/76}$ Se- <sup>82</sup> S	$\frac{0.25}{\text{S shale s}}$ S shale so Se from the spike. $\varepsilon^{76/77}Se$ (%) -1.654 -2.853	$\frac{0.17}{\text{standard}}$ $\delta^{82/77}\text{Se}$ $\frac{2^{2}82}{78} = 2 \text{ se}$ $4 = 0.416$ $3 = 0.522$	$\frac{3.}{\text{s analy}}$ . Unles	$\frac{10}{7 \text{ sed in}}$	2 se 0.276 0.393	y. * ind ults we	licate
<sup>7*</sup> Se- <sup>78</sup> Se Table 2. 1 Sample SGR-1b SGR-1b <i>GKF01</i> <i>Depth (n</i>	Results ratios co obtained Forma	mean s of the two onverted to l using a 7	$     \frac{16}{\text{vo USG}}     \text{ o } \delta^{82/76} S^{78} Se^{-82} S^{78} Se^{-82} S^{78} Se^{-82} S^{78} Se^{-82} S^{78} Se^{-82} Se^{$	$\frac{0.25}{\text{S shale s}}$ Se from the spike. $\varepsilon^{76/77}Se$ $\frac{(\%\infty)}{-1.65^2}$ $-2.853$	$\frac{0.17}{\text{standard}} \delta^{82/77} \text{Se}$	$\frac{3.}{\text{s analy}}$ . Unless	$\frac{10}{\text{/sed in}}$ s indice $7Se_{\frac{82}{78}}$ $\frac{7}{78}e_{\frac{82}{78}}$	this stud rated, res 2 se 0.276 0.393	y. * ind ults we	licate
<sup>7*</sup> Se- <sup>78</sup> Se Table 2. 1 Sample SGR-1b SGR-1b <i>GKF01</i> <i>Depth (n</i> 832 1	<ul> <li><i>Results</i></li> <li><i>Results</i></li> <li><i>ratios cc</i></li> <li><i>obtained</i></li> <li><i>Forma</i></li> </ul>	mean s of the two onverted to l using a 7	$\frac{16}{\text{vo USG}}$ o $\delta^{82/76}$ Se- <sup>82</sup> S $\Delta^{33}$ (%)	$\frac{0.25}{\text{S shale s}}$ Se from le spike. $\varepsilon^{76/77}Se$ $(\%\infty)$ -1.654 -2.853	$\frac{0.17}{\text{standard}}$ $\delta^{82/77}\text{Se}$ $\frac{2}{78} = 2 \text{ se}$ $\frac{2}{78} = 0.410$ $0.522$	$\frac{3}{5}$ s analy Unless $\frac{\varepsilon^{82/7}}{(\%)^{5}} = -0.$	$\frac{10}{\text{/sed in}}$ s indice $7Se_{\frac{82}{78}}$ $\frac{7}{8}e_{\frac{82}{78}}$ $\frac{500}{496}$	this stud eated, res 2 se 0.276 0.393	y. * ind ults we	licate re
<sup>7*</sup> Se- <sup>78</sup> Se Table 2. 1 Sample SGR-1b SGR-1b SGR-1b <i>GKF01</i> <i>Depth (n</i> 832.1 982	<ul> <li>spike</li> <li>Results</li> <li>ratios cc</li> <li>obtained</li> <li>Forma</li> <li>n)</li> <li>U. N</li> <li>U. N</li> </ul>	mean s of the two onverted to l using a <sup>7</sup> ation	$     \begin{array}{r}             16 \\             vo USG \\             vo \delta^{82/76} \\             \sqrt{8} Se^{-82} S \\             ^{78} Se^{-82} S \\             ^{78} Se^{-82} S \\             ^{33} S \\             (\%) \\             4.5 \\             4.5 \\             4.6 \\             4.5 \\             4.6 \\             4.5 \\             4.6 \\             4.5 \\             4.6 \\             4.5 \\             4.5 \\             4.5 \\             4.6 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             4.5 \\             5. \\             5. \\           $	$\frac{0.25}{\text{S shale s}}$ Se from be spike. $\varepsilon^{76/77}Se$ (%) -1.654 -2.853 -1.489 -1.654	$     \begin{array}{r}       0.17 \\       standard \\       \delta^{82/77} Se     \end{array} $ $     \begin{array}{r}       2_{82} \\       2_{78} \\       2_{78} \\       0.410 \\       3_{78} \\       0.522 \\     \end{array} $	$3.$ s analy Unles $\varepsilon^{82/7}$ $\varepsilon^{82/7}$ $(\%)$ 5 -0. 2 -0. 0 -0. 4 -0.	$\frac{10}{\text{/sed in}}$ s indice $7Se_{\frac{82}{78}}$ $\frac{5600}{496}$ 208 543 460	this stud ated, res 2 se 0.276 0.393 0.373 0 370	y. * ind ults we	licate
<sup>74</sup> Se- <sup>78</sup> Se Table 2. 1 Sample SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b	n) Nesults ratios cc obtained Forma U. N L. N L. N	mean s of the two onverted to l using a <sup>7</sup> ation ation Nauga Fm. Nauga Fm.	$     \begin{array}{r}                                     $	$\begin{array}{r} 0.25 \\ \hline S \text{ shale s} \\ \hline S \text{ shale s} \\ \hline S \text{ from } \\ e \text{ spike.} \\ \hline \varepsilon^{76/77} Se \\ \hline (\%\infty) \\ \hline -1.654 \\ -2.853 \\ \hline -1.489 \\ -1.651 \\ -2.200 \end{array}$	$\begin{array}{r} 0.17 \\ \text{standard} \\ \delta^{82/77} \text{Se} \\ \hline \\ \frac{2^{82}}{78} & 2 \text{ se} \\ \hline \\ 0.410 \\ 3 & 0.522 \\ \hline \\ 0 & 0.720 \\ 0.894 \\ \hline \\ 0 & 0.300 \\ \hline \end{array}$	$\frac{3.}{\text{s analy}}$ . Unless $\frac{\epsilon^{82/7}}{(\%)^{5} - 0.2}$	$\frac{10}{7}$ sed in sindices indices in	this stud ated, res 2 se 0.276 0.393 0.373 0.370 0 270	y. * ind ults we	licate
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Vauga Fm.	$     \begin{array}{r}             16 \\             vo USG \\             0 \delta^{82/76} \\             5 \\             \sqrt{^{78}} Se^{-82} \\             \Delta^{33} S \\             (\%)             4.5 \\             4.6 \\             6.9 \\             1.7             7         $	$\begin{array}{r} 0.25 \\ \hline S \text{ shale s} \\ \hline S \text{ shale s} \\ \hline S \text{ spike.} \\ \hline \varepsilon  $	$\begin{array}{c} 0.17 \\ \text{standard} \\ \delta^{82/77} \text{Se} \\ \end{array}$	3. s analy Unles $\varepsilon^{82/7}$ (%) $\varepsilon^{82/7}$ (%) $\varepsilon^{2}$ -0. 0 -0. 1 -0. 0 -0. 1 -0. 0 -0. 1 -0. 0 -0. 1 -0. 0 -0. 1 -0. 0 -0. 1 -0.	$\frac{10}{7}$ sed in sindices indices $7Se_{\frac{82}{78}}$ (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (2008) (200	this stud ated, res 2 se 0.276 0.393 0.373 0.370 0.270 0.149	y. * ind ults we	licate
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<sup>7*</sup> Se- <sup>78</sup> Se Table 2. 1 Sample SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1b SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1C SGR-1	<ul> <li><i>Results</i></li> <li><i>Results</i></li> <li><i>ratios cc</i></li> <li><i>obtained</i></li> <li><i>Forma</i></li> <li><i>N</i></li> <li><i>U. N</i></li> <li><i>L. N</i></li> <li><i>L. N</i></li> <li><i>Mon</i></li> <li><i>Lokan</i></li> </ul>	mean s of the two onverted to l using a <sup>7</sup> ation ation Nauga Fm. Nauga Fm. Nauga Fm. teville Fm. mmona Fm	$     \begin{array}{r}             16 \\             vo USG \\             0 \delta^{82/76} \\             \sqrt{3} \\             28 \\             \sqrt{3} \\             (\%)             4.5 \\             4.6 \\             6.9 \\             -1.7 \\             . 4.7         $	$\begin{array}{r} 0.25 \\ \hline S \text{ shale s} \\ \hline S \text{ shale s} \\ \hline S \text{ spike.} \\ \hline \varepsilon  $	$\begin{array}{r} 0.17 \\ \text{standard} \\ \delta^{82/77} \text{Se} \\ \hline \\ 282 \\ 78 \\ 2 \text{ se} \\ \hline \\ 282 \\ 78 \\ 2 \text{ se} \\ \hline \\ 0.410 \\ 3 \\ 0.522 \\ \hline \\ 0.894 \\ \hline 0.894 \\ $	3. s analy Unles $\varepsilon^{82/7}$ (%) $\varepsilon^{82/7}$ (%) $\varepsilon^{2}$ -0. 0. $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $-0.$ $0.$ $0.$ $0.$ $0.$ $0.$ $0.$ $0.$	$\frac{10}{7} \text{ sed in}$ ss indice $7 \text{ Se}_{\frac{82}{78}}$ $\frac{7}{78} \text{ se}_{\frac{82}{78}}$ $\frac{5}{400}$ $\frac{5}{496}$ $\frac{5}{433}$ $\frac{6}{433}$ $\frac{3}{361}$	this stud ated, res 2 se 0.276 0.393 0.373 0.370 0.270 0.149 0.340 2 sd 0.274	y. * ind ults we	licate re

Table 3. Se mass independent fractionation isotope ratios.  $\Delta^{33}$ S data from Zerkle et al.

806	(2012).	
807		
808		
809	<u>Appendix</u>	
810	Interference corrections for 78-82 double spiked samples	
811	After on-peak blank subtraction the ion beam intensities <sup>76</sup> Tot, <sup>77</sup>	Tot, <sup>78</sup> Tot and <sup>82</sup> Tot
812	at masses 76, 77, 78 and 82 respectively are taken to be the sums	of the contributions
813	from: i) Se and Ge atomic ions and ii) SeH and AsH molecular io	ons, i.e.,
814		
815	$^{76}$ Se = $^{76}$ Tot - $^{76}$ Ge - AsH (1)	
816	$^{77}$ Se = $^{77}$ Tot - $^{76}$ SeH (2)	
817	$^{78}\text{Se} = ^{78}\text{Tot} - ^{77}\text{SeH}$ (3)	
818		
819	<sup>82</sup> Se suffers no contributions from these species. With the assumption	otion of a single
820	instrumental mass bias parameter, $\beta$ (see Rudge et al., 2009), while	ich applies to all
821	species, the ion beam intensities <sup>76</sup> Ge, SeH and AsH in the above	equations may be
822	written in terms of the measured intensities of <sup>73</sup> Ge, <sup>82</sup> SeH and A	s thus,
823		,
821	$^{76}S_{e} = {}^{76}T_{o}t {}^{73}C_{o}[{}^{76}C_{o} / {}^{73}C_{o}]_{e}{}^{\beta pGe} A_{e} h_{o}{}^{\beta(pAsH-p_{83}^{SeH})}$	(A)
02 <del>4</del> 075	Se = 10i - 0e[0e/0e]e - As ne	(4)
823	and the set of s	
826	$^{\prime\prime}Se = ^{\prime b}Tot - ^{\prime b}Se \cdot he^{\beta(p_{77} - p_{83} - )}$	(5)
827		
828	$^{78}Se = ^{78}Tot - ^{77}Se \cdot he^{\beta(p_{78}^{SeH} - p_{83}^{SeH})}$	(6)
829		
830	where	
831		
832	$h = {}^{82}SeH/{}^{82}Se$	(7)
833		
834	$p^{Ge} = \ln\{mass(^{76}Ge)/mass(^{73}Ge)\}$	(8)
835		
836	$p^{AsH} = ln \{mass(AsH)/mass(As)\}$	(9)
837		
838	$p_{83}^{SeH} = \ln \{ mass(^{82}SeH) / mass(^{82}Se) \}$	(10)
839		· · /
840	$p_{77}^{SeH} = \ln \{ mass(^{76}SeH) / mass(^{76}Se) \}$	(11)
841		× ,
842	$p_{78}^{\text{SeH}} = \ln \{ \text{mass}(^{77}\text{SeH}) / \text{mass}(^{77}\text{Se}) \}$	(12)
843		
844	[ <sup>76</sup> Ge/ <sup>73</sup> Ge] is the true Ge isotopic abundance ratio, assumed kno	wn.
845		
846	Let,	
847	$m_1(\beta) = {}^{77}Se/{}^{76}Se$	(13)
848		()
849	$m_2(\beta) = {^{78}Se}/{^{76}Se}$	(14)
850	m <sub>2</sub> (p) 50, 50	(* ')
851	$m_2(\beta) = {}^{82}Se^{/76}Se$	(15)
857	$m_3(p)$ by by	(13)
002		

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 Using equations 4–6 we may write down expressions for these three isotopes ratios in terms of measured intensities, known constants and the instrumental mass bias,  $\beta$ . For the measurement of a mixture of sample and double-spike tracer we may write down three mixing relations,  $m_i(\beta)e^{-\beta P_i} = \lambda T_i + (1-\lambda)n_ie^{-\alpha P_i}$ , (i = 1,2,3) (16)

where  $P_i = \ln \{mass(^{XX}Se)/mass(^{76}Se)\}$  and xx=77, 78 and 82 for i = 1, 2, and 3respectively, T<sub>i</sub> is the isotopic ratio of the double spike tracer,  $\alpha$  is the fractionation (parameter) of the sample relative to the reference ratio,  $n_i$ , and  $\lambda$  is the mixing parameter. These are the well-known sample-tracer mixing equations (see the Double Spike Toolbox (Rudge et al., 2009) for these well-established definitions and equations) with the mixture ratio,  $m_i$ , now a function of  $\beta$  as a result of the interference correction. we can solve these three equations for the three unknowns  $\alpha$ ,  $\beta$  and  $\lambda$  by means of standard numerical methods. The reported  $\delta^{82/76}$ Se is then given by, 

 $\delta^{82/76} Se = (e^{-\alpha P_3} - 1) \times 10^3 \tag{17}$ 



