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## Self-powered SBD solar-blind photodetector fabricated on the single crystal of $\beta$ -Ga<sub>2</sub>O<sub>3</sub>

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A Schottky barrier diode (SBD) solar-blind photodetector was fabricated based on the single crystal  $\beta$ -Ga<sub>2</sub>O<sub>3</sub>. Cu and Ti/Au were deposited on the top and bottom surface of Ga<sub>2</sub>O<sub>3</sub> as Schottky and ohmic contacts, respectively. The SBD exhibits a higher rectification ratio of up to  $5 \times 10^7$  at  $\pm 2$  V. The photoresponse spectra show a maximum responsivity at 241 nm and a cutoff wavelength of 256 nm. The solar-blind/ultraviolet and solar-blind/visible rejection ratio can reach a high level of up to 200 and 1000, respectively. It is interesting that the device has a clear response to the solar-blind wavelength at zero bias, which confirms it can be used as a self-powered solar-blind photodetector.

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### Introduction

The solar radiation in the range of 200–280 nm cannot reach the earth's surface due to the absorption by ozone so that the ultraviolet (UV) light in this range is appropriate for communication as a 'black background'.<sup>1,2</sup> If a photodetector working in this range has no response to solar radiation, the photodetector is called a solar-blind photodetector. Nowadays the solar-blind photodetector has been widely applied in many fields such as missile warning, fire alarms, ozone monitoring, under water and space communication, and so on.<sup>3–5</sup>

Currently, the materials suitable for fabricating solar-blind photodetectors include AlGa<sub>N</sub>, MgZnO and Ga<sub>2</sub>O<sub>3</sub>. Both AlGa<sub>N</sub> and MgZnO are ternary alloys, which require complex processes to obtain high Al<sup>1</sup> and Mg<sup>6</sup> content materials.  $\beta$ -Ga<sub>2</sub>O<sub>3</sub> is one of the best materials for solar-blind photodetector fabrication due to its wide bandgap of 4.9 eV ( $\approx 253$  nm), which locates at the center position of the solar-blind wavelength. More importantly, large size and high quality Ga<sub>2</sub>O<sub>3</sub> single crystals can be prepared through conventional methods such as floating zone (FZ),<sup>7</sup> edge-defined film-fed growth (EFG)<sup>8</sup> and Czochralski (CZ)<sup>9</sup> methods. Ga<sub>2</sub>O<sub>3</sub> solar-blind photodetectors are now mainly fabricated on thin films with a variety of film growth technique such as PLD,<sup>10</sup> MBE,<sup>11</sup> MOCVD<sup>12</sup> and radio-frequency magnetron sputtering,<sup>13</sup> and on nanostructures *i.e.* nanowires,<sup>14</sup> rods,<sup>15</sup> belts<sup>16</sup> and arrays.<sup>17</sup> However, defects introduced during these growth methods can trap and de-trap

the photogenerated carriers within the deep levels, which could cause photocurrent persistence for an extended period (from approximately a few minutes to hours) after the excitation light is shut off, and this phenomenon is called the persistent photoconductivity (PPC) effect.<sup>16,18</sup> Single crystal material can reduce the defects density to a very low level and increase the device response speed. And eliminating the influences of defects<sup>19</sup> and narrow band gap heterojunction substrate,<sup>20</sup> the response range of the single crystal Ga<sub>2</sub>O<sub>3</sub> photodetector can be limited sharply to solar-blind wavelength. Moreover, the photodetector based on a single crystal can be operated with a low dark current density, which can increase the device sensitivity. Besides, when prospecting the application of avalanche photodetector, single crystal material would be a better choice for the high resistance to avalanche current due to the perfect crystalline quality.

Traditional photodetectors typically require an external bias as the driving force to prevent the recombination of photo-generated carriers.<sup>21</sup> However, the external power sources limit their independence in application as the size of the power sources always far exceed that of detectors. Nowadays, it is highly desirable for the photodetector to function independently of external power sources. It has been common in the previous reports to call devices of this type "self-powered".<sup>18,22,23</sup> Self-powered detectors decrease energy consumption and are more environmentally friendly. Combined with wireless technology, self-powered detectors present greater adaptability and sustainability, even used in unattended hazardous atmosphere or harsh environment. Self-power can be achieved by photovoltaic effect by means of p–n junction, heterojunction and Schottky junction. When UV lights incidents into the devices, built-in electric field in these structures can separate photogenerated electron–hole pairs and output the electrical signal.<sup>24,25</sup>

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In this work, the self-powered Ga<sub>2</sub>O<sub>3</sub> Schottky barrier diode (SBD) solar-blind photodetector was fabricated with the deposition of Cu and Ti/Au on the top and bottom surfaces of Ga<sub>2</sub>O<sub>3</sub> single crystal substrate, acting as Schottky and ohmic contact respectively. And the current–voltage (*I*–*V*) and photoresponse characteristics of the photodetector were investigated.

## Experimental

The commercial (201)-orient β-Ga<sub>2</sub>O<sub>3</sub> substrate with 680 μm thickness was used to fabricate the SBD by cutting along the [010] and [102] orientation into a 5 × 5 mm<sup>2</sup> piece. The unintentional doped Ga<sub>2</sub>O<sub>3</sub> crystal is n-type with a carrier concentration of about 10<sup>17</sup> cm<sup>-3</sup>. The samples were firstly cleaned using methanol (5 min)/acetone (5 min)/methanol (5 min)/deionized water (5 min) successively. Next, the samples were etched in the solution with H<sub>2</sub>SO<sub>4</sub> : H<sub>2</sub>O<sub>2</sub> : H<sub>2</sub>O = 4 : 1 : 1 for 5 min. Finally, the samples were dipped in 90 °C deionized water until the temperature cooling down to room temperature. The metal deposition was performed by vacuum thermal evaporation method. Ohmic contact electrode was fabricated firstly with Ti/Au (20 nm/200 nm) deposition on the bottom surface of the Ga<sub>2</sub>O<sub>3</sub> substrate followed by 500 °C annealing in nitrogen for 3 min. Then Cu (20 nm) was deposited on the top surface of the Ga<sub>2</sub>O<sub>3</sub> substrate with 150 °C annealing in nitrogen for 5 min as the Schottky contact electrode. The schematic structure of the detector is shown in the inset of Fig. 1.

X-ray diffraction (XRD) measurement was carried out to determine crystalline quality using Bruker D8 DISCOVER diffractometer. The *I*–*V* characteristics for the SBD were measured using Keithley 4200-SCS semiconductor characterization system. Time-resolved photocurrent response

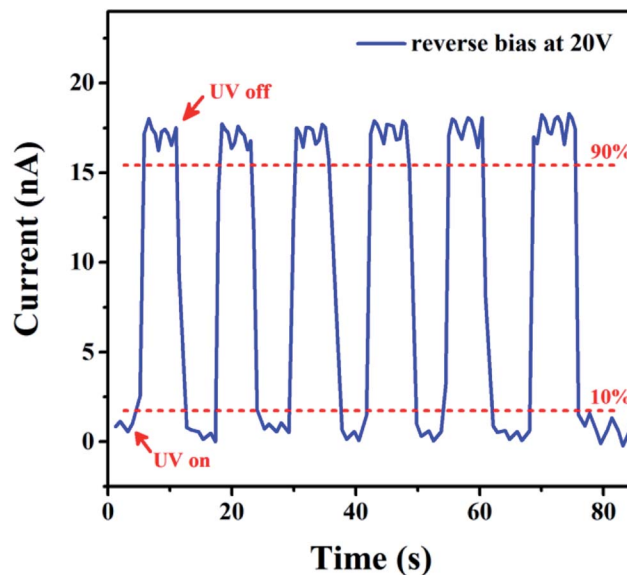


Fig. 2 Time-resolved photocurrent spectrum with the UV light (253 nm) of photodetector at –20 voltage.

spectrum was conducted under 254 nm light with Keithley 4200-SCS semiconductor characterization system. The photo-response spectra of the photodetectors were measured in a SPEX scanning monochromator employing a 150 W Xe lamp as the illumination source. The spectrum of the Xe lamp can cover from 200 nm to over 1000 nm, and the responsivity spectrum was obtained by measuring the photocurrent (calibrated with a standard Si photodiode) under the illumination of the Xe lamp spectrum from 200 nm to 500 nm using the scanning monochromator.

## Results and discussion

XRD rocking curve measurement was carried out to determine the crystalline quality of the single crystal Ga<sub>2</sub>O<sub>3</sub>. The full width at half maximum (FWHM) values of (603) diffraction peak is 42.2 arcsec (figure not shown here), which confirms good crystalline quality. The *I*–*V* characteristics of the photodetector were obtained at dark condition, as shown in Fig. 1. Linear characteristic for the bottom electrode is shown in the inset, which indicates ohmic contacts between Ga<sub>2</sub>O<sub>3</sub> and Ti/Au electrode. In Fig. 1, the forward current is linear in the semilogarithmic scale at low forward bias voltages and deviates from linearity due to the effect of series resistance *R*<sub>s</sub>. The reverse current increases slightly, which is probably caused by barrier height inhomogeneities.<sup>26</sup> The forward current can reach above milliamp order of magnitude and the reverse current is about sub-nanoamp order of magnitude. The SBD exhibits a higher rectification ratio at ±2 V up to 5 × 10<sup>7</sup>. Assuming that thermionic emission is the dominant current transport mechanism in the SBD, and taking into account the effect of series resistance *R*<sub>s</sub>, the diode equations can be written as eqn (1):<sup>27–29</sup>

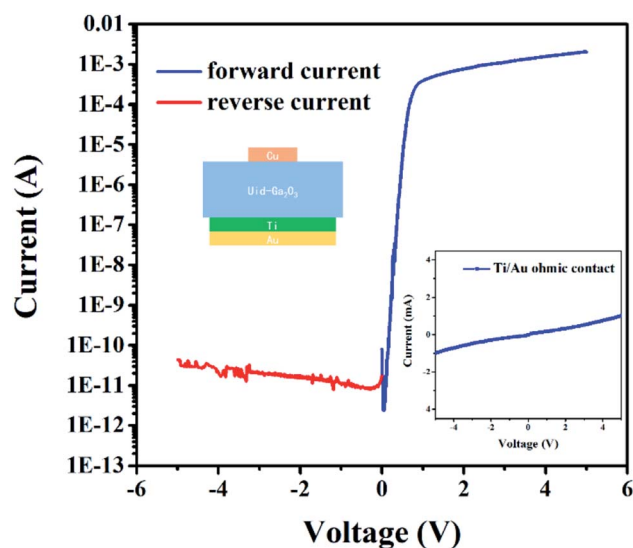


Fig. 1 *I*–*V* characteristics of the Ga<sub>2</sub>O<sub>3</sub> solar-blind photodetector at dark condition. The inset depicts the ohmic contact characteristic on the bottom side of the single crystal Ga<sub>2</sub>O<sub>3</sub>.



$$I = I_s \exp\left(\frac{q(V - IR_s)}{nKT}\right) = A^{**} T^2 \exp\left(-\frac{q\phi_b}{kT}\right) \exp\left(\frac{q(V - IR_s)}{nKT}\right) \quad (1)$$

where  $I$  is current,  $I_s$  is saturation current density,  $V$  is applied forward bias voltage,  $K$  is Boltzmann constant,  $q$  is elementary charge,  $n$  is ideality factor,  $A^{**}$  is effective Richardson constant,  $\phi_b$  is the barrier height and  $T$  is temperature. The effective Richardson constant  $A^{**}$  is calculated to be  $41.04 \text{ A cm}^{-2} \text{ K}^{-2}$  at room temperature,<sup>30</sup> using electron effective mass  $m^* = 0.342m_0$ ,<sup>31</sup> and free electron Richardson constant  $A^* = 120 \text{ A cm}^{-2} \text{ K}^{-2}$ . Fitting this model to the linear range of the semi-log plot of  $I$  versus  $V$ , values  $\phi_b$  is  $0.95802 \text{ eV}$  and  $n$  is  $1.47378$  at room temperature. The results are very close to the  $\text{Cu}/\text{Ga}_2\text{O}_3$  Schottky diode based on  $\text{Ga}_2\text{O}_3$  thin film.<sup>26</sup>

The response speed is an important parameter for photo-detector, which can be used to evaluate the detection performance to the fast-varying signal. Time-resolved photocurrent response spectrum was measured with on/off switching of  $254 \text{ nm}$  light under a working bias of  $-20 \text{ V}$ , as shown in Fig. 2. The 10–90% rise time is estimated to be  $0.729 \text{ s}$  and 90–10% decay time to be  $1.209 \text{ s}$  and no long time PPC effect is observed due to the high crystalline quality. This PPC phenomenon is generally found in the  $\text{Ga}_2\text{O}_3$  thin film and nanostructure photodetectors.<sup>32,33</sup> The on and off state current for six cycles remain the same level, which indicates the high reversibility and high reproducibility of the  $\text{Ga}_2\text{O}_3$  solar-blind photodetector.

Fig. 3(a) is the photoresponse spectra under different reverse voltages up to  $35 \text{ V}$ . The responsivity at peaks increases obviously at low bias range, while it increases slowly at high bias range due to the recombination as the photogenerated carriers transport through the single crystal. The responsivity of the photodetector can be calculated by eqn (2):<sup>34</sup>

$$R = \frac{I_{\text{photo}} - I_{\text{dark}}}{S \times P} \quad (2)$$

where  $R$  is the responsivity of the photodetector,  $I_{\text{photo}}$  is the photo current,  $I_{\text{dark}}$  is the dark current,  $S$  is the effective

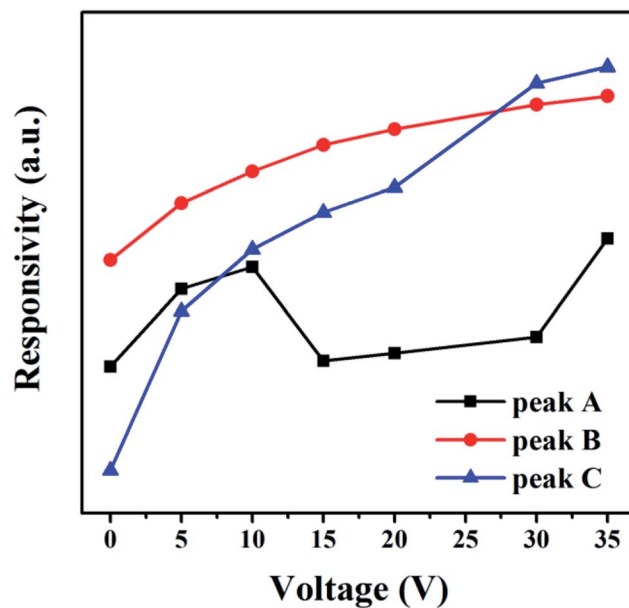


Fig. 4 Intensity of three peaks in the photoresponse spectra at different bias.

illuminated area and  $P$  is the light intensity. The relationship between the responsivity and the external quantum efficiency is given by eqn (3):<sup>29</sup>

$$\eta = \frac{hcR}{e\lambda} \quad (3)$$

where  $h$  is Planck's constant,  $c$  is the velocity of light,  $e$  is the basic electron charge, and  $\lambda$  is the incident light wavelength. The device responds obviously at zero bias, as shown in Fig. 3(b), which demonstrates that the device can be used as a self-powered solar-blind photodetector. A sharp cutoff wavelength defined as the ratio between the maximum and the natural constant ( $e \approx 2.718$ ) is located at  $256 \text{ nm}$  (close to the bandgap of  $\beta\text{-Ga}_2\text{O}_3$ ), and the predominant peak at  $241 \text{ nm}$ , both of which have blueshift than the previous report<sup>19,20</sup> due to the low defects concentration of the single crystal substrate. The

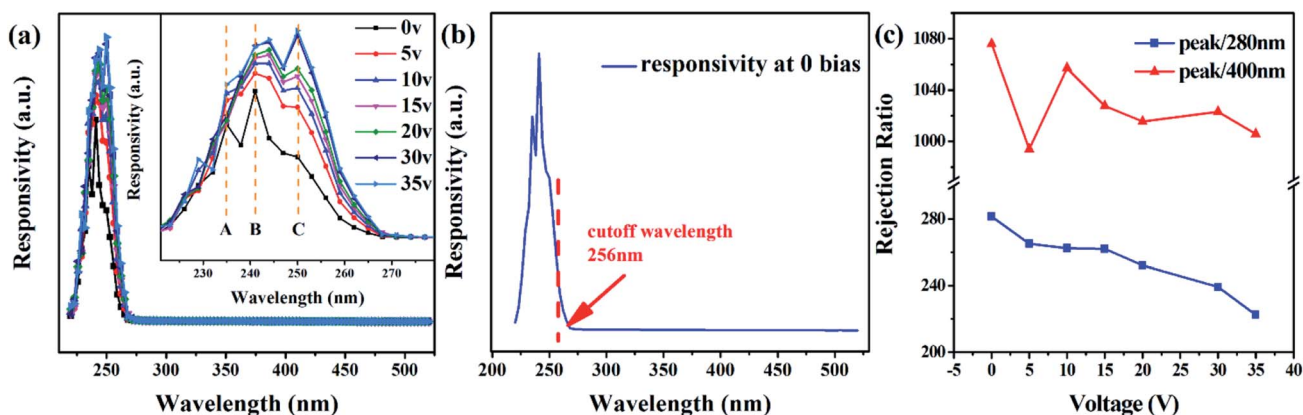


Fig. 3 Photoresponse characteristics for the photodetector: (a) responsivity at different bias, (b) responsivity at 0 bias, (c) solar-blind/visible and solar-blind/UV rejection ratio at different bias.



solar-blind to visible and UV rejection ratio (defined as the ratio of responsivity at peak to 400 and 280 nm) at different reverse bias can reach more than 1000 and 200 respectively, both of which present maximum ratio at zero bias as shown in Fig. 3(c).

The responsivity at solar-blind wavelength range exhibit three peaks, as shown in the inset of Fig. 3(a). All of them changes in different trends as the reverse bias increases, as shown in Fig. 4. This phenomenon can be explained by the different absorption mechanisms of Ga<sub>2</sub>O<sub>3</sub> to different wavelengths of UV light. According to Medvedeva's calculation,<sup>35</sup> states of s, p and d of the cations in Ga<sub>2</sub>O<sub>3</sub> all contributes to the conduction band. Due to the different degree of the orbitals hybridization, the electrons have different velocities on various orbitals. We assume that UV light with different wavelengths excited electrons to different orbitals, which caused different variation trends as the increasing bias. A further investigation will be required to obtain a better understanding of the absorption mechanisms.

## Conclusions

Cu SBD solar-blind photodetector has been successfully fabricated based on the single crystal β-Ga<sub>2</sub>O<sub>3</sub>. The SBD exhibited a higher rectification ratio up to  $5 \times 10^7$  at ±2 V. The photo-response spectrum shows a maximum responsivity at 241 nm and a cutoff wavelength of 256 nm. The low dark current and sharp cutoff wavelength result from the high crystalline quality. The device has a clear response to solar-blind wavelength at zero bias, which confirms it can be used as a self-powered solar-blind photodetector.

## Conflicts of interest

There are no conflicts to declare.

## Acknowledgements

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